



1 **Stratigraphic templates for ice core records of the past 1.5** 2 **million years**

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10 **Abstract.** The international ice core community has a target to obtain continuous ice cores stretching back as far
11 as 1.5 million years. This would provide vital data (including a CO₂ profile) allowing us to assess ideas about
12 the cause of the Mid-Pleistocene Transition (MPT). The European Beyond EPICA project and the Australian
13 Million Year Ice Core project each plan to drill such a core in the region known as Little Dome C. Dating the
14 cores will be challenging, and one approach will be to match some of the records obtained with existing marine
15 sediment datasets, informed by similarities in the existing 800 kyr period. Water isotopes in Antarctica have
16 been shown to closely mirror deepwater temperature, estimated from Mg/Ca ratios of benthic foraminifera, in a
17 marine core on the Chatham Rise near to New Zealand. The dust record in ice cores resembles very closely a
18 South Atlantic marine record of iron accumulation rate. By assuming these relationships continue beyond 800
19 ka, our ice core record could be synchronised to dated marine sediments. This could be supplemented, and allow
20 synchronisation at higher resolution, by the identification of rapid millennial scale-events that are observed both
21 in Antarctic methane records and in emerging records of planktic oxygen isotopes and alkenone sea surface
22 temperature (SST) from the Portuguese Margin. Although published data remain quite sparse, it should also be
23 possible to match ¹⁰Be from ice cores to records of geomagnetic palaeointensity and authigenic ¹⁰Be/⁹Be in
24 marine sediments. However, there are a number of issues that have to be resolved before the ice core ¹⁰Be record
25 can be used. The approach of matching records to a template will be most successful if the new core is in
26 stratigraphic order, but should also provide constraints on disordered records, if used in combination with
27 absolute radiogenic ages.

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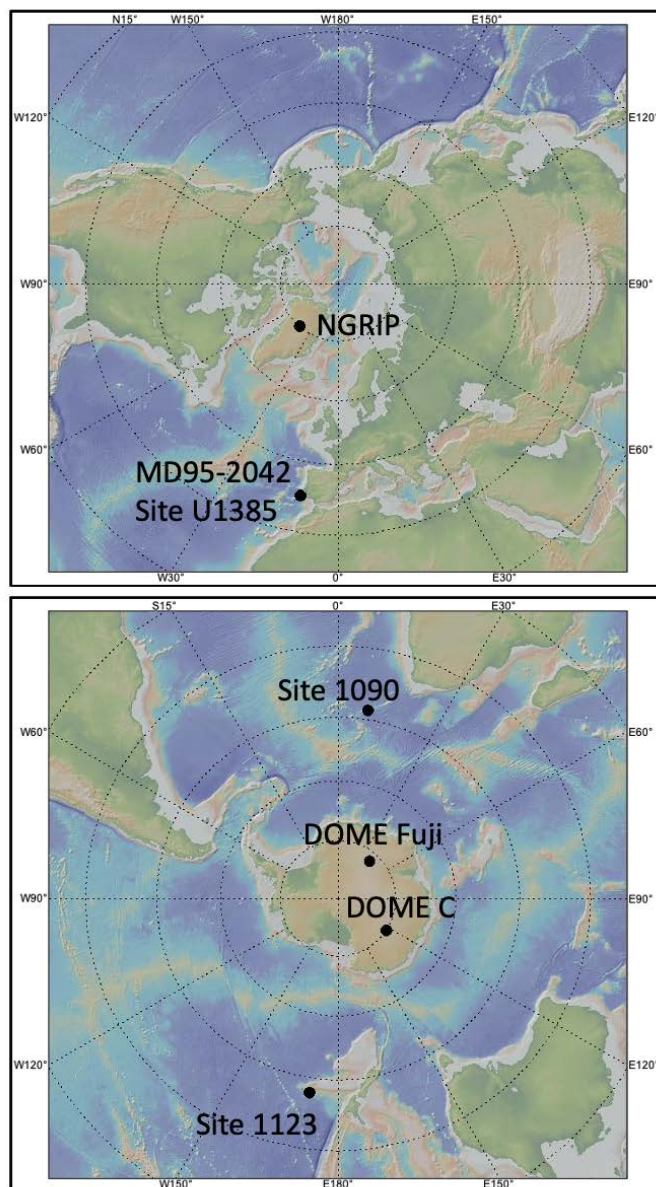
29 **1. Introduction**

30 Ice cores have provided iconic records of changes in atmospheric composition and climate over
31 glacial/interglacial cycles, with Antarctic datasets extending, so far, 800 kyr into the past (e.g. Bereiter et al.,
32 2015; Jouzel et al., 2007; Wolff et al., 2010). While this illustrates what is often referred to as the “100 kyr
33 world”, marine (e.g. Lisiecki and Raymo, 2005) and terrestrial records indicate that a different style and strength
34 of glacial cycle existed earlier in the Pleistocene, during the “41 kyr world”. The causes of the so-called mid-
35 Pleistocene Transition (MPT) remain hotly debated (Clark et al., 2006), with changes in CO₂ concentration or
36 changes in the nature of the ice/rock interface underlying continental ice sheets often invoked.

37 Some of the issues surrounding these debates could be resolved if an ice core record, extending beyond the MPT
38 and including records of past greenhouse gas concentrations, could be obtained. It has therefore become a key
39 target of the ice core community to find a location to drill a core reaching as far back as 1.5 Ma (Fischer et al.,
40 2013). Several projects to obtain such a core are partially underway, including the European Beyond EPICA
41 project which plans to drill between 2021 and 2025 at a site known as Little Dome C (LDC). This site is only
42 about 30 km from the site of the EPICA Dome C drilling (Fig. 1) that reached 800 ka, but is located on top of a
43 subglacial highland, thus avoiding basal melting that led to loss of the oldest ice at Dome C. The Australian
44 Million Year Ice Core (MYIC) project is targeting the same region of Antarctica.

45 A major challenge is to date such a core. Recently greenhouse gas concentrations were reported for ice as old as
46 2 Ma at a blue ice location of Allan Hills, Antarctica (Yan et al., 2019). While this provided tantalising
47 snapshots of atmospheric composition, the dating was too imprecise to assign data unequivocally to particular
48 parts of glacial cycles, or even to specific cycles. While this is a particular issue for discontinuous records such
49 as those from blue ice, dating is also likely to be a major problem for a “standard” core, even assuming it is
50 complete and continuous.

51 A number of methods can be used to try and date the ice older than 800 ka. As with the blue ice, absolute ages
52 may be estimated from radiometric methods, including ⁸¹Kr decay (Buizert et al., 2014; Crotti et al., 2021), and
53 the growth in atmospheric concentration with time of ⁴⁰Ar (Bender et al., 2008; Yan et al., 2019), but both of
54 these methods currently have large error bars at ages of 1 million years or more. The decay of cosmogenic
55 isotopes (using the ratio of ¹⁰Be/³⁶Cl to remove production rate variations) also has potential, but issues with
56 ³⁶Cl loss at low accumulation rate sites (Delmas et al., 2004) have to be solved and the dating accuracy will
57 likely be similar to the one using ⁸¹Kr and ⁴⁰Ar.



58

59 Figure 1. Location map. The maps show the locations of the ice and marine cores shown in this paper. Colours
60 represent topography and bathymetry. Figure made with GeoMapApp (www.geomapapp.org) / CC BY.



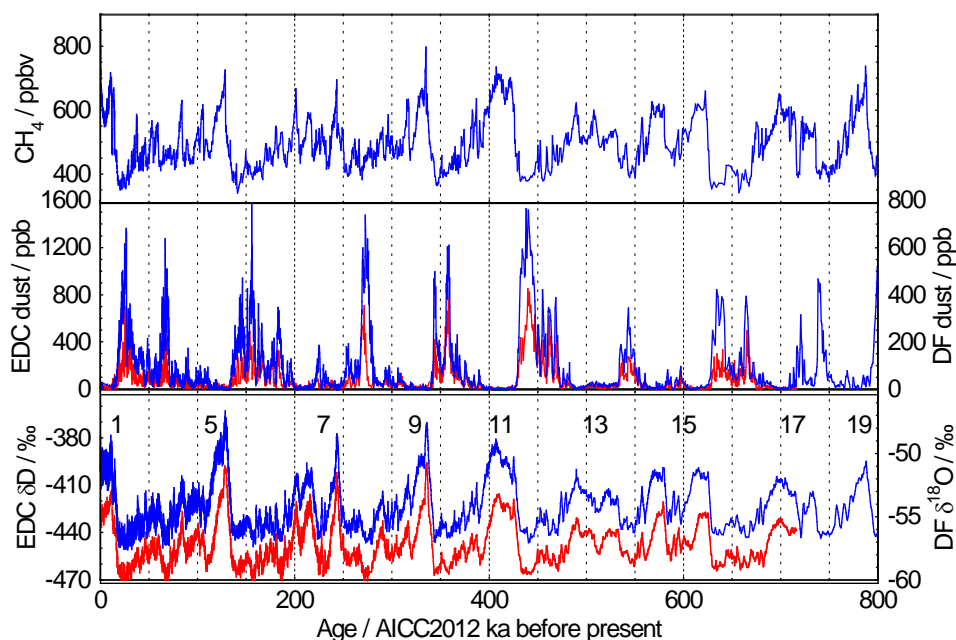
61 The main technique for dating the ice <800 ka in age has been to combine an estimate of past snow
62 accumulation rate and thinning with a range of fixed points that tie the ice core to known ages (Bazin et al.,
63 2013). These fixed points include the signal of low intensity of the geomagnetic dipole field associated with
64 polarity reversals recorded as increases in ^{10}Be deposition (Raisbeck et al., 2006), and various orbital tuning
65 targets including the ratio of O_2/N_2 (Kawamura et al., 2007). These methods, as well as the radiometric ones,
66 will certainly be applied to the new 1.5 Ma projects. However, diffusion, lack of resolution and disturbed ice
67 flow, with the possibility of folded ice near the bed, as has been seen at deep ice core sites in Greenland
68 (Grootes et al., 1993; NEEM Community Members, 2013), mean that further stratigraphic methods to date the
69 core may be needed.

70 One additional option is to create templates to which the records generated in the new projects can be matched.
71 The orbital targets (for tuning of O_2/N_2 and $\delta^{18}\text{O}_{\text{atm}}$) used to construct the 800 ka age model (Bazin et al., 2013)
72 are simple examples of the use of such templates, and will not be included here as they are very straightforward
73 to construct. Marine and terrestrial records that are rather well-dated extend beyond 1.5 million years. As an
74 example many marine records have been mapped, using benthic isotopes, onto the LR04 marine stack (Lisiecki
75 and Raymo, 2007), whose age uncertainty at 1.5 Ma is estimated at 6 kyr, or the more recent Prob-Stack (Ahn et
76 al., 2017).

77 In this paper, we consider which ice core parameters may have analogues in the marine record that could be
78 used as templates onto which a future ice core could be mapped. We focus particularly on the EPICA Dome C
79 ice core (EDC), because its close proximity to the planned ice cores at LDC leads us to expect a similar signal in
80 most parameters. We use ice core datasets which have already been shown to closely mirror a particular marine
81 record over the past 800 kyr. We consider the mechanistic basis for such agreement and whether it is likely to
82 apply through the MPT to 1.5 Ma. We then present “predictions” of what some parameters might look like in
83 the new ice core, which can be used as both a test of integrity and continuity, and as a first dating tool for the
84 core.

85 2. EPICA Dome C ice core records

86 In the following sections, we will consider possible analogues for 4 ice core parameters (Fig. 2). The water
87 isotope record ($\delta^{18}\text{O}$ and δD) is the most basic climate parameter (Jouzel et al., 2007) recorded in the ice,
88 generally considered to represent temperature at the ice core site. Dust is the insoluble component of impurities
89 trapped in the ice, and represents terrestrial material from the southern continents. Both dust and water isotopes
90 display particularly strong changes over glacial cycles with more subdued millennial scale variations. Methane
91 is the one component in the ice core record that displays abrupt events, parallel to the rapid Dansgaard-Oeschger
92 events seen in Greenland ice cores. ^{10}Be (not shown in Fig. 2) is the cosmogenic isotope most commonly
93 measured in ice, and its production is controlled by changes in Earth’s and the Sun’s magnetic fields which also
94 influence cosmogenic isotopes archived in other material. These 4 components will be considered in more
95 detail in the following sections.



96
97 Figure 2. Ice core data over the past 800 kyr. Ice core records covering the past 800 kyr from Dome C (blue) and
98 from the last 720 kyr from Dome Fuji (red). Top panel: methane (Louergue et al., 2008); middle panel: dust
99 (Kawamura et al., 2017; Lambert et al., 2008); lower panel: water isotopes (Jouzel et al., 2007; Kawamura et al.,
100 2017), with interglacial marine isotope stage numbers marked.

101

102 Although we are specifically aiming here to create a template for the European or Australian drilling at LDC, we
103 note that in most details, the features and relative changes seen for water isotopes, dust and ^{10}Be are expected to
104 be similar across the East Antarctic plateau. This is illustrated in Fig. 2, where we have plotted $\delta^{18}\text{O}$ and dust
105 concentration from Dome Fuji (Fig. 1) (Kawamura et al., 2017) along with δD (Jouzel et al., 2007) and dust
106 concentration (Lambert et al., 2008) from EDC, all plotted on the AICC2012 age scale. The absolute level of
107 dust concentrations varies spatially across the Antarctic plateau, being dependent on travel distance from the
108 main Patagonian dust source region (Fischer et al., 2007a) and higher at Dome Fuji than at Dome C. This
109 difference is explained in part because of the different analysis method used for Dome F and Dome C that
110 includes different size ranges, but in any case the pattern is almost identical on multimillennial timescales;
111 methane is of course expected to show the same concentrations across Antarctica. Our templates for LDC are
112 therefore likely to serve as equally valid for other sites across East Antarctica.

113 3. Water isotopes

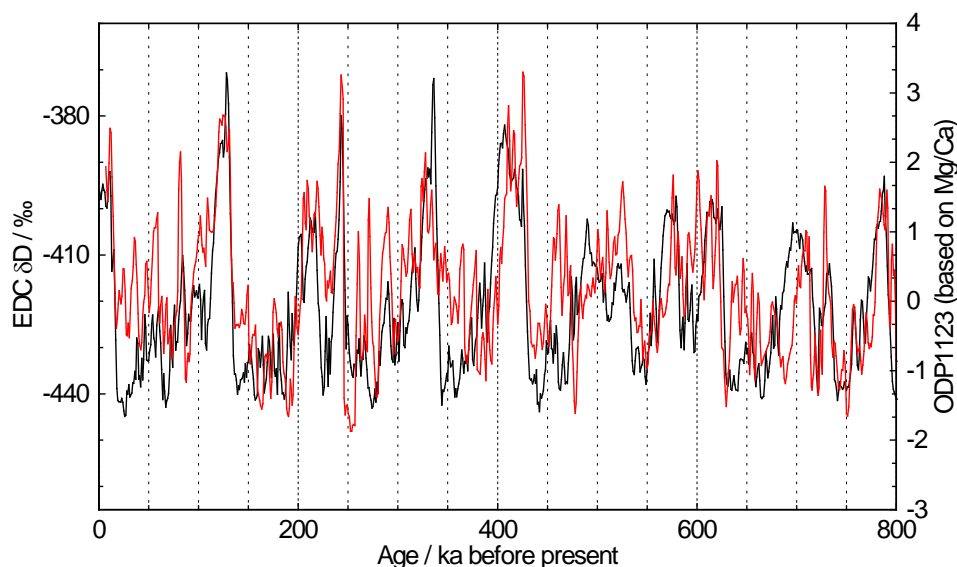
114 Water isotopes (δD , $\delta^{18}\text{O}$) in ice cores are generally taken to represent the temperature at the ice core site,
115 although the reality is actually much more complicated than that (Buizert et al., 2021; Jouzel et al., 1997). It
116 therefore makes sense to look for a potential marine analogue that also records mainly temperature. Although
117 water isotope records from ice are sometimes plotted along with oxygen isotope records from marine cores, the



118 latter reflect a combination of temperature and ice volume (as well as local salinity effects), and so the
119 variability of the two records may not be comparable on glacial-interglacial scales. A commonly used
120 geochemical temperature sensor in marine cores is the ratio of Mg/Ca in foraminifera.

121 Planktic Mg/Ca records, covering 800 ka and more, are available from a number of marine sites and should
122 reflect sea surface temperatures (SSTs) (e.g. Shakun et al., 2015). However while we expect some match
123 between Antarctic temperatures and those from the high southern latitudes, we would expect most other sites to
124 display a rather different pattern owing to the operation of the bipolar seesaw (e.g. Barker et al., 2011).
125 Elderfield et al. (2012) noticed a striking similarity between the deepwater temperature inferred from Mg/Ca of
126 benthic foraminifera at Ocean Drilling Program (ODP) site 1123 (Fig. 1), on the Chatham Rise east of New
127 Zealand, and the temperature inferred from δD in the EDC ice core. They hypothesised that this is because
128 deepwater temperature, particularly in the South Pacific, reflects the temperature of sinking surface waters and
129 of Antarctic and proximal air temperature. Mean ocean temperature (determined by analysing noble gas ratios
130 in ice cores) also shows a very similar pattern to Antarctic surface temperature across the last two glacial
131 terminations (Baggenstos et al., 2019; Bereiter et al., 2018; Shackleton et al., 2020), which supports the
132 interpretation. In recent years, deep water temperatures covering at least 1.5 Ma have been obtained from two
133 other sites in the North Atlantic (Sosdian and Rosenthal, 2009) and North Pacific (Ford and Raymo, 2019), but
134 given their location they are not expected to reflect Antarctic climate, thus leaving only ODP site 1123 as a
135 suitable comparator.

136 In Figure 3, we compare the record of δD from EDC with that of benthic Mg/Ca from site 1123 over the last 800
137 ka. The Mg/Ca record is presented as converted to temperature and interpolated (Elderfield et al., 2012), and is
138 on the LR04 age model (Lisiecki and Raymo, 2005), while the ice core data is on AICC2012 (Bazin et al.,
139 2013). By using a suitable scaling to overlay the two records we can compare their fidelity to each other and
140 observe the extended Mg/Ca record as a possible template for δD over 1.5 Ma.



141

142 Figure 3. Ice core and marine sediment data reflecting temperature for the past 800 kyr. EDC deuterium (black,
143 AICC2012 age scale) (Jouzel et al., 2007). ODP site 1123 deepwater temperature (red, LR04 age model), based
144 on Mg/Ca (Elderfield et al., 2012).

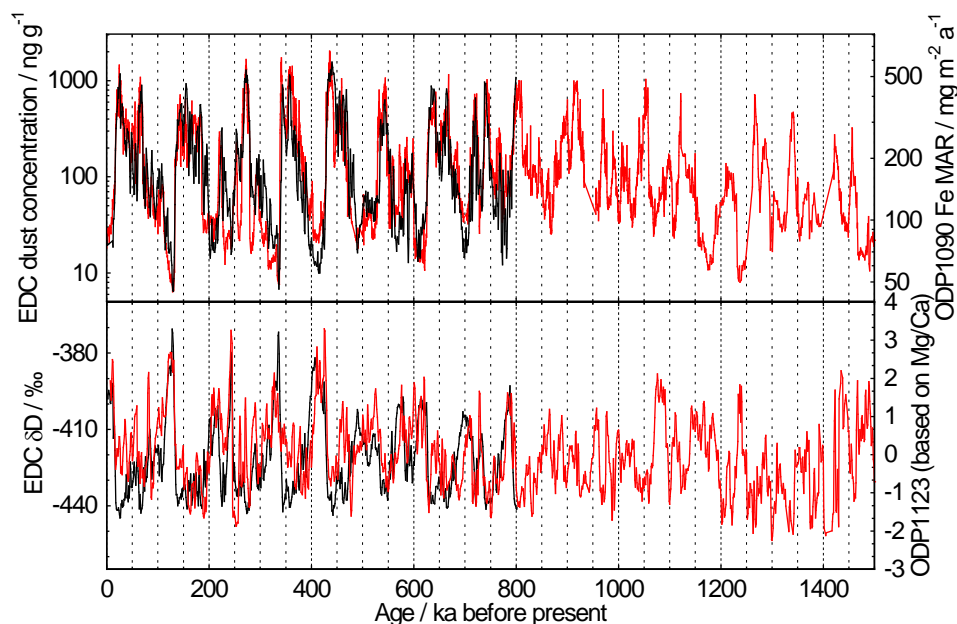
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146 The similarity between the two records is strong at the orbital timescale where both the shape and the relative
147 amplitude of each glacial cycle is the same in the two records. However, there are significant mismatches at the
148 shorter, multimillennial, timescale. Some very prominent millennial-scale AIM (Antarctic isotopic Maximum)
149 events in the ice core record are very weak, or in some cases not clearly resolved, in the marine record. Some of
150 the issues may actually be related to temporal synchronisation, and perhaps to the resolution of the marine
151 record. But still, it would be hard to use the marine record as a template for an ice core record between 450 and
152 550 ka (MIS 13). This is a concern because some of the sections of Site 1123 beyond 800 ka have a similar
153 nature to that section.

154 Despite these concerns over the fidelity of the marine record as a predictor of the ice core isotope signal, we
155 would expect the similarity to continue provided deepwater temperature at high southern latitudes continues to
156 be driven by surface temperatures around Antarctica before the MPT. What could disrupt such a link would be
157 significant changes in ocean circulation and in the reach of different water masses. Such changes may well have
158 occurred over the MPT (Ford and Raymo, 2019), and one suggestion is that they are related to a hypothesised
159 change in the Antarctic Ice Sheet (Raymo et al., 2006) from largely terrestrial to marine-based. While such
160 changes would certainly have impacted the supply of water affected by Antarctic surface temperatures to the
161 deep ocean, the proximity of site 1123 to Antarctica makes it unlikely that a southern influence was completely
162 absent at that time. We therefore see it as likely that the site 1123 Mg/Ca record extended to 1.5 Ma (Fig. 4)
163 does serve as an approximate template for at least the glacial/interglacial variability in Antarctic temperature and
164 therefore LDC deuterium. However, we accept the possibility that the exact nature of the relationship between



165 the two records could have differed in the early part of the period from that observed after 800 ka, and indeed
166 should a mismatch be found in the ice core record it will provoke reconsideration of the assumptions made here.



167
168 Figure 4. Ice core records to 800 ka and marine records to 1500 ka. Lower panel: EDC deuterium
169 (Jouzel et al., 2007) and Mg/Ca-based deepwater temperature (red) from site ODP1123 (Elderfield et al., 2012).
170 Upper panel: EDC dust (black) (Lambert et al., 2008) and Fe MAR from ODP site 1090 (red) (Martinez-Garcia
171 et al., 2011).

172

173 It would obviously be beneficial to search for other marine analogues of Antarctic temperature. The similarity of
174 benthic oxygen isotopes in cores on the Portuguese Margin to Antarctic temperature was noted previously,
175 albeit on a very short time period (Shackleton et al., 2000). The extension of this record, which is underway
176 (Birner et al., 2016) would provide a much better resolved record, with clear millennial scale signals, and its
177 applicability as a template could be assessed. The caveat is that the underpinning reason for similarity of a
178 benthic isotope record controlled by several factors (ice volume, temperature, hydrography and water mass
179 changes) with Antarctic temperature is unclear, making it difficult to assess the likelihood that the relationship
180 persisted before the MPT.

181 4. Dust

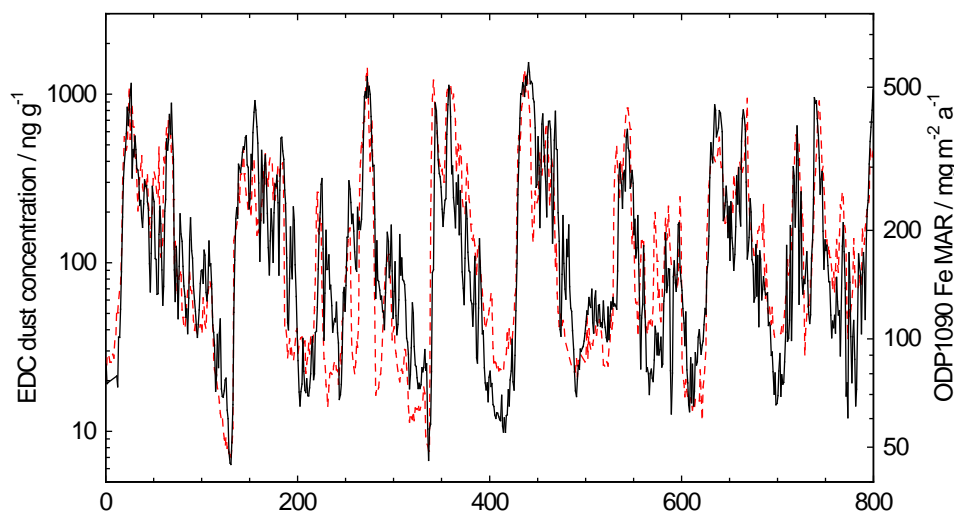
182 Terrestrial dust is measured directly in ice cores, as insoluble particle numbers and sizes which can be converted
183 to mass concentrations, and indirectly in marine sediments through the concentrations or ratios of elements that
184 are mainly (at appropriate sites) of windborne terrestrial origin. The 800 ka record of dust concentration in the
185 EDC ice core (Lambert et al., 2008) shows strong glacial-interglacial cycles, with high concentrations of dust in



186 glacial periods, and some multimillennial scale variability. Elemental and isotopic analysis indicates that the
187 dust mainly originates from South American sources, particularly in Patagonia (Delmonte et al., 2008). As a
188 result we would expect a close relationship between dust arriving in Antarctica and dust deposited onto the
189 South Atlantic during the early stages of the path to Antarctica.

190 ODP site 1090 (Fig. 1) is ideally located to sample dust during its transport in the westerly wind belt from the
191 Patagonian sources towards Antarctica. Martinez-Garcia et al. (2011) noted that different dust proxies in the
192 sediment core from site 1090 matched well with each other over 4 Ma, and with EDC dust flux over 800 ka.
193 Here we compare their preferred dust proxy (mass accumulation rate of iron, Fe MAR) with the dust
194 concentration at EDC (Lambert et al., 2008). We use concentration rather than flux because this is what we will
195 be able to measure in the deeper parts of the LDC core – the flux is a derived quantity that requires knowledge
196 of the snow accumulation rate. It is therefore a fairer test to assess the similarity of the measured quantity
197 (concentration) to the marine target.

198 In Fig. 5 we compare the two records over the last 800 ka. Note that glacialials have high values of dust, that both
199 records have been smoothed to 1 ka averages, and are plotted on log scales. With the appropriate scaling, the
200 agreement between the two records is remarkable. This applies both to the consistent amplitude relationship, to
201 the shape of glacial-interglacial cycles, and to the identification of almost every multimillennial scale peak in
202 both records. The section from 450-550 ka, which was problematical in the isotope records discussed above,
203 shows a good match.



204
205 Figure 5. Ice core and marine sediment dust data for the past 800 kyr. EDC dust (black) (Lambert et al., 2008);
206 Fe MAR from ODP site 1090 (red) (Martinez-Garcia et al., 2011). Data have been smoothed to 1 kyr averages
207 and the marine data were aligned (Martinez-Garcia et al., 2011) to the ice core age model.

208

209 The agreement is made more surprising by the fact that the dynamic range of the two records is very different:
210 the marine dust, being geographically closer to the dust production in Patagonia than is the long-range



211 transported ice core record, varies by a factor 10 (minimum in MIS 5e, maximum in MIS 13), while the ice core
212 dust concentration varies by a factor >100 (factor 200 between MIS 5e and MIS 13). The range of dust flux at
213 EDC would be about a factor 50, because the snow accumulation rate is about four times higher in MIS 5e than
214 in MIS 13. This implies that the causes of dust variability are split into two halves: a factor of about 10 is due
215 mainly to changes at or near the source of the dust, another factor of about 5 is due to changes in lifetime during
216 the long meridional journey to Antarctica. This has been discussed several times before (Fischer et al., 2007b;
217 Lambert et al., 2008; Markle et al., 2018; Petit and Delmonte, 2009; Wolff et al., 2010) and although the
218 different approaches led to somewhat different amplification factors by dust source and transport processes, the
219 comparison shows that solutions that match the available data must consider changes both in source and in
220 lifetime.

221 This implies that the extended marine dust record (Fig. 4) could be an excellent template for the dust record
222 expected in the LDC ice cores. The part of the variance that is based on changes at or near the source should
223 remain, whatever occurred across the MPT. The second part of the variability, arising from changes in aerosol
224 lifetime over the Southern Ocean, has been in phase with changes at the source over the last 800 ka. This could
225 in theory have altered if there were major changes in atmospheric circulation across the MPT. Nonetheless, it
226 seems likely that the basic glacial-interglacial pattern, as well as the imprint of millennial scale change will have
227 persisted.

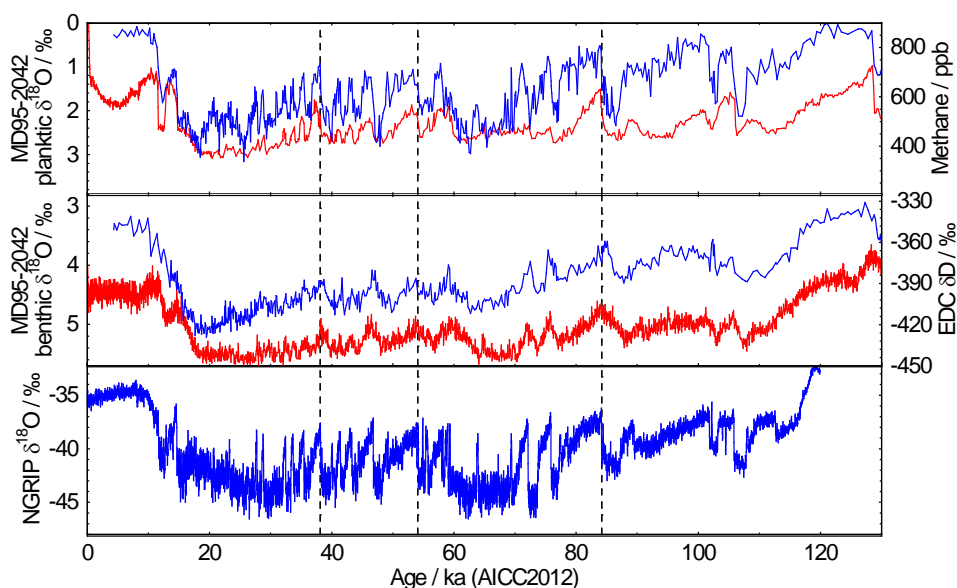
228 5. Methane as a pattern for Dansgaard-Oeschger variability

229 While the EDC water isotope and dust records show strong variability, particularly on orbital timescales, that
230 can be used for pattern matching, their variations tend to be smooth, so that correlation is clear but imprecise.
231 Records containing the imprint of Dansgaard-Oeschger (D-O) events have the capacity to identify sharp time
232 points, and therefore to give much closer synchronisation, and many more clear tie points. Using the model of
233 the bipolar seesaw, it is possible to rather convincingly reproduce D-O events from the Antarctic isotope record,
234 to produce what is known as the synthetic Greenland record (GL_T-syn) (Barker et al., 2011). However, the
235 synthetic record can never have the sharpness of the original signal and in particular for ice older than 800 kyr
236 diffusion in the ice may have smoothed the higher frequency climate signal in the water isotope record. The only
237 record in Antarctic ice that does retain the character of the D-O events is the methane record.

238 Over the last glacial cycle, every significant D-O event recorded in the Greenland ice core record (North
239 Greenland Ice-Core Project (NorthGRIP) Members, 2004) is also seen in the EDC methane record (Loulergue et
240 al., 2008) (Fig. 6). The same pattern of abrupt climate change is seen in many other northern hemisphere
241 climate records, with a particularly faithful representation observed in planktonic oxygen isotope and alkenone
242 SST data from marine sediment cores from the Portuguese Margin (Govin et al., 2014; Shackleton et al., 2000).
243 Note that the benthic $\delta^{18}\text{O}$ from the Portuguese Margin strongly resembles the water isotope record from
244 Antarctica (here, δD from EDC) and that the phasing of planktic and benthic $\delta^{18}\text{O}$ on the Iberian Margin is the
245 same as that seen between CH_4 and δD in the Antarctic ice core record. This pattern has been interpreted as
246 being indicative of a thermal bipolar seesaw, and offers another signature for matching ice core and marine
247 records.



248 While the planktonic oxygen isotope and alkenone SST records reproduce the NGRIP (Greenland) ice core
249 isotopic record well in terms both of shape and amplitude, the methane record is less easy to match to the marine
250 record. This is because the amplitude of methane change in comparison to isotopic change (in either the
251 Greenland ice or North Atlantic marine record) is very variable. For example Greenland Interstadial (GI) 19 (at
252 73 ka) is very strong in the isotope records but shows a methane amplitude of only 70 ppb (Baumgartner et al.,
253 2014), and a sensitivity (methane jump/Greenland temperature change) less than a third of some other events.
254 This means that, in an unknown section of older core, we could only expect to make unequivocal matches for
255 some D-O events.



256
257 Figure 6. Sharp millennial scale features in the last glacial cycle. Bottom panel: NGRIP $\delta^{18}\text{O}$ showing the
258 pattern of Dansgaard-Oeschger events (North Greenland Ice Core Project Members, 2004). Middle panel:
259 Benthic $\delta^{18}\text{O}$ (blue) from site MD95-2042 (Govin et al., 2014; Shackleton et al., 2000) and the deuterium
260 record from the Antarctic EDC ice core (Jouzel et al., 2007). Upper panel: Planktonic $\delta^{18}\text{O}$ (blue) from site
261 MD95-2042 (Govin et al., 2014; Shackleton et al., 2000) showing the same pattern as Greenland $\delta^{18}\text{O}$; methane
262 from the Antarctic EDC ice core (red) (Loulergue et al., 2008) showing a more subdued version of the same
263 variability. Vertical lines mark examples of the sharp onsets of three interglacials (Greenland Interstadial (GI) 8,
264 14 and 21).

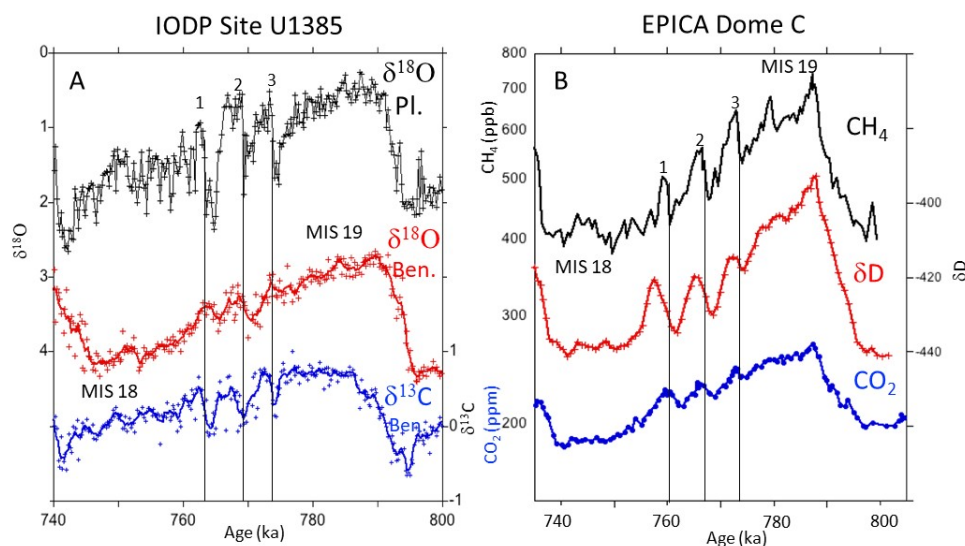
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266 High resolution isotopic data and alkenone SST collected at site U1385 (Fig. 1), extending to 1.45 Ma (Hodell et
267 al., 2015), located close (25 km) to core MD95-2042 (discussed above and shown in Fig. 6) indicate that events
268 of a D-O nature extend throughout (Birner et al., 2016). Thus the planktonic isotope and SST records from that
269 site, soon to be published, should serve as a regional template for D-O variability. Using it with the methane ice
270 core record makes the assumption that the teleconnection between North Atlantic climate variability and the



271 (predominantly tropical) methane sources (Bock et al., 2017) remained intact before the MPT. This could be
272 tested if East Asian speleothem records extended deeper in time than is currently the case (Cheng et al., 2016).

273 As an example of the potential for this method, we examine the relationship between the oldest part of the EDC
274 record (MIS 19) and the equivalent data from site U1385 (Fig. 7). Here we can clearly identify the three strong
275 millennial events on the MIS 19/18 boundary in both the marine and ice core record, with the sharp onsets in
276 planktonic $\delta^{18}\text{O}$ (marine) and methane (ice) and the more symmetric change in benthic $\delta^{18}\text{O}$ (marine) and δD
277 (ice). Carbon cycle data in both records also show the signature of the events.



278
279 Figure 7. (A) Planktonic $\delta^{18}\text{O}$ (black), benthic $\delta^{18}\text{O}$ (red) and $\delta^{13}\text{C}$ (blue) over the MIS19-18 transition at Site
280 U1385 (Sánchez Goñi et al., 2016) compared to (B) CH_4 (black), δD (red) and atmospheric CO_2 (blue) in the
281 EPICA Dome C ice core. Three strong millennial events (labeled 1-3) occur on the MIS19-18 transition that are
282 recorded in both the marine sediment and ice cores. Vertical dashed lines are drawn at the abrupt transitions
283 from cold stadials to warmer interstadial conditions. Note that the phasing of ice core CH_4 and δD is not quite as
284 expected from later time periods (Fig. 6) and may reflect uncertainty in Δ -age (the age difference between the
285 ice and gas records).

286

287 The variable amplitude of methane peaks relative to North Atlantic records may make it harder to use than some
288 other records. Nonetheless the simplicity of the match at ~ 770 ka suggests that methane in ice for the most
289 prominent millennial-scale features, used in a complementary way with other records, and matched against the
290 Portuguese Margin datasets, will provide a viable way of aligning the marine and ice records rather precisely at
291 least for the cycles immediately below 800 ka. In the highly thinned ice over 1.2 Ma old, where there may be



292 >10 ka/m of ice (Lilien et al., 2021), the use of high-resolution continuous online laser spectrometric
293 measurement techniques to measure CH₄ (Chappellaz et al., 2013; Rhodes et al., 2015) should still allow
294 resolution of millennial features provided diffusion of methane (Bereiter et al., 2014) and of δD (Pol et al.,
295 2010) is limited.

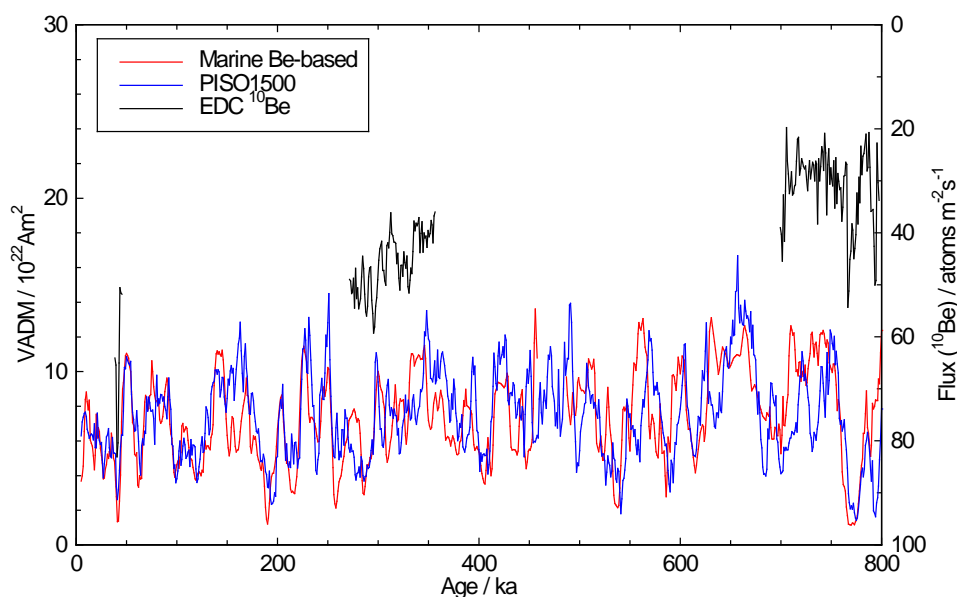
296 6. ¹⁰Be

297 The production of ¹⁰Be in the atmosphere, and its subsequent deposition to Antarctic snow, is controlled by the
298 flux of cosmic rays, which in turn is influenced by the solar magnetic field (showing solar cycles), and on longer
299 timescales by changes in Earth's magnetic field. As examples, centennial scale variations in ¹⁰Be in ice over the
300 last 14 kyr can be matched to variations in ¹⁴C (Muscheler et al., 2014), while the Laschamp magnetic excursion
301 at about 41 kyr BP (Raisbeck et al., 2017) and the Brunhes-Matuyama magnetic reversal at about 780 kyr BP
302 (Raisbeck et al., 2006) are easily identified in ice cores. However there are, as with all aerosol-bound proxies,
303 atmospheric transport influences on the relative amount of produced ¹⁰Be that is transported to Antarctica.
304 Additionally, in the central East Antarctic plateau the concentration of ¹⁰Be shows a very clear imprint of
305 climate that is mainly removed by calculating the flux. We will therefore have to independently estimate the
306 snow accumulation rate in order to use any ¹⁰Be template for dating.

307 In the marine record, the strength of Earth's magnetic field is imprinted in records of geomagnetic
308 palaeointensity. A number of reconstructions have been made using individual cores, but a carefully constructed
309 stack from different sites is especially valuable. The PISO-1500 stack of relative palaeointensity (RPI)
310 (Channell et al., 2009) is particularly widely used, and could serve as a template for long-term variations in ice
311 core ¹⁰Be. In theory an even more direct comparator would be an index derived from the authigenic ¹⁰Be/⁹Be
312 ratios in marine sediments (Simon et al., 2018; Simon et al., 2016). Measurements extend beyond 2 Ma, and
313 show a good correlation with the RPI (Channell et al., 2009). However because detailed ¹⁰Be data exist only for
314 a very few cores, the RPI might be considered a more robust dataset at this stage.

315 Both RPI and ¹⁰Be/⁹Be show the strong features that we know have been seen in the ice core record: in
316 particular the Laschamp excursion and the Brunhes-Matuyama boundary. It has been reported that the PISO-
317 1500 stack shows a good correlation with the unpublished record of ¹⁰Be flux from 200-800 ka (Cauquoin,
318 2013), with a correlation coefficient reported as r=0.62 after the timescales have been aligned. Unfortunately,
319 we can only show the comparison for the few published sections of ice (Fig 8).

320 The extended datasets, both of palaeointensity (Channell et al., 2009) and authigenic ¹⁰Be/⁹Be ratios (Simon et
321 al., 2018) should therefore be useful templates with which to compare the ¹⁰Be data obtained from the LDC ice
322 core. In Fig. 9 we show these two datasets, as an indication of what a ¹⁰Be flux record from the new core should
323 show. Note that uncorrected ¹⁰Be/⁹Be data automatically include the degree of decay (1.39 Myr half-life) that
324 will also apply in the ice core, whereas the PISO1500 do not include that decay. The paleointensity lows
325 associated with polarity reversals in particular (Fig. 9) should be quite prominent in the ¹⁰Be record (analogous
326 to the Brunhes/Matuyama boundary). Some of the other prominent excursions events should also be captured.

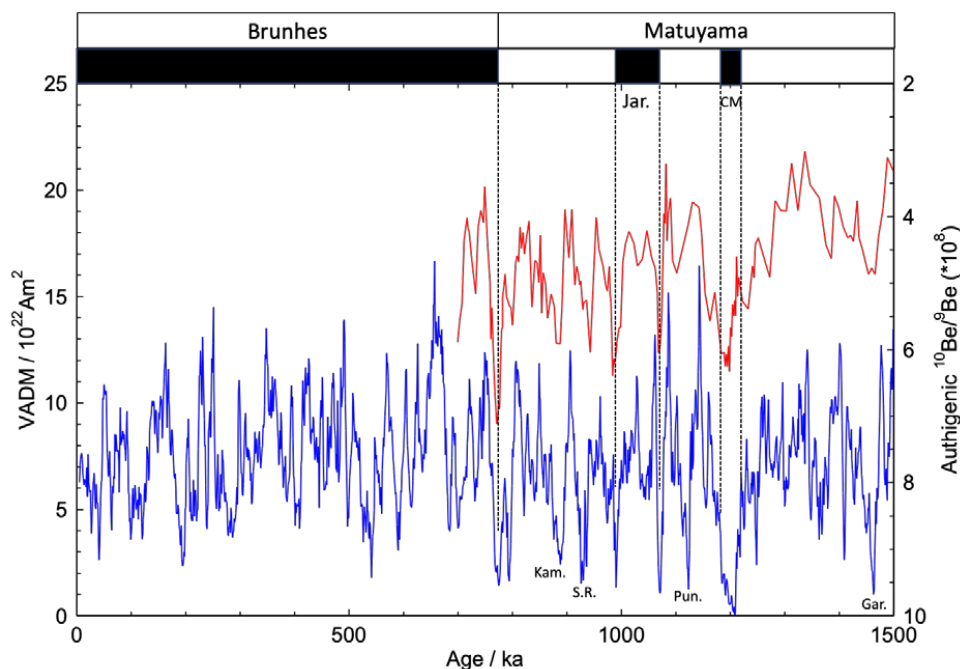


327

328 Figure 8. Palaeointensity and ^{10}Be data for the last 800 kyr. Virtual axial dipole moment (VADM) from the
329 PISO1500 palaeointensity stack (blue) (Channell et al., 2009); VADM derived from an authigenic $^{10}\text{Be}/^9\text{Be}$
330 ratio stack (Simon et al., 2016) using an empirical calibration (red); published ^{10}Be fluxes from the EDC ice
331 core (black, right axis) (Cauquoin et al., 2015; Raisbeck et al., 2017; Raisbeck et al., 2006).

332

333 There are two aspects that degrade the ability of ^{10}Be alone to provide a dating template. The first is that it is the
334 ^{10}Be flux that resembles marine data, and we will only have measurements of ^{10}Be concentration. This issue
335 applies also to dust (as discussed above), but the dynamic range between glacial and interglacial for dust is so
336 great (factor 10 in marine sediments, higher still in ice) that the influence of accumulation rate changes is second
337 order and does not mask the signal that is common between ice and marine sediments. For ^{10}Be the range of the
338 data (factor 2 between low and high) is similar to the range of accumulation rates, meaning that the
339 concentration is equally influenced by the cosmogenic production rate and the snow accumulation rate. By itself
340 the ^{10}Be concentration will be hard to place onto the template.



341

342 Figure 9. Palaeointensity and authigenic $^{10}\text{Be}/^9\text{Be}$ from marine sediments for the last 1.5 Myr. Virtual axial
343 dipole moment (VADM) from the PISO1500 palaeointensity stack (blue) (Channell et al., 2009); authigenic
344 $^{10}\text{Be}/^9\text{Be}$ (decay-corrected) from core MD97-2143 (red) (Simon et al., 2018). Each of the polarity reversals
345 (Brunhes, Jaramillo, Cobb Mountain) is associated with a palaeointensity low. Other prominent excursions in
346 the Matuyama Chron are labeled: Kam = Kamikatsura; S.R. = Santa Rosa; Pun = Punaruu; Gar = Gardar
347 (Channell, 2017).

348

349 There is one possible way to deal with this issue. The accumulation rate for the EDC core was actually a product
350 derived from the age modelling, but based on a prior where the accumulation rate was assumed to be directly
351 related to the temperature and hence to the water isotope ratios. If we assume that that relationship was
352 unchanged over 1.5 million years then the best fit values from the 800 kyr of EDC could be used, along with
353 water isotope ratios measured at LDC to estimate the accumulation rate for each depth and therefore calculate a
354 flux of ^{10}Be . This will have considerable uncertainties but is likely to allow identification of the main features in
355 the expected ^{10}Be record.

356 An additional problem is the one encountered when the Brunhes-Matuyama section of the EDC ice core was
357 analysed (Raisbeck et al., 2006), that ^{10}Be in deeper ice shows spikes that appear to be inhomogeneous across
358 the core and may be associated with high concentrations of dust and other chemical concentrations. The spikes
359 have been tentatively ascribed to a concentration effect where ^{10}Be becomes associated with dust particles which
360 also seem to clump together into aggregates in the deeper ice (de Angelis et al., 2013). For the Brunhes-
361 Matuyama section of the EDC ice core, the spikiness in ^{10}Be was bypassed using median concentrations



362 (Raisbeck et al., 2006), and it may be that such a strategy will continue to work in older ice. However, further
363 work is needed to understand the conditions that lead to this effect.

364 **7. Discussion and conclusion**

365 We have presented templates for what an undisturbed (i.e., where time is monotonic with depth) ice core from
366 LDC might be expected to show. The marine dust record (represented here by Fe MAR at ODP site 1090) could,
367 with reasonable assumptions, be an excellent template for the LDC dust record. The Mg/Ca data from site 1123,
368 matched against the LDC water isotope record, could provide limited validation. The methane data, matched
369 against D-O variability at site U1385 may be capable of adding some sharper tie points in a record that has
370 already been matched to first order. ^{10}Be concentration, converted to an estimated flux using water isotope data,
371 should be a useful additional validation, particularly in identifying the major features with low Vertical Axial
372 Dipole Moment (VADM) and expected high ^{10}Be concentration.

373 It will be a greater challenge to use these records to aid the age modelling if the record is disturbed, with folds or
374 missing ice, as has often been the case with ice near the bed of ice sheets (e.g. NEEM Community Members,
375 2013). In that case, one cannot rely on the shape of the signal to identify the time period represented. Instead, we
376 are dependent on using the absolute values – for example finding a time period where the values in the templates
377 are all consistent with the measured values. The derived ^{10}Be production may be particularly important in this
378 case, because it is independent of climate and, thus, may provide a more robust age assignment compared to the
379 other templates considered here, which are highly correlated on glacial/interglacial time scales. This will have to
380 be done with considerable caution, given the uncertainties involved in the assumptions about the unchanged
381 relationship between the measured values and their marine equivalents over time. Finally an age model for the
382 new core will of course also use other data, including those from gas measurements ($\delta^{18}\text{O}_{\text{atm}}$ and O_2/N_2 , which
383 can be matched to calculated orbital targets), and any radiometric absolute ages that can be obtained from the
384 limited ice volumes available.

385 **Data availability**

386 All the datasets shown in this paper have already been published elsewhere, as indicated by the relevant
387 references.

388 **Author contribution**

389 All authors conceived the idea for this paper. EW prepared the first draft and all authors reviewed and edited the
390 text.

391 **Competing interests**

392 The authors declare that they have no conflict of interest.

393 **Acknowledgments**

394 This publication was generated in the frame of Beyond EPICA. The project has received funding from the
395 European Union's Horizon 2020 research and innovation programme under grant agreement No. 815384
396 (Oldest Ice Core). It is supported by national partners and funding agencies in Belgium, Denmark, France,



397 Germany, Italy, Norway, Sweden, Switzerland, The Netherlands and the United Kingdom. Logistic support is
398 mainly provided by PNRA and IPEV through the Concordia Station system. The opinions expressed and
399 arguments employed herein do not necessarily reflect the official views of the European Union funding agency
400 or other national funding bodies. This is Beyond EPICA publication number XX. This publication is also
401 associated with the Million Year Ice Core (MYIC) Project of the Australian Antarctic Program (AAP). We
402 thank Alexander Cauquoin and Grant Raisbeck for advice about the published ^{10}Be data. EW was supported by
403 a Royal Society Professorship. HF acknowledges the long-term financial support of ice core science by the
404 Swiss National Science Foundation. TvO acknowledges support by the Australian Government Department of
405 Industry, Science, Energy and Resources (grant no. ASCI000002).

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