



# 1 A theory of glacial cycles: resolving Pleistocene puzzles

2 Hsien-Wang Ou<sup>1</sup>

3 <sup>1</sup> Department of Earth and Environmental Sciences, Lamont-Doherty Earth Observatory of Co-  
4 lumbia University, Palisades, NY10964, USA (retired)

5 *Correspondence to:* Hsien-Wang Ou ([hsienou0905@gmail.com](mailto:hsienou0905@gmail.com))

6

7 **Abstract.** Since the summer surface air temperature that regulates the ice margin is anchored on  
8 the sea surface temperature, we posit that the climate system constitutes the intermediary of the  
9 orbital forcing of the glacial cycles. As such, the relevant forcing is the annual solar flux ab-  
10 sorbed by the ocean, which naturally filters out the precession effect in early Pleistocene but  
11 mimics the Milankovitch insolation in late Pleistocene. For a coupled climate system that is  
12 inherent turbulent, we show that the ocean may be bistable with a cold state defined by the freez-  
13 ing point subpolar water, which would translate to ice bistates between a polar ice cap and an ice  
14 sheet extending to mid-latitudes, enabling large ice-volume signal regardless the forcing ampli-  
15 tude so long as the bistable thresholds are crossed. Such thresholds are set by the global convec-  
16 tive flux, which would be lowered during the Pleistocene cooling, whose interplay with the ice-  
17 albedo feedback leads to transitions of the ice signal from that dominated by obliquity to the  
18 emerging precession cycles to the ice-age cycles paced by eccentricity. Through a single dy-  
19 namical framework, the theory thus may resolve many long-standing puzzles of the glacial cy-  
20 cles.

## 21 1 Introduction



22           In late Pleistocene, the global ice volume exhibits pronounced variation at orbital fre-  
23   quencies (Hays et al. 1976). As Antarctica is largely iced over since about ten million years  
24   ago (Berger 1979), the ice signal reflects mainly that of the northern ice sheet, whose correla-  
25   tion with the Milankovitch insolation supports the latter’s control of the ice volume (Milan-  
26   kovitch 1941). The linkage however is necessarily nonlinear since the 100-ky eccentricity con-  
27   tains little power, yet it dominates the ice-age cycles of the late Pleistocene, the so-called “100-  
28   ky problem” (Elkibbi and Rial 2001). Equally puzzling, the ice signal exhibits mainly the  
29   obliquity periodicity in the early Pleistocene despite the greater precession amplitude (Raymo  
30   and Nisancioglu 2003), and then the mid-Pleistocene transition (MPT, all acronyms are listed  
31   in Appendix A) from the obliquity- to the eccentricity-dominated ice cycles is not accompanied  
32   by appreciable change in the Milankovitch insolation (Clark et al. 2006). Given the distinct-  
33   ness of these features, they should emerge from fundamental physics of the climate-cryosphere  
34   system, which is yet to be delineated.

35           The above observations have weeded out some previous resolutions of the 100-ky prob-  
36   lem, as briefly recounted below. Early attempts have invoked internal oscillations of the ice  
37   sheet due to mass-balance feedback or isostatic adjustment (Weertman 1976; Imbrie and Im-  
38   brie 1980; Oerlemans 1982; Pollard 1983; Pelletier 2003), but these are broadband processes,  
39   which would elevate the low-frequency variance but not produce a spectral peak coincidental  
40   with the eccentricity (Imbrie et al. 1993). Others have invoked stochastic or chaotic dynamics  
41   in conjunction with the orbital forcing (Wunsch 2003; Huybers 2009), which however may be  
42   at odds with the observation that ice ages seem to always terminate at rising eccentricity



43 (Raymo 1997; Kawamura et al. 2007; Lisiecki 2010), suggesting a more deterministic process  
44 (Meyers and Hinnov 2010). Both internal oscillations and stochastic dynamics may not ex-  
45 plain why they are not operative in early Pleistocene when the northern hemisphere glaciation  
46 has already set in (Ravelo et al. 2004) and yet 100-ky cycles are absent. Then there are con-  
47 ceptual or dynamical-system models (Saltzman et al. 1984; Ghil 1994; Paillard 1998;  
48 Tziperman et al. 2006; Imbrie et al. 2011; Crucifix 2013; Daruka and Ditlevsen 2016), which  
49 can be tuned to replicate the observed signals, but since key parameters are not linked to meas-  
50 urable quantities, these models are mostly unfalsifiable as more parameters invariably lead to  
51 better fit, which also has limited prognostic utility since free parameters may not be assumed  
52 fixed for a changed forcing scenario.

53 A more palpable paradigm is that the ice sheet is bistable (Weertman 1976), so the hys-  
54 teresis can be triggered by eccentricity-modulated forcing to produce the 100-ky cycle, and if  
55 the hysteresis thresholds is lowered by the putative decrease of the atmospheric CO<sub>2</sub> during the  
56 Pleistocene, one has a possible explanation of the MPT (Berger et al. 1999; Calov and Ga-  
57 nopolski 2005; Abe-Ouchi et al. 2013). But there is no evidence of such CO<sub>2</sub> trend (Honisch et  
58 al. 2009), which moreover is likely a response than a cause of the climate change (Petit et al.  
59 1999; Siegenthaler et al. 2005; Honisch et al. 2009), and then the atmospheric CO<sub>2</sub> in any event  
60 has only minor effect on the temperature (Broecker and Denton 1989; Petit et al. 1999), the  
61 reason that the model-produced temperature range is small compared with the observed one  
62 (Abe-Ouchi et al. 2013). One serious difficulty of the above hysteresis is that one of the bi-  
63 states is an ice-free state, so the ice signal is minimal prior to the MPT (Berger and Loutre



64 2010, Fig. 2), which is at odds with the substantial obliquity signal observed in early Pleisto-  
65 cene (Letreguilly et al. 1991; Raymo and Nisancioglu 2003; Lisiecki and Raymo 2007). We  
66 shall argue later (Sect. 3.3) that an ice-free state is untenable in the Pleistocene, so the hystere-  
67 sis paradigm, if it were to apply, must involve different bistates, which as we shall see can be  
68 engendered by the ocean.

69 Not sufficiently acknowledged in above studies is the central role played by the ocean.  
70 Physically, the summer surface air temperature (SAT) that controls the ice margin is anchored  
71 by the sea surface temperature (SST), so the ocean must constitute the primary entry of the ra-  
72 diative forcing into the climate-cryosphere system (Broecker and Denton 1989). Observation-  
73 ally, the meridional overturning circulation (MOC), the SST and the SAT all covary strongly  
74 during glacial cycles (Ruddiman et al. 1986; Imbrie et al. 1992, Fig. 6; Keigwin et al. 1994,  
75 Fig. 2; Lisiecki et al. 2008; Nie et al. 2008), which arguably precede the ice-volume signal by  
76 several millennia, suggesting a causal linkage (Petit et al. 1999; Shackleton 2000; Medina-  
77 Elizalde and Lea 2005). In addition, it is noted that abrupt climate changes involve MOC  
78 jumping between modes (Broecker et al. 1985; Alley et al. 2000), a bistability that has been  
79 demonstrated by coupled climate models (Manabe and Stouffer 1988) whose dynamical basis  
80 however remains unclear. To remedy this shortfall, this author (Ou 2018) has recently ex-  
81 tended Stommel's (1961) ocean-only model to include the atmospheric coupling, which reveals  
82 bistates that depend on the forcing timescale. For sub-millennial timescale, they are the ones



83 seen in coarse-grain numerical models, but for orbital timescales, the bistates are set by an en-  
84 tropy principle of the nonequilibrium thermodynamics (NT) and hysteresis can be triggered by  
85 the modulated orbital forcing.

86 Justified on both physical and observational grounds, we posit therefore that the cou-  
87 pled climate system constitutes the intermediary of the orbital forcing of the glacial cycles. Di-  
88 ametrically opposite to numerical models that often add poorly constrained physical elements  
89 to improve the simulations, our approach seeks to isolate minimal physics that may account for  
90 the observed phenomenon. As organized below, our theory consists of two parts: the elucidation  
91 of the inner working of a coupled climate system, and the filtering of the orbital forcing  
92 through this system to generate glacial cycles; the two parts are contained in Sect. 2 and 3, re-  
93 spectively. In Sect. 4, we summarize the essence of the theory and discuss how it may resolve  
94 many Pleistocene puzzles of the glacial cycles.

## 95 **2 Coupled climate model**

96 In striving for minimal physics, we consider a model configuration of the North Atlan-  
97 tic as sketched in Fig. 1, for which both ocean and atmosphere are divided into warm and cold  
98 masses by mid-latitude fronts and an ice sheet may form on the adjacent continental strip ter-  
99 minating at the Arctic Ocean (all symbols are defined in Appendix B). The ocean is heated by  
100 the annual absorbed shortwave (SW) flux, which heats the overlying air by the convective and



101 net longwave (LW) fluxes, the latter assumed spatially uniform. Since glacial cycles are domi-  
102 nated by the subpolar temperature, the model variables pertain to the cold-box deviations from  
103 the global-means, the latter assumed known.

## 104 2.1 Regime diagram

105 The derivation of the model is provided in Ou (2018) to which the readers are referred  
106 for details. Sufficing for the present purpose, we shall summarize the model solution and its  
107 underlying physics via a regime diagram shown in Fig. 2 whereby the cold-box deviations  
108 from the global-means are plotted against the MOC ( $K$ ). All variables have been nondimen-  
109 sionalized (hence primed) whose scaling definitions and their standard values are listed in Ap-  
110 pendix B. This regime diagram is drawn for a forcing of  $q' = 1$  and a global convective flux  
111 of  $\bar{q}'_c = .75$ .

112 For an infinite  $K$ , the ocean is homogeneous, so the cold-box SST deficit ( $T'$ ) is zero  
113 and the cold-box SAT ( $T'_a$ ) is colder by the global convective flux ( $\bar{q}'_c$ ). As  $K$  decreases, the  
114 subpolar water cools, which cools the overlying air and induces an atmospheric heat transport,  
115 the latter in turn weakens the ocean heat transport (the total heat transport is fixed by  $q'$ ) hence  
116 reduces the steepness of the temperature curves. There is however an upper limit to the atmos-  
117 pheric heat transport since the convective flux (the spacing between two temperature curves)  
118 cannot be negative, which occurs at

$$119 \quad T' = 2\bar{q}'_c. \quad (1)$$



120 Beyond this “convective bound”, the atmospheric heat transport has saturated, and the two  
121 temperature curves would merge to steepen at the same faster rate (inverse in  $K$ ).

122 Since the moisture transport is proportional to the atmospheric heat transport on ac-  
123 count of the Clausius-Clapeyron Equation (Ou 2007), the increasing atmospheric heat transport  
124 with decreasing  $K$  implies a salinity deficit ( $S'$ ) that increases faster than temperature deficit  $T'$   
125 before the convective bound, but at the same rate afterward when the atmospheric heat  
126 transport has saturated. The disparate slopes of the two curves result in a density surplus ( $\rho' =$   
127  $T' - S'$ ) that has opposite slopes straddling the convective bound, the latter thus divides the cli-  
128 mate regime into warm and cold branches, a robust outcome of the atmospheric coupling.

129 To specify the climate state from the continuum of these curves, one needs a constraint  
130 on the MOC, which is customarily assumed linear in the density surplus (Stommel 1961;  
131 Marotzke and Stone 1995) as indicated in the thick dashed line. The proportional constant (the  
132 inverse of the slope) is referred as the admittance and the line, the admittance line whose inter-  
133 sects with the density curve then specify the climate state. For the admittance chosen, the  
134 ocean has two stable states (solid ovals) enclosing an unstable saddle point (open oval), a bista-  
135 bility that is predicated on the convective bound hence the ocean/atmosphere coupling.

136 The presence of the cold state has been demonstrated by coupled numerical models  
137 (Manabe and Stouffer 1988), which indeed is characterized by vanishing convective flux over  
138 the subpolar water (their Fig. 17), as predicted by the convective bound. As its further compu-  
139 tational support, coupled models have shown hysteresis when the freshwater flux is perturbed



140 (Rahmstorf et al. 2005), which can be readily gleaned from the regime diagram. In these nu-  
141 merical models, which do not resolve eddies, the admittance depends sensitively on the diapyc-  
142 nal diffusivity --- a highly uncertain property of the ocean, which is in effect finely tuned to  
143 replicate the observed state (Rahmstorf et al. 2005). Moreover, the hysteresis caused by chang-  
144 ing radiative flux would be of the opposite sign of --- hence plays no part in --- the glacial cy-  
145 cles: a stronger radiative flux, for example, would raise the temperature curve (hence lower the  
146 density curve) to cause transition to the cold glacial instead of the warm interglacial. The cul-  
147 prit lies in the fixed admittance line of a *laminar* ocean, which can be justified only for short-  
148 term sub-millennial climate changes (Ou 2018), but must be relaxed for the orbital forcing of a  
149 *turbulent* ocean, as discussed next.

## 150 **2.2 Ocean bistates**

151 Differing from the laminar ocean of coarse-grain numerical models, the actual MOC is  
152 subjected to microscopic fluctuations associated with random eddy exchanges across the sub-  
153 tropical front. Applying the fluctuation theorem (Crooks 1999), Ou (2018) deduces that the  
154 admittance is not fixed but would evolve on millennial “entropy adjustment” time toward the  
155 maximum entropy production (MEP). The latter thus is a veritable generalization of the sec-  
156 ond fundamental law to nonequilibrium thermodynamics (NT, Ozawa et al. 2003), which has  
157 been applied previously to climate theories (Kleidon 2009). We blur the admittance line by a  
158 shaded cone to signify fluctuations, which thus would slowly pivot by increasing entropy pro-  
159 duction until it attains maximum marked by solid ovals.



160 As derived in Ou (2018), the warm MEP state is given by

161  $(T', K) = (q', 1/2),$  (2)

162 which defines our interglacial. As a crude check, applying standard parameters (Appendix B)  
163 yields a subpolar SST of  $6^{\circ}\text{C}$  and MOC of 14 Sv (for a basin width of  $6 \times 10^3 \text{ km}$ ), not unlike  
164 the present interglacial (Peixoto and Oort 1992; Macdonald 1998). It should be noted that  
165 since our MOC does not depend on the uncertain diapycnal diffusivity, it is more robust than  
166 that produced from coarse-grain numerical models.

167 The cold MEP on the other hand is characterized by low-salinity subpolar water at the  
168 freezing point  $T'_f$ , which is consistent with the observed one during last glacial maximum  
169 (LGM, CLIMAP 1976; Duplessy et al. 1992) hence referred as the glacial state. It is signifi-  
170 cant that although the subpolar water is at the freezing point, it remains ice-free; this is because  
171 the sea ice would curb the ocean heat loss to weaken the MOC hence the entropy production,  
172 in contradiction to the MEP. This deduction however applies only for timescales long com-  
173 pared with the millennial entropy adjustment time, so it does not preclude the extensive sea-ice  
174 formed at the onset of the Heinrich events or during the termination of ice ages (Broecker  
175 1994; Denton et al. 2010). Incidentally, since the sea ice has little thermal inertia hence can be  
176 present *only* when the water is at the freezing point, it can play no active role in regulating the  
177 glacial cycles, as previously conjectured (Gildor and Tziperman 2003).

178 **2.3 Hysteresis**



179           With the bistable MEP shown in Fig. 2, one readily discerns possible hysteresis when  
180 the orbital forcing varies, as illustrated in Fig. 3. Suppose one is at the warm state (solid cir-  
181 cles), a reduction in the absorbed solar flux would cool the temperature, as indicated by the  
182 solid arrows and open circles. When the temperature cools to below the convective bound (the  
183 horizontal dashed line marked  $2\bar{q}'_c$ ), the warm MEP no longer exists, and the ocean would en-  
184 ter the cold branch to propel toward the cold MEP. With (1) and (2), this cold transition  $q'_c$  oc-  
185 curs at

$$186 \quad q'_c \equiv 2\bar{q}'_c, \quad (3)$$

187 a function only of the global convective flux, an external parameter. Now suppose one is at the  
188 cold MEP of freezing point (solid squares), then a rising absorbed solar flux would raise the  
189 temperature curve to propel the cold state, as indicated by dashed arrows and open squares.  
190 The flattening of the admittance line combined with random fluctuations would vault the cli-  
191 mate state into the warm branch, followed by its propelling toward the warm MEP. As a con-  
192 servative upper bound on the warm transition  $q'_w$ , one may set a level admittance line to yield  
193 (Ou 2018, from his Eqs. 8 and 12)

$$194 \quad q'_w \equiv (1 + \mu)\bar{q}'_c, \quad (4)$$

195 where  $\mu$  is related to the moisture content of the atmosphere with a standard value of about .3.  
196 The above two thresholds define the bistable interval, which thus is a function only of the  
197 global convective flux. As the latter is decreasing during the Pleistocene cooling, its interplay



198 with the ice-albedo feedback and forcing modulation is seen later to produce varied glacial cy-  
199 cles.

### 200 **3 Glacial cycles**

201 To translate the possibility of climate hysteresis to particulars of glacial cycles, we need  
202 to first determine the relevant orbital forcing, as discussed next.

#### 203 **3.1 Orbital forcing**

204 Because the thermal inertia of the ocean has integrated the seasonal cycles, the relevant  
205 orbital forcing is the annual absorbed SW flux integrated over the subpolar water. Being the  
206 annual mean, the forcing is dominated by that over high latitudes, which is about an order  
207 greater than that over the tropics (Berger et al. 2007, Fig. 2.7); and then over high latitudes, the  
208 forcing is dominated by the summer insolation due both to the vanishing hence unvarying win-  
209 ter insolation (Berger 1988, Fig. 18b; Tricot and Berger 1988, Fig. 4b) and to the ice-albedo  
210 feedback that amplifies the summer forcing. Here we must stress the necessity of the ice-al-  
211bedo feedback in instituting the precession forcing without which its annual absorbed flux is  
212 zero because of the Kepler's law. To avert this stricture, Huybers (2006) has posited a higher  
213 but unspecified melt threshold in late Pleistocene, but with the forcing being the absorbed flux,  
214 the ice-albedo feedback provides a palpable and quantifiable effect, as estimated next.

215 In early Pleistocene before the activation of the ice-albedo feedback, the forcing would  
216 be limited to the obliquity on account of the Kepler's law. In late Pleistocene when the ice



217 sheet may grow to mid-latitudes, the ice-albedo range can be as large as 0.3 (CLIMAP 1976,  
218 Fig. 1), which would incur a range in the absorbed flux of  $150 W \cdot m^{-2}$  given the summer in-  
219 solation of  $500 W \cdot m^{-2}$ . Adding the obliquity range of  $40 W \cdot m^{-2}$  (Imbrie et al. 1993, Fig.  
220 1), the total range is  $190 W \cdot m^{-2}$ , which would be halved to  $95 W \cdot m^{-2}$  for the annual mean.  
221 Noting that integration over the subpolar water has allayed the latitudinal difference of the  
222 obliquity and precession forcing, and reduced by half the excess atmospheric attenuation over  
223 the global mean to about 10% (Tricot and Berger 1988, Fig. 1), hence both are neglected.

224 With the above estimates, we see that the late Pleistocene forcing is comparable to the  
225 Milankovitch insolation (Imbrie et al. 1993, Fig. 1), which thus may be taken as its proxy. We  
226 should stress that our use of the Milankovitch insolation is not because of its direct effect on  
227 the summer SAT or ablation, as widely applied, but because it mimics the annual absorbed flux  
228 that drives the SST, on which the summer SAT anchors.

### 229 **3.2 Summer SAT**

230 Since ice ablation is controlled by the summer SAT (Pollard 1980), the latter needs to  
231 be linked to the annual SAT determined from our climate model (Fig. 2). The increasing conti-  
232 nentality with latitudes induces strong latitudinal variation of both the annual and the seasonal  
233 SAT (Oerlemans 1980; Donohoe and Battisti 2013), which on the other hand is smoothed by  
234 the atmospheric motion. Because of these complications, it would seem intractable to deduce  
235 the summer SAT from the heat balance, and indeed its calculation by numerical models typi-  
236 cally involves adjusting the eddy diffusivity to replicate the observed SAT (Oerlemans 1980;



237 North et al. 1983). To circumvent such empirical tuning that necessarily degrades its progn-  
238 sis, we discern nonetheless a robust constraint on the summer SAT from observations (Peixoto  
239 and Oort 1992, Fig. 7.5), namely, it is near the freezing point at the edge of the Arctic Ocean,  
240 which can be reasoned below on physical grounds.

241 We first argue that the summer air cannot be consistently warmer than the freezing  
242 point since it would melt the perennial ice to contradict its emergence since about 3 million  
243 year ago (Clark 1982; Ravelo et al. 2004). We then argue that the summer air may not be con-  
244 sistent colder than the freezing point either since the Arctic Ocean would then freeze over,  
245 which contradicts an ice-free North Atlantic on account of the MEP (Sect. 2.2) and its entry  
246 into the Arctic Ocean that would maintain open coastal water.

247 At the southern end of the subpolar water, because of the much reduced continentality  
248 hence seasonal cycle, the summer SAT can be approximated by the annual SAT determined  
249 from our climate model. Having pinned down the two end points, we then invoke the atmos-  
250 pheric dynamics to connect them with a straight line, as seen in Fig. 4 (the thick solid line for  
251 the interglacial). The summer SAT would pivot about the freezing point at its northern end by  
252 the orbital forcing (the shaded cone) whereas for the glacial state, it would assume a uniform  
253 freezing point throughout the subpolar region hence overlay with the abscissa, a deduction that  
254 is consistent with LGM simulations (Clark et al. 1999, Fig. 2).

### 255 **3.3 Ice margin**



256 Assuming a constant lapse rate, the summer SAT profile shown in Fig. 4 also repre-  
257 sents the snowline whose intersect with the southern face of the ice sheet defines the equilib-  
258 rium line (EL). To determine the equilibrium line altitude (ELA)  $h_0$  hence the ice margin, we  
259 need to consider both the ice dynamics and the mass balance and, given the myriad processes  
260 involved, we shall retain only the minimal physics. For the ice dynamics, we assume a perfect  
261 plasticity with yield stress  $\tau_i$ , so the momentum balance yields (see Van der Veen 2013)

$$262 \quad d(h^2) = -c \cdot dx \quad (5)$$

263 where  $h$  is the height of the southern face of the ice sheet and  $c \equiv (g\rho_i)^{-1}(2\tau_i)$  with  $g$ , the  
264 gravitational acceleration and  $\rho_i$ , the ice density. We assume an ablation rate given by  $\nu T_i$   
265 where  $T_i$  is the ice surface temperature (above the freezing point) and  $\nu$ , an empirical constant  
266 (Pollard 1980), and let  $l_0$  and  $l$  be the  $x$ -coordinates of the EL and ice margin, respectively,  
267 then the annual ablation is, invoking (5),

$$268 \quad A_b = \int_{l_0}^l \nu T_i dx$$
$$269 \quad \approx \frac{\nu\gamma}{c} \int_{h_0}^0 (h_0 - h) d(h^2)$$
$$270 \quad = \frac{\nu\gamma}{3c} h_0^3, \quad (6)$$

271 which thus depends strongly on the ELA. The accumulation  $A_c$  is simply the moisture flux  
272 crossing the EL, which can be linked to the energy flux and the local moisture content --- both



273 are strongly constrained by the EL being at the freezing point (Ou 2007). For this reason, the  
274 accumulation would be largely insulated from the varying surface climate associated with the  
275 glacial cycles and even further removed from the changing latitudinal gradient of the summer  
276 insolation (Raymo and Nisancioglu 2003). Equating the ablation and accumulation, we derive  
277 the ELA

$$278 \quad h_0 \approx \left(\frac{3cA_c}{\nu\gamma}\right)^{1/3}, \quad (7)$$

279 which is estimated next.

280 Given the above constraint on the moisture flux by the freezing-point EL, we shall ap-  
281 proximate it by the summer moisture flux into the Arctic Ocean as the latter is rimmed by the  
282 freezing temperature. Based on Peixoto and Oort (1992, Fig. 12.21), this moisture flux is esti-  
283 mated to be  $A_c = 2 \times 10^5 m^2/y$ . Setting additionally  $\tau_i = 1$  bar (Van der Veen 2013) so that  
284  $c = 22 m$ ,  $\nu = 2 m (y \text{ } ^\circ\text{C})^{-1}$  (Pollard 1980) and  $\gamma = 6 \text{ } ^\circ\text{C km}^{-1}$ , we estimate  $h_0 = 1 km$ ,  
285 which is like that observed over the current Greenland ice sheet (Oerlemans 1991). This ELA  
286 implies an ice margin aligned with the summer isotherm of  $6 \text{ } ^\circ\text{C}$ , which may explain why the  
287 Greenland is largely ice-covered as this isotherm is located near its southern edge (North et al.  
288 1983, their Fig. 5b). North et al. (1983) has pegged the ice margin at the summer  $0 \text{ } ^\circ\text{C}$  iso-  
289 therm, which is deficient since there would be no summer ablation to counter the yearly accu-  
290 mulation, and then it would imply an ice-free Greenland, contradicting the observed one. Alt-  
291 hough the parameters in (7) are uncertain, the sensitivity is somewhat dampened by the 1/3



292 power, so our pegging of the ice margin with the summer isotherm of  $O$  ( $10^{\circ}\text{C}$ ) should gener-  
293 ally apply.

294 With the above estimates, the interglacial ice sheet (the medium shade in Fig. 4) ex-  
295 tends about  $1/3$  toward the subtropical front and the ice margin would vary linearly with the  
296 forcing (the dark cone), which we shall categorize as the polar ice cap, such as the present  
297 Greenland ice sheet. For the glacial state, the summer SAT is at the freezing point throughout  
298 the subpolar region, so the ice sheet would extend to the subtropical front (lightly shaded) cor-  
299 responding to the Laurentide ice sheet (LIS). It is seen therefore that the ocean bistates would  
300 translate to ice sheet of vastly different sizes, resulting in a large ice-volume signal regardless  
301 the forcing range so long as the bistable thresholds are crossed.

302 The large ice signal thus differs fundamentally from that associated with the inherent  
303 ice bistates of polar ice cap and ice-free state. And then with the surface snowline hovering  
304 around the edge of the Arctic Ocean (Sect. 3.2), it is readily seen from Fig. 4 that the ice-free  
305 state is unstable: a slight southward migration of the snowline would produce a finite polar ice  
306 cap via the mass-balance feedback, yet its erasure requires a warming of about  $6^{\circ}\text{C}$  accompa-  
307 nied by a northward snowline migration of  $O$  (1000 km) (the dashed line), an unlikely pertur-  
308 bation. This deduction of a stable polar ice cap is consistent with numerical studies, which  
309 show that the current Greenland ice sheet would melt away only with more than  $5^{\circ}\text{C}$  warming  
310 (Letreguilly et al. 1991).

### 311 **3.4 Mid-Pleistocene transition (MPT)**



312           The MPT from the obliquity- to the eccentricity-dominated ice-volume cycles has  
313 posed a significant challenge to the astronomical theory since the Milankovitch insolation dis-  
314 plays no discernible change. Besides the decreasing CO<sub>2</sub> trend that we have critiqued in Sect.  
315 1, the changing substrate geology has also been postulated (Clark et al. 2006); both however  
316 represent extraneous physics to the Pleistocene cooling, which is likely tectonic in origin (Rud-  
317 diman and Raymo 1988). Instead, as a direct consequence of the Pleistocene cooling, we argue  
318 that the global convective flux would be lowered. This is because the cooling implies a drier  
319 air hence a smaller downward LW flux (Ou 2001, Fig. 2), which then requires a smaller global  
320 convective flux for the ocean heat balance --- all else being equal. Significantly, the global  
321 convective flux is precisely what sets the bistable interval of our climate, and we shall see its  
322 interplay with the ice-albedo feedback provides a natural account of the MPT. There are of  
323 course no proxy data for the convective flux, so its reduction during the Pleistocene cooling  
324 can only be supported by the seeming robust physics and its potency in explaining the MPT.

325           Based on the discussion to follow, we show the Pleistocene evolution of the forcing and  
326 ice signals in Fig. 5 (time proceeds to the left), which consists of three stages and their transi-  
327 tions. The three stages correspond roughly to the three periods depicted in Imbrie et al. (1993,  
328 their Fig. 3) and the two transitions, the early and middle Pleistocene transitions identified by  
329 Lisiecki and Raymo (2007). For illustrative purpose, the forcing is represented by a shaded en-  
330 velop (neglecting the obliquity and precession periods) centered on the thick dashed line (re-  
331 ferred as the mean forcing) and the forcing markers along the right ordinate are merely indica-  
332 tive. The vertical bars are bistable intervals spanned by the cold and warm thresholds given in



333 (3) and (4), which are slowly approaching the horizontal axis as the global convective flux de-  
334 creases during the Pleistocene cooling.

335 At Stage 1 in the warm early Pleistocene, there is little ice-albedo feedback to effectuate  
336 the precession forcing (Sect. 3.1), so the ice signal is simply that of the polar ice cap line-  
337 arly perturbed by the obliquity forcing. This Stage 1 can be identified with the time span be-  
338 fore 1.5 Ma (million years ago), which thus is dominated by interglacial cycles at the 41-ky  
339 obliquity period. The continuing cooling would enhance the ice-albedo feedback hence the  
340 precession forcing, both attaining maxima when the deepest precession trough  $q'_{max}$  exceeds  
341 the cold threshold (3) to generate the glacial state. This being the precondition of the full-  
342 fledged precession forcing that defines Stage 2, we thus set (from Eq. (3))

$$343 \quad \bar{q}'_c = q'_{max}/2 \quad (8)$$

344 as a crude marker for the early Pleistocene transition (EPT) from Stage 1 to 2. As the glacial  
345 state is not yet generated prior to Stage 2, the ice signal varies linearly with the forcing, so the  
346 EPT can be identified with the time span of 1.5-1 Ma based on the observed ice signal (Imbrie  
347 et al. 1993, Fig. 3). Since the ice-albedo feedback primarily depresses the precession troughs  
348 but not its peaks, the precession broadening of the forcing envelop should manifest in the deep-  
349 ening of its mean, as indicated in the figure. This should reflect in the SST and ice volume,  
350 which indeed is a pronounced feature in observations (Imbrie 1993, Fig. 3; McClymont et al.  
351 2013, Fig. 8).



352           Stage 2 is defined by full-fledged precession forcing (hence modulated by eccentricity)  
353 when the bistable centerline still lies below the mean forcing. Although glacial states are now  
354 generated by deep precession troughs during high eccentricity, they are invariably nullified by  
355 the next precession peaks, and the phase span of this bistate oscillation is lighter shaded to  
356 symbolize the presence of substantial ice sheet. Outside this phase span, the precession  
357 troughs no longer clear the cold threshold so there is only interglacial ice signal (dark-shaded)  
358 varying linearly with the precession forcing. We shall identify Stage 2 with the time span of 1  
359 to .7 Ma, which thus is characterized by the emerging glacial/interglacial (G/IG) cycles at the  
360 21-ky precession period.

361           The continuing cooling causes the bistable centerline to rise above the mean forcing,  
362 which defines Stage 3. There are again bistate oscillations during high eccentricity indicated  
363 by the lighter shade, outside of which however the precession peaks no longer clear the warm  
364 transition, so the glacial state would persist through the low eccentricity to allow the full  
365 growth of the ice sheet to mid-latitudes. This prolonged glacial state of extensive ice sheet de-  
366 fines the ice age, which is symbolized by the unshaded forcing envelop. Stage 3 thus is domi-  
367 nated by ice-age cycles at the 100-ky eccentricity period, which corresponds to the observed  
368 time span of .5 Ma to the present. It is seen from the figure that since the ice age terminates  
369 and commences at the same warm threshold, there is no need to invoke differing physics for  
370 their occurrences, as suspected previously (Broecker et al. 1985; Raymo 1997).

371           The MPT from Stage 2 to 3 thus spans the time interval of .7-.5 Ma when the bistable  
372 centerline crosses the mean forcing, which can be seen from (3)-(4) to be given by



$$373 \quad \bar{q}'_c = \frac{2}{3+\mu} \bar{q}^t. \quad (9)$$

374 While the precession is not of zero period, nor are the transitions sharply defined, which may  
375 spread over several hundred thousand years, so the above criteria may provide crude markers  
376 of the two transitions. For a cursory check of these criteria, we set a mean forcing of  $100 \text{ W} \cdot$   
377  $\text{m}^{-2}$  and a forcing amplitude of  $50 \text{ W} \cdot \text{m}^{-2}$ , then EPT and MPT would be marked by global  
378 convective fluxes of  $75 \text{ W} \cdot \text{m}^{-2}$  and  $61 \text{ W} \cdot \text{m}^{-2}$ , respectively. Since the global convective  
379 flux prior to the Pleistocene should be like the present interglacial hence of  $O(100 \text{ W} \cdot \text{m}^{-2})$ ,  
380 and given the Pleistocene cooling of order  $10^\circ\text{C}$  (Ruddiman and Raymo 1988, Fig. 3), the  
381 downward LW flux as well as the global convective flux can be reduced by  $50 \text{ W} \cdot \text{m}^{-2}$  (Ou  
382 2001, Fig. 2), so the above criteria of the EPT and MPT are readily met to support their expla-  
383 nation by the model.

384 The deduced three stages represent a shift of the power spectra from that dominated by  
385 the obliquity to the emergence of the precession to that dominated by the eccentricity, as in-  
386 deed seen in the observed ones (Imbrie et al. 1993, Fig. 3; Berger 1988, Fig. 16; McClymont et  
387 al. 2012, Fig. 6, top panel). The emergence of the precession signal in Stage 2 would shorten  
388 the interglacial and enhance the saw-tooth asymmetry, which are among the defining features  
389 of the observed EPT (Lisiecki and Raymo 2007).

### 390 **3.5 Timeseries**



391 For a visualization of the temporal signals, we next present timeseries calculated from  
392 the model for Stage 2 and 3 (no need to show Stage 1 characterized by interglacial signals lin-  
393 ear in the obliquity forcing). We have argued in Sect. 3.1 that with the full operation of the  
394 ice-albedo feedback hence precession forcing, the model forcing can be approximated by the  
395 Milankovitch insolation, which is set to

$$396 \quad q' = \bar{q}^t + \sum_{i=1}^3 a_i \cos \omega_i t. \quad (10)$$

397 where the time-mean forcing (the cold-box deficit) is  $\bar{q}^t = 100 W \cdot m^{-2}$ , obliquity ( $i = 1$ ) has a  
398 period of 41 ky and amplitude  $a_1 = 10 W \cdot m^{-2}$ , and precessions ( $i = 2$  and 3) have periods  
399 of 18.5 and 23 ky with amplitudes  $a_2 = a_3 = 20 W \cdot m^{-2}$ , respectively, which renders a 21 ky  
400 precession modulated by 95 ky eccentricity. The total range of our forcing thus is  $100 W \cdot$   
401  $m^{-2}$ , as is the observed Milankovitch insolation (Berger et al. 1996, Fig. 3a).

402 With the above forcing, our climate model discussed in Sect. 2 would produce (time-  
403 varying) equilibrium SST and ice margin, which are now subscripted “ $e$ ” for distinction. To  
404 calculate the timeseries, we apply the relaxation equations:

$$405 \quad dT'/dt = (T_e' - T')/\tau_T, \quad (11)$$

406 and

$$407 \quad dl/dt = (l_e - l)/\tau_l \quad (12)$$



408 where the time constant for temperature is the entropy adjustment time set at 1 ky and  $\tau_l$  is the  
409 time constant for the ice margin, which we distinguish between its advance and retreat (Weert-  
410 man 1964). The ice advance is limited by the accumulation: for a snowfall of 0.3 m/y for ex-  
411 ample (Ohmura and Reeh 1991), to build up an ice sheet to 3 km high takes about 10 ky, which  
412 is thus set to be the advance time constant. The ice retreat on the other hand can be much  
413 faster: for 2 degrees warming for example, the melt rate is 4 m/y (Pollard 1980), which is an  
414 order greater than the accumulation, we thus set the retreat time constant to be 1 ky. The relax-  
415 ation equations being linear, using different time constants merely affects the lag of the curves  
416 but produces no material difference.

417 We show in Fig. 6 timeseries and power spectra of the forcing (solid line), the subpolar  
418 SST (dashed) and ice margin (dotted) for Stage 2 and 3 of Fig. 5 (the MATLAB script is pro-  
419 vided in Appendix C). The initial condition is the warm MEP state and integration is carried  
420 forward for 400 ky; since the glacial cycles are largely repetitive, we plot only the last cycle,  
421 the power spectra are however calculated for the full 400-ky timeseries. The upper axis repre-  
422 sents the global-mean absorbed flux and SST. The forcing, being referenced to the former, is  
423 expressed in its temperature equivalent ( $100 \text{ W} \cdot \text{m}^{-2}$ , for example, would convert to  $8 \text{ }^\circ\text{C}$ , see  
424 Appendix B), the global-mean SST is set to  $14 \text{ }^\circ\text{C}$ , the ice margin is its fractional extension  
425 into the subpolar region, and the shaded bar indicates the bistable interval.

426 It is seen that the forcing timeseries resembles the observed Milankovitch insolation  
427 (Berger et al. 1996, Fig. 3a) and expectedly contains no power at the eccentricity period. The  
428 timeseries for Stage 2 (Fig. 6a) show that only one precession trough during high eccentricity



429 has exceeded the cold threshold to generate the glacial state characterized by freezing-point  
430 SST and an ice sheet extending about half-way into the subpolar. Other than this single glacial  
431 episode lasting half the precession period, the rest of the timeseries are the interglacial SST that  
432 tracks the forcing with slight delay and negligible ice-cap cycles. Given the short duration of  
433 the glacial state, the SST and ice-margin spectra show no appreciable power at the eccentricity  
434 period, consistent with the observed spectra (Imbrie et al. 1993, Fig. 3).

435         The timeseries of Stage 3 differs qualitatively from that of Stage 2. There are episodes  
436 of interglacial during high precession peaks, which however always revert to glacial state at the  
437 next precession trough and then there is only glacial state spanning the low eccentricity, its  
438 long duration allows the ice sheet to grow to mid-latitudes. Although the SST and ice-margin  
439 spectra retain the precession and obliquity peaks as Stage 2, they show a strong eccentricity  
440 peak absent in Stage 2. This sharp contrast is consistent with the observed spectra (Imbrie et  
441 al. 1993, Fig. 3).

442         The ice signal bears sufficient resemblance to the last ice age cycle to allow marking of  
443 the corresponding marine isotope stages (MIS), as indicated in the figure, whose observational  
444 features thus may be interpreted by the model physics. According to our model, the cold sub-  
445 stages are characterized by freezing-point subpolar water, which is consistent with the ob-  
446 served expansion of the polar watermass and appearance of the polar species (McManus et al.  
447 1994). Being a glacial state, the ice growth to the mid-latitudes is only limited by the duration  
448 of the half precession period, which has nonetheless reached half-way to the subtropical front.  
449 This modelled ice sheet is consistent with its observational estimate (Ruddiman et al. 1980;



450 Chapman and Shackleton 1999), which is also supported by Ice-rafted debris (IRD) events pre-  
451 conditioned on a large ice sheet (McManus et al. 1994). In our interpretation, all substages are  
452 generically similar with the glacial at the cold substages reversed by the next precession peak  
453 to the interglacial warm substages, as seen in their comparable temperature (Berger 1979, Fig.  
454 8). It is the MIS 4 that represents the onset of the ice age (Ruddiman et al. 1980) as the suc-  
455 ceeding precession peak fails to clear the warm threshold, resulting in prolonged coldness and  
456 an ice sheet extending to mid-latitudes as manifested in the LIS. It is noted that the ice margin  
457 is saw-toothed even within one precession trough due solely to the disparate advance and re-  
458 treat rates, and this asymmetry is strongly amplified for the ice ages due to the ice growth  
459 through the low eccentricity before the abrupt ice retreat.

#### 460 **4 Discussion**

461 The central tenet of our theory is that the ocean is the intermediary of the orbital forcing  
462 of the global ice, as strongly argued by Broecker and Denton (1989). Since the ocean is heated  
463 by the annual absorbed flux integrated over the subpolar water, it naturally filters out latitudi-  
464 nal difference of the obliquity and precession forcing --- except the latter would become effec-  
465 tive when the ice-albedo feedback is activated during the Pleistocene cooling. As such, the  
466 forcing is dominated by the obliquity component in the early Pleistocene, but can be approxi-  
467 mated by the Milankovitch insolation in the late Pleistocene. The use of the latter in our model  
468 thus is not because of its direct effect on the summer SAT and ablation, but because it mimics  
469 the late-Pleistocene forcing of the ocean.



470           While there can be inherent bistability of finite ice sheet and ice-free state (Weertman  
471 1961), we argue that the ice-free state is untenable in Pleistocene with the emergence of the  
472 Arctic perennial ice about 3 Ma, as also attested by the current Greenland ice sheet and the  
473 substantial obliquity signal even in the early Pleistocene. Rather, we posit that bistability of  
474 the ice simply reflects that of the subpolar ocean, which in a coupled and NT climate system  
475 may exhibit bistable warm and freezing-point temperature. Through its effect on the summer  
476 SAT that controls the ice margin, this bistability would translate to that of the ice characterized  
477 by polar ice cap and an ice sheet extending to mid-latitudes. The vast difference of these ice  
478 bistates produces strong ice signal regardless the forcing perturbation so long as the bistable  
479 thresholds are crossed.

480           The bistable interval of the coupled climate is linked to the global convective flux,  
481 which would be lowered during the Pleistocene cooling that produces drier atmosphere; its in-  
482 terplay with the ice-albedo feedback leads to three stages of the ice cycles, as well discerned in  
483 observations (Imbrie et al. 1993; Lisiecki and Raymo 2007). In the warm early Pleistocene be-  
484 fore appreciable ice-albedo feedback hence the precession forcing, the ice cycles are simply  
485 that of the polar ice cap perturbed linearly by the obliquity (Stage 1). The continuing cooling  
486 would enhance the ice-albedo feedback hence the precession forcing; while the precession  
487 troughs during high eccentricity may induce the glacial state, it is nullified by the next preces-  
488 sion peak, resulting in G/IG cycle at the precession period (Stage 2). With further cooling, the  
489 precession peaks may no longer clear the warm threshold, so the glacial state would last  
490 through the low eccentricity, resulting in ice-age cycles paced by the eccentricity (Stage 3).



491 The transitions between the three stages correspond to the EPT and MPT discerned in Lisiecki  
492 and Raymo (2007), which are now assigned specific markers that can be crossed during the  
493 Pleistocene cooling.

494 It should be noted that in our formulation, there are only bistable glacial and interglacial  
495 states, the ice age is merely the glacial state that lasts through low eccentricity, there is no need  
496 for a third full-glacial state as posited by Paillard (1998), who has not provided dynamical ba-  
497 sis for such a state nor the transition rules among his tri-states. Since an interactive MOC is  
498 key to the ocean bistability, numerical models that fix the SST or assume a slab ocean (North  
499 et al. 1983; Abe-Ouchi et al. 2013) obviously cannot capture the ocean effect on the climate.  
500 While coarse-grained coupled models have produced ocean hysteresis (Rahmstorf et al. 2005),  
501 it is opposite in sign to that of the glacial cycle, the reason being, without resolving eddies, the  
502 MOC is constrained by a fixed diapycnal diffusivity (Sect. 2.1). Deprived of the proper ocean  
503 hysteresis, numerical calculations of the glacial cycles are compelled to prescribe a CO<sub>2</sub> trend  
504 or an orbital-period CO<sub>2</sub> in augmenting the glacial signal (Berger et al. 1999; Ganopolski and  
505 Calov 2011; Willeit et al. 2019) --- both need further justification (Sect. 1). To properly con-  
506 strain the MOC requires resolving ocean eddies, which poses a daunting challenge to numeri-  
507 cal models because of the long time-integration needed, but a phenomenological approach of  
508 coding the entropy production tendency, say, via a variable eddy diffusivity, may remain feasi-  
509 ble.

## 510 **5 Resolving glacial puzzles**



511 Our theory provides a single dynamic framework that may resolve seemingly unrelated  
512 Pleistocene puzzles of the glacial cycles, as further expounded below. The reason that there is  
513 only 41-ky obliquity cycles in the early Pleistocene is because, without the ice-albedo feed-  
514 back, the precession has no effect on the annual absorbed flux on account of the Kepler's law,  
515 our theory thus may resolve the "41-ky" problem. That such flux being the relevant forcing  
516 has rendered moot some previous solutions to the 41-ky problem, such as Raymo and Ni-  
517 sancioğlu (2003) or Raymo and Huybers (2008). The MPT to the 100-ky ice-age cycle occurs  
518 when the Pleistocene cooling has allowed the glacial state to last through the low eccentricity  
519 hence paced by the latter; its strong signal is caused by disparate ice bistates between the polar  
520 ice cap and an ice sheet extending to mid-latitudes; our model thus may resolve the "100-ky"  
521 problem.

522 So long as the ice-age pacing is enabled by the shorter period 100-ky eccentricity, the  
523 strength of the ice-age cycle is no longer affected by the 400-ky eccentricity even though it has  
524 greater amplitude (Berger and Loutre 1991, Fig. 4a), the theory thus may resolve the "400-ky"  
525 problem (Imbrie et al. 1993; Berger and Loutre 2010). In fact, it is seen from Fig. 5 that a  
526 smaller eccentricity would produce a longer-lasting ice age to augment the ice signal, which  
527 may explain why the 100-ky signal is gaining strength when the eccentricity is decreasing  
528 (Clark et al. 1999, Fig. 6b) or why the lower eccentricity at Stage 11 is accompanied by higher  
529 100-ky signal; the latter often dubbed the "Stage-11" problem (Imbrie et al. 1993, Fig. 2).

530 Since the onset and termination of the ice ages are threshold phenomena, both can be  
531 off by one precession period depending on the precise timing, the ice-age cycles thus may vary



532 between 80- and 120-ky (Raymo et al. 1997) to resolve the “variable termination” problem.  
533 Since the summer insolation anomalies are out-of-phase between hemispheres, their synchro-  
534 nous glacial cycles have posed a significant puzzle (Broecker and Denton 1989), but since the  
535 relevant orbital forcing is the annual absorbed flux, it has naturally removed this hemispheric  
536 difference. Then with the northern ice sheet dominating the response, it would feed back onto  
537 the global balance to synchronize the Antarctic climate, as suggested by the latter’s slight lag  
538 (a few millennia) from the Milankovitch insolation (Kawamura et al. 2007); the model thus  
539 may possibly resolve this “polar synchronization” problem.

#### 540 **Appendix A: Acronyms**

541	EL	Equilibrium line
542	ELA	Equilibrium-line altitude
543	EPT	Early Pleistocene transition
544	G/IG	Glacial/interglacial
545	IRD	Ice-rafted debris
546	Ka	Thousand years ago
547	Ky	Thousand years
548	LGM	Last glacial maximum
549	LIS	Laurentide ice sheet
550	LW	Long-wave
551	Ma	Million years ago
552	MEP	Maximum entropy production



553	MIS	Marine isotope stage
554	MOC	Meridional overturning circulation
555	MPT	Mid-Pleistocene transition
556	NT	Nonequilibrium thermodynamics
557	SAT	Surface-air temperature
558	SST	Sea-surface temperature
559	SW	Short-wave

#### 560 **Appendix B: Symbols and standard values**

561	$A_b$	Ablation rate
562	$A_c$	Accumulation rate ( $= 2 \times 10^5 m^2/y$ )
563	$C_{p,o}$	Specific heat of ocean ( $= 4.2 \times 10^3 J Kg^{-1} K^{-1}$ )
564	$g$	Gravitational acceleration ( $= 9.8 m \cdot s^{-2}$ )
565	$h$	Ice-surface height
566	$h_0$	ELA
567	$K$	Mass exchange rate of MOC
568	$[K]$	Scale of $K$ ( $= \alpha^* L (2\rho_o C_{p,o})^{-1} = 4.5 m^2/s$ )
569	$l$	$x$ -coordinate of ice margin
570	$l_e$	Equilibrium $l$
571	$l_0$	$x$ -coordinate of ELA
572	$L$	Latitudinal span of cold box ( $= 3 \times 10^3 km$ )



573	$q'$	Cold-box deficit of absorbed solar flux
574	$\bar{q}^t$	Long-term mean of $q'$ ( $= 100 W \cdot m^{-2}$ )
575	$[q']$	scale of $q'$ ( $= \bar{q}^t = 100 W \cdot m^{-2}$ )
576	$\bar{q}'_c$	Global convective flux
577	$a_i$	Amplitudes of Milankovitch insolation
578	$q'_c$	Cold-transition threshold
579	$q'_w$	Warm-transition threshold
580	$S'$	Cold-box salinity deficit
581	$[S']$	Scale of $S'$ ( $= \alpha[T']/\beta=1.79$ )
582	$\bar{T}$	Global-mean SST ( $= 14^0C$ )
583	$T'$	Cold-box SST deficit
584	$[T']$	Scale of $T'$ ( $= [q']/\alpha^* = 8^0C$ )
585	$T'_a$	Cold-box SAT deficit (from global-mean SST)
586	$T'_e$	Equilibrium $T'$
587	$T'_f$	Freezing-point temperature
588	$T'_i$	Ice surface temperature
589	$\alpha$	Thermal expansion coefficient ( $= 1.7 \times 10^{-4} \cdot ^0C^{-1}$ )
590	$\alpha^*$	Surface transfer coefficient ( $= 12.5 W \cdot m^{-2} \cdot ^0C^{-1}$ , Ou 2018)
591	$\beta$	Saline contraction coefficient ( $= 7.6 \times 10^{-4}$ )
592	$\gamma$	Lapse rate ( $= 6^0C/km$ )
593	$\rho'$	Cold-box density surplus



- 594  $[\rho']$  Scale of  $\rho'$  ( $= \rho_o \alpha [T'] = 1.36 \text{ Kg} \cdot \text{m}^{-3}$ )  
595  $\rho_i$  Ice density ( $= 0.9 \times 10^3 \text{ Kg} \cdot \text{m}^{-3}$ )  
596  $\rho_o$  Ocean density ( $= 10^3 \text{ Kg} \cdot \text{m}^{-3}$ )  
597  $\tau_i$  Yield stress ( $= 1 \text{ bar}$ )  
598  $\tau_l$  Ice-sheet time constant ( $= 1/10 \text{ ky}$  for retreat/advance)  
599  $\tau_T$  MEP-adjustment time ( $= 1 \text{ ky}$ )  
600  $\mu$  Moisture-content parameter ( $= 0.3$ )  
601  $\nu$  Ice-melt parameter ( $= 1.6 \text{ m} \cdot \text{y}^{-1} \cdot ^\circ\text{C}^{-1}$ )

## 602 **Appendix C: MATLAB script**

```
603 % assign parameters
604 dt=1;tmax=400;t=(0:dt:tmax);m=length(t);m2=m/2;
605 fre1=2*pi/41;fre2=2*pi/18.5;fre3=2*pi/23;
606 amp1=0.8;amp2=1.6;amp3=1.6;
607 pha1=0;pha2=0;pha3=0;
608 dtemp1=4.5;qmean=8;dtemp2=2*dtemp1;
609 tempf=14;temprange=10;
610 tausst=1;tauac=10;tauab=1;
611 gamma=0.3;
612 % set arrays
613 sste=zeros(1,m);sst=zeros(1,m);
614 icee=zeros(1,m);ice=zeros(1,m);
615 qprime=zeros(1,m);q=zeros(1,m);
616 % initialize with interglacial state (ig)
617 ig=true;
618 for i=1:(m-1)
619     %insolation
620     qprime(i)=amp1*cos(fre1*t(i)+pha1)...
621         +amp2*cos(fre2*t(i)+pha2)...
622         +amp3*cos(fre3*t(i)+pha3);
623     q(i)=qmean-qprime(i);
624     %calculate equilibrium temperature and ice margin
625     if ig
626         sste(i)=q(i);
627         slt1=q(i)/2+dtemp1;
```



```
628     icee(i)=1-(tempf-slt1)/temprange;
629     if q(i)>=dtemp2
630         ig=false;
631     end
632 end
633 if ~ig
634     sste(i)=tempf;
635     icee(i)=1;
636     if q(i)<=(1+gamma)*dtemp1
637         ig=true;
638     end
639 end
640 %time integration using runge-kutta scheme
641 sst(1)=sste(1);ice(1)=icee(1);
642 sst(i+1)=runge(sste(i),sst(i),dt,tausst);
643 ice(i+1)=rungeice(icee(i),ice(i),dt,tauac,tauab);
644 end
645 %rescale ice margin
646 ice5=5*(1-ice);
647 ice10=10*ice;
648 q=tempf-q;
649 sst=tempf-sst;
650 figure
651 % plot timeseries
652 subplot(2,1,1)
653 plot(t,q,'-k',t,sst,'--k',t,ice5,':k')
654 axis([60,170,0,tempf])
655 set(gca,'XDir','reverse');
656 %legend('q','sst','ice5','location','southeast')
657 title({'gla6a:','amp1=',num2str(amp1),'amp2=',num2str(amp2),...
658     ',amp3=',num2str(amp3),'pha1=',num2str(pha1),...
659     ',pha2=',num2str(pha2),'pha3=',num2str(pha3)];...
660     ['dtemp1=',num2str(dtemp1),'qmean=',num2str(qmean),...
661     ',temprange=',num2str(temprange),'tauac=',num2str(tauac),'tauab=',num2str(tauab),...
662     ',gamma=',num2str(gamma)]})
663 xlabel('t (ky)');ylabel('temp')
664 % plot power spectra
665 yq=fft(qprime);
666 yq=fftshift(yq);
667 sstprime=sst-mean(sst);
668 ysst=fft(sstprime);
669 ysst=fftshift(ysst);
670 iceprime=ice10-mean(ice10);
```



```
671 yice=fft(iceprime);
672 yice=fftshift(yice);
673 f=(-m/2:m/2-1)/(dt*m);
674 powerq=2*abs(yq).^2/(m*m);
675 powersst=2*abs(ysst).^2/(m*m);
676 powerice=2*abs(yice).^2/(m*m);
677 subplot(2,1,2)
678 plot(f,powerq,'-k',f,powersst,'--k',f,powerice,':k')
679 axis([0,0.1,0,3])
680 legend('q','sst','ice')
681 xlabel('freq (cycles/ky)');ylabel('power')
682
683 % runge-kutta scheme for sst
684 function y=runge(x,r,dt,tau)
685     heating=@(r) (x-r)/tau;
686     k1 = heating(r);
687     k2 = heating(r+0.5*dt*k1);
688     k3 = heating(r+0.5*dt*k2);
689     k4 = heating(r+dt*k3);
690     y = r+1/6*dt*(k1+2*k2+2*k3+k4);
691
692 % runge-kutta scheme for ice margin
693 function y=rungeice(x,r,dt,tauac,tauab)
694     msign=x-r;
695     if msign<0
696         tau=tauab;
697     else
698         tau=tauac;
699     end
700     mbalance=@(r) (x-r)/tau;
701     k1 = mbalance(r);
702     k2 = mbalance(r+0.5*dt*k1);
703     k3 = mbalance(r+0.5*dt*k2);
704     k4 = mbalance(r+dt*k3);
705     y = r+1/6*dt*(k1+2*k2+2*k3+k4);
706
```

707 **References**



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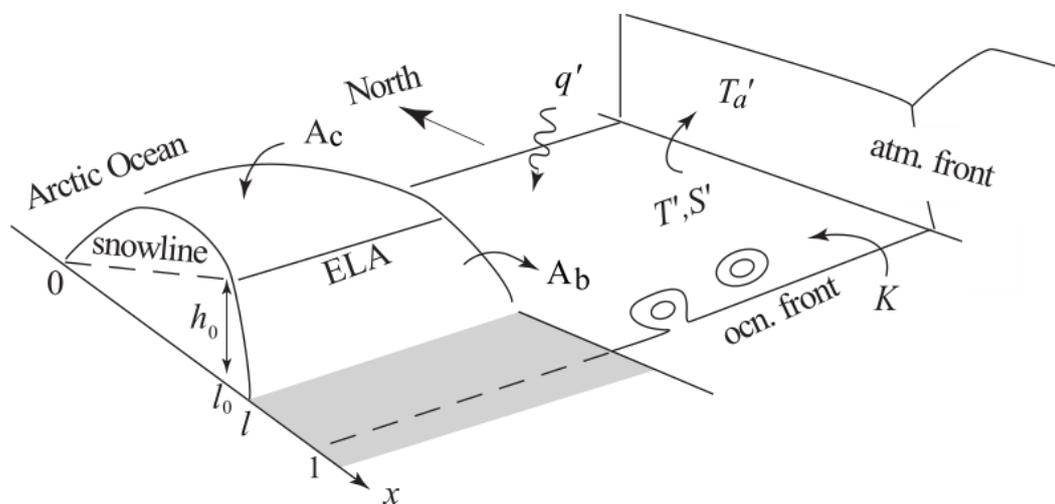


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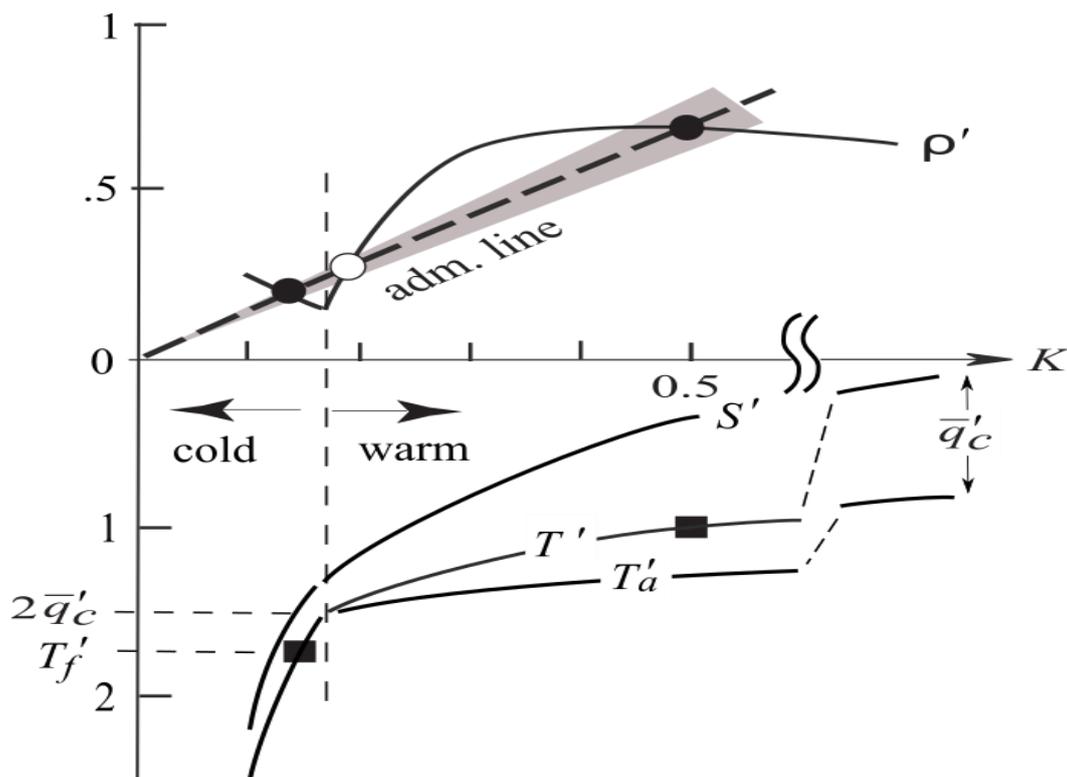
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933 **Figure 1:** The model configuration of coupled ocean/atmosphere composed of warm and cold  
934 boxes aligned at mi-latitudes and an ice sheet on a continental strip terminated at the Arctic  
935 Ocean. The prognostic variables include the cold-box deviations from global means, the  
936 MOC, and the ice margin (all symbols are listed in Appendix B)

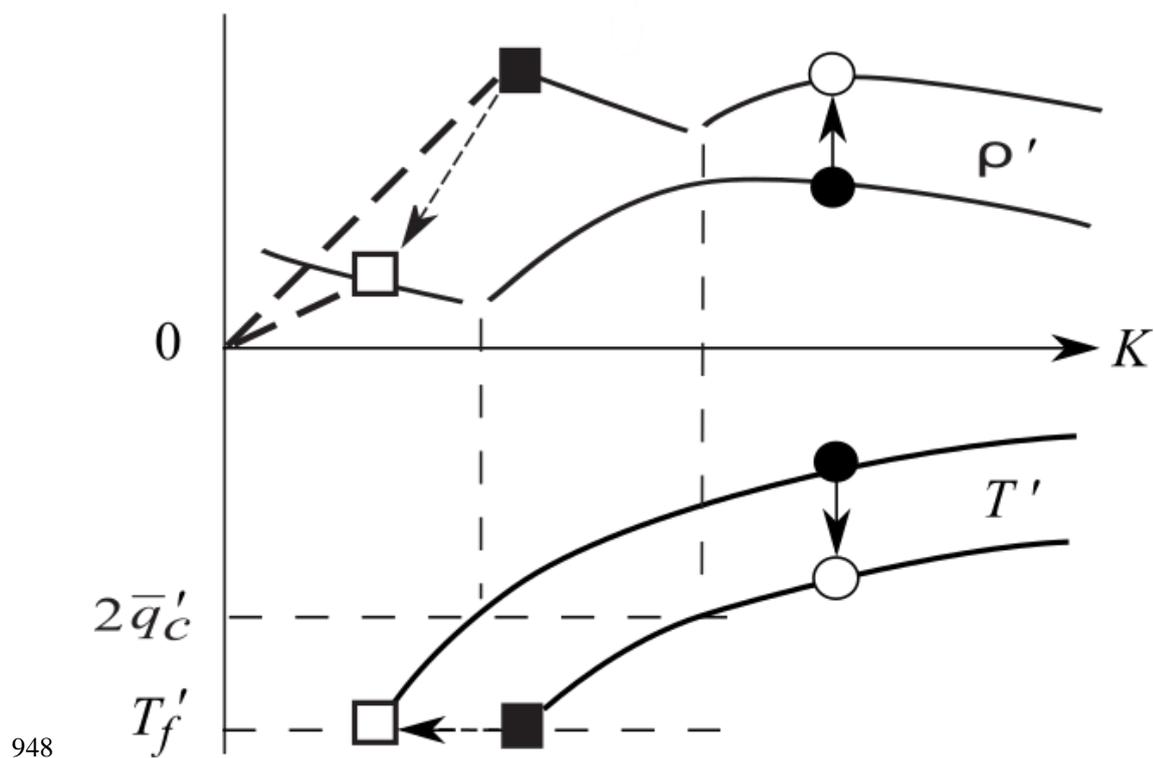
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939 **Figure 2:** The regime diagram in which the cold-box deviations (solid lines) from the global-  
 940 means (the horizontal axis) are plotted against the MOC ( $K$ ). The vertical dashed line marks  
 941 the convective bound when the convective flux vanishes, which divides the warm and cold  
 942 branches. The intersects of the admittance line (thick dashed) with the density curve specifies  
 943 the climate state (solid ovals). Subjected to fluctuations (shaded cone), the admittance line  
 944 would pivot toward the MEP states (solid rectangles), which define the interglacial and glacial  
 945 states, the latter characterized by the freezing-point but ice-free subpolar water. The graph is  
 946 for the case of  $q' = 1$ ,  $\bar{q}'_c = .75$ , and  $T'_f = 1.75$

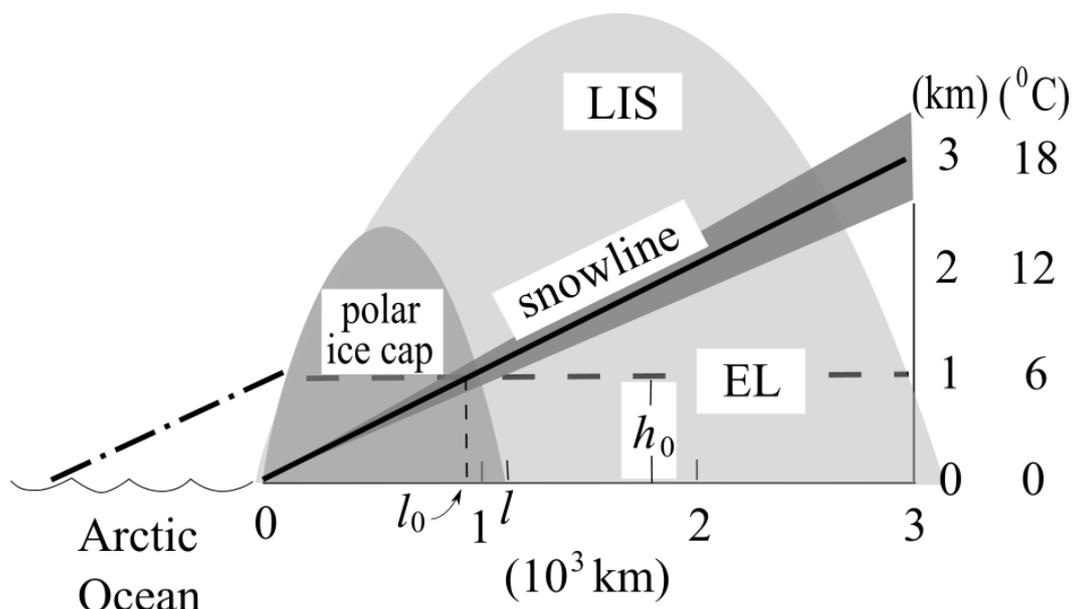
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949 **Figure 3:** The evolution of the warm MEP when  $q'$  increases (solid arrows from solid to open  
950 circles) and the cold MEP when  $q'$  decreases (dashed arrows from solid to open squares). The  
951 thick dashed lines are the admittance lines associated with the cold MEP

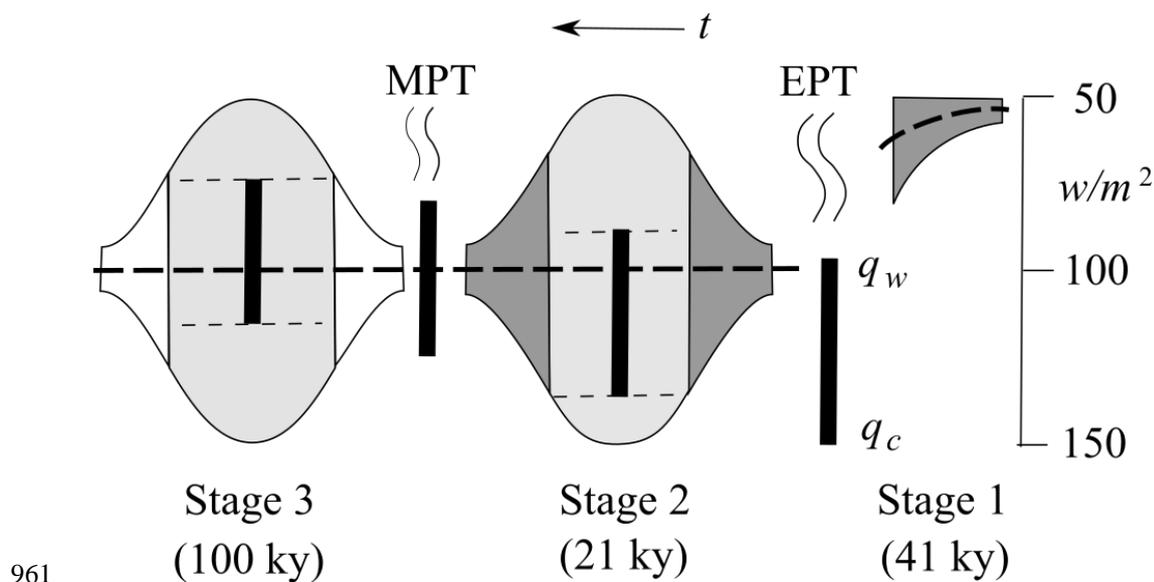
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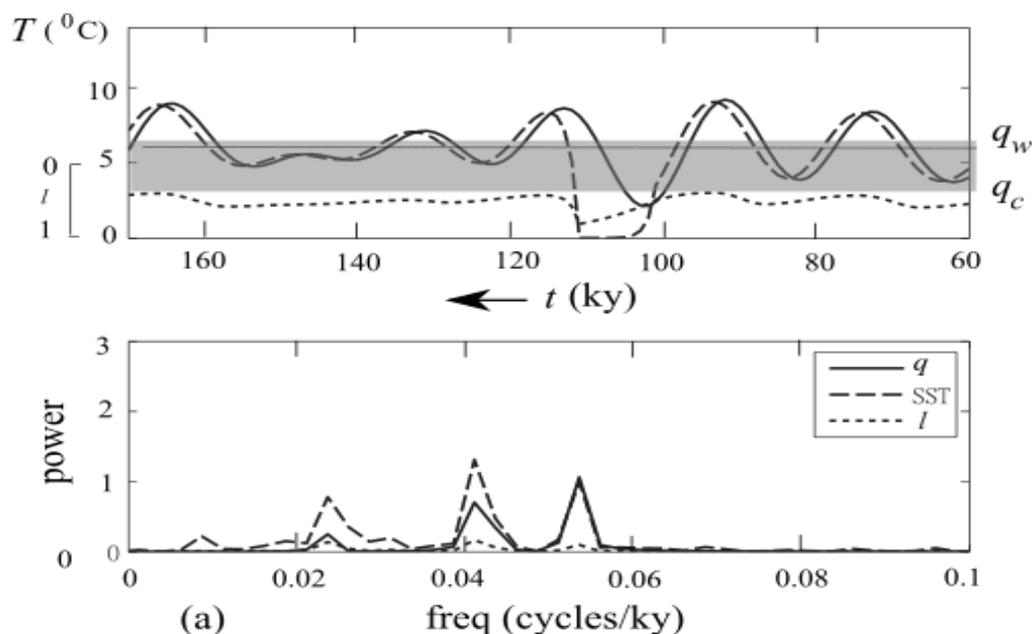
954 **Figure 4:** The summer SAT and snowline (aligned in the solid line) over the subpolar region  
 955 during the interglacial, the dark cone indicates their perturbation by the orbital forcing. The  
 956 ELA ( $h_0$ , dashed line) specifies the margin of the polar ice cap (medium shade). The dash-dot-  
 957 ted line marks the snowline when the polar ice cap may transition to the ice-free state with  
 958 strong warming. During the glacial, the summer SAT is at the freezing point (hence aligned  
 959 with the abscissa) and the ice sheet extends to the subtropical front (light shade)

960



961  
962 **Figure 5:** The evolution of the forcing envelop and ice signals during Pleistocene cooling,  
963 which consists of three stages and their transitions. The vertical bars are bistable intervals  
964 spanned by the cold ( $q_c$ ) and warm ( $q_w$ ) thresholds, which rise due to the Pleistocene cooling.  
965 Stage 1 is dominated by the interglacial cycles (hence dark-shaded) at the 41-ky obliquity pe-  
966 riod. Stage 2 sees the emergence of the G/IG cycles (hence light-shaded) at the 21-ky preces-  
967 sion period. Stage 3 is dominated by the ice-age cycles (hence unshaded) at the 100-ky eccen-  
968 tricity period

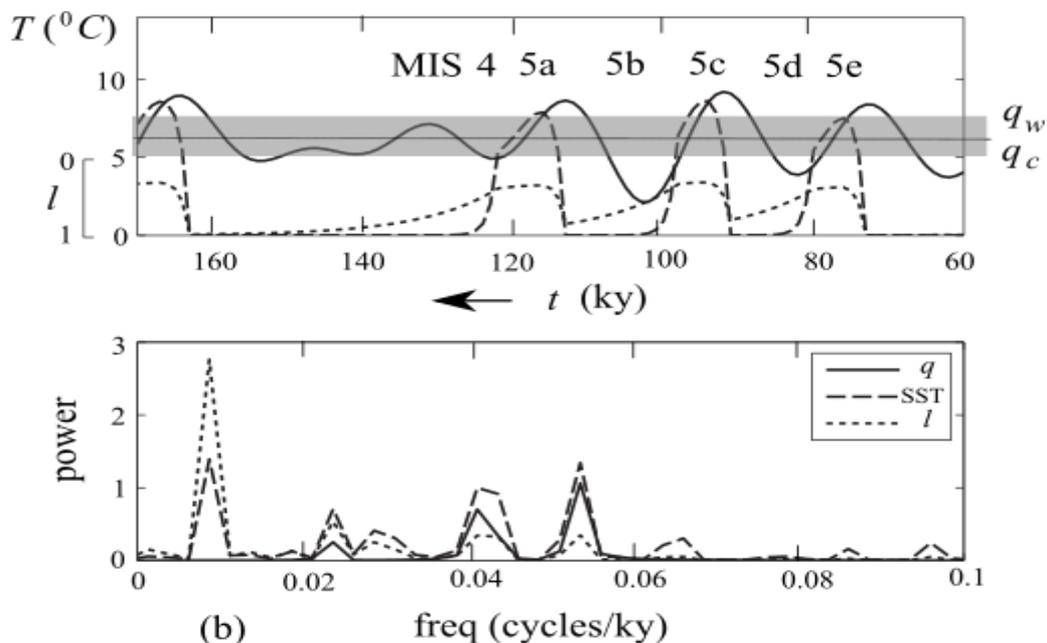
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971 **Figure 6(a):** Timeseries and power spectra of the forcing ( $q$ , solid lines in equivalent tempera-  
972 ture), subpolar SST (dashed, with a global-mean of  $14^{\circ}\text{C}$ ) and ice margin ( $l$ , dotted, in frac-  
973 tional extension into the subpolar). The thin horizontal line is the time-mean forcing and the  
974 bistate interval (shaded bar) is that of Stage 2 shown in Fig. 5, which allows the generation of  
975 the glacial state during high eccentricity, but the SST and ice-margin spectra remain dominated  
976 by the precession

977



978

979 **Figure 6(b):** Same as Fig. 6a but for Stage 3 when the bistate interval (shaded bar) is further  
980 raised by Pleistocene cooling. There are both glacial and interglacial states during high eccen-  
981 tricity corresponding to the labelled marine isotope stages, but only the glacial state spanning  
982 the low eccentricity, allowing the full growth of the ice sheet. The SST and ice-margin spectra  
983 exhibit a strong peak at the eccentricity period despite its absence in the forcing spectrum

984