

Cryptotephra in the East Antarctic Mount Brown South ice core

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Abstract.

Ice cores contain stratified layers of impurities scavenged from the atmosphere, which are a vital tool for investigating the Earth system. Reconstructing past eruption records by way of ice core tephrochronology can help us understand ash dispersal, atmospheric circulation processes, and the impacts of volcanic eruptions on climate. This study presents the coastal East Antarctic Mount Brown South (MBS, 69.11°S, 86.31°E; 2084 m ASL) ice core as an untapped tephrochronological archive. We utilize a novel cryptotephra sampling plan, integrating ice core data, HYSPLIT air parcel trajectories, and known eruption records, and identify two distinct cryptotephra horizons at ~13.3 and ~17.9 m depth in the MBS-Alpha ice core. We also find sparse ~~tephra grains~~ [cryptotephra grains from various unidentified sources](#) throughout the core. Through geochemical characterization with electron probe microanalysis (EPMA), we correlate the two cryptotephra horizons with the 1991 eruption of Cerro Hudson and the continuous eruptions of Mt. Erebus throughout the mid-1980s. The volcanic horizons identified here underscore the role of MBS in extending the regional volcanic record, helping to constrain ice core dating efforts, and enhancing understanding of volcanic ash dispersal to East Antarctica.

1 Introduction

Ice core records are important archives of the changing ~~earth systems~~ [Earth system](#). They record recent climate events in detail, including volcanic eruption histories, and in some cases the strength and source of such eruptions (Svensson et al., 2020; Sigl et al., 2014; Lowe, 2011; Abbott et al., 2024; Narcisi et al., 2012; Kurbatov et al., 2006; Dunbar et al., 2003). Ice sheets and ice caps can contain both soluble (e.g. volcanic sulfate) and insoluble [volcanic products](#) (e.g. ash) ~~volcanic products~~, deposited on the ice, providing valuable information on past volcanic eruption histories, ~~and analyzing~~. [Analyzing](#) these products in ice cores can deepen our understanding of volcanic climate forcing, and help to refine ice core chronologies (Gao et al., 2008; Castellano et al., 2004; Sigl et al., 2014, 2015; Lohmann and Svensson, 2022; Lin et al., 2022).

Tephrochronology, the use of tephra ([used here to refer to any material ejected during explosive volcanic eruptions](#)), ~~and in ice core studies, typically ash sized~~ and [cryptotephra \(any tephra not visible to the naked eye, due to either sparseness or small grain size\)](#) as isochronous horizons in stratigraphic archives such as ice cores and sediment records can be used to date archives,

synchronize geographically distinct records, and inform about past volcanic events (Lowe, 2011; Geyer et al., 2023; Cook et al., 2022). While volcanic horizons can be identified by soluble tracers ~~like volcanic sulfate or acidity~~, including volcanic sulfate, acidity, and non-sea-salt conductivity, characterization and geochemical fingerprinting of tephra and cryptotephra can be used to identify the specific volcanic sources for ~~eruptions events~~ (Cook et al., 2018; Lin et al., 2022; Svensson et al., 2020; Geyer et al., 2023; Lowe et al., 2024; Cook et al., 2018; Winstrup et al., 2019; Svensson et al., 2020; Lin et al., 2022; Geyer et al., 2023).

Ice core records are often geographically distal or ultra-distal (up to thousands of kilometers away or more) archives of volcanic ash, and the tephra found are typically micrometer-scale volcanic glass shards from volcanic ash fallout transported long distances by atmospheric circulation processes (Lowe, 2011; Geyer et al., 2023). Locating and identifying cryptotephra ~~(tephra not visible to the naked eye due to size or sparseness)~~ in ice cores is especially difficult, as volcanic markers such as sulfate, conductivity, or acidity do not always co-occur with tephra deposits from the same eruption, and comprehensive sampling must be undertaken to produce a full tephrochronological framework (Cook et al., 2018; Lin et al., 2022; Abbott et al., 2024; Lowe, 2011; Narcisi et al., 2010; Cook et al., 2022; Basile et al., 2001).

Tephra previously identified in other studies of Antarctic ice cores and blue ice areas come from local and regional sources, including Antarctic ~~(Narcisi et al., 2006; Basile et al., 2001; Abbott and Davies, 2012)~~ (Narcisi et al., 2006; Basile et al., 2001; Abbott and Davies, 2012) and sub-Antarctic island volcanoes (Basile et al., 2001; Abbott et al., 2024; Narcisi et al., 2012). Additionally, tephra found in Antarctic ice and snow has been correlated to large eruptions from ultra-distal ~~eruptions e.g. Aotearoa~~ /sources e.g. Aotearoa New Zealand (~~Taupo~~ Taupō; Dunbar et al. (2017)), South America (Cerro Hudson, Puyehue-Cordón Caulle; Abbott et al. (2024); Narcisi et al. (2012); Koffman et al. (2017)), and possibly even Mexico (El Chichón; Palais et al. (1992)).

Many efforts in recent years have been undertaken to improve Antarctic tephrochronologies, using ice cores from across the continent including Talos Dome (Narcisi et al., 2012), Siple Dome (Kurbatov et al., 2006), Vostok (Basile et al., 2001; Narcisi et al., 2010), and others (Abbott et al., 2024; Narcisi et al., 2005, 2010). Ice cores spanning millenia can be used to produce tephrochronologies ~~spanning~~ covering the Holocene or Last Glacial period, often with an aim to constrain timings of changes to global climate (Lin et al., 2022; Lohmann and Svensson, 2022; Cook et al., 2022; Castellano et al., 2004; Abbott and Davies, 2012). ~~Short-ice~~ Ice cores from higher ~~resolution sites~~ accumulation sites such as WAIS-Divide and Law Dome, on the other hand, can provide important means for investigating volcanic eruptions and climate in detail in the more recent past (Abbott et al., 2024; Plunkett et al., 2023; Piva et al., 2023; Narcisi and Petit, 2021; Sigl et al., 2014, 2013; Narcisi et al., 2012; Gao et al., 2008).

~~A number of volcanoes with the potential to transport volcanic material to Antarctica have been active during the satellite era (1978 to present) (Global Volcanism Program, 2024).~~ Volcanic products most commonly ~~seen in Antarctic~~ identified in Antarctic ice core records include those from Antarctic (e.g. Mt. Erebus) and Sub-Antarctic ~~Island~~ island volcanoes (including South Sandwich and South Shetland Islands), as well as lower latitude volcanoes in Chile and ~~Aoteroa/Aotearoa~~ Aotearoa New Zealand (Narcisi et al., 2012; Dunbar et al., 2017; Koffman et al., 2017). A number of volcanoes with the potential to disperse volcanic material to Antarctica have been active during the satellite era (1979 to present) (Global Volcanism Program, 2024). Such

satellite era eruption events are typically well observed through comprehensive monitoring programs and satellite remote sensing (Global Volcanism Program, 2024; Francis et al., 1996; Poland et al., 2020).

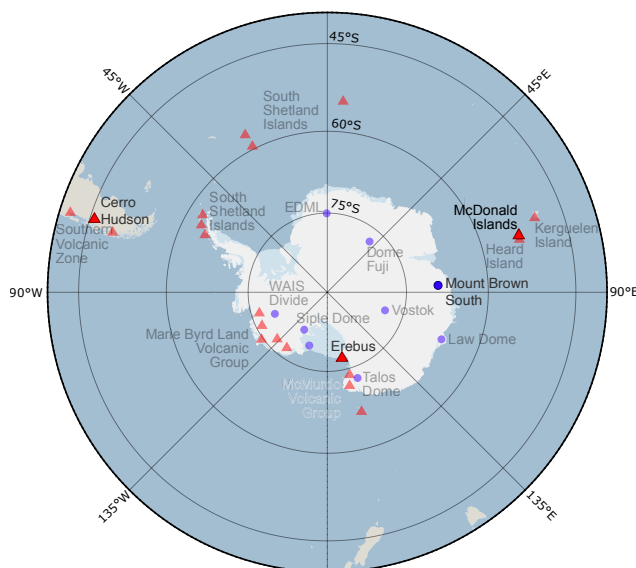


Figure 1. Map of Antarctica showing the location of the MBS ice core site, as well as volcanic regions of interest relevant to this study (basemap provided by the SCAR Antarctic Digital Database, accessed via the Norwegian Polar Institute’s Quantarctica package (Matsuoka et al., 2018)).

60 The Mount Brown South (MBS) ice cores, comprising ~~comprising a deep an intermediate~~ ice core (295-290 m) and ~~3~~-three surface cores (~20 - ~~25-26~~ m), were drilled in coastal East Antarctica during the 2017-2018 austral summer field season (69.11°S, 86.31°E, 2084 m ASL; Fig. 1). The MBS cores provide a new, high resolution climate archive spanning 1137 years, with shallow cores providing duplicate records from ~~the satellite era~~-1979 to 2017 (Vance et al., 2024a).

The MBS site was selected for its teleconnections and strong climatological link to the Southern Indian Ocean, providing
65 a millennial-length past climate record for a region underrepresented in the existing array of Antarctic ice core records (Vance et al., 2016, 2024a). Subsequent analyses have shown that East Antarctic ice core records (from MBS as well as the Law Dome ice core) provide insight into past climate conditions, ~~including capturing~~. ~~East Antarctic records have preserved~~ anthropogenic changes to atmospheric greenhouse gases (Etheridge et al., 1996), El Niño-Southern Oscillation (Crockart et al., 2021) and Australian hydroclimate and bushfire conditions (Udy et al., 2022, 2024), as well as regional events including atmospheric
70 rivers and extreme precipitation events (Jackson et al., 2023; Gkinis et al., 2024b; Zhang et al., 2023).

Due to ~~the coastal Antarctic site~~s coastal Antarctic location, as well as the known teleconnections across the region (Vance et al., 2016; Crockart et al., 2021), we propose that MBS is well situated ~~amongst Antarctic ice cores among the Antarctic ice core array~~ to receive and store wind-blown volcanic ash, making it a useful tephrochronological archive. We present an investigation of cryptotephra in the MBS ice cores, in order to assess potential preservation of volcanic ash, and better charac-

75 terize the tephra transport pathways in the region. This study focuses on the satellite era record from MBS (~~1978 to present~~1979
to 2017) as a first investigation of cryptotephra in MBS. Our characterization of the cryptotephra horizons together with at-
mospheric modelling and satellite derived data will inform future design and interpretation of tephrochronology paleo-record
studies of MBS and ice cores across Antarctica.

2 Materials and methods

80 2.1 Mount Brown South ice cores

The MBS-Main core (4.25-294.785 m depth) is supported by three surface firn cores: MBS-Alpha (surface to 20.41 m depth),
-Bravo (surface to 20.225 m depth), and -Charlie (surface to 25.89 m depth). Mean accumulation rates derived by annual
layer counting throughout the satellite era ~~is 0.298~~are $0.309 \pm 0.07 \text{ m yr}^{-1}$ IE0.08 m yr⁻¹ ice equivalent (IE) for the main core
and ~~0.309~~ $0.298 \pm 0.08 \text{ m yr}^{-1}$ IE0.07 m yr⁻¹ IE for the MBS-Alpha core (Crockart et al., 2021). This correlates with the ERA5
85 estimated site annual accumulation of $0.302 \pm 0.05 \text{ m yr}^{-1}$ IE (Crockart et al., 2021). The climatology and site conditions of
Mount Brown South are discussed in detail in Vance et al. (2016, 2024a), Crockart et al. (2021) and Jackson et al. (2023).

The MBS-Main and MBS-Alpha cores have been analyzed for both trace impurities and stable water isotopes ($\delta^{18}\text{O}$) (Gkinis
et al., 2024a, b; Jackson et al., 2023). Measurements were performed on discrete samples at 3 cm resolution, using ion chro-
matography (chemistry/trace impurities, Vance et al. (2024b); Harlan et al. (2024b)) and cavity ring down spectroscopy (stable
90 water isotopes, Moy et al. (2024); Gkinis et al. (2024b)) at the Institute for Marine and Antarctic Studies at the University
of Tasmania and Physics of Ice, Climate and Earth, Niels Bohr Institute, at the University of Copenhagen. Additionally, the
MBS-Main core was analyzed using continuous flow analysis (CFA) for trace impurities (Harlan et al., 2024b) and stable water
isotopes (Gkinis et al., 2024b), also at the University of Copenhagen.

Limited sample material remains of the MBS-Main core throughout the satellite era depths, after re-sampling was required
95 due to challenges with contamination during discrete chemistry sampling. The outer edge pieces remaining after preparation
of samples for CFA remain available to use, but the sample size is small ($\sim 2\text{cm}^2$). The MBS-Alpha core, had a much larger
remaining sample area usable for this study ($> 10\text{cm}^2$). Part of the MBS-Alpha core was damaged in a freezer failure, resulting
in some sections with one slightly melted edge along the length of the core. The trace chemistry samples from the damaged
cores appear to be reliable (see Vance et al. (2016)) and the damaged edges of the core have been removed. While the core is
100 not suitable for analyses where high temperature or slight meltwater migration would compromise the results, the cores are
still able to be used for microparticle analysis, including tephra studies.

The chemistry and stable water isotopes for the MBS-Main and MBS-Alpha cores have been analyzed and compared across
cores (Crockart et al., 2021). The cores align well, and robust chronologies have been produced for both cores (Crockart et al., 2021; Vance
. The MBS-Main and surface cores have been dated using independent layer counting and volcanic matching. A thorough
105 description of the process of chronology development and the resulting age scale (MBS2023) is presented in Vance et al. (2024a)
. The MBS2023 chronology provides depths for annual horizons for the main and surface cores (Vance et al., 2024b). Thanks

to the robust chronology development efforts for these cores, we are able to cross-match depths across the MBS ice core array, allowing us to make use of the multiple cores to provide larger volumes of ice core material for sampling.

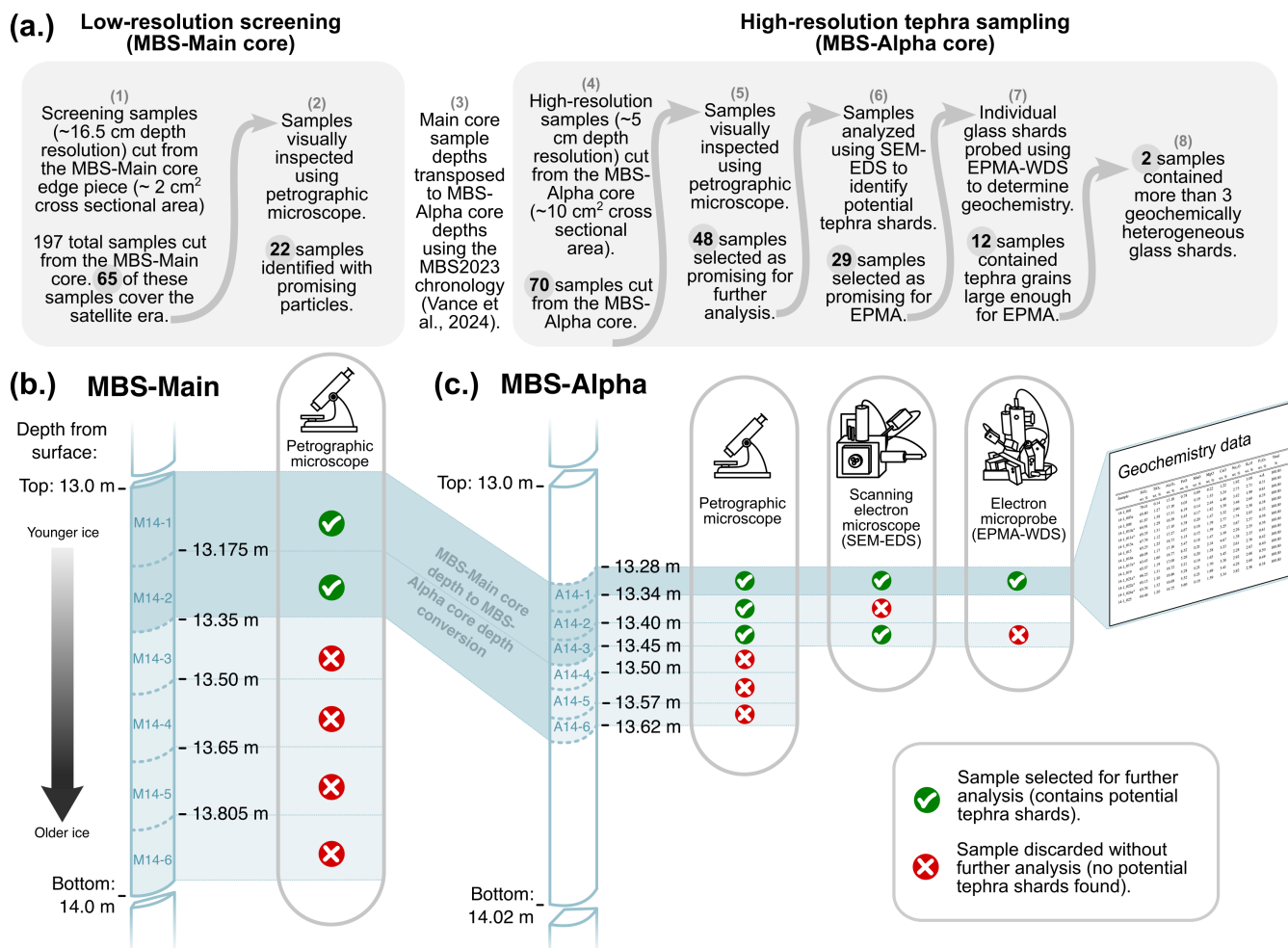


Figure 2. a. Schematic describing two-phase sampling strategy. A first low-resolution set of "screening samples" were prepared from the **MBS-main** **MBS-Main** core, followed by a resampling from the MBS-Alpha core, due to larger available sample volume. Samples from the MBS-Alpha core were assessed based on visual examination using a petrographic microscope, scanning electron microscopy by energy dispersive spectroscopy (SEM-EDS) with automated mineralogy (AMICS), backscatter electron (BSE) imaging, and finally geochemistry was measured using EPMA-WDS Electron probe microanalysis (EPMA). Examples shown of sampling procedure for selected samples from 13 - 14 m depth from the MBS-Main (b.) and MBS-Alpha (c.) cores.

2.2 Atmospheric circulation modeling

110 The Hybrid Single Particle Lagrangian Integrated Trajectory model (HYSPPLIT; Stein et al. (2015)) was used as a basis for
sample selection to investigate potential sources of volcanic glass shards identified at the MBS site. Trajectories were used
both to guide the sampling strategy and in the correlation of the cryptotephra horizons found in MBS-Alpha. We generated
six-hourly ten day back trajectories originating from 1500 m AGL at the MBS site (to avoid trajectories hitting the ground; Fig.
1), during the time periods associated with abundant tephra shards (>4 shards in one sample) to identify potential tephra depths
115 in the core. For ease of trajectory generation and repeatability, we used the PySPLIT package (Warner, 2018). NCEP/NCAR
reanalysis data (Kalnay et al., 1996) was used for the meteorology conditions. Individual trajectories were clustered using the
built-in clustering algorithm in HYSPPLIT. The aim of the algorithm is to identify a number (user-defined) of maximally distinct
clusters while minimizing variability within each cluster (Stein et al., 2015).

Lagrangian transport models like HYSPPLIT are some of the most commonly used tools for investigating particle atmospheric
120 transport but should be considered with caution. As back trajectory timelines become longer (as in the case of the 10-day
trajectories used here), the potential for error due to atmospheric noise introduced into the model increases. Despite the potential
for the influence of such noise in the model systems, we find HYSPPLIT to be useful as a confirmatory tool, used in parallel
with geochemical analysis and existing satellite observations of volcanic ash transport.

2.3 Ice core sampling

125 ~~Sample~~ For the first low-resolution screening process, sample depths were selected from MBS-Main in order to target vol-
canic material entrained in mid-latitude moisture from the southern Indian Ocean, which is the origin of significant snowfall
accumulation to MBS (Jackson et al., 2023; Udy et al., 2021; Vance et al., 2024a). This could include material from volcanoes
in the Kerguelen Plateau region (Heard and McDonald Islands) ~~with the aim to better understand volcanic fertilization over~~
~~the Kerguelen Plateau~~ or from sources further afield. In order to refine our sampling strategy, we targeted our sampling toward
130 towards air masses that passed over the Heard and ~~MaeDonald-McDonald~~ McDonald Islands region prior to arriving at MBS (based on
HYSPPLIT air parcel back trajectory modeling ~~-, described in~~ section 2.2). ~~This forms part of a broader project investigating~~
~~iron fertilization in the Southern Ocean, however, for the purposes of this study, we consider tephra layers of any origin.~~
~~Sampling was conducted from the MBS-Main and MBS-Alpha cores (Fig. 2) depth ranges covering the satellite era (down to~~
~~~20 meters).~~

135 As this is the first investigation of cryptotephra in the MBS ice cores, for this study we focus on the satellite era (~~1978-present~~ 1979-present).  
This allows us to utilize atmospheric circulation modeling to assess volcanic ash transport, and to consider satellite volcanic  
observations to inform our volcanic matching efforts. Additionally, knowledge of the magnitude and timing of eruption events  
identified in satellite era ice cores can help ~~tune fine-tune~~ future efforts to reconstruct events identified from volcanic horizons  
in the more distant past. Furthermore, from a practical perspective, the duplicate cores covering this period ~~allow~~ allows for  
140 larger sample volumes for analysis, increasing the likelihood of capturing sparse tephra in a given sample.

First, a broad low resolution screening of ~~target~~ the targeted depths was conducted from the ~~MBS-Main~~ MBS-Main core  
based on the atmospheric modelling results and using samples ranging from 15-18 cm in length (six samples per one-meter-  
long ice core segment, Fig. 2b, measured to the nearest half-centimeter), and cut with a band saw from the outer edge of the core

( $\sim 2 \text{ cm}^2$  cross sectional area). With a minimum annual layer thickness throughout the satellite era of 0.256 m firn depth (mean  
145  $0.501 \pm 0.137 \text{ m}$ , based on [the MBS2023 chronology](#) (Vance et al., 2024b)), this sample size represents at least sub-annual  
sampling resolution. ~~MBS-Main-MBS-Main~~ core samples were melted in ~~polycarbonate bottles~~ [clean, new polycarbonate  
bottles before being transferred to clean new 10 ml centrifuge tubes](#) and centrifuged (5 minutes, 3000 rpm). ~~Resulting-The~~  
[resulting](#) concentrate was evaporated on frosted glass slides and [the remaining](#) sample material was coated in low-viscosity  
epoxy resin (Logitech type 301 2-part epoxy resin). For efficiency [and because of difficulties resulting from the preparation of](#)  
150 [the slides](#), these samples were [only](#) investigated using optical microscopy ~~and~~, [and samples depths were](#) prioritized based on  
[the](#) presence or absence of potential tephra grains. ~~Out-of-Of~~ a total of 65 exploratory samples from the ~~MBS-Main-MBS-Main~~  
core, 22 samples were selected based on visual inspection using optical microscopy for re-sampling of the corresponding depth  
ranges in the MBS-Alpha core (Fig. 2a. 1-2).

~~Second, a higher resolution sampling was conducted from the MBS-Alpha core (due to larger availability of sample material~~  
155 ~~remaining~~; [As these samples were prepared from a narrow outer edge piece from the ice core, the small cross-sectional  
area meant that robust sample cleaning/decontamination procedures were not feasible for the MBS-Main samples. We were  
therefore not able to eliminate the potential for contamination by transfer from one core depth to another. General lab contamination  
debris also potentially obscured tephra and limited our ability to form robust identifications. Additionally, due to the sample  
preparation methods used for the MBS-Main samples \(uneven surface of the glass slides used\), we were not able to reliably](#)  
160 [fully section the very fine tephra grains for further microanalysis and were thus required to rely only on optical microscopy  
for these samples. This necessitated secondary sampling from the MBS-Alpha core, of which there was more sample volume  
remaining](#) (Fig. 2c). The [MBS-Main samples were valuable in guiding our sample selection for MBS-Alpha, allowing very  
efficient, targeted use of the core.](#)

[The second phase of sampling was able to be conducted at a higher resolution due to the available sample volume available](#)  
165 [from MBS-Alpha. The](#) depths from the 22 selected ~~MBS-Main-MBS-Main~~ core samples were transposed to the corresponding  
MBS-Alpha core depths (based on the annual horizons presented in Vance et al. (2024a), Fig. 2 a. 3), resulting in 70 sample  
depths ranging from 4-8 cm in length (Fig. 2a. 5). Samples were measured to the nearest half-centimeter and cut using a thin  
bladed pull-saw. [Careful decontamination procedures were followed during the cutting and preparation of the MBS-Alpha ice  
samples to prevent transfer of sample material from one sample to the next. MBS-Alpha samples were cut and decontaminated](#)  
170 [under a laminar flow hood in the cold lab. The cutting surface and blades were cleaned between each individual  \$\sim 5 \text{ cm}\$  sample,  
and both the ceramic blade used and the cutting surface \(a clean, new plastic cutting board\) were changed and washed with  
ultrapure water between samples from each meter of core \(e.g. a freshly washed blade and cutting board were used for samples  
from MBS-Alpha 14, and then washed before decontamination of samples from MBS-Alpha 15\).](#)

After  $\sim 1 \text{ mm}$  of material from each of the outer edges was removed using a ceramic blade, the samples had a cross sectional  
175 area of  $\sim 10 \text{ cm}^2$ . Samples were melted at ambient temperature in rinsed sterile Whirl-Pak bags ~~and~~. [Melted sample material  
was transferred to acid-washed 15 ml centrifuge tubes, and Whirl-Pak bags were rinsed at least twice with ultrapure water to  
minimize the possibility of sample material being left behind in the bags. Samples were then](#) centrifuged (5 minutes, 3000  
rpm) ~~in acid-washed 15 ml centrifuge tubes. Resulting~~ [and the resulting](#) concentrate was pipetted ~~onto~~ [into wells created by](#)

180 placing cleaned 25 mm acrylic rings on polyimide (Kapton) tape adhered to flat glass plates, and set to evaporate on a hot plate at 60 °C ~~and backfilled within 25 mm acrylic rings~~. The centrifuge tubes were subsequently rinsed at least twice with ultrapure water, and the rinse water was added to the same sample area to evaporate. Each sample well was then backfilled with low-viscosity epoxy (Struers EpoFix) to create round resin mounts.

The sample mount surfaces were polished using 1 ~~mieron-~~μm aluminum oxide polishing compound to remove surface resin and expose any tephra grains present in the mounts, then cleaned in an ultrasonic bath to remove excess polishing compound. 185 Samples were examined using transmitted and reflected light using a petrographic microscope, and mounts identified as containing tephra grains of suitable size and quality for future ~~EPMA (electron probe microanalysis (EPMA); e.g. with large enough surface area for the EPMA beam accounting for mineral inclusions and/or vesicles; larger than ~5 μm on the shortest axis.~~ Fig. 2a. 6) were carbon-coated in preparation for ~~backscatter electron (BSE-) imaging, BSE imaging, scanning electron microscopy~~ by energy dispersive spectroscopy (SEM-EDS, FEI MLA 650 ESEM), and electron probe microanalysis (EPMA; EPMA (2a. 190 7-9) at the Central Science Laboratory at the University of Tasmania.

## 2.4 Sub-annual age determination

~~The MBS Main and surface cores have been dated using independent layer counting and volcanic matching (Vance et al., 2024a). The MBS2023 chronology provides depths for annual horizons for the main and surface cores (Vance et al., 2024b). As the focus of this study is on satellite-era tephrochronology, and the~~ As the annual layer thickness of the MBS cores is suitable for sub-annual sampling, we are able to provide much higher resolution samples, when compared to many longer-term tephrochronologies of Antarctic and Greenlandic deep ice cores, which are often not able to produce annual resolution sampling (e.g. Cook et al. (2022); Narcisi et al. (2005, 2012)). 195

Due to the high resolution of our samples, it is helpful for source identification to be able to provide sub-annual dates for our sample depths, however a sub-annual chronology has not been published for the MBS cores. ~~MBS is characterized by variability in accumulation (both between years and within years; Vance et al. (2024a); Crockart et al. (2021); Jackson et al. (2023)) associated with extreme precipitation events at the site (Jackson et al., 2023). This variability limits the accuracy of a Vance et al. (2024a) characterize the seasonality of impurity flux to the MBS site, used here to infer sub-annual age scale derived by linear interpolation between annual horizons. Assessment of sub-annual dating for particular tephra sample depths relies on subjective interpretation of the relative position of the sample depths in comparison with seasonally-varying chemical species measured in the ice core and/or water isotope measurements (Vance et al., 2024b; Moy et al., 2024). In an effort to ages.~~ 200

To estimate sub-annual ages for the samples in this study, we use rely on variations in austral summer peaking sulfate-on-chloride sulfate-to-chloride ratio ( $\text{SO}_4^{2-}/\text{Cl}^-$ ) and  $\delta^{18}\text{O}$  (Vance et al., 2024b; Moy et al., 2024) and austral winter peaking sodium ( $\text{Na}^+$ ) (Vance et al., 2024b) to approximate the time of year to which the samples may correspond. However, MBS is characterized by variability in accumulation (both between years and within years; Vance et al. (2024a); Crockart et al. (2021); Jackson et al. (2023) ) associated with extreme precipitation events at the site (Jackson et al., 2023). This variability limits the accuracy of a sub-annual age scale derived by linear interpolation, and therefore absolute sub-annual dating should be considered with caution. 210

## 2.5 Geochemical analysis of tephra

The mounted volcanic glass shards were analyzed using EPMA at the Central Science Laboratory (CSL) at the University of Tasmania (Fig. 2) in 2023 and early 2024. Analyses were conducted on a JEOL JXA-8530F Plus field emission microprobe with 5-five wavelength dispersive spectrometers. Single point analyses used a 2  $\mu\text{m}$  beam diameter with 4 nA beam current, 15 kV accelerating voltage. Analytical conditions were chosen to balance minimizing beam damage, particularly Na ion migration, while maintaining a beam size suitable for our smallest ( $\leq 5 \mu\text{m}$ ) shards. ~~Concentrations of 14~~

~~EPMA analytical conditions were selected in collaboration with the instrument scientists, with the aim of obtaining as robust analytical totals as possible on our small glass shards while minimizing alkali ion migration and maintaining consistent analytical conditions across all measurements. While larger a beam size could have been used for a few of the larger samples analyzed using the broad-beam overlap method (Iverson et al., 2017), we chose to prioritize consistent analytical techniques, using the same beam size for all samples analyzed in order to enable better comparison across samples and obtain results from the largest number of glass shards possible in our samples.~~

~~Concentrations of 12 major and minor elements-element oxides~~ were measured in the individual shards ( $\text{SiO}_2$ ,  $\text{TiO}_2$ ,  $\text{Al}_2\text{O}_3$ ,  $\text{FeO}$ ,  $\text{MnO}$ ,  $\text{MgO}$ ,  $\text{CaO}$ ,  $\text{Na}_2\text{O}$ ,  $\text{K}_2\text{O}$ ,  $\text{P}_2\text{O}_5$ ,  $\text{SO}_3$ , F). A series of mineral standards were used for instrument calibration, a rhyolitic glass secondary standard (VG-586 reference standard from the Smithsonian Institution Department of Mineral Sciences (VG-568, NMNH 72854; Jarosewich et al. (1980)) was used to validate accuracy and precision of measurements between measurement sessions. See Supplementary information for full details on the EPMA measurements and standards.

~~Data-The dataset~~ was detection limit filtered (99% confidence), and erroneous analyses of non-volcanic material (e.g. quartz grains or other mineral particles in the samples) were removed from the dataset. ~~The majority of the analyses have analytical total oxide values of >90%, however due to the very small size of the shards (<5m), some analyses resulted in low analytical totals. Iverson et al. (2017) show with their "broad-beam overlap" method that analytical totals as low as 67% are statistically similar to the same analyses with higher analytical totals, with only a slight decrease in precision. Narcisi et al. (2019) show reliable results with analytical totals as low as 60%. Following this approach, analyses with analytical totals of at least 50% are presented here, however the few data with totals below 60% are considered with some caution. Data was~~ Data were corrected for Cl concentrations and normalized to 100% anhydrous based-on-the as recommended by Iverson et al. (2017) prior to interpretation. Full geochemical results and secondary standards measurements can be found in the supplementary materials.

## 2.6 Atmospheric circulation modeling

~~The Hybrid-Single-Particle-Lagrangian-Integrated-Trajectory model (HYSPLIT; Stein et al. (2015)) was used to investigate potential sources of volcanic glass shards identified at the MBS site. We generated six-hourly ten-day back trajectories originating from 1500 m AGL at the MBS site (Fig. 1), during the time periods associated with abundant tephra shards (>4 shards in one sample). For ease of trajectory generation and repeatability, we used the PySPLIT package (Warner, 2018). NCEP/NCAR reanalysis data (Kalnay et al., 1996) was used for the meteorology conditions. Individual trajectories were~~

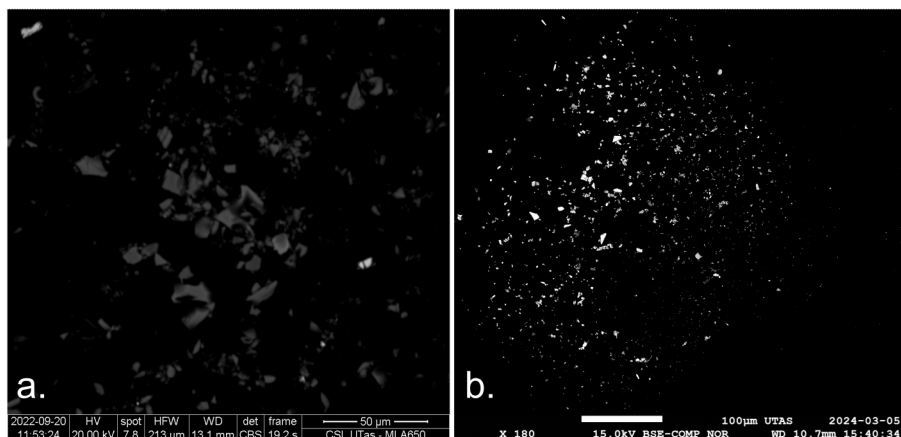
245 clustered using the built-in clustering algorithm in HYSPLIT. The aim of the algorithm is to identify a number (user-defined) of maximally distinct clusters while minimizing variability within each cluster (Stein et al., 2015).

### 3 Results

**Table 1.** Summary of [the 12 MBS-Alpha ice core](#) samples containing volcanic glass shards [analyzed by EPMA-WDS](#). Sample depths were measured to the nearest half-centimeter from the top of each ice core segment during sampling in the freezer laboratory and assimilated with core top depths; measurement error associated with core length estimates are in line with information provided in Vance et al. (2024a). Ages provided here are approximated based on MBS2023 chronology (Vance et al., 2024a). Composition classifications are based on total [alkalis vs silica-alkali-silica](#) (TAS) diagram (Le Bas et al., 1986), the number of shards of each composition is given in parentheses. Further discussion of proposed volcanic sources (bold text in table) can be found in section 4.

| Sample ID   | Sample depth range    | Approx. date    | Total shard count | Composition ( <i>Le Bas et al., 1986</i> )                        | Proposed source     |
|-------------|-----------------------|-----------------|-------------------|-------------------------------------------------------------------|---------------------|
| 7-6         | 6.39-6.45 m           | Mid 2004        | 3                 | Phonolite (1), Dacite (2)                                         | -                   |
| 7-7         | 6.45-6.51 m           | Mid 2004        | 1                 | Rhyolite (1)                                                      | -                   |
| 8-5         | 7.98-8.05 m           | Mid 2001        | 3                 | Rhyolite (43)                                                     | -                   |
| 9-10        | 8.94-9.00 m           | Early 2000      | 3                 | Rhyolite (1), Dacite (1),<br>Basaltic Andesite (1)                | -                   |
| 13-3        | 12.79-12.85 m         | Mid 1992        | 1                 | Rhyolite (1)                                                      | -                   |
| <b>14-1</b> | <b>13.28-13.34 m</b>  | <b>Mid 1991</b> | <b>13</b>         | <b>Dacite (12), Andesite (1)</b>                                  | <b>Cerro Hudson</b> |
| 16-4        | 15.93-16.00 m         | Early 1987      | 2                 | Rhyolite (2)                                                      | -                   |
| 17-1        | 16.00-16.045 m        | Early 1987      | 1                 | Rhyolite (21)                                                     | -                   |
| 17-5        | 16.20-16.25 m         | Mid 1986        | 1                 | Trachyte/Trachydacite (1)                                         | -                   |
| <b>17-9</b> | <b>16.87-16.915 m</b> | <b>Mid 1985</b> | <b>10</b>         | <b>Phonolite (10)</b>                                             | <b>Mt. Erebus</b>   |
| 18-1        | 17.00-17.05 m         | Mid 1985        | 1                 | Trachyandesite (1)                                                | -                   |
| 18-5        | 17.24-17.30 m         | Early 1985      | 3                 | Trachyandesite (1),<br>Trachyte/Trachydacite (1),<br>Rhyolite (1) | -                   |

250 [Glass-Potential glass](#) shards were identified ~~using a petrographic microscope~~ in 48 out of the 70 samples prepared from the MBS-Alpha core [using a petrographic microscope](#). Those 48 samples were inspected using ~~backscatter electron-BSE~~ imaging and SEM-EDS with automated mineralogy [at 5  \$\mu\$ m resolution](#), and 29 were [identified as containing glass shards](#) selected for further analysis using EPMA. [All samples contained relatively abundant material of non-volcanic origin, including mineral grains, diatom fragments, and infrequently, identifiable laboratory contamination \(glove fibers, dust, etc.\)](#) The samples [not selected for further analysis either contained mineral grains previously mis-identified \(e.g. quartz\) or grains too small or positioned too deep in the resin to have been appropriately sectioned in polishing.](#) These 29 samples were analyzed with



**Figure 3.** Example backscatter electron images of sample material found in the MBS-Alpha core in the 13.28-13.34 m (a) and 16.87-16.915 m (b) samples.

255 EPMA, producing reliable results for tephra grains in 12 of those 29 samples contained glass shards (Fig. 2). The 17 discarded samples contained grains with complex morphology (no surface large enough for reliable geochemical measurements using EPMA (Fig. 2) probe positioning), grains that had not been fully sectioned due to their size, or shards that were thin enough that the interaction volume of the probe beam included resin below the shard. A summary of all glass shards identified-volcanic glass shards geochemically characterized in the MBS-Alpha core is reported in Table 1 and further discussion of analytical limitations is provided in Section 3.3.)

### 260 3.1 Tephra morphology

Glass shards are most abundant in the 13.28-13.34 and 16.87-16.915 m depth samples. The 13.28-13.34 m sample has the largest glass shards (Fig. 3), ranging from 10-20 microns-μm along the longest axis. All shards are angular, while the larger shards in this sample have cusped, bubble-wall, or y-junction morphologies and smaller shards appear more blocky or platy. Glass shards in the 16.87-16.915 m sample are typically smaller, <10 microns, and-μm, with only one shard <10 μm (~15  
 265 microns-μm) across the longest axis. These shards have simpler angular morphology, and include blocky and platy shapes, some have fluted or cusped edges. The remaining samples contain glass shards ranging from <5 to 15 microns μm, with mostly simple angular morphologies.

### 3.2 Geochemical composition

Major element oxides of all tephra identified in the MBS-Alpha core are provided in Table 2. The MBS-Alpha glass shards have  
 270 SiO<sub>2</sub> values ranging from 51.70 to 77.31 wt.%, with 1.38-9.12 wt.% Na<sub>2</sub>O, and 0.28-6.0 wt.% K<sub>2</sub>O (Fig. 4, Table 2). When plotted on the total alkali-silica diagram (TAS diagram; Le Bas et al. (1986)) TAS diagram (Le Bas et al., 1986), compositions



**Table 2.** Major element oxide concentrations of volcanic glass shards identified in the MBS-Alpha core. Colored symbols correspond with the legend in Figure 4. Analyses were conducted on a JEOL JXA-8530F Plus field emission microprobe with five WDS spectrometers. Single point analyses used a 2  $\mu\text{m}$  beam diameter with 4 nA beam current, 15 kV accelerating voltage. Data are normalized to 100% using the Cl-adjusted analytical total, following the "broad beam overlap" method introduced in Iverson et al. (2017). "n.d." indicates measured values below instrument detection limit. For shards with multiple point measurements, average of measurements is indicated by "a" in sample name, and  $n$  indicates number of points measured per shard.

| Sample    |   | SiO <sub>2</sub> | TiO <sub>2</sub> | Al <sub>2</sub> O <sub>3</sub> | FeO   | MnO   | MgO   | CaO   | Na <sub>2</sub> O | K <sub>2</sub> O | P <sub>2</sub> O <sub>5</sub> | Total  | $n$      |
|-----------|---|------------------|------------------|--------------------------------|-------|-------|-------|-------|-------------------|------------------|-------------------------------|--------|----------|
|           |   | wt. %            | wt. %            | wt. %                          | wt. % | wt. % | wt. % | wt. % | wt. %             | wt. %            | wt. %                         | %      |          |
| 7-6_002*  | ✕ | 66.16            | 0.09             | 20.07                          | 1.00  | n.d.  | 0.14  | 4.94  | 6.53              | 0.69             | 0.11                          | 100.00 | <u>1</u> |
| 7-6_013   | ✕ | 65.13            | 1.63             | 14.78                          | 8.61  | n.d.  | 0.47  | 2.55  | 5.44              | 0.98             | 0.32                          | 100.00 | <u>1</u> |
| 7-6_014   | ✕ | 59.08            | 0.99             | 19.51                          | 4.27  | 0.23  | 0.81  | 1.34  | 7.61              | 5.50             | 0.29                          | 100.00 | <u>1</u> |
| 7-7_006   | + | 77.16            | 0.30             | 13.11                          | 0.22  | n.d.  | 0.17  | 0.94  | 2.06              | 6.00             | 0.00                          | 100.00 | <u>1</u> |
| 8.5_005   | ● | 77.31            | 0.13             | 13.11                          | 0.40  | 0.06  | 0.07  | 0.51  | 3.96              | 4.49             | 0.03                          | 100.00 | <u>1</u> |
| 8.5_007*  | ● | 77.10            | 0.04             | 13.18                          | 0.43  | n.d.  | 0.00  | 0.48  | 4.36              | 4.36             | 0.05                          | 100.00 | <u>1</u> |
| 8.5_010   | ● | 73.97            | 0.34             | 14.60                          | 1.43  | n.d.  | 0.43  | 1.53  | 3.22              | 4.51             | 0.06                          | 100.00 | <u>1</u> |
| 9-10_014  | ▲ | 56.70            | 1.12             | 21.70                          | 7.56  | n.d.  | 2.20  | 5.13  | 2.58              | 2.53             | 0.43                          | 100.00 | <u>1</u> |
| 9-10_017  | ▲ | 75.26            | 0.08             | 15.38                          | 0.47  | n.d.  | 0.40  | 0.19  | 6.71              | 1.47             | 0.00                          | 100.00 | <u>1</u> |
| 9-10_021  | ▲ | 67.32            | 0.39             | 12.13                          | 8.00  | 0.17  | 6.83  | 1.47  | 1.44              | 2.08             | 0.15                          | 100.00 | <u>1</u> |
| 13-3_011  | ▼ | 69.48            | 0.59             | 15.77                          | 2.65  | n.d.  | 0.85  | 2.64  | 3.43              | 4.29             | 0.13                          | 100.00 | <u>1</u> |
| 14-1_003a | ★ | 65.80            | 1.27             | 17.16                          | 5.05  | 0.19  | 1.53  | 3.24  | 2.73              | 2.71             | 0.31                          | 100.00 | <u>2</u> |
| 14-1_006  | ★ | 61.87            | 1.38             | 17.31                          | 6.19  | 0.14  | 2.44  | 4.48  | 3.42              | 1.99             | 0.61                          | 100.00 | <u>1</u> |
| 14-1_010a | ★ | 64.91            | 1.29             | 16.59                          | 5.45  | 0.17  | 1.82  | 3.30  | 3.46              | 2.60             | 0.35                          | 100.00 | <u>3</u> |
| 14-1_011a | ★ | 65.35            | 1.31             | 17.10                          | 5.38  | 0.20  | 1.47  | 3.32  | 2.90              | 2.58             | 0.38                          | 100.00 | <u>2</u> |
| 14-1_013a | ★ | 67.79            | 1.12             | 17.27                          | 4.67  | 0.16  | 1.39  | 2.77  | 1.74              | 2.83             | 0.32                          | 100.00 | <u>2</u> |
| 14-1_015  | ★ | 65.23            | 1.25             | 16.73                          | 5.15  | 0.15  | 1.59  | 3.25  | 3.67              | 2.57             | 0.36                          | 100.00 | <u>1</u> |
| 14-1_016a | ★ | 66.05            | 1.17             | 17.16                          | 5.47  | 0.18  | 1.47  | 3.39  | 2.26              | 2.55             | 0.36                          | 100.00 | <u>2</u> |
| 14-1_017a | ★ | 63.43            | 1.60             | 16.77                          | 6.52  | 0.21  | 2.34  | 4.67  | 1.38              | 2.33             | 0.61                          | 100.00 | <u>2</u> |
| 14-1_019  | ★ | 65.57            | 1.19             | 17.00                          | 5.26  | 0.20  | 1.58  | 3.23  | 2.61              | 2.76             | 0.42                          | 100.00 | <u>1</u> |
| 14-1_021a | ★ | 66.23            | 1.31             | 16.72                          | 5.21  | 0.19  | 1.65  | 3.45  | 2.28              | 2.63             | 0.40                          | 100.00 | <u>2</u> |
| 14-1_022a | ★ | 65.15            | 1.30             | 16.86                          | 5.28  | 0.21  | 1.70  | 3.38  | 2.92              | 2.66             | 0.50                          | 100.00 | <u>2</u> |
| 14-1_024a | ★ | 63.76            | 1.32             | 16.69                          | 5.52  | 0.21  | 1.69  | 3.41  | 4.28              | 2.60             | 0.49                          | 100.00 | <u>2</u> |
| 14-1_025  | ★ | 64.48            | 1.20             | 16.75                          | 5.60  | 0.19  | 1.59  | 3.34  | 3.82              | 2.56             | 0.34                          | 100.00 | <u>1</u> |
| 16-4_014  | ▶ | 76.90            | 0.06             | 13.77                          | 0.42  | 0.10  | 0.07  | 0.48  | 3.75              | 4.36             | n.d.                          | 100.00 | <u>1</u> |
| 16-4_024  | ▶ | 73.11            | 0.03             | 16.05                          | 2.50  | 0.04  | 0.02  | 0.68  | 3.22              | 4.34             | n.d.                          | 100.00 | <u>1</u> |
| 17-1_014  | ◀ | 75.83            | 0.13             | 18.81                          | 0.72  | n.d.  | 0.33  | 0.67  | 3.06              | 0.28             | 0.02                          | 100.00 | <u>1</u> |

continued on following page

continued from previous page

| Sample    |    | SiO <sub>2</sub> | TiO <sub>2</sub> | Al <sub>2</sub> O <sub>3</sub> | FeO   | MnO   | MgO   | CaO   | Na <sub>2</sub> O | K <sub>2</sub> O | P <sub>2</sub> O <sub>5</sub> | Total  | n |
|-----------|----|------------------|------------------|--------------------------------|-------|-------|-------|-------|-------------------|------------------|-------------------------------|--------|---|
|           |    | wt. %            | wt. %            | wt. %                          | wt. % | wt. % | wt. % | wt. % | wt. %             | wt. %            | wt. %                         | %      |   |
| 17-5_081* | ■  | 61.27            | 0.14             | 25.36                          | 1.26  | n.d.  | n.d.  | 1.78  | 6.25              | 3.92             | 0.00                          | 100.00 | 1 |
| 17-9_005  | ★  | 57.02            | 1.01             | 20.14                          | 5.04  | 0.27  | 0.99  | 1.91  | 7.88              | 5.22             | 0.26                          | 100.00 | 1 |
| 17-9_008  | ★  | 58.76            | 0.97             | 19.97                          | 3.99  | 0.13  | 0.80  | 1.00  | 8.87              | 5.11             | 0.29                          | 100.00 | 1 |
| 17-9_009* | ★  | 57.10            | 0.98             | 19.80                          | 5.81  | 0.20  | 0.97  | 1.80  | 7.04              | 5.76             | 0.31                          | 100.00 | 1 |
| 17-9_010  | ★  | 57.32            | 1.03             | 19.68                          | 5.30  | 0.29  | 0.86  | 1.92  | 7.69              | 5.49             | 0.32                          | 100.00 | 1 |
| 17-9_011  | ★  | 56.48            | 0.98             | 20.16                          | 4.73  | 0.31  | 0.93  | 1.76  | 9.12              | 5.11             | 0.25                          | 100.00 | 1 |
| 17-9_013  | ★  | 56.44            | 1.09             | 20.06                          | 5.64  | 0.33  | 0.92  | 1.85  | 8.10              | 5.19             | 0.29                          | 100.00 | 1 |
| 17-9_018  | ★  | 57.38            | 0.90             | 20.12                          | 5.04  | 0.20  | 0.96  | 1.68  | 8.56              | 4.87             | 0.18                          | 100.00 | 1 |
| 17-9_019  | ★  | 56.31            | 1.06             | 20.01                          | 5.18  | 0.25  | 0.87  | 1.95  | 8.46              | 5.32             | 0.28                          | 100.00 | 1 |
| 17-9_020  | ★  | 57.64            | 0.93             | 19.94                          | 4.83  | 0.25  | 0.84  | 1.67  | 8.73              | 4.78             | 0.22                          | 100.00 | 1 |
| 17-9_032  | ★  | 56.86            | 1.04             | 20.09                          | 5.18  | 0.31  | 0.86  | 1.76  | 8.24              | 5.25             | 0.23                          | 100.00 | 1 |
| 18-1_014  | ⊕⊙ | 51.70            | 2.94             | 14.15                          | 11.16 | 0.16  | 3.09  | 7.72  | 4.38              | 3.72             | 0.94                          | 100.00 | 1 |
| 18-5_004  | ◆  | 65.13            | 0.12             | 20.50                          | 0.22  | n.d.  | 0.00  | 1.82  | 8.32              | 3.97             | n.d.                          | 100.00 | 1 |
| 18-5_013  | ◆  | 72.81            | 0.39             | 15.29                          | 1.76  | 0.17  | 0.38  | 1.10  | 4.08              | 4.12             | n.d.                          | 100.00 | 1 |
| 18-5_020* | ◆  | 53.09            | 1.03             | 18.21                          | 13.84 | 2.37  | 0.54  | 4.28  | 5.06              | 0.55             | 0.86                          | 100.00 | 1 |

"a" indicates values averaged over multiple analyses on a single glass shard; \* indicates analytical totals below 60%.

### 295 3.3 Analytical limitations

EPMA analytical conditions were selected and analyses completed prior to the publication of the recommendations for best practices for EPMA analysis of fine grained tephra (Innes et al., 2024). Our analytical conditions produced meaningful results, however some limiting aspects require consideration.

300 Many of the analyses have analytical total oxide values of >90%, due to the very small size of the shards (<5µm). However, approximately half of the sample analyses resulted in lower analytical totals. Of the 42 EPMA data measurements presented here, five have analytical totals below 60% (indicated in Table 1). We attribute low analytical totals to the following scenarios (1) glass shards that are very thin (either naturally, or due to the amount of polishing required), resulting in an interaction volume depth that includes the resin substrate below the shards and/or (2) larger shards with complex morphology causing the resin to be not uniformly removed from the surface of the shards, resulting in some resin being included in the beam area. This  
305 could have been resolved by additional polishing, however this risked fully polishing away some of the smaller shards in the sample. As a result, we chose to accept slightly lower totals in order to include analyses of as many shards as possible.

Iverson et al. (2017) show with their "broad beam overlap" method that analytical totals as low as 67% are statistically similar to the same analyses with higher analytical totals, with only a slight decrease in precision. Narcisi et al. (2019) show reliable results with analytical totals as low as 60%. Innes et al. (2024) report good accuracy despite somewhat lower precision

310 for EPMA analyses of very fine grained tephra shards resulting in analytical totals ranging from as low as 35% up to 101 wt.%. Following this approach, analyses with analytical totals of at least 50% are presented here. Only one sample with <60% total (shard 17-9\_009) is included in the horizons that we pursue source correlations for. As the major element compositions of this shard fall within the range of the other glass shards in this sample, we have chosen to include it in our analyses.

315 The probe beam used in EPMA has the potential to damage the sample by inducing migration of alkali ions (typically Na<sub>2</sub>O; Gedeon et al. (2000); Humphreys et al. (2006); Kuehn et al. (2011)). Analytical conditions for EPMA were selected to minimize potential migration of alkali ions for small glass shards and repeated analysis of secondary standards did not show a loss of sodium over time (see supplementary materials). This loss of sodium is due to probe beam damage in cases of repeated analysis in a small area or single glass shard. While most of the glass shards in our sample were too small to fit multiple analyses, some of the glass shards in the 13.28-13.34 m sample were large enough for multiple analyses. Data from these analyses are presented in Table 2 as the average of multiple analyses (see supplement for individual point measurements).  
320 Some of these glass shards did show decreases in Na<sub>2</sub>O over multiple analyses (see Figure S.1 in supplementary materials) and consequently we have not relied exclusively on Na<sub>2</sub>O values for our tephra source identification.

### 3.4 Sample age estimates

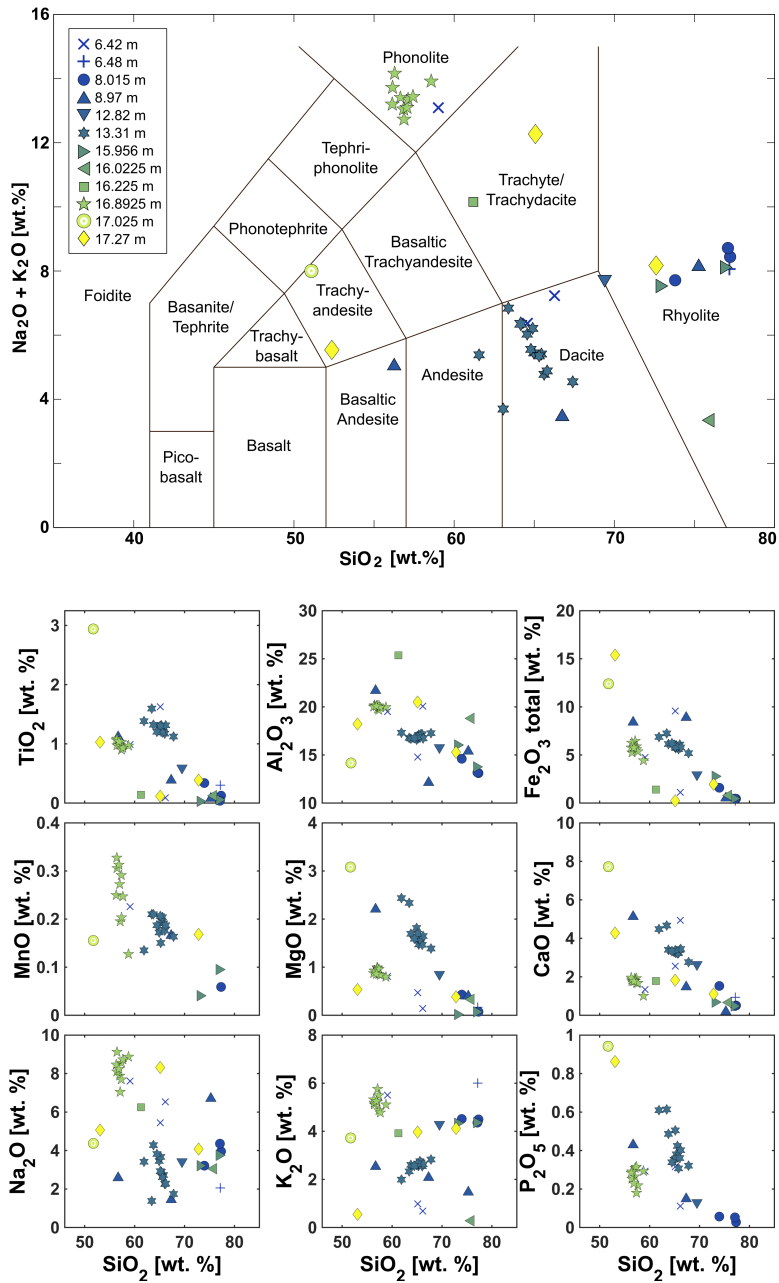
325 The 13.28-13.34 m sample falls between the 1991 and 1992 annual horizons (13.761 and 13.099 m respectively, MBS2023; Vance et al. (2024a)). In Figure 5, sample depth ranges are shown alongside seasonally varying isotope and chemistry species measured in the MBS-Alpha core. It is notable that during the development of the MBS2023 chronology, it was determined that trace chemistry was as reliable as  $\delta^{18}O$ , and at times more so, in determining annual horizon positions in the MBS records (see section 4.3 in Vance et al. (2024a)). It can be seen in Figure 5 that the 13.28-13.34 m sample falls towards the end of the austral winter sodium peak, but just at the start of the austral summer peak in sulfate-on-chloride, indicating that the glass shards were likely deposited with snow that fell in mid to late 1991. It can also be seen that the sample falls just at the beginning of a significant peak in non-sea-salt sulfate, likely volcanic in origin.  
330

The 16.87-16.915 m sample falls between the 1985 and 1986 annual horizons (17.282 and 16.679 m respectively). This sample depth coincides with the end of the austral winter sodium peak and the start of the the sulfate-on-chloride peak, indicating that this sample likely represents snow from the middle of the year (late austral winter to early spring) 1985.

335 For clarity, we will hereafter refer to these two samples by their year of origin, according to the MBS2023 chronology; 1991 for the 13.28-13.34 m sample, and 1985 for the 16.87-16.915 m sample.

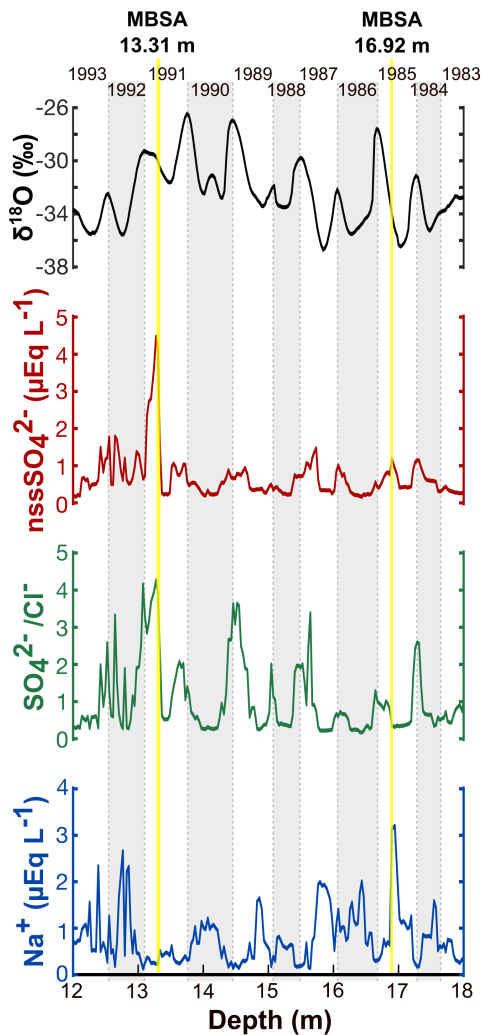
## 4 Volcanic source identification

The majority of tephra-bearing sample depth ranges from MBS-Alpha contain sparse and/or heterogeneous tephra shard compositions. However, when plotted on the total alkalis vs. silica diagram (TAS; Le Bas et al. (1986)) TAS diagram (Le Bas et al., 1986)  
340 , the 1991 and 1985 samples contain abundant glass shards (13 and 10 shards respectively), each with largely homogeneous compositions (Fig. 4). We therefore consider these two samples as discrete cryptotephra horizons in the MBS-Alpha core. The



**Figure 4.** Total alkalis vs silica alkali-silica (TAS, Le Bas et al. (1986)) diagram and variation diagrams for all individual glass shards identified in the MBS-Alpha core. Data are normalized to 100% (anhydrous) using the Cl-adjusted analytical total, after Iverson et al. (2017).

remaining samples each contain three or fewer glass shards. Some of these samples do cluster together across sample depths (e.g. cluster of rhyolitic deposits, Fig. 4), however as these shards comprise a wide range of ice core depths, they cannot be



**Figure 5.** Selected seasonally varying glaciochemical species (non-seasalt sulfate ( $\text{nssSO}_4^{2-}$ ), calculated using the measured  $\text{Na}^+$  to remove the sea salt sulfate component, following Plummer et al. (2012), ratio of sulfate to chloride ( $\text{SO}_4^{2-}/\text{Cl}^-$ ), and sea salts-sodium ( $\text{Na}^+$ )) and isotope ( $\delta^{18}\text{O}$ ) measured from the MBS-Alpha core with tephra horizon depths indicated (yellow bars). Annual horizons (MBS2023 January 1 year boundaries, after Vance et al. (2024a)) indicated by vertical dashed lines.

correlated as the product of a single eruption event. As none of the remaining samples (aside from the 13.28-13.34 m and  
 345 16.87-16.915 m horizons) include more than three shards, for this study, we primarily focus our source attribution efforts on  
 groups with more than three shards of similar composition to ensure reliable correlations (similar to the approach taken by  
 Cook et al. (2022)).

## 4.1 Rhyolite group

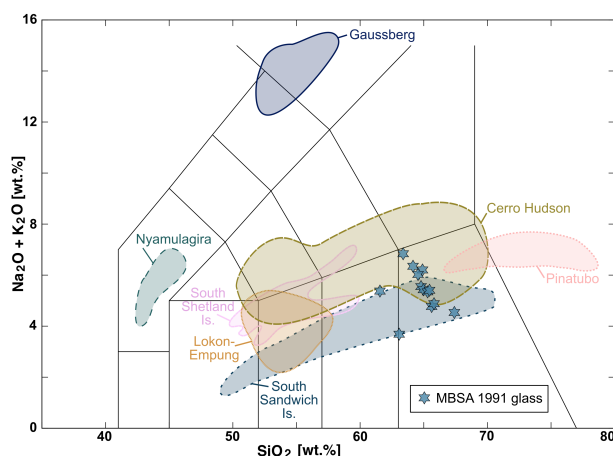
A ~~handful~~ small number of shards originating from a range of sample depths (from 6.48 to 17.27 m depth), form a compositionally similar group of rhyolites. Despite originating from a broad range of depths (and therefore ages), if treated as a group, the composition of these shards cluster together across most major element oxides (Fig. 4). The composition of this cluster is similar to compositions determined for volcanic glass shards produced in the 1991 eruption of Pinatubo (Luhr and Melson, 1996; Tamura and Nakagawa, 2023). However none of the depths where glass shards of this composition were found correspond within reasonable dating error to the 1991 eruption of Pinatubo, where a sulfate peak is often used as a dating horizon for 1991/1992 in other ice cores (Plummer et al., 2012; Crockett et al., 2021; Vance et al., 2024a).

The rhyolitic glass shard compositions fall within the range of compositions of sparse mid- to late-Holocene rhyolite shards reported in the B53 and B54 ice cores, which Abbott et al. (2024) correlate with shards identified in the Talos Dome ice core. Narcisi et al. (2012) attribute these Talos Dome rhyolite shards to various extra-Antarctic sources (including Andean and New Zealand sources). Del Carlo et al. (2018) have correlated similar high-K rhyolites with either "extremely evolved" products of West Antarctic intraplate volcanoes (including McMurdo group and Marie Byrd Land volcanoes, not active during the satellite era), but also show similarities to Antarctic samples correlated with South American (Aguilera) as well as South Shetland Islands volcanoes. Additionally, these rhyolites also show some compositional similarity to the rhyolitic component of a sample of ash collected from the crater rim of Mt. Erebus in 2000 (Silaev et al., 2020). Due to the multiplicity of sources attributed to tephra with similar compositions, as well as their sparseness in the MBS-Alpha core, we cannot confidently attribute a single source to this compositional group of rhyolites.

An alternative interpretation of such sparse and heterogeneous samples is to suppose they represent a fraction of the background "dust" signal often seen in ice core analyses (Plunkett et al., 2020; Hutchison et al., 2024; Vallelonga and Svensson, 2014; Delmonte et al., 2013). This background signal could include re-mobilized volcanic ash from a variety of sources, transported from nearby ice-free areas or other extra-Antarctic sources (Delmonte et al., 2013). Typically volcanic glass shards transported by wind display evidence of reworking e.g. abraded margins (Dunbar et al., 2003). ~~Although textural~~ Textural alteration depends on the duration and conditions of erosion and weathering glass shards are exposed to. ~~While~~ all glass shards ~~in this group chosen for analysis~~ have angular shard-like morphologies that we interpret to be primary textures (see supplementary materials figure S.2 for an example BSE image of the 8.015 m sample), we cannot rule out the possibility that these glass shards come from re-mobilized material.

While robust contamination preventative precautions were taken during sampling, there is some possibility that some of these unidentified rhyolites (as well as the other unidentified sparse glass shards) could have been the result of contamination from unknown lab sources or between samples. We cannot definitively rule out this as a possibility, however we consider it unlikely, as the sparse shards have been characterized as tephra both based on visual inspection and geochemical composition, and all sample-contact surfaces were rigorously cleaned between preparation of each sample.

The presence of a substantial spike in ~~non-sea-salt-sulfate (likely volcanic)~~  $\text{nssSO}_4^{2-}$  (Fig. 5, Vance et al. (2024b)), likely of volcanic origin, co-occurring with the 1991 sample bearing multiple glass shards suggests that the tephra found could be associated with the volcanic event that produced the sulfate signal in the ice core. Based on ~~our understanding of~~ the age of the sample, we consider the two primary candidates for the source of the 1991 tephra horizon to be the eruptions of Pinatubo (Philippines, ~~VEI~~ volcanic explosivity index (VEI) 6) and Cerro Hudson (Chile, VEI 5), both occurring in mid-1991 (Global Volcanism Program, 2024). However we also investigate all other significant (VEI 3 or greater) known tropical and Southern Hemisphere eruptions from 1991 (Nyamulagira, Democratic Republic of Congo and Lokon-Empung, Indonesia). Additionally, we include other Antarctic and sub-Antarctic volcanic sources (Gaussberg, South Sandwich Islands, and South Shetland Islands), as although there are no known eruptions of these volcanoes during 1991, they are known or speculated possible sources of tephra (primary or reworked) identified in other Antarctic ice cores (Geyer et al., 2023; Narcisi and Petit, 2021).



**Figure 6.** Total ~~alkali vs silica~~ alkali-silica diagram (Le Bas et al., 1986) showing the MBS-Alpha 1991 glass together with representative data for possible eruption sources: Holocene products of Cerro Hudson (Haberle and Lumley, 1998), Nyamulagira (Aoki et al., 1985), Lokon-Empung (Dmitrieva et al., 2023), Gaussberg (Salvioli-Mariani et al., 2004), Pinatubo (Luhr and Melson, 1996; Tamura and Nakagawa, 2023), South Sandwich and South Shetland Islands (Narcisi et al. (2005), and references therein).

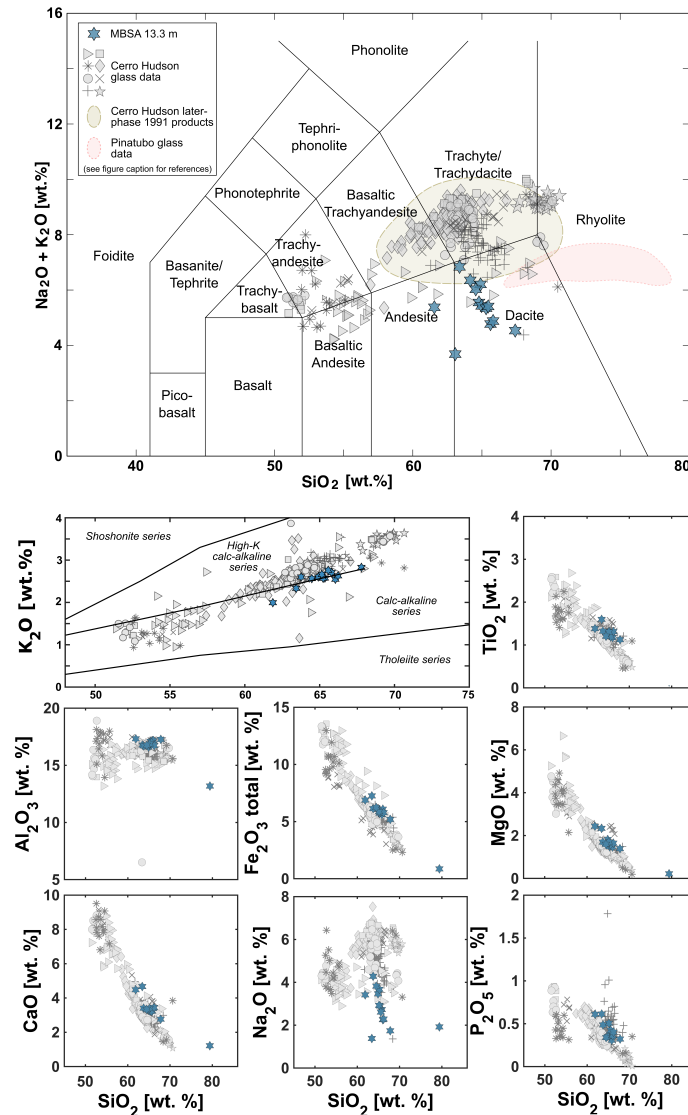
It is clear from the TAS diagram (Fig. 6) that products of Nyamulagira and Lokon-Empung do not correlate with glass from the MBS-Alpha 1991 horizon, and neither do the products of Gaussberg or the South Shetland Islands. We therefore rule these out as potential sources. While the South Sandwich Islands do appear to correlate somewhat better on the TAS diagram, when compared with literature values (Fig. 6; Abbott et al. (2024); Pearce et al. (1995)), the 1991 tephra is characterized by significantly higher  $\text{K}_2\text{O}$  values (1.99-3.09 wt.%) than South Sandwich Islands (<1.5 wt.%  $\text{K}_2\text{O}$ ) for similar  $\text{SiO}_2$

ranges (Abbott et al., 2024; Pearce et al., 1995). This combined with ~~no~~ the lack of recorded eruptions of the South Sandwich Island volcanoes during this time period, leads us to rule out South Sandwich Islands as a potential source.

We are left with ~~Cerro Hudson, and Pinatubo, both major~~ Pinatubo (June 15, 1991 eruptions (VEI 5 and 6, respectively),  
400 and Cerro Hudson (August 8-14) as potential sources for the glass in the MBS-Alpha 1991 horizon. While we are able to approximate general sub-annual ages of the individual samples, due to the intra-annual variability seen in MBS accumulation (see section ??), the two-month difference between the Pinatubo and Cerro Hudson eruptions is below our ability to resolve based on dating uncertainty and the size of the sample (~6 cm). Since we cannot rely on sample age, we focus on geochemistry and atmospheric transport modeling to guide our source identification. The products of the 1991 eruption of Pinatubo are  
405 primarily rhyolitic (Fig. 7), and have much lower  $\text{TiO}_2$  (<0.52 wt.%) and FeO (<3.01 wt.%; (Luhr and Melson, 1996; Tamura and Nakagawa, 2023)) than both the MBS-Alpha 1991 glass and the majority of the Cerro Hudson products. Based on the difference in  $\text{TiO}_2$  and FeO, we rule out Pinatubo as a potential source candidate for the dacite cluster in the 1991 tephra. The geochemical compositions of the glass shards in the 1991 sample show a strong similarity to the Holocene volcanic products of the Cerro Hudson volcano (Fig. 7).

410 The 1991 eruption of Cerro Hudson was one of the 20th century's largest explosive eruptions, producing 43 km<sup>3</sup> (bulk volume) of tephra (Kratzmann et al., 2009). Beginning with a phreatomagmatic event on August 8, 1991, with a VEI of 3. A second phase of the eruption began on August 12 with a Plinian eruption with VEI of 5, which produced the majority of the tephra ejected during the eruption (Kratzmann et al., 2009, 2010b). The eruption ejected ash as high as >18 km into the stratosphere, and ash dispersal was well documented spanning thousands of kilometers to the southeast (Doiron et al., 1991; Constantine et al.

415 ~  
The major elements of the 1991 dacite tephra match well with Holocene volcanic products of Cerro Hudson ~~-(Fig. 7; variation diagrams against MgO included in supplementary materials, Figure S.3).~~ This is seen particularly with the later products from the 1991 eruption, as the composition shifted from basaltic-andesite towards large volumes of trachyandesite to rhyo-dacite products (Fig. 7, Wilson et al. (2011); Naranjo S. et al. (1993); Kratzmann et al. (2009)). The dacite glass from the  
420 ~~13.3-1991~~ tephra horizon show a slightly less alkaline composition than some Cerro Hudson products, including glass shards from the 1991 eruption recovered from visible tephra deposits in lake and terrestrial cores from southern Chile (Streeter et al. (2024), Fig. 7). This lower alkalinity is driven largely by a lower concentration of  $\text{Na}_2\text{O}$ , while the remaining major element oxides show similar concentrations to the reported literature values for Cerro Hudson eruption products (Fig. 7; Fernandez and Bitschene (1993); Haberle and Lumley (1998); Gutiérrez et al. (2005); Kratzmann et al. (2009); Del Carlo et al. (2018);  
425 Panaretos et al. (2021); Abbott et al. (2024); Streeter et al. (2024)). ~~While this~~ The lower  $\text{Na}_2\text{O}$  value could be ~~due to in part due to the~~ minimal migration of sodium during EPMA analysis (Gedeon et al., 2000; Humphreys et al., 2006; Kuehn et al., 2011); we consider this unlikely, as analytical conditions for EPMA were selected to minimize potential migration of alkali ions for small glass shards, and repeated analysis of secondary standards did not show a loss of sodium over time (see supplementary materials described in section 3.3 above. Despite the potential  $\text{Na}_2\text{O}$  loss, the majority of our analyses are within the range of  
430 concentrations of major element oxide concentrations seen in the literature for Cerro Hudson (Fig. 7). Based the geochemistry



**Figure 7.** Total ~~alkali vs silica~~ alkali-silica diagram (Le Bas et al. (1986)) and individual major element variation diagrams showing the 13.3 tephra horizon from MBS-Alpha (teal hexagram), together with literature values from various Holocene eruptions of Cerro Hudson (grey circles (Fernandez and Bitschene, 1993), triangles (Haberle and Lumley, 1998), crosses (Gutiérrez et al., 2005), squares (Kratzmann et al., 2010b)(Kratzmann et al., 2010a), plusses (Del Carlo et al., 2018), stars (Panaretos et al., 2021), asterisks (Abbott et al., 2024), and diamonds (Streeter et al., 2024)). Shaded-Green shaded area indicates characteristic compositions of volcanic products from only the later phase of the 1991 eruption of Cerro Hudson (Naranjo S. et al., 1993; Wilson et al., 2011; Kratzmann et al., 2009). Pink shaded area indicates characteristic compositions of volcanic glasses from the 1991 eruption of Mt. Pinatubo (Luhr and Melson, 1996; Tamura and Nakagawa, 2023). Discrimination lines on K<sub>2</sub>O diagram after Peccerillo and Taylor (1976).

and chronology of this cryptotephra horizon, we propose that the MBS-Alpha 1991 ~~cryptotephra-tephra~~ originate from the 1991 eruption of Cerro Hudson.

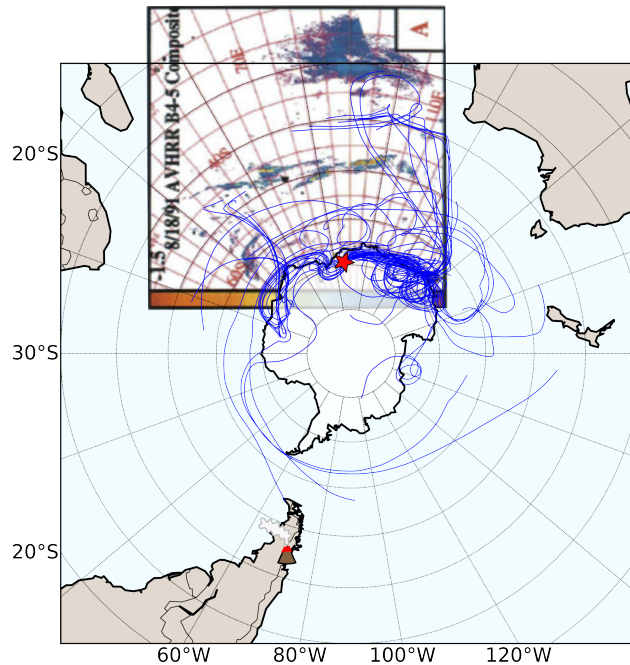
#### 4.2.1 1991 Horizon: Sulfate and tephra deposition timing

The sample depth of the 1991 tephra horizon coincides with the start of a significant increase in the ~~non-sea-salt-sulfate~~ (nss-  
435  $\text{SO}_4^{2-}$ ) chemistry of the MBS-Alpha core (Fig. 5). It is common in ice core tephra studies to see a tephra layer deposited before the corresponding peak in nss- $\text{SO}_4^{2-}$  (~~Burke et al., 2019~~)([Burke et al., 2019](#); [Koffman et al., 2013](#)). The delay between deposition of volcanic ash and volcanic-produced sulfate is variable, based on a combination of distance from the source volcano, scale and plume height of eruptive event, atmospheric transport conditions (Abbott et al., 2024; Plunkett et al., 2023; Koffman et al., 2017; Burke et al., 2019), as well as potential diffusion within the ice (Rhodes et al., 2024). A peak in sulfate  
440 that occurs without substantial lag following tephra deposition may indicate more rapid tropospheric transport and shorter residence time of sulfate aerosols, however the relationship is complicated in cases of long range transport from tropical eruptions (Plunkett et al., 2023). For the 520 BCE HW<sub>6</sub> eruption of Cerro Hudson, Abbott et al. (2024) ~~identify-identified~~ a time lag of less than one year between elevated microparticle levels (indicative of volcanic ash/tephra deposition) and the peak in sulfate deposition in the East Antarctic Plateau B53 ice core.

445 In the case of the 1991 horizon, however, the sulfate spike is assumed to be a combined signal from both the Pinatubo eruption (June 12-15, 1991) and the Cerro Hudson eruption (August ~~8-128-15~~, 1991). The 1991 eruption of Pinatubo ejected an approximated 18 - 20 megatons of  $\text{SO}_2$  into the atmosphere (Guo et al., 2004; Pitts and Thomason, 1993), the Cerro Hudson eruption two months later only produced an estimated 1.5 - 4 megatons of  $\text{SO}_2$  (Doiron et al., 1991; Constantine et al., 2000; Case et al., 2024). The  $\text{SO}_2$  produced in volcanic eruptions is oxidized through atmospheric processes, becoming  $\text{SO}_4^{2-}$   
450 before deposition onto the snow surface, where it is measurable in the ice core (Burke et al., 2019). The order-of-magnitude greater  $\text{SO}_2$  injection from Pinatubo prevents us from being able to disentangle the Hudson  $\text{SO}_2$  signal from bulk ice core measurements without analysis of sulfur isotopes (Burke et al., 2019). Despite the difference in  $\text{SO}_2$  injection, the Cerro Hudson eruption is estimated to have produced a slightly larger volume of tephra-fall deposits than Pinatubo: 2.7 km<sup>3</sup> and 1.8 - 2.2 km<sup>3</sup> dense-rock equivalent, respectively (Guo et al., 2004; Paladio-Melosantos et al., 1996). ~~Additionally, it~~ It is not  
455 uncommon for multiple peaks to be seen in ice core sulfate measurements in years following a major eruption, which can be due to seasonal deposition of sulfate or confounding eruptions (Legrand and Wagenbach, 1999; Guo et al., 2004). Due to the confounding sulfate signals, however, without sulfur isotopic measurements, interpretation of the time-lag between Cerro Hudson tephra deposition and bulk sulfate spike is not possible.

#### 4.2.2 1991 Horizon: Atmospheric circulation

460 ~~The~~ ~~Because the~~ ash dispersal and sulfate aerosol cloud produced by the 1991 eruption of Cerro Hudson has been well studied (Doiron et al., 1991; Schoeberl et al., 1993; Constantine et al., 2000; Kratzmann et al., 2010b; Case et al., 2024), ~~we are~~ able to use the observed atmospheric transport of Cerro Hudson ash in addition to HYSPLIT trajectory modeling. Use of infrared advanced very high resolution radiometer (AVHRR) and total ozone mapping spectrometer (TOMS) satellite data ~~can~~



**Figure 8.** Map showing ten-day back trajectories originating every six hours (44 total trajectories) from the MBS site (indicated by red star), from 12 to 22 August, 1991. An overlay of the Inset AVHRR data in blue and yellow shading (image adapted from Constantine et al. (2000) ) figure shows the location of the Cerro Hudson ash cloud on 18 August, 1991. Cerro Hudson indicated by volcano symbol. (AVHRR figure from Constantine et al. (2000)).

be was used to investigate the volcanic ash and sulfate aerosol distribution, respectively, resulting from the 1991 Cerro Hudson eruption (Doiron et al., 1991; Constantine et al., 2000; Carn et al., 2003). The majority of the ash cloud produced during the Cerro Hudson eruptions was injected near the tropopause, and 90% of the ash cloud was observed to have settled out within a few days (Constantine et al., 2000). AVHRR data show that some part of the ash cloud was detected as far as Australia within five to six days of the August 15 eruptive phase (Constantine et al., 2000). AVHRR images from 18 August, 1991 show part of the Cerro Hudson volcanic ash cloud positioned over the Kerguelen Plateau, north of the MBS site, lying along the 50°S parallel. On the same day, a smaller fragment of the ash cloud appears to be situated north-west of the MBS site, at approximately 60°S, 45°E (Constantine et al. (2000), Fig 8).

It has been established that MBS is climatologically linked to the Southern Indian Ocean (Vance et al., 2016), and atmospheric back trajectory modeling shows that the MBS region regularly receives air masses passing meridionally via the Southern Indian Ocean, and across the Kerguelen Plateau (Jackson et al., 2023). Ten-day HYSPLIT back trajectories computed every 6 hours originating from MBS for 12 - 22 of August, 1991 (dates chosen to coincide with the Cerro Hudson eruption and observed transport of the ash cloud) show multiple trajectories following a path that could reasonably transport Cerro Hudson ash to MBS (Fig. 8). Cluster analysis on 6-hourly trajectories, at 1-hour intervals (resulting in 5 clusters) for

the month of August 1991 shows that clusters representing 44% of the trajectories follow a route that would pass through the August 18 Cerro Hudson ash cloud. These transport setups would readily transport very fine grained ash (like the glass shards identified in the MBS-Alpha 1991 sample) towards the MBS site.

HYSPLIT trajectory evidence for long-range atmospheric transport of volcanic ash from the 1991 eruption of Hudson volcano to the MBS site together with our geochemical evidence for similarities between the MBS-Alpha 1991 ~~cryptotephra~~ tephera and Cerro Hudson 1991 tephra provide robust support for our interpretation that the cryptotephra in this layer was produced during the 1991 eruption of Cerro Hudson.

### 485 4.3 1985 Horizon

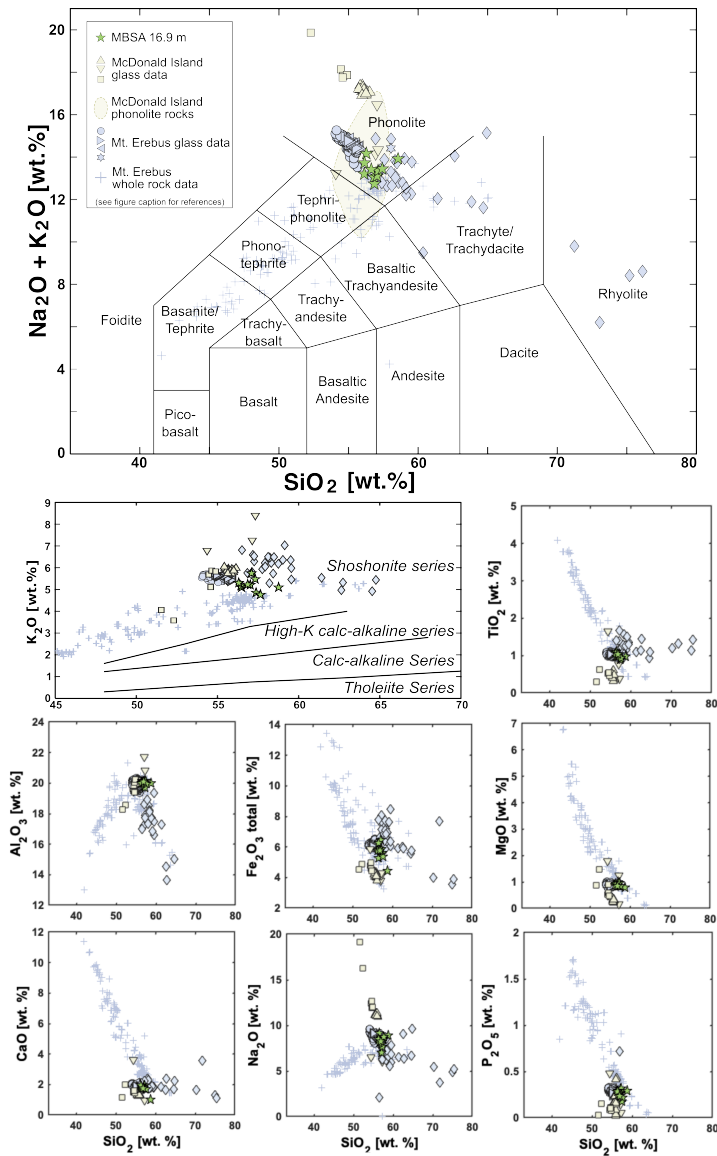
There are no globally significant ~~volcanoes~~ volcanic eruptions on the order of magnitude of Cerro Hudson or Pinatubo recorded in 1985 to help narrow the search for the source of the glass shards identified in the 1985 ~~cryptotephra~~-horizon. The phonolitic composition of the glass shards ~~-, however, are is~~ is uncommon enough to narrow down the candidates for volcanic source matching. To our knowledge, there are only two phonolitic volcanoes active during the last 1000 years (Global Volcanism Program, 2024). These two volcanoes are the McDonald Islands volcano, located near Heard Island on the Kerguelen Plateau, ~2000 km NNW of MBS, and Mount Erebus, Ross Island, Antarctica, ~2500 km SSE of MBS.

There are no known observations of volcanic activity from the McDonald Islands from their discovery in 1874 until 1997. Since 1997, only two eruptive events (Stephenson et al., 2005) have been described, with ongoing hydrothermal activity (Spain et al., 2020). However, the eruptive history of the islands is poorly known due to their extreme remoteness, and near-constant cloud cover in the region making satellite observation difficult for most of the year (Fox et al., 2021).

Mount Erebus has been continuously active since at least 1972 ~~-, with~~ with a maximum VEI of 2, and periods of increased activity from 1984 through 1985 (Global Volcanism Program and McClelland, 1986). Reports from nearby Scott Base describe frequent activity, including the detection of strombolian eruptions, audible explosions, and incandescent ejecta seen from McMurdo Sound ~~-, and~~ as far away as Butter Point (70 km from the volcano) (Global Volcanism Program and McClelland, 1984).

The phonolite tephra identified in the 1985 horizon show similarities to the composition of volcanic glasses from both Erebus and McDonald Islands (Fig. 9a).

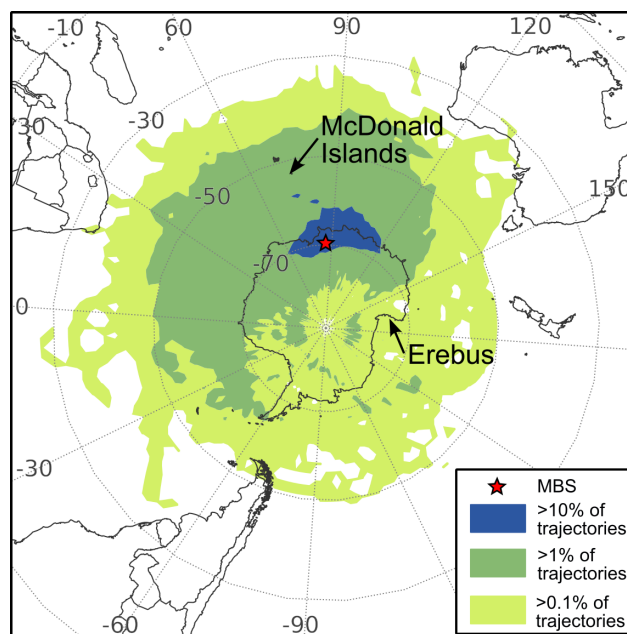
Few analyses of McDonald Islands volcanic material exist in the published literature (Barling et al., 1994; Leach et al., 2016). Leach et al. (2016) investigated an ~~obsidian-rich~~ obsidian-rich phonolitic pumice washed up on a beach in the Chatham Islands, Aotearoa, and characterized it as a product of McDonald Islands, likely from the 1997 eruption. Due to the rigorous matching efforts of Leach et al. (2016), and to expand the available match dataset to include recent eruptive products, we include the Chatham Island sample as McDonald for comparison. Analyses from MBS-Alpha 1985 show a less alkaline composition than that of McDonald Islands, largely driven by a lower concentration of Na<sub>2</sub>O (Fig. 9). Additionally, MBS-Alpha 1985 has higher concentrations of TiO<sub>2</sub>, FeO, and MnO, and generally appears distinct from the literature values for McDonald Islands (Barling et al., 1994; Leach et al., 2016).



**Figure 9.** Total alkali vs silica alkali-silica diagram (Le Bas et al. (1986)) and individual major element variation diagrams showing the 1985 tephra horizon from MBS-Alpha (green pentagrams, this study), together with literature values for eruptive products from Mt. Erebus (blue right pointing triangles (Kelly et al., 2008), left pointing triangles (Iverson et al., 2014), circles (Harpel et al., 2008), diamonds (Silaev et al., 2020), hexagrams (Narcisi et al., 2012), and pluses (Kyle, 1990; Kyle et al., 1992; Martin et al., 2021)) and McDonald Islands (squares (Leach et al. (2016) ; McDonald Islands samples), upward pointing triangles (Leach et al. (2016) obsidian floater), and downward pointing triangles (Barling et al., 1994)). Shaded area in TAS diagram shows McDonald Islands phonolite rock compositions (reproduced from Clarke et al. (1983) figure Figure 3). Discrimination lines on  $\text{K}_2\text{O}$  diagram after Peccerillo and Taylor (1976).

The variation diagrams for  $K_2O$ ,  $TiO_2$ ,  $Al_2O_3$ ,  $Fe_2O_3$ ,  $MgO$ , and  $CaO$  show that the MBS-Alpha 1985 glass correlates more strongly with literature values for Mt. Erebus (Kelly et al., 2008; Iverson et al., 2014; Harpel et al., 2008) than those of the McDonald Islands, despite the slightly higher  $SiO_2$  (Kelly et al., 2008; Harpel et al., 2008; Narcisi et al., 2012; Iverson et al., 2014) (Fig. 9). Comparison glass data for Mt. Erebus are sourced from englacial (blue ice and ice core) tephra correlated with Erebus (Harpel et al., 2008; Narcisi et al., 2012; Iverson et al., 2014), lava bombs ejected from 1972-2004 collected near the flanks of Erebus (Kelly et al., 2008), and ash sampled near the crater rim in 2000 (Silaev et al., 2020). While some of the literature shows very stable, homogeneous populations for the products of Mt. Erebus (Iverson et al., 2014), analysis of the 2000 eruption crater rim ash has more highly variable compositions, ranging from phonolitic to trachytic to rhyolitic. Additionally, when whole rock samples are considered (Kyle, 1990; Kyle et al., 1992; Martin et al., 2021), the 1985 tephra fall well within the range of  $SiO_2$  compositions expected from Erebus.  $MgO$  bivariate diagrams were also considered (supplementary materials Figure S.4), and the MBS-Alpha 1985 tephra cluster more strongly with Mt. Erebus products than with those of the McDonald Islands.

#### 4.3.1 1985 Horizon: Atmospheric circulation



**Figure 10.** Trajectory frequencies for daily 10-day trajectories from May to December 1985. Frequencies calculated as sum of trajectories passing over a 2-degree gridded area, normalized by the total number of trajectories (856 total trajectories).

Using our general age assessment of mid-late 1985 (Fig. 5) as a starting point, we use HYSPLIT to evaluate potential transport to MBS. Because we do not have an exact eruption date from which to run specific forward trajectories, we use a trajectory frequency analysis here based on a wide date range aimed at encompassing the approximated age of the cryptotephra horizon. Analysis of six-hourly ten day back trajectories for May - December 1985 (856 total trajectories) shows that the MBS

site received air masses from a wide range of sources during this period (Fig. 10), including the areas around McDonald Islands and Erebus. Back trajectories from MBS passing over McDonald Islands are more frequent (>1% of trajectories), though some trajectories do pass over Erebus (>0.1%).

530 Potential transport of any volcanic material produced in the Kerguelen Plateau region is supported by our atmospheric trajectory modeling, showing that a significant proportion of trajectories pass over the Kerguelen Plateau en route to MBS. These trajectories likely represent meridional transport of warm air that frequently bring significant amounts of precipitation to the MBS site (Jackson et al., 2023; Vance et al., 2024a). Transport from Erebus likely follows the polar easterlies commonly seen in the region, which align to the mean wind direction at MBS (Vance et al., 2024a).

535 Our HYSPLIT trajectory analysis indicates that the MBS-Alpha 1985 tephra could have been transported from a variety of sources, including Mt. Erebus and the McDonald Islands. However, the combination of the major element ~~geochemistry~~ geochemistry presented here and the known volcanic history of the two potential sources lead to our assessment that the 1985 tephra most likely originated from Mt. Erebus. The potential impact of the identification of Mt. Erebus tephra on our understanding of Antarctic atmospheric transport setups underscores the need for further investigation of the MBS-Alpha 1985  
540 tephra, including trace element analysis.

## 5 Discussion

### 5.1 ~~Ice-core dating~~ MBS as cryptotephra archive

Here we have identified and described two cryptotephra horizons in the MBS-Alpha ice core which we interpret to be the products of eruptions of Mt. Erebus (1985) and Cerro Hudson (1991). Volcanic sulfate and elevated aerosol particle counts  
545 have been previously identified in Antarctic ice congruent with the timing of the 1991 eruption of Cerro Hudson (Evangelista et al., 2022; Legrand and Wagenbach, 1999). The MBS 1991 horizon is, to our knowledge, the first geochemically linked identification of Cerro Hudson 1991 glass tephra shards identified in Antarctic ice.

~~The identification of~~ Our discovery of these cryptotephra horizons highlights both the importance of developing new sampling strategies for locating cryptotephra in ice cores and the potential of the 1991 Cerro Hudson eruption in the MBS-Alpha  
550 ~~ice core has potential implications for ice core dating of both the MBS ice core array and other East Antarctic coastal East Antarctic site as a millennial-length archive for future tephrochronology studies.~~

~~Tephra and cryptotephra have not been reported for many coastal East Antarctica ice cores. Volcanic events are commonly used to provide concrete tie-points in ice core dating efforts, however many of these tie-points are based on  $\text{nssSO}_4$  measurements, rather than volcanic ash layers. Volcanic sulfate aerosols can take months to years to be deposited at a distal ice core site, with elevated sulfate levels in~~ This is because visible tephra layers are uncommon; cryptotephra in these cores tend to be small and sparse, and traditional techniques for identification of cryptotephra require time-consuming, comprehensive sampling of hundreds of meters of ice core material (Cook et al., 2022). Here, we have developed an efficient sampling strategy for the MBS ice core, targeting depths where we expected cryptotephra to be found by using the atmospheric transport characteristic of the site together with existing volcanic records and ice core chemistry ~~often apparent for multiple years following the eruption~~

560 event that produced the sulfate (Lin et al., 2022; Svensson et al., 2020; Plunkett et al., 2023; Burke et al., 2019). Additionally, peaks in volcanic sulfate in ice cores have been shown to be altered through diffusion within the ice as well as layer thinning (Rhodes et al., 2024). Elevated sulfate levels caused by the 1991 eruptions of Cerro Hudson and Pinatubo were measured in aerosol measurements at coastal Antarctic stations and in snowfall across the high Antarctic Plateau over multiple years, until at least the end of 1993 (Legrand and Wagenbach, 1999). Guided by identification of episodes of meridional transport from the southern Indian Ocean region to the MBS site, our method demonstrates the potential of this type of sampling strategy in successfully locating sparse cryptotephra in other coastal Antarctic ice core locations (such as Roosevelt Island; Winstруп et al. (2019)) where sampling based on ice core chemistry alone is made challenging by intermittent precipitation or high background levels of marine aerosols.

This delay in sulfate deposition and multiple sulfate peak phenomenon, as well as within-core migration of sulfate, can cause discrepancies in ice core dating (Plummer et al., 2012; Sigl et al., 2014, 2015; Crockett et al., 2021). The inability to distinguish the source of nssSO<sub>4</sub> deposition between multiple potential source volcanoes without sulfur isotope analysis adds an additional layer of complexity to volcanic matching in ice core dating (Lin et al., 2022; Burke et al., 2019). In the dating of snowpit data from the Dome Fuji site in East Antarctica (Dronning Maud Land), Oyabu et al. (2023) are unable to use the Cerro Hudson/Pinatubo volcanic horizon as a dating tie-point due to the presence of three nssSO<sub>4</sub> peaks within a ~0.5 m interval, likely due to seasonal variability in volcanic sulfate deposition, but possibly also from multiple volcanic eruptions. The presence of volcanic material from Mt. Erebus and Cerro Hudson in the MBS-Alpha ice cores indicates the potential for the preservation of broader Antarctic and ex-Antarctic sources for cryptotephra in East Antarctic ice cores and the MBS ice core array in particular. We expect that other cryptotephra layers exist in the MBS-Alpha ice core, transported atmospheric pathways not highlighted here. Cryptotephra from large, satellite era eruptions outside of our targeted source area (e.g. Puyehue-Cordón Caulle, El Chichón) may exist in MBS-Alpha but were not captured by our method. This is the first tephrochronological investigation of MBS ice cores, and provides insight into a region underrepresented in the Antarctic ice core array. Future tephra studies, either from the remaining MBS-Main archive, or from a new core drilled in a similar location, would greatly increase our understanding of the volcanic archive potential of marginal East Antarctic sites. While the available remaining volume of MBS-Main archive is small, (2 cm<sup>2</sup> cross section, a volume often used in tephrochronology studies (?)), we propose that our findings here justify using an atmospheric-transport focused sampling to develop a tephrochronology of the deeper section of MBS-Main.

Due to rapid tropospheric transport, however, volcanic ash can be advected to

## 5.2 Tephra transport pathways

590 Successful preservation of tephra in an ice core site on relies on an explosive eruption producing an ash plume coinciding with favorable atmospheric circulation patterns to transport the ash to the ice core site. Multiple volcanic factors influence this process including the height of the ash plume, the volume of ash produced, duration of eruption and the location of the volcano relative to the order-of-days-to-weeks. For this reason, tephrochronology can be a more reliable means of locating volcanic tie-points in ice core dating (Geyer et al., 2023; Cook et al., 2022; Lin et al., 2022). Geochemical confirmation of

595 ~~Cerro Hudson as the source of the 1991 horizon confirms the accuracy of deposition site (Lowe, 2011). The size of the eruption and the dating of the MBS-Alpha ice core, as the 1991 sample is appropriately linked to mid-1991 in the current MBS chronology (MBS2023, Vance et al. (2024a)). As the time period around 1991 has previously provided challenges with ice core dating (Plummer et al., 2012; Crockett et al., 2021), the ability to assign a specific date to the 1991 cryptotephra horizon identified here could prove instrumental in verifying future ice core dating efforts, and enable the more accurate characterization of climate proxy-signal links in ice core research.~~

600 Several globally significant volcanic eruptions have been identified in the sulfate and/or conductivity record of the 1137-year record of MBS (Vance et al., 2024a; Harlan et al., 2024a). These eruptions have informed the chronology and synchronization of MBS with other Antarctic ice cores (WAIS, Law Dome, and Roosevelt Island, Vance et al. (2024a)) The identification here of extra-Antarctic (Cerro Hudson) cryptotephra in MBS confirms the potential to locate and identify (crypto)tephra from other location of the volcano have typically been the criteria guiding the sampling of Antarctic ice core for tephra and cryptotephra,  
605 with the expectation that only volcanoes erupting in Antarctica or large, globally significant eruptions in the deeper ice of MBS. Identification of any of the volcanic events proposed by Vance et al. (2024a) as volcanic tie-points would have critical implications for refinement of the chronology of MBS and other Antarctic ice cores spanning similar time periods.

### 5.3 MBS as tephra archive

610 outside of Antarctica (VEI>5) are likely to be preserved in Antarctic ice cores (Del Carlo et al., 2018). Here we have ~~identified and described two cryptotephra horizons in~~ expanded our sampling strategy to include atmospheric trajectory analysis as a method to successfully sample cryptotephra in ice core. MBS is well suited to test our strategy because we have a strong understanding of the synoptic setups that govern atmospheric transport to the ice core site (Vance et al., 2016; Udy et al., 2021; Jackson et al.

615 The atmospheric linkage of coastal East Antarctica and the MBS-Alpha ice core, with 12 dacitic shards identified in 1991, and 10 phonolitic shards in 1985. These tephra horizons highlight the importance of the coastal East Antarctic site as a potential millennial-length archive for future tephrochronology studies. The atmospheric processes that are necessary to transport volcanic material from Mt. Erebus and Cerro Hudson indicate the variability of potential source regions from which MBS could receive and preserve volcanic ash. When considering daily ten day back trajectories over the satellite era, it is apparent that the site receives air masses from almost the entire circumpolar region.

620 ~~When studies were made to select the site for the East Antarctic ice core that became MBS~~ MBS region with the southern mid-latitudes is well documented (Crockett et al., 2021; Udy et al., 2021; Jackson et al., 2023; Vance et al., 2024a). When the MBS ice core site was selected, one of the selection criteria was the prevalence of teleconnections to lower latitudes, especially capturing interconnectedness with large scale modes of climate variability (including the Southern Annular Mode and the Indian Ocean Dipole; Vance et al. (2016); Crockett et al. (2021); Udy et al. (2021)), however with thorough Vance et al. (2016)  
625 ). The existing body of knowledge about the role of meridional transport (Udy et al., 2021, 2022; Jackson et al., 2023) and the observed coastal easterlies at the MBS site (Vance et al., 2024a) support MBS as a potential cryptotephra archive.

~~The back trajectory analysis, it is clear that the performed in the planning of our study provides evidence that climatological transport pathways linking to MBS go beyond its connection with the southern Indian Ocean, and connect the MBS site with a broad range of high latitude air masses. Together with the evidence of very long-range transport of volcanic material (on the order of thousands to tens of thousands of kilometers), this predicts a great opportunity for further comprehensive tephra sampling from the full length of the ~300 m (1137 years) of MBS cores to develop a coastal East Antarctic tephrochronology framework covering the last millennium. Such a tephrochronology has the potential to capture products of large scale volcanic events throughout this period, including eruptions such as Krakatau and Tarawera (1884), Tambora (1815), Kuwae (1454), and the 1259 unknown eruption.~~

### 635 **5.3 Tephra transport pathways**

~~Recent investigations into weather systems like air masses. This is consistent with recent investigations that demonstrate that atmospheric rivers and extreme precipitation across the Antarctic continent (Wille et al., 2019, 2021; Inda-Díaz et al., 2021; Baiman et al., 2021) show that these systems bring significant heat and precipitation from the sub-tropics and mid-latitudes to polar sites like MBS over short timescales. The coastal Antarctic site that makes (Wille et al., 2019, 2021; Inda-Díaz et al., 2021; Baiman et al., 2023; MacLennan et al., 2023) . The characteristics that made MBS so well suited to be as a high resolution record of climate variability (Vance et al., 2016; Crockart et al., 2021; Jackson et al., 2023) also increases increase its chance of preserving tephra from lower latitudes if when eruptions coincide with these kinds of atmospheric events. Ice core sites like MBS might preserve tephra from older eruptions which, as with Cerro Hudson, have been previously thought to be unlikely candidates for the East Antarctic site. Our results demonstrate tephra transport to MBS on the scale of thousands of kilometers by both westerly (from Cerro Hudson) and easterly (from Mt. Erebus) winds.~~

645 ~~The presence of Erebus tephra at MBS also provides important insights into Antarctic atmospheric transport.~~

#### 5.2.1 Meridional transport of Cerro Hudson tephra

The identification of Cerro Hudson tephra in MBS ice verifies the proposed meridional link from MBS to the southern Indian Ocean (Vance et al., 2016). With the Cerro Hudson ash cloud observed over the southern Indian Ocean on August 18, 1991 (Constantine et al., 2000), we interpret that at least some of that ash cloud must have been transported meridionally southwards for deposition at MBS to occur. This meridional transport is validated by Hysplit modelling for the period spanning the 1991 horizon, and Neff and Bertler (2015) find that atmospheric transport from Patagonia to the Southern Ocean and Antarctica is highly efficient. This type of meridional transport is especially prevalent during high precipitation days at MBS (Jackson et al., 2023; Vance et al., 2024a). These events are likely to bring any airborne volcanic ash to the ice core site, where it can be deposited by the wet-deposition of tropospheric aerosols characteristic of the MBS site (Crockart et al., 2021).

Cryptotephra from Holocene eruptions of Cerro Hudson have previously been proposed in Antarctic ice cores (Kurbatov et al., 2006; Narasimhan et al., 2018). However, some correlations have been refuted, and transport of tephra from low latitudes to Antarctica has been deemed unfeasible based on proposed inability of air masses to penetrate the Antarctic polar jet stream (Del Carlo et al., 2018). Our identification of tephra from the 1991 Cerro Hudson eruption in MBS contradicts this claim, and supports studies on the

660 prevalence of rapid, meridional moisture transport and precipitation to coastal East Antarctica (Wille et al., 2019, 2021; Jackson et al., 2023  
. Our results confirming meridional transport support the proposed correlations of other Antarctic ice cores with South American  
volcanic sources (e.g. Kurbatov et al. (2006); Narcisi et al. (2010, 2012)).

### 5.2.2 Easterly transport of Mt. Erebus tephra

Transport from Erebus contradicts ~~to~~ the general understanding that the majority of atmospheric transport to Antarctica follows  
665 the circumpolar westerly winds (Neff and Bertler, 2015). As evidenced by the Hysplit trajectory modeling (Fig. 10), transport  
from Erebus to MBS occurs relatively infrequently, and primarily in line with the Antarctic polar easterly winds, ~~which are~~  
linked to the southern edge of synoptic scale mid-latitude low pressure systems. These polar easterlies ~~can be are~~  
seen in the mean zonal winds, in line with the snow features seen and conditions experienced during drilling at the MBS site (Vance  
et al., 2024a). The identification of MBS-Alpha 1985 tephra as originating from Erebus poses these polar easterly winds as a  
670 potentially overlooked mechanism for particle transport to coastal East Antarctic sites.

~~Despite our assessment that the~~

### 5.2.3 Further implications

Ice core sites like MBS might preserve tephra from eruptions which (as with Cerro Hudson and Erebus) had been previously  
thought to be unlikely candidates for the marginal East Antarctic site. Our results demonstrate tephra transport to MBS is viable  
675 on the scale of thousands of kilometers by both westerly (from Cerro Hudson) and easterly (from Mt. Erebus) winds. The two  
cryptotephra horizons identified in the MBS-Alpha core originate from two very distinct source regions. This, combined with  
the lack of evidence for tephra from other large or more proximal eruptions highlights the importance of favorable atmospheric  
conditions in the preservation of tephra in ice cores.

While we have identified the 1985 sample originates from Erebus, the prevalence of HYSPLIT trajectories from McDonald  
680 Islands, and our understanding of the transport and precipitation sources via the Southern Indian Ocean leave open a possibility  
for meridional transport along this pathway to bring volcanic material from the McDonald Islands and nearby Heard Island to  
MBS. Transport from McDonald Islands would be supported by the tephra as originating from Erebus, HYSPLIT trajectories  
from the Heard and McDonald Islands region are more frequent (Fig. 10). The same processes that transported the 1991 tephra  
from Cerro Hudson ~~;~~ warm airmasses could transport ash from McDonald Islands, given an eruption of suitable size. This is  
685 seen in the warm air masses regularly transported along this meridional pathway, bringing substantial amounts of snow to the  
MBS site (Jackson et al., 2023). ~~Investigations into bubble free layer features in ice cores like MBS and the Law Dome ice~~  
~~core have been studied based on their potential to record temperature changes during and after large events transporting vapor~~  
~~from low latitudes (Zhang et al., 2023). Further development of this method could be used to help understand specific moisture~~  
~~transport conditions in cases where bubble free layers fall within tephra sample depths, including the 16.87-16.915 m depth~~  
690 ~~(1985) sample, in which one such bubble free layer feature has been identified (at 16.88 m depth).~~ The lack of identifiably  
Heard and McDonald Islands tephra in MBS-Alpha indicates that the either recorded satellite-era eruptions of Heard and  
McDoanld Islands were not substantial enough to be transported, or did not occur when transport conditions were favorable.

### 5.3 Ice core chronology refinement

695 The identification of the 1991 Cerro Hudson eruption in the MBS-Alpha ice core has potential implications for ice core dating of both the MBS ice core array and other East Antarctic ice cores. Volcanic events are commonly recognized in ice cores via  $\text{nssSO}_4$  (volcanic sulphate) anomalies. Deposition of volcanic sulfate from stratospheric eruptions is relatively spatially homogeneous, and measurable in ice cores using standard ice core chemistry analyses (Lin et al., 2022). These volcanic sulfate records have long been used in the development of robust ice core chronologies, allowing for the synchronization of ice core  
700 records across hemispheres (?).

The use of volcanic sulfate alone in identifying volcanic tie points at annual resolution, however, is not without challenges. Without sulfur isotope analysis, it can be difficult to determine the specific source of  $\text{nssSO}_4$  deposition between multiple potential source volcanoes (Lin et al., 2022; Burke et al., 2019). Additionally, volcanic sulfate in ice cores can be altered through diffusion within the ice (Rhodes et al., 2024). Sulfate aerosols from distal sources can take months to years to be  
705 deposited, with elevated sulfate levels in ice core chemistry often apparent for multiple years following an eruption (Lin et al., 2022; Svensson et al., 2022). For example, elevated sulfate concentrations caused by the 1991 eruptions of Cerro Hudson tephra in MBS ice verifies the proposed meridional link from MBS to the southern Indian Ocean. While there is satellite image evidence of and Pinatubo were found in aerosol measurements at coastal Antarctic stations and in snowfall across the Antarctic Plateau with peaks measured until at least the end of 1993 (Legrand and Wagenbach, 1999). These factors can make defining the precise timing of a volcanic  
710 event challenging (Plummer et al., 2012; Sigl et al., 2014, 2015; Crockart et al., 2021). In the dating of a snowpit record from the Dome Fuji site in East Antarctica Oyabu et al. (2023) were unable to precisely define the Cerro Hudson/Pinatubo volcanic horizon as a dating tie-point due to the presence of three  $\text{nssSO}_4$  peaks within a  $\sim 0.5$  m interval, likely due to seasonal variability in volcanic sulfate deposition, but possibly also from multiple volcanic eruptions.

Because individual tephra shards can often be geochemically linked to a specific eruption, tephrochronology can help  
715 with the refinement of ice core chronologies. Rapid tropospheric transport can carry tephra to an ice core site on the order of days to weeks (Lowe, 2011). If tephra can be identified, it can be used to add precision to sulfate-based chronologies (Geyer et al., 2023; Cook et al., 2022; Lin et al., 2022). Geochemical confirmation of Cerro Hudson as the source of the Cerro Hudson ash cloud sitting over the southern Indian Ocean on August 18, 1991 (Constantine et al., 2000), the presence of tephra from the ash cloud implies that at least a small fraction of the ash cloud must have been transported meridionally southwards  
720 for deposition at MBS to occur. This meridional transport is also validated by Hysplit modelling for the period spanning the 1991 horizon, and Neff and Bertler (2015) find that atmospheric transport from Patagonia to the Southern Ocean and Antarctica is highly efficient. These meridional wind patterns are especially prevalent during high precipitation days at MBS (Jackson et al., 2023; Vance et al., 2024a). These high-precipitation events are likely to bring any airborne volcanic ash to the site, and drive wet deposition of tropospheric aerosols, the characteristic mode of deposition at the MBS site (Crockart et al., 2021)  
725 confirms the accuracy of the dating of the MBS-Alpha ice core, as the 1991 sample is appropriately linked to mid-1991 in the current MBS chronology (MBS2023, Vance et al. (2024a)). As the time period around 1991 has previously provided challenges

with ice core dating (Plummer et al., 2012; Crockart et al., 2021), the ability to assign a specific date to the 1991 cryptotephra horizon identified here could prove instrumental in verifying future ice core dating efforts, and enable the more accurate characterization of climate proxy-signal links in ice core research.

730 ~~Previous tephra from Holocene eruptions of Cerro Hudson have been proposed in~~ Several globally significant volcanic eruptions have been identified in the sulfate and/or conductivity record of the MBS-Main (Vance et al., 2024a; Harlan et al., 2024a). These eruptions have informed the chronology and synchronization of MBS with other Antarctic ice cores (Kurbatov et al., 2006; Nareisi et al., 2010, 2012). ~~However, the identification of Cerro Hudson in the Talos Dome ice core by Nareisi et al. (2012) has been refuted, and potential for transport of tephra from low latitudes to Antarctica has been deemed unfeasible based on proposed inability of air masses to~~  
735 ~~penetrate the Antarctic polar jet stream (Del Carlo et al., 2018). Our identification of tephra from the 1991 Cerro Hudson eruption in MBS stands in opposition of this claim, and supports studies on the prevalence of rapid, meridional moisture transport and precipitation to coastal East Antarctica (Wille et al., 2019, 2021; Jackson et al., 2023; Turner et al., 2019). Our results confirming meridional transport support the proposed correlations of other East (WAIS, Law Dome, and Roosevelt Island, Vance et al. (2024a))~~ The identification here of extra-Antarctic (Cerro Hudson) cryptotephra in MBS confirms the potential to locate and identify  
740 (crypto)tephra from other globally significant eruptions in the deeper ice of MBS. Identification of any of the volcanic events proposed by Vance et al. (2024a) as volcanic tie-points would have important implications for refinement of the chronology of MBS and other Antarctic ice cores with South American volcanic sources (e.g. Kurbatov et al. (2006); Nareisi et al. (2010, 2012)) spanning similar time periods.

## 6 Conclusions

745 We developed a sample depth selection strategy for identifying potential ice core depths containing cryptotephra in an East Antarctic ice core using a combination of ice core chemistry, known eruption events, and atmospheric transport modeling. A thorough sampling of the satellite-era Mount Brown South Alpha ice core ~~yielded >40~~ using this method yielded 42 volcanic glass shards from 12 ~~samples~~ sample depths. Of these, we present correlations for two samples, each with ~~homogeneous compositional clusters of~~ largely homogeneous compositions. The samples, comprising 10 and 14 glass shards, 13 glass shards  
750 respectively, are proposed to originate from eruptions of Erebus (active throughout the early to mid-1980s) and the 1991 eruption of Cerro Hudson. The identification of two tephra-cryptotephra layers in a core spanning only ~20 m (40 years) is notable important for future studies due to the relative sparseness of tephra previously identified in East Antarctic ice. This work is a proof of concept for locating and identifying cryptotephra based on this type of targeted sampling, rather than a more time consuming comprehensive screening of an entire ice core.

755 Identification of tephra at 1985, and its correlation with Mt. Erebus, suggests that the continuing low-level (VEI 1 or 2), Strombolian eruptions occurring at the site have the ability to produce ash that can be transported thousands of kilometers given favorable atmospheric conditions. ~~Given~~ With appropriate transport conditions, volcanic material can be transported along the polar easterly winds prevalent at MBS, despite the more frequent events in which warm moist air is advected southwards bringing extreme precipitation to coastal East Antarctica.

760 Identifying tephra from the 1991 Cerro Hudson eruption in MBS is especially significant for its relevance to ice core dating and core synchronization. The presence of Cerro Hudson tephra shards at 13.28-13.34 m, correlated to 1991, allows definitive confirmation that the MBS-Alpha core is well dated. As this sulfate spike is typically used as a tie point for the early-mid 1990s in Antarctic ice cores, this tephra can confirm or correct the estimate of the important Pinatubo dating horizon. In cases where multiple sulfate peaks are seen around the eruption year, finding Cerro Hudson ash in the core ~~in~~at a depth corresponding to  
765 mid-1991 removes any doubt about the timing of sulfate deposition at this site. Additionally, this tephra provides evidence of extra-Antarctic volcanic ash passing the Antarctic circumpolar wind belt to be deposited in Antarctica, previously thought to be improbable.

Our findings position ~~coastal East Antarctica and~~ the MBS site (coastal East Antarctica more broadly) as an important potential archive for future tephrochronological work, due to its high ~~resolution~~accumulation and teleconnections to such a  
770 diverse array of ~~low~~lower latitude source regions. We suggest ~~a future systematic~~future targeted sampling of the ~~MBS-main~~full MBS-Main core for the development of a comprehensive tephrochronology of coastal East Antarctica. Such a millennial-length ~~tephrochronology~~tephrochronology would provide a record from a climatologically important region, previously underrepresented in the existing East Antarctic ice core array.

*Data availability.* The MBS-Alpha Geochemistry datasets produced in this study are available as part of the downloadable supplementary  
775 files accompanying this manuscript. The MBS2023 chronology (Vance et al., 2024b) is available at <http://dx.doi.org/doi:10.26179/352b-6298>. MBS surface core chemistry and water isotope datasets (Moy et al., 2024; Crockart et al., 2021) are available at <http://dx.doi.org/doi:10.26179/372t-4q89> and <http://dx.doi.org/doi:10.4225/15/58eedf6812621>.

*Author contributions.* MH led the study, including developing the sampling plan and writing the manuscript. MH, TV, JF, and HAK conceived the study. MH and EC prepared ice core samples for analysis at the University of Copenhagen with the help of AS. MH performed  
780 sample analysis at the Central Science Laboratory at the University of Tasmania. MH, JF, EC, and AS contributed to geochemical characterization of samples. TV leads the Mount Brown South Ice Core project. All coauthors contributed to writing the manuscript.

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795 The Scientific Colour Maps (Crameri, 2023) are used in this study to prevent visual distortion of the data and exclusion of readers with colour-vision deficiencies (Crameri et al., 2020).

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