

Millennial variations of atmospheric CO₂ during the early Holocene (11.7–7.4 ka)

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Abstract. We present a new high-resolution record of atmospheric CO₂ from the Siple Dome ice core, Antarctica over the early Holocene (11.7–7.4 ka) that quantifies natural CO₂ variability on millennial timescales under interglacial climate conditions. Atmospheric CO₂ decreased by ~10 ppm between 11.3 and 7.3 ka. The decrease was punctuated by local minima at 11.1, 10.1, 9.1 and 8.3 ka with amplitude of 2–6 ppm. These variations correlate with proxies for solar forcing and local climate in the South East Atlantic polar front, East Equatorial Pacific and North Atlantic. These relationships suggest that weak solar forcing changes might have impacted CO₂ by changing CO₂ outgassing from the Southern Ocean and the East Equatorial Pacific and terrestrial carbon storage in the Northern Hemisphere over the early Holocene.

20 1 Introduction

Future climate and ecosystem changes due to the continual increase of atmospheric carbon dioxide concentrations caused by human activities are inevitable (IPCC, 2013). Understanding the links between the carbon cycle and climate become important for accurate projection of future climate change. Atmospheric CO₂ is controlled by carbon exchange with ocean and land reservoirs, and increased CO₂ in the future and consequent changes in the earth system will in turn impact CO₂ levels via

25 feedbacks (Friedlingstein et al., 2006). Due to the limited duration of direct measurements of atmospheric CO₂, which only started in 1957 (Keeling, 1960), our understanding of the carbon cycle dynamics is limited on longer time scales. Air bubbles occluded in Antarctic ice cores allow us to reconstruct ancient air and may help us better understand the mechanisms that control atmospheric CO₂ (Ahn and Brook, 2008, 2014; Bereiter et al., 2012; Higgins et al., 2015; Lüthi et al., 2008; Marcott et al., 2014; Nehrbass-Ahles et al., 2020; Petit et al., 1999).

30 Understanding the carbon cycle during interglacial periods is particularly useful because climate boundary conditions are similar to those of the near future. Previous work on late Holocene CO₂ records shows centennial CO₂ variability linked with

climate, but the control mechanisms remain unclear, in part due to the potential mixture of natural and anthropogenic sources and sinks (Ahn et al., 2012; Bauska et al., 2015; Etheridge et al., 1996; Goosse, 2010; Indermühle et al., 1999; Rubino et al., 2013; Ruddiman, 2003, 2007). By contrast, CO₂ records for the early Holocene (11.7 to 7.3 ka) should reflect only natural
35 CO₂ variability due to a smaller human population (Ruddiman, 2003).

The early Holocene (11.7–7.0 ka), is known as a relatively stable period in comparison with glacial periods. Several authors have linked centennial to millennial variability in the early Holocene to changes in solar forcing, including studies of the eastern equatorial Pacific (Marchitto et al., 2010), North Atlantic (Bond et al., 2001) and the Southern Ocean (Nielsen et al., 2004) with responses in proxy records at ~11.1, 10.1, 9.1 and 8.3 ka linked to solar variability (Bond et al., 2001; Marchitto et
40 al., 2010). A weaker (stronger) solar activity has been linked with increased (decreased) ice-rafted debris in North Atlantic (Bond cycle), dominant El-Niño-like conditions (La Niña-like conditions) in the eastern equatorial Pacific, weaker (stronger) Asian monsoons, expansion (reduction) of sea ice in the Southern Ocean and colder (warmer) sea surface temperature in the Southern Ocean (Bond et al., 2001; Marchitto et al., 2010; Nielsen et al., 2004; Reimer et al., 2004; Vonmoos et al., 2006). However, it is not clear what mechanisms are involved (Bond et al., 2001; Darby et al., 2012; Marchitto et al., 2010).

45 Atmospheric CO₂ on millennial time scales is mainly controlled by exchange with oceanic reservoirs and terrestrial carbon stocks. Existing atmospheric CO₂ records from EPICA Dome C (Dome C) show little variability of atmospheric CO₂ on millennial time scales from 10.9 to 7.3 ka (Monnin et al., 2001; Monnin et al., 2004). However, high-frequency signals might be muted due to gas trapping processes at this low-accumulation site (Spahni et al., 2003).

In this study, we measured 99 samples of atmospheric CO₂ with ages between 11.7 and 9.0 ka from the Siple Dome ice core.
50 This new record complements the existing Siple Dome CO₂ record for 9.0–7.3 ka (Ahn et al., 2014). With this record, we investigate the relationship between atmospheric CO₂ and climate variations on centennial and millennial time scales. Siple Dome benefits from an accumulation rate 4.2 times higher than at EDC and 1.8 times higher than at Taylor Dome (Table 1). A conservative estimate for the width of the gas age distribution in the Siple Dome record gives ~42 years for the early
55 Holocene (Ahn et al., 2014). Thus, the Siple Dome ice core allows high temporal resolution and higher quality gas data with a more precise age scale and signals that are much less muted by the gas trapping process. The temporal resolution on average during the early Holocene reaches ~30 yr as compared to ~80 yr in the EDC record.

2 Methods

2.1 CO₂ measurements

247 individual ice samples from 99 depth intervals were measured by needle cracker dry extraction and gas chromatography
60 methods at Seoul National University (SNU) (see Figure S1 in SI (Supplementary Information)). We adopted the well-established measurement methods from Oregon State University (OSU) (Ahn et al., 2009) with minor modifications including sharpening of the tips of ice-crushing pins to increase the gas extraction efficiency, and use of a newer model Agilent 7890 gas chromatograph (GC).

Briefly, ice samples were cut and trimmed carefully with a band saw in a -21°C walk-in freezer at SNU. All visible cracks were removed to eliminate potential CO_2 alteration by trapping modern air. An ice sample of $\sim 8\text{--}10$ g was placed in a double walled vacuum chamber maintained at about -35°C using cold ethanol circulation between the walls of the chamber while flowing ultra-pure of N_2 gas (99.9999%) into the chamber. The ice sample was crushed in the cooled chamber by 91 steel needles moving straight up to down using a linear motion (bellows) vacuum feedthrough. The liberated air from the ice was collected for 3 min in a sample tube in a cryogenic system maintained at 11 K. The CO_2 mixing ratio was determined by the Agilent 7890A GC equipped with a flame-ionization detector, using a Ni catalyst which converts CO_2 to CH_4 before measurement. Sample air was injected into a stainless steel sample loop and the extracted air from each ice sample was analysed twice. The GC system was calibrated daily with a standard air tank (293.25 ppm CO_2 , WMOX2007 mole fraction scale, calibrated by US National Oceanic and Atmospheric Administration, Global Monitoring Division). To examine the linearity of the GC, ice samples from five different depth intervals (CO_2 concentrations of 239–251 ppm) were analysed with two different air standards (188.9 and 293.3 ppm CO_2 , respectively). The average difference in the results using the different standards was 0.4 ± 0.9 ppm (1σ) (Table S1 [in SI](#)).

2.2 Age scale of the Siple Dome ice core records

The Siple Dome samples are placed on the improved Siple Dome chronology developed by Yang et al. (2017), which is aligned with the Greenland Ice Core Chronology, 2005 (GICC05) using the synchronization of CH_4 and $\delta^{18}\text{O}_{\text{atm}}$ time series. Abrupt CH_4 changes have been shown to be synchronous within about 50 years with abrupt climate changes in Greenland during the last glacial period (Baumgartner et al., 2014; Rosen et al., 2014). Using this principle, abrupt changes in the composite Siple Dome CH_4 data were aligned with abrupt changes in $\delta^{18}\text{O}_{\text{ice}}$ from the NGRIP ice core (North Greenland Ice Core Project members, 2004; Rasmussen et al., 2006) at the 8.2 ka event and end of the Younger Dryas (Yang et al., 2017). For the time period of 11.64–8.10 ka, ages were updated from the original chronology of Severinghaus et al. (2009) by interpolating the age offsets at the tie points (Yang et al., 2017). For the time intervals outside of 11.64–8.10 ka, the age difference was set constant with the difference at the closest tie point. The modified gas ages are younger than the Severinghaus et al. (2009) ages by less than ~ 110 years.

~~2.3 Correlation coefficients between CO_2 and climate proxies~~

~~In order to assess the relationship between CO_2 and climate, we calculate correlations and estimate leads and lags between CO_2 and proxies of solar activity as well as climate proxies thought to be themselves related to solar activity. The Pearson correlation coefficient r is commonly used to verify relationships between variables. However, r does not take chronological uncertainty into account. As such, we apply a Monte Carlo procedure to estimate the correlations between CO_2 and climate proxies. In the procedure, we adjust the chronologies of the two series, within their chronological uncertainties, and re-calculate~~

r. We do so 1,000 times for each pair, allowing us to calculate a mean correlation coefficient that is more representative of the relationship between time-uncertain series.

We use a similar method to calculate the significance of this correlation against a random red-noise process. At each of the 1,000 steps, we use an AR(1) model (lag-1 auto-regression) to fit data the series. We then use these AR(1) characteristics to randomly generate two synthetic series with the same red noise. Then, we calculate the percentage of correlations between the randomized synthetic series that are lower than the correlation coefficients of the real series to assess the significance of the correlation.

Finally, we can calculate the maximum correlation lag between the two series. At each step in the 1,000-iteration Monte Carlo procedure, we calculate the lag which gives the maximum correlation by shifting one of the series by 10-year increments, for constant lags between -200 (a CO₂ lead) and 200 years (a CO₂ lag). Then, we make a histogram of the calculated maximum correlation lags, from which the mode can be selected as an approximation of the phasing between the two series. We also report the maximum correlation between the two series, but note that this is not representative of the actual relationship (it simply gives an idea of how much the correlation can be improved by adding a lag between the two series).

3 Results

3.1 The new high-resolution CO₂ record during the early Holocene

We obtained 99 data points that cover 622.14–539.06 m at SNU, corresponding to 11.7–9.0 ka (Figure 1). To extend the record to 7.4 ka, we made a composite dataset using a previous CO₂ record from the Siple Dome ice core covering 9.0–7.4 ka measured by the needle cracker system at OSU (Ahn et al., 2014) (Figure 1). Between 2 and 6 replicates (2.6 and 2.4 on average for SNU and OSU data, respectively) from individual depth intervals were analysed. The standard error of the mean of replicates from the same depth interval was 0.8 and 0.5 ppm on average for SNU and OSU data, respectively. The sampling resolution is ~30 years for 11.7–9.0 ka and ~15 years for 9.0–7.3 ka.

To make a composite record of atmospheric CO₂, we tested for bias between the two data sets. Siple Dome samples from 7 depth intervals between 538.55–490.16 samples were analysed at both laboratories (Ahn et al., 2014). The SNU measurements were higher than the OSU measurements by 0.3±0.7 ppm (1σ) on average, indicating that the SNU and OSU results agree well (Table S2 in SI). The small offset of 0.3 ppm was added to OSU data before combining them with the SNU results.

3.2 Comparison with existing CO₂ records for the early Holocene

The new atmospheric CO₂ record from Siple Dome was compared to the existing CO₂ data from Dome C measured using the needle cracker at University of Bern (UB) (Monnin et al., 2001; Monnin et al., 2004) and the existing CO₂ data from the WAIS Divide ice core measured by the needle cracker at OSU (Marcott et al., 2014) (Figure S2a-in-SI). On multi-millennial time scales, the baseline levels of the Siple Dome and WAIS Divide CO₂ records (Marcott et al., 2014) are higher than those from

Dome C (Flückiger et al., 2002; Monnin et al., 2004) and Taylor Dome (Andermühle et al., 1999) records (Figure S2 in SI2).

125 The CO₂ offset between the Dome C and Siple Dome ice cores is 3–6 ppm (Figure 2b).
The offset between Siple Dome CO₂ data in this study and other CO₂ data sets ~~could~~ might be related to differences in the analytical methods used to make the measurements. To examine the inter-laboratory analytical offset, several Taylor Dome ice samples were analysed at OSU (Ahn et al., 2014). The OSU results were higher than those at UB by 1.5 ppm on average. Taking the analytical offset between OSU and SNU of 0.3±0.7 ppm (1σ) into consideration, the 3–6 ppm CO₂ offset between

130 the Siple Dome record (measured at OSU and SNU) and Dome C or Taylor Dome (measured at UB) cannot be entirely attributed to experimental offset.

~~CO₂ records can be contaminated by the in situ production of CO₂ caused by carbonate acid reactions (Anklin et al., 1997; Barnola et al., 1995; Delmas, 1993; Neftel et al., 1988; Smith et al., 1997a; Smith et al., 1997b). Calcium carbonate (CaCO₃), hydrogen peroxide (H₂O₂) and formaldehyde (HCHO) in ice cores can cause carbonate acid reactions, leading to large scattering of atmospheric CO₂ data (Smith et al., 1997b). Antarctic ice cores have relatively low concentrations of hydrogen peroxide and carbonates and lower site temperatures compared to Greenlandic ice cores, which reduces the risk of CO₂ contamination (Tschumi and Stauffer, 2000). It is estimated that the in situ production of CO₂ for Antarctic ice cores is smaller than 1.5 ppm (Bereiter et al., 2009). In situ production of CO₂ cannot be ruled out but the effect should not greatly impact the offset between records from the different ice cores.~~

135 ~~Oxidation of organic compounds (e.g., 2H₂O₂ + HCHO → 3H₂O + CO₂) can also produce CO₂ in ice (Tschumi and Stauffer, 2000). The Dome C site is located further from the ocean than Siple Dome and we therefore expect lower organic content in the Dome C ice. Concentrations of organic compounds at our sampling depths are not available. However, the concentration of oxidant H₂O₂ on the top 2.5–100 m in the Siple Dome core is below the detection limit of 0.02 μM (McConnell, 1997), although 0.02 μM H₂O₂ still has potential to produce CO₂ and can increase the mixing ratio in bubbles by 5 ppm given sufficient supply of organic compounds (Ahn et al., 2004).~~

140 ~~discussed CO₂ offsets during the Holocene (11.7–0 ka). For the early Holocene, five datasets were compared: CO₂ from the WAIS Divide ice core measured by the needle cracker at OSU (Marcott et al., 2014); CO₂ from the Siple Dome ice core measured by the needle cracker at OSU (Ahn et al., 2014), CO₂ from Law Dome measured by sublimation at UB (Eggleston, 2015); CO₂ from Talos Dome measured by the Centrifugal Ice Microtome (CIM) at UB; and CO₂ from Dome C measured by~~

150 ~~the needle cracker at UB (Monnin et al., 2001; Monnin et al., 2004). The individual datasets show similar patterns of CO₂ variability over the early Holocene. Offsets among ice cores are lower during the late Holocene but become larger during the early Holocene. During the early Holocene, the largest offset was between Dome C and Law Dome, up to 12 ppm. On the other hand, the baseline level of the Siple Dome CO₂ record is similar to that of the WAIS Divide CO₂ and Talos Dome CO₂ records. CO₂ levels are lower for ice core sites far away from the Antarctic coast, thus we might hypothesize that the level of~~

155 ~~CO₂ concentration could be related to the in situ conditions (e.g. temperature and accumulation rate and the amount of organic compounds) (Nehrbass-Ahles, 2017). However, the offset among records are relatively small.~~

To compare the new record to the existing records on millennial time scales, we calculate the Pearson correlation coefficient between Siple Dome CO₂ and existing CO₂ records. The offsets between existing CO₂ records and our data are also calculated (Figure 2). For these calculations, we use 250-yr running means of CO₂ records.

160 The Correlation coefficient between Siple Dome CO₂ and WAIS divide CO₂ is 0.7 ($p < 0.001$). However, the CO₂ offset between Dome C record and Siple Dome record is quite random (Figure 2B) because of scattering in the WAIS Divide CO₂ record during the early Holocene period. The WAIS Divide CO₂ data during the early Holocene was reconstructed from the ice just below the bubble clathrate transition zone (BTCZ). Previous studies raised an issue about the possibility of high frequency noise of atmospheric CO₂ record in the ice just below the BTCZ (Lüthi et al., 2010; Shackleton et al., 2019). This
165 phenomenon might be related to gas fractionation effect because of clathrate layering during bubble-clathrate transformation. Gas content starts to be fractionated in the BCTZ because of the differential permeation of gas species when bubbles have transformed to clathrates. CO₂ concentration in the first layer of clathrates is more enriched with higher bubble-to-clathrate permeation rates. Below the BCTZ, gas content slowly homogenizes again through molecular diffusion (Bereiter et al., 2009), which can cause high frequency noise to the ice below the BCTZ. Thus, the WAIS Divide CO₂ data is not sufficient to discuss
170 millennial variabilities of the early Holocene.

We observe that CO₂ data sets from Siple Dome and Dome C share similar trends in CO₂ variations despite the CO₂ offset in longer term means of 3–8 ppm. The CO₂ record from the Siple Dome is highly correlated with the CO₂ record from Dome C ($r = 0.89$, $p < 0.001$). The CO₂ offset between Dome C record and Siple Dome record decreases continuously from 11.7 ka to 7 ka with small variations at around 9.3 and 8.3 ka (Figure 2). The small variations of Dome C CO₂ record (1.4 ppm, compared
175 to 3.0 ppm for Siple Dome) can be explained by the lower sampling resolution (~80 yrs for Dome C vs. ~20 yrs for Siple Dome) and a stronger damping effect on CO₂ concentration change at Dome C due to the slower gas trapping process at Dome C (Spahni et al., 2003).

The millennial CO₂ variations in the ice cores could be attributed to different degrees of in-situ CO₂ production in ice. The in-situ production of CO₂ caused by carbonate-acid reactions (Anklin et al., 1997; Barnola et al., 1995; Delmas, 1993; Neftel et al., 1988; Smith et al., 1997a; Smith et al., 1997b) and oxidation of organic acids (Tschumi and Stauffer, 2000). Although
180 Antarctic ice cores have relatively low concentrations of carbonates and lower site temperatures compared to Greenlandic ice cores (Tschumi and Stauffer, 2000), it is estimated that the in-situ production of CO₂ for Antarctic ice cores is smaller than 1.5 ppm (Bereiter et al., 2009). If the chemical alteration is the main cause of the millennial-scale CO₂ variations, we may expect to observe CO₂ age offsets among different cores because of dissimilar ice age-gas age differences. However, no available
185 data set supports this possibility.

To further evaluate the in-situ CO₂ production, we considered potential reactions. First, we compared the CO₂ with non-sea-salt Ca (nssCa) content in the ice to check the carbonate-acid reaction in the ice. The concentration of nssCa is mainly controlled by dust delivery but it also can be produced partially by the carbonate-acid reaction in ice. Thus, we examined the concentration of nssCa ion in the Siple Dome and Dome C ice. The nssCa records do not correlate well with the filtered
190 millennial CO₂ variations in both Siple Dome ($r = -0.33$) and Dome C ($r = 0.15$) records during the early Holocene (Figures

S2 and Figure S3 in SI). In addition, the nssCa trends in Dome C and Siple Dome ice do not agree (Figures S2 and Figure S3 in SI), but millennial CO₂ variations do. Second, we checked the CO₂ production by oxidation of organic compounds (e.g., 2H₂O₂ + HCHO → 3H₂O + CO₂) in ice (Tschumi and Stauffer, 2000). The Dome C site is located further from the ocean than Siple Dome and we therefore expect lower organic content in the Dome C ice. Concentrations of organic compounds at our sampling depths are not available. However, the concentration of oxidant H₂O₂ on the top 2.5–100 m in the Siple Dome core is below the detection limit of ~0.02 μM (McConnell, 1997), although 0.02 μM H₂O₂ still has potential to produce CO₂ and can increase the mixing ratio in bubbles by 5 ppm given sufficient supply of organic compounds (Ahn et al., 2004). In summary, the existing Dome C CO₂ record covering the early Holocene share similar trends in the Siple dome CO₂ record despite an offset in longer term means of a few ppm. We note that CO₂ offsets of several ppm among different ice cores are common features in different time intervals such as the last millennium (Ahn et al., 2012; Monnin et al., 2004; Rubino et al., 2019; Siegenthaler et al., 2005) and Marine Isotope Stage 3 (Ahn et al., 2008; Bereiter et al., 2012) although they share the same trends of CO₂ change on multi-centennial to multi-millennial time scales. Thus, it is likely that the millennial CO₂ variations during the early Holocene in the Siple Dome and Dome C cores reflect atmospheric CO₂ changes.

3.3 Atmospheric CO₂ variations on the millennial time scale during the early Holocene

Figure 1 shows the atmospheric CO₂ record from Siple Dome during the early Holocene. Atmospheric CO₂ increased by ~8 ppm between 11.7 and 11.3 ka and then decreased by ~10 ppm from 10.9 to 7.3 ka. The rapid CO₂ increase at 11.7–11.3 ka might be associated with abrupt warming in the North Atlantic Greenland and parts of the Northern Hemisphere and abrupt strengthening of Atlantic Meridional Overturning Circulation at the end of the last glacial termination (Marcott et al., 2014; Monnin et al., 2001). The long term CO₂ trend is generally similar to that of the major water isotope (δD) variations in Antarctic ice cores reflecting Antarctic temperature variations (Figure S4 in SI).

The Siple Dome CO₂ record shows millennial variability of ~2–6 ppm with local minima at 11.1, 10.1, 9.1 and 8.3 ka (Figure 1). These variations resemble variability in other paleoclimate records that has been linked to solar cycle variations on these time scales, as described above (Figure S5 in SI) (Figures 3 and S5).

To examine the relationship between atmospheric CO₂ and the other paleoproxy data sets on millennial time scales, the Siple Dome CO₂ record was smoothed and high pass filtered at 1/1800 yr due to two necessities. First, it is likely that high-frequency variabilities of atmospheric CO₂ record (decadal-scale variations and centennial-scale variations) are high frequency noise of atmospheric CO₂ record. Thus, we smoothed data sets to eliminate high-frequency variability. Before making a 250-yr running mean, we made a 1-yr interpolation, because sample spacing between data points covering the early Holocene is not constant. Second, to eliminate multi-millennial drift of CO₂ record, the data was high pass filtered at 1/1800 yr, following previous methods by Bond et al. (2001) and Marchitto et al., (2010). The proxy records were also processed in the same way as the CO₂ record to remove high-frequency variability and long-term drift.

We evaluated uncertainties of the smoothed and high pass filtered CO₂ record using Monte Carlo simulation. Random sampling was made from a probability distribution for each measured value and its standard deviation. We repeated this series of simulations 10,000 times, which is shown as 2σ in Figure 1 (see SI for detailed information).

225 We calculated correlation coefficients between the filtered CO₂ and climate proxy series to understand their relationship with atmospheric CO₂ (Figure 3, see SI for methods). Correlation coefficients, their significance, and maximum correlation lags are shown in Figure 4 and Table 2. The composite CO₂ record from the Siple Dome is anti-correlated with the stacked IRD record in the North Atlantic (Bond et al., 2001) ($r = -0.49 \pm 0.1$, CO₂ time lag of 120 ± 155 years), SST record in the eastern equatorial Pacific indicating El Niño-like or La Niña-like conditions ($r = -0.41 \pm 0.13$, CO₂ time lag of 50 ± 219 years) (Marchitto et al., 2010), and sea ice in the Southern Ocean ($r = -0.35 \pm 0.17$, CO₂ time lag of 190 ± 228 years) (Nielsen et al., 2004). On the other hand, the CO₂ record is positively correlated with summer sea-surface temperature (SSST) in the Southern Ocean ($r = 0.35 \pm 0.17$, CO₂ time lag of 52 ± 228 years) (Nielsen et al., 2004). These results may imply a tentative link between atmospheric CO₂ variations and climate change on millennial time scales. The time lags might be caused by age uncertainties of the proxy records and/or response time of atmospheric CO₂ to climate change (Bauska et al., 2015; Bereiter et al., 2012; 235 Carvalhais et al., 2014).

Interestingly, the highest anti-correlations we find are between the Siple Dome CO₂ record and the ¹⁴C production rate ($r = -0.49 \pm 0.12$, CO₂ time lag of -20 ± 148 years) and ¹⁰Be flux ($r = -0.52 \pm 0.08$, CO₂ time lag of 110 ± 63 years). This suggests that CO₂ and solar activity co-vary on millennial time scales (Figure 4 and Table 2). Given these observations, solar activity might be linked to the atmospheric CO₂ variations by the response of carbon cycle to climate change during the early Holocene (11.7–7.0 ka) (Figure 4 and Table 2, Figure 2 and Figure 3). 240

4 Discussion

4.1 Possible carbon cycle control mechanisms in the Early Holocene

Understanding a link between climate variations and solar activity on millennial time scales during the early Holocene is important to decipher carbon cycle mechanisms. However, the climate mechanisms have not yet been deciphered. A possible mechanism is that changes of solar activities may impact on stratospheric ozone concentrations, which can change stratospheric and tropospheric circulation patterns (Meehl et al., 2009). Higher solar activity may enhance the precipitation in the Intertropical Convergence Zone (ITCZ) and South Pacific Convergence Zone (SPCZ) (Meehl et al., 2009; van Loon et al., 2007). Consequently, the intensified moisture at those areas would increase trade wind strength and upwelling in the East Equatorial Pacific region. These conditions would lead to La Niña like climate states on millennial time scale (Marchitto et al., 2010). This change in the East Equatorial Pacific might have affected the North Atlantic (Darby et al., 2012). 245 250

If the CO₂ variations we observe are affected by solar variabilities via climate, a number of mechanisms could be involved, including the terrestrial or marine carbon cycles, or both.

We discuss three possibilities here. First, a close relationship between CO₂ and climate proxies in Antarctica (Jouzel et al., 2007) on multi-millennial time scales (Figure S4 ~~in SI~~) suggests that CO₂ variations on these time scales might be principally controlled by Southern Ocean processes. Atmospheric CO₂ can be controlled by temperature and salinity in the ocean (the solubility pump); solubility of CO₂ is greater in cooler and fresh surface waters (Broecker, 2002; Takahashi et al., 1993). The formation of deep water occurs in polar regions with high water density, where surface waters are cold, thus, the oceanic uptake of atmospheric CO₂ through this mechanism is stronger in polar regions (Sigman and Boyle, 2000). We observed a tentative link between atmospheric CO₂ and summer sea surface temperature (SSST) from the polar front region of the South East Atlantic on millennial time scales (Nielsen et al., 2004), which implies that lower SSST in the Southern Ocean ~~might have~~ lead to ~~the~~ reduction of atmospheric CO₂.

Increased sea ice extent might have blocked release of CO₂ from CO₂-rich deep water to the atmosphere, and therefore decreased atmospheric CO₂ concentration as previously suggested for glacial-interglacial CO₂ variations (Stephens and Keeling, 2000). Our Siple Dome CO₂ record is negatively correlated with the sea ice extent in the Southern Ocean, although the sea ice extent reconstruction shown in Figure 3 represents only the east Atlantic region of the Southern Ocean.

Oceanic processes associated with El Niño-like and La Niña-like climate variation could also impact the carbon cycle. Marine sediment cores from the East Equatorial Pacific show that solar activity proxies are well correlated with El Niño-like and La Niña-like climate variations in the East Equatorial Pacific SST proxy record (Marchitto et al., 2010). The East Equatorial Pacific is the region where CO₂-rich deep water upwells. Increased upwelling during La Niña -like conditions and resulting increased CO₂ outgassing have been suggested for the CO₂ increase during the last deglaciation (Kubota et al., 2014). Siple Dome CO₂ is anti-correlated with SST in the East Equatorial Pacific on millennial time scales (Figure 2), which may imply that La Niña-like climate can lead to higher CO₂ values.

Terrestrial carbon is involved with photosynthesis and respiration in plants, and with soil respiration (microbial and root respiration). Thus, terrestrial carbon is mostly controlled by temperature and precipitation (Davidson et al., 2000; Mielnick and Dugas, 2000). On multi-millennial time scales, when temperature in Greenland increases from 10.9 to 7.4 ka, atmospheric CO₂ decreases. Expansion of vegetation in the Northern Hemisphere may partially contribute to the decrease in atmospheric CO₂ (Indermühle et al., 1999).

A recent high resolution study for the last 1,200 ~~yearyrs~~ shows that centennial CO₂ variability was mainly controlled by terrestrial carbon, most likely in the high latitude of the Northern Hemisphere (Bauska et al., 2015). The stacked IRD from the North Atlantic may be used for an indicator of cool conditions in the North Atlantic (Bond et al., 1992; Bond et al., 2001). The strong relationship between IRD and atmospheric CO₂ indicates that colder climate in the North Atlantic may lower atmospheric CO₂ by impacting terrestrial carbon stocks during the early Holocene.

$\delta^{18}\text{O}_{\text{ice}}$ from the North Greenland Ice Core Project (NGRIP) ice core (Rasmussen et al., 2006) indicating temperature in Greenland also reveal millennial local minima at similar time intervals as those of CO_2 (~11.4, 10.9, 10.2, 9.3 and 8.2 ka), however, atmospheric CO_2 and temperature in Greenland are mismatched at the earliest early Holocene and ~8.2 ka. Thus, there is no significant linear relationship between CO_2 and temperature in Greenland on millennial time scales, and our calculation indicates that CO_2 leads temperature in Greenland on millennial time scales, though the correlation is still too small to assume any relationship ($r = 0.21 \pm 0.07$, CO_2 time lag of -130 ± 63 years).

Temperature in Greenland during the early Holocene might be partially influenced by the internal climate system or/and by low-latitude solar forcing indirectly. Two main cooling events in Greenland are recorded at ~11.4 and ~8.2 ka (Rasmussen et al., 2007). The well-known 8.2 ka cooling event is mainly influenced by the collapse of the Laurentide ice sheet (Merz et al., 2015) rather than by solar forcing; when temperature was colder in Greenland at ~11.4 ka, solar forcing was higher, not reaching a minimum at until ~11.2 ka. It is also elusive whether solar forcing has an influence on climate in Greenland at ~11.4 ka (Mekhaldi et al., 2020). In short, a linkage between atmospheric CO_2 and climate change during the early Holocene remains uncertain due to insufficient paleoclimate records and model simulations.

A positive correlation between solar forcing and atmospheric CO_2 is observed during the Little Ice Age (LIA). There are two periods in which sunspots were exceedingly rare. During the Maunder sunspot minimum (1647–1715 CE), total solar irradiance (TSI) was reduced by $0.85 \pm 0.16 \text{ W m}^{-2}$. Atmospheric CO_2 records from Antarctic ice cores commonly show a decrease trend during this period (Ahn et al., 2012; Monnin et al., 2004; Siegenthaler et al., 2005; Rubino et al., 2019). During the Spörer Minimum (1450–1550 CE), TSI record during this period also shows a decrease trend. However, atmospheric CO_2 decrease is not significant in Law Dome and EPICA Dronning Maud Land (EDML) records (Monnin et al., 2004; Siegenthaler et al., 2005; Rubino et al., 2019), while WAIS divide ice record shows a decrease during this period (Ahn et al., 2012) (Figure S7 in SI).

~~presented proxy data from the coast of Alaska reflecting the Arctic Oscillation (AO) over the past 9.0 ka. This study found 1,500-year cycles of AO estimated using a proxy of sea ice drift in the Arctic Ocean, which is similar to the cycles of IRD in the North Atlantic. But there is no significant linear relationship between AO and solar activity on millennial time scales during the 9.0 ka record. Solar activity may not be a direct forcing of the AO during the Holocene (Darby et al., 2012). AO variations might be driven by the internal climate system or affected by low latitude solar forcing indirectly. Thus, it is unclear that Greenland temperature variations may cause atmospheric CO_2 variations on millennial time scales during the early Holocene.~~

310 5 Conclusion

In this study, we present a 30 yr-resolution CO_2 record during the early Holocene. Our data show that millennial atmospheric CO_2 variability of 2–6 ppm correlates with several climate proxies such as IRD in the North Atlantic, sea ice extent in the

Southern Ocean, El Niño-~~like~~ or La Niña-~~like~~ conditions in the East Equatorial Pacific, all of which appear to coincidentally occur with~~be temporally linked to~~ solar activity minima (Bond et al., 2001; Marchitto et al., 2010; Nielsen et al., 2004; Reimer et al., 2004; Vonmoos et al., 2006). The relationships with the proxies are consistent with changes in several different mechanisms that could impact atmospheric CO₂ on millennial time scales including changing CO₂ outgassing from the Southern Ocean and the East Equatorial Pacific, and changing terrestrial carbon storage in the Northern Hemisphere. Our new observations may improve our understanding of the relationship between interglacial climate and carbon cycles on millennial time scales in the absence of anthropogenic CO₂ perturbations. ~~The existing climate proxy data and carbon cycle models are not sufficient to clearly decipher the relationship. Further study should focus on clearly deciphering the millennial CO₂ control mechanisms with improved paleo proxy records and carbon cycle models.~~

Data availability. All data will be available on PANGAEA (Paleoclimatology database websites).

Author contributions. The research was designed by JS, JA and EB. The CO₂ measurements were performed by JS with contributions from HL and JA. The data analyses were led by JS and JCB with contributions from JMS and JA. JS wrote the manuscript with inputs from all authors.

Competing interests. The authors declare that they have no conflict of interest.

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Table 1. Glaciological characteristics of Antarctic ice cores.

Core name	Mean Annual Temperature (°C)	Mean Accumulation Rate as Water Equivalent (g cm ⁻² yr ⁻¹ as water equivalent)	References
Siple Dome	-25.4	12.4	Hamilton (2002); Severinghaus et al. (2001); Taylor et al. (2004)
Taylor Dome	-42	7	Waddington and Morse (1994)
EPICA Dome C	-54	3	Schwander et al.(2001); EPICA Dome C 2001-02 Science and Teams (2002); Tabacco et al.(1998)
WAIS Divide	-31	20	Banta et al.(2008); Morse et al.(2002)

Table 2. Correlation between Siple Dome CO₂ record and climate proxy records. Column A shows correlation coefficients between CO₂ and proxies with CO₂ time lags. Column B shows correlation coefficients between CO₂ and proxies without CO₂ time lag. “With MC” are mean values from the simulations taking age uncertainties into account. “Without MC” is the classic calculation of correlation, without taking age uncertainty into account. Significance of the lag correlations was assessed against 1,000 repetitions of the lag correlation calculation using synthetic data stochastically generated to have the same red noise characteristics as the original series.

<u>Proxy records</u> <u>(Reference)</u>	<u>A: Correlation between CO₂ and proxies with CO₂ time lag (yrs)</u>				<u>B: Correlation between CO₂ and proxies without CO₂ time lag</u>	
	<u>With MC</u>		<u>Without MC</u>		<u>With MC</u>	<u>Without MC</u>
	<u>r</u> <u>(p-value)</u>	<u>Time lag</u>	<u>r</u> <u>(p-value)</u>	<u>Time lag</u>	<u>r</u> <u>(p-value)</u>	<u>r</u> <u>(p-value)</u>
<u>CO₂ - ¹⁴C production rate</u> <u>Marchitto et al.(2010);</u> <u>Reimer et al.(2004)</u>	<u>-0.49± 0.12</u> <u>(0.3192)</u>	<u>-20±148</u>	<u>- 0.76</u> <u>(0.0003)</u>	<u>50</u>	<u>-0.48</u> <u>(0.007)</u>	<u>-0.70</u> <u>(< 0.001)</u>
<u>CO₂ - ¹⁰Be flux from</u> <u>Greenland ice core</u> <u>Finkel and Nishiizumi (1997);</u> <u>Marchitto et al. (2010);</u> <u>Vonmoos et al. (2006)</u>	<u>-0.52± 0.08</u> <u>(0.2847)</u>	<u>110±63</u>	<u>- 0.61</u> <u>(0.0087)</u>	<u>110</u>	<u>-0.29</u> <u>(0.05)</u>	<u>-0.32</u> <u>(< 0.001)</u>
<u>CO₂ - IRD from the North</u> <u>Atlantic region</u> <u>Bond et al. (2001);</u> <u>Marchitto et al. (2010)</u>	<u>-0.49± 0.1</u> <u>(0.3084)</u>	<u>120±155</u>	<u>- 0.73</u> <u>(0.0009)</u>	<u>170</u>	<u>-0.33</u> <u>(0.05)</u>	<u>-0.21</u> <u>(< 0.001)</u>
<u>CO₂ - SST from eastern</u> <u>equatorial Pacific</u> <u>Marchitto et al. (2010)</u>	<u>-0.40± 0.13</u> <u>(0.337)</u>	<u>50±219</u>	<u>- 0.61</u> <u>(0.009)</u>	<u>80</u>	<u>-0.38</u> <u>(0.04)</u>	<u>-0.55</u> <u>(< 0.001)</u>
<u>CO₂ - Sea ice in the Southern</u> <u>Ocean</u> <u>Nielsen et al. (2004)</u>	<u>-0.35± 0.17</u> <u>(0.2899)</u>	<u>190±228</u>	<u>- 0.57</u> <u>(0.0151)</u>	<u>100</u>	<u>-0.24</u> <u>(0.17)</u>	<u>-0.48</u> <u>(< 0.001)</u>
<u>CO₂ - SST in the Southern</u> <u>Ocean</u> <u>Nielsen et al. (2004)</u>	<u>0.35± 0.17</u> <u>(0.3070)</u>	<u>52±228</u>	<u>0.57</u> <u>(0.0144)</u>	<u>30</u>	<u>0.35</u> <u>(0.06)</u>	<u>0.56</u> <u>(< 0.001)</u>
<u>CO₂ - NGRIP δ¹⁸O</u> <u>Rasmussen et al. (2006)</u>	<u>0.21± 0.07</u> <u>(0.2684)</u>	<u>-130±63</u>	<u>0.11</u> <u>(0.3411)</u>	<u>270</u>	<u>0.09</u> <u>(0.5)</u>	<u>0.06</u> <u>(0.2)</u>

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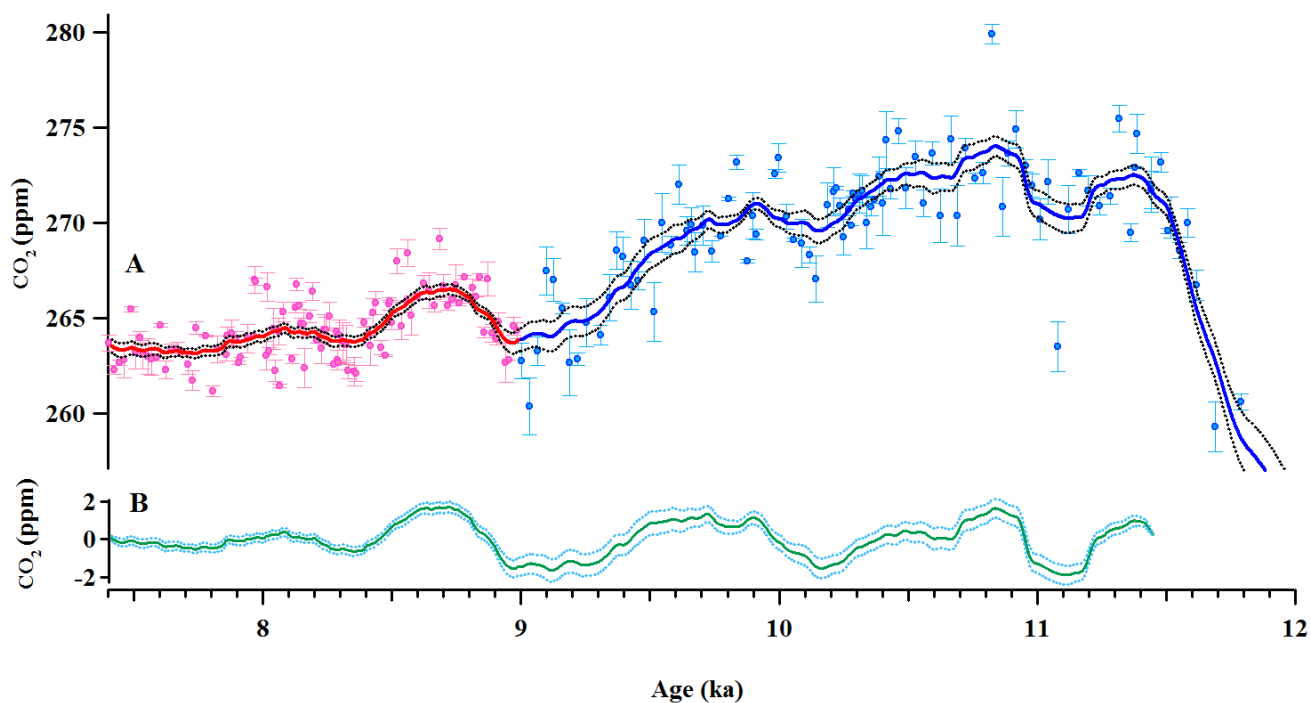


Figure 1. High-resolution atmospheric CO₂ records obtained from Siple Dome ice core, Antarctica during the early Holocene. A. Pink and blue circles are Siple Dome ice core records obtained at Oregon State University (Ahn et al., 2014) and Seoul National University (this study), respectively. Lines represent 250-yr running means and dotted lines, 2 σ uncertainties calculated from Monte Carlo simulation. For the simulation, we produced 1000 different sets of CO₂ concentrations which vary randomly with Gaussian propagation in their uncertainties. B. Green line indicates 250-yr running means of the original Siple Dome CO₂ data processed by high-pass filtering at 1/1800 year⁻¹. Blue line indicates 2 σ uncertainties of calculated from Monte Carlo simulation.

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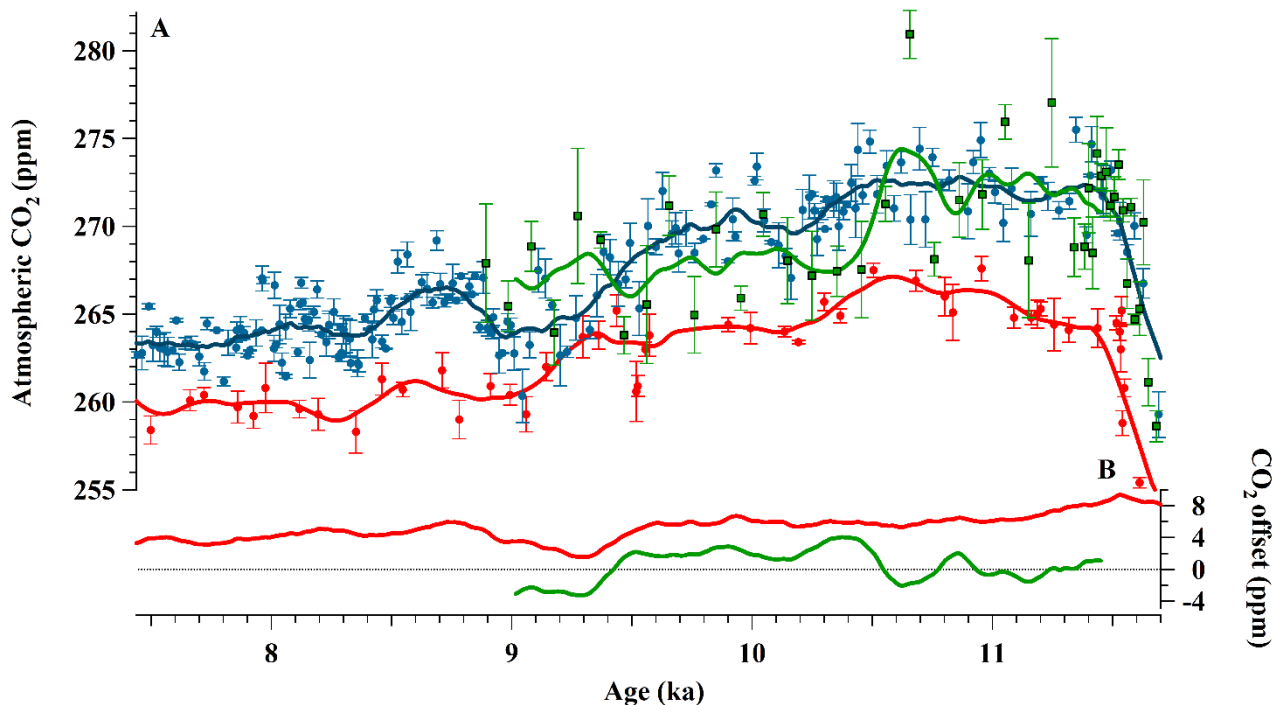


Figure 223. A: Atmospheric CO_2 records. Red dots: Atmospheric CO_2 record from Dome C ice core. Red line: 250-yr running means of atmospheric CO_2 record from Dome C ice core. Blue dots: Atmospheric CO_2 record from Siple Dome ice core. Blue line: 250-yr running means of atmospheric CO_2 record from Siple Dome ice core. Green dots: Atmospheric CO_2 record from WAIS Divide ice core. Green line: 250-yr running means of atmospheric CO_2 record from WAIS Divide ice core. B: CO_2 offset between Siple Dome CO_2 record and other published CO_2 records. Red line: CO_2 offset between Siple Dome CO_2 record and Dome C CO_2 record. Green line: CO_2 offset between Siple Dome CO_2 record and WAIS divide CO_2 record.

Figure 223. A: Atmospheric CO_2 records. Red dots: Atmospheric CO_2 record from Dome C ice core. Red line: 250-running mean of atmospheric CO_2 record from Dome C ice core. Blue dots: Atmospheric CO_2 record from Siple Dome ice core. Blue line: 250-running mean of atmospheric CO_2 record from Siple Dome ice core. Green dots: Atmospheric CO_2 record from WAIS Divide ice core. Green line: 250-running mean of atmospheric CO_2 record from WAIS Divide ice core. B: CO_2 offset between Siple Dome CO_2 record and other published CO_2 records. Red line: CO_2 offset between Siple Dome CO_2 record and Dome C CO_2 record. Green line: CO_2 offset between Siple Dome CO_2 record and WAIS divide CO_2 record.

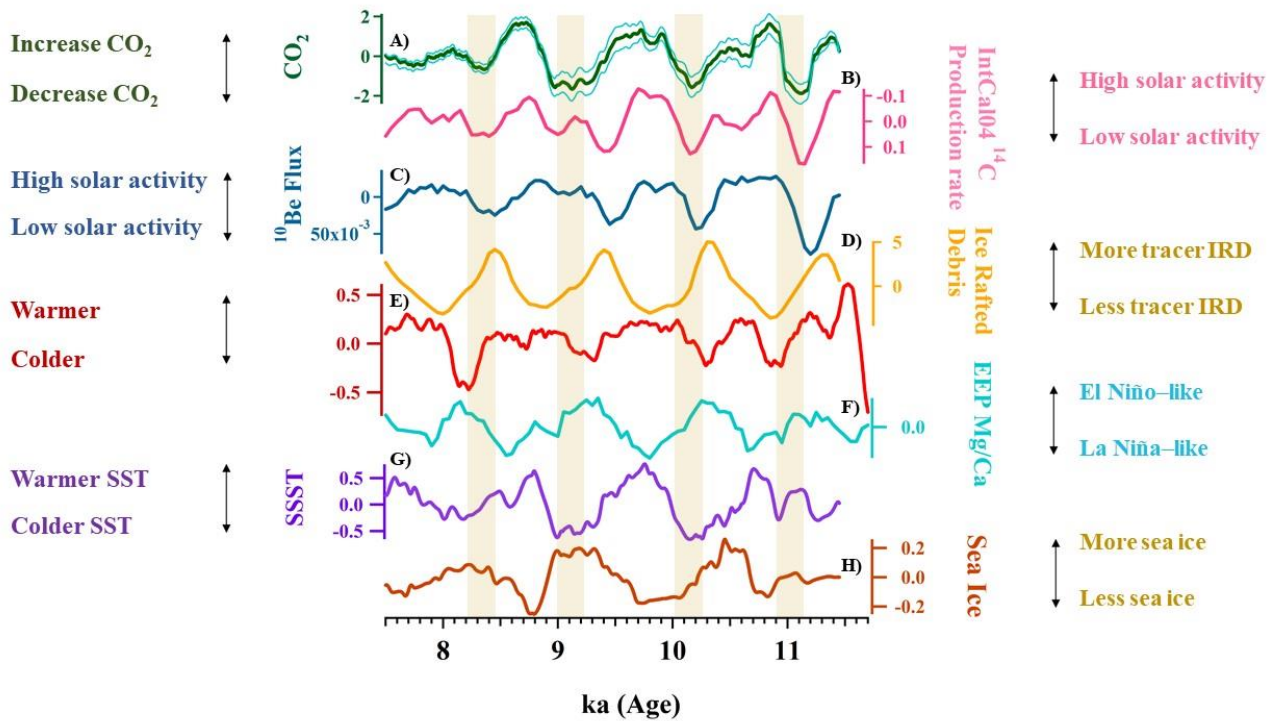


Figure 3. Comparison of atmospheric CO₂ with climatic proxy records over the early Holocene. The records were smoothed at ~250 years and high-pass filtered at 1/1800 year⁻¹. A) Atmospheric CO₂ record from Siple Dome (in this study). Dotted lines, 2σ uncertainties calculated from Monte Carlo simulation. B) ¹⁴C production rate from IntCal04 Δ¹⁴C data (Marchitto et al., 2010; Reimer et al., 2004). C) ¹⁰Be flux record from ice core on the GICC05 timescale (Finkel and Nishiizumi, 1997; Marchitto et al., 2010; Rasmussen et al., 2006; Vonmoos et al., 2006). D). IRD stacked records from the North Atlantic regions on untuned calibrated ¹⁴C age model (Bond et al., 2001; Marchitto et al., 2010). E) North Greenland Ice Core Project (NGRIP) ice core isotope ratio on the GICC05 timescale (Rasmussen et al., 2006). F) Sea surface temperature from the eastern equatorial Pacific indicating El Niño-like or La Niña-like conditions (Marchitto et al., 2010). The data was radiocarbon dated by accelerator mass spectrometry (AMS), which was recalibrated by the Marine09 calibration curve (Reimer et al., 2009). G) Sea surface temperature from the Polar Front of the Southern Ocean on the chronology of Mortyn et al. (2003) (Nielsen et al., 2004). H) Sea ice presence from the Polar Front of the Southern Ocean on the chronology of Mortyn et al. (2003) (Nielsen et al., 2004).

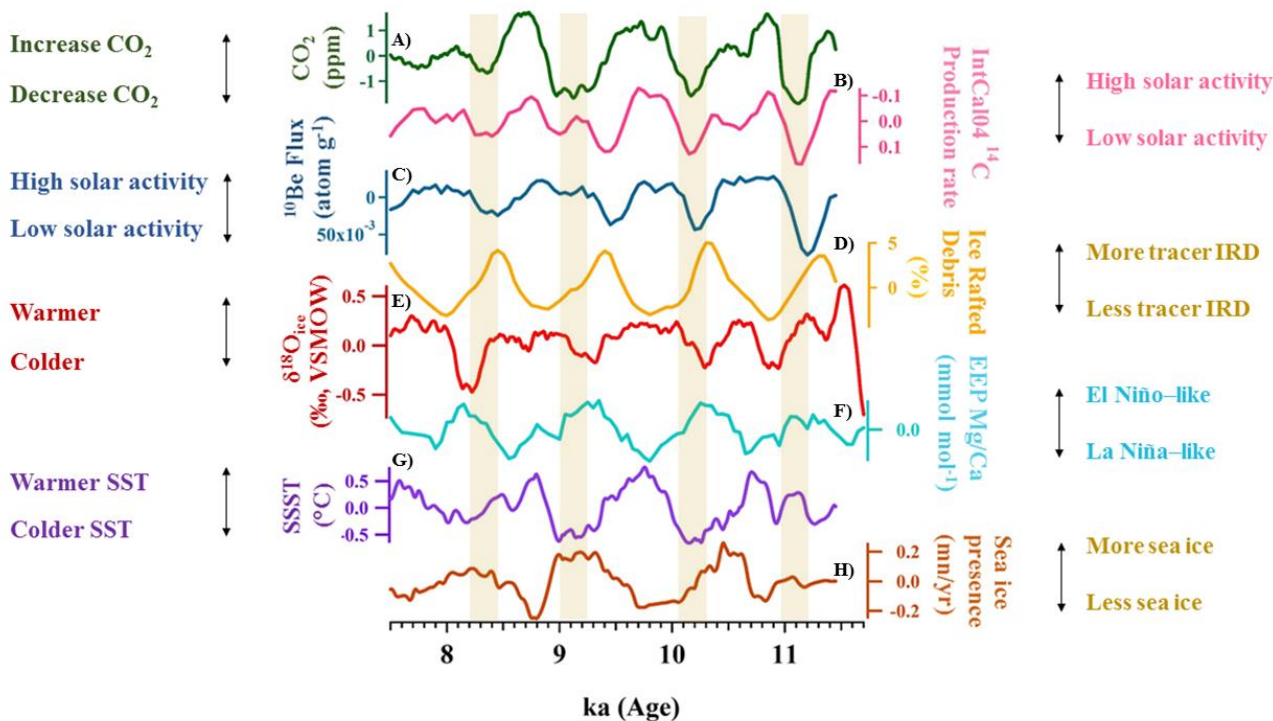
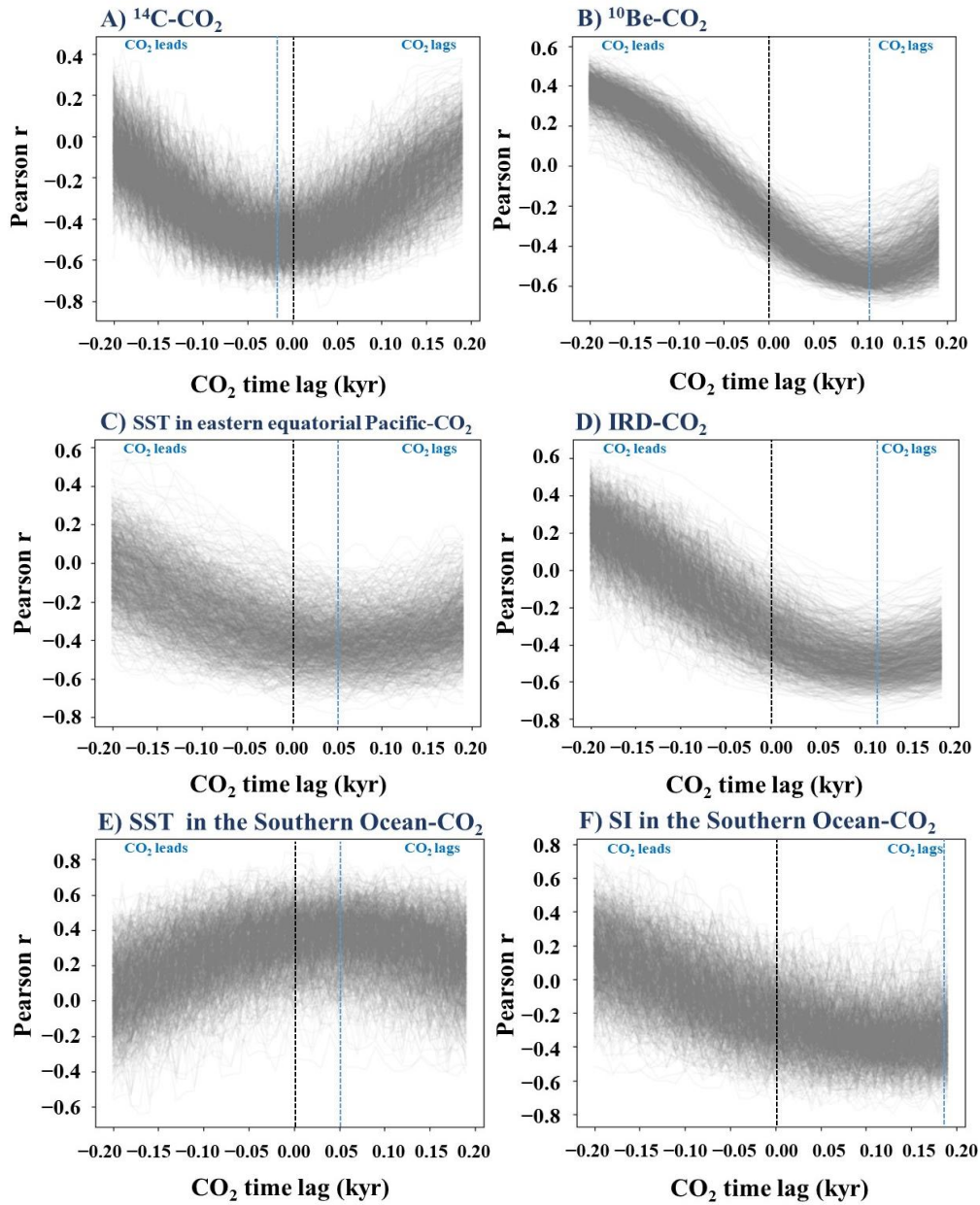


Figure 32. Comparison of atmospheric CO_2 with climatic proxy records over the early Holocene. The records were smoothed at ~ 250 years and high pass filtered at $\frac{1}{1800}$ year $^{-1}$. A) Atmospheric CO_2 record from Siple Dome (in this study). B) ^{14}C production rate from tree ring (Marchitto et al., 2010; Reimer et al., 2004). C) ^{10}Be flux record from ice core (Finkel and Nishiizumi, 1997; Marchitto et al., 2010; Vonmoos et al., 2006). D) IRD stacked records from the North Atlantic regions (Bond et al., 2001; Marchitto et al., 2010). E) North Greenland Ice Core Project (NGRIP) ice core isotope ratio (Rasmussen et al., 2006). F) Sea surface temperature from the eastern equatorial Pacific indicating El Niño-like or La Niña-like conditions (Marchitto et al., 2010). G) Sea surface temperature from the Polar Front of the Southern Ocean (Nielsen et al., 2004). H) Sea ice presence from the Polar Front of the Southern Ocean (Nielsen et al., 2004).



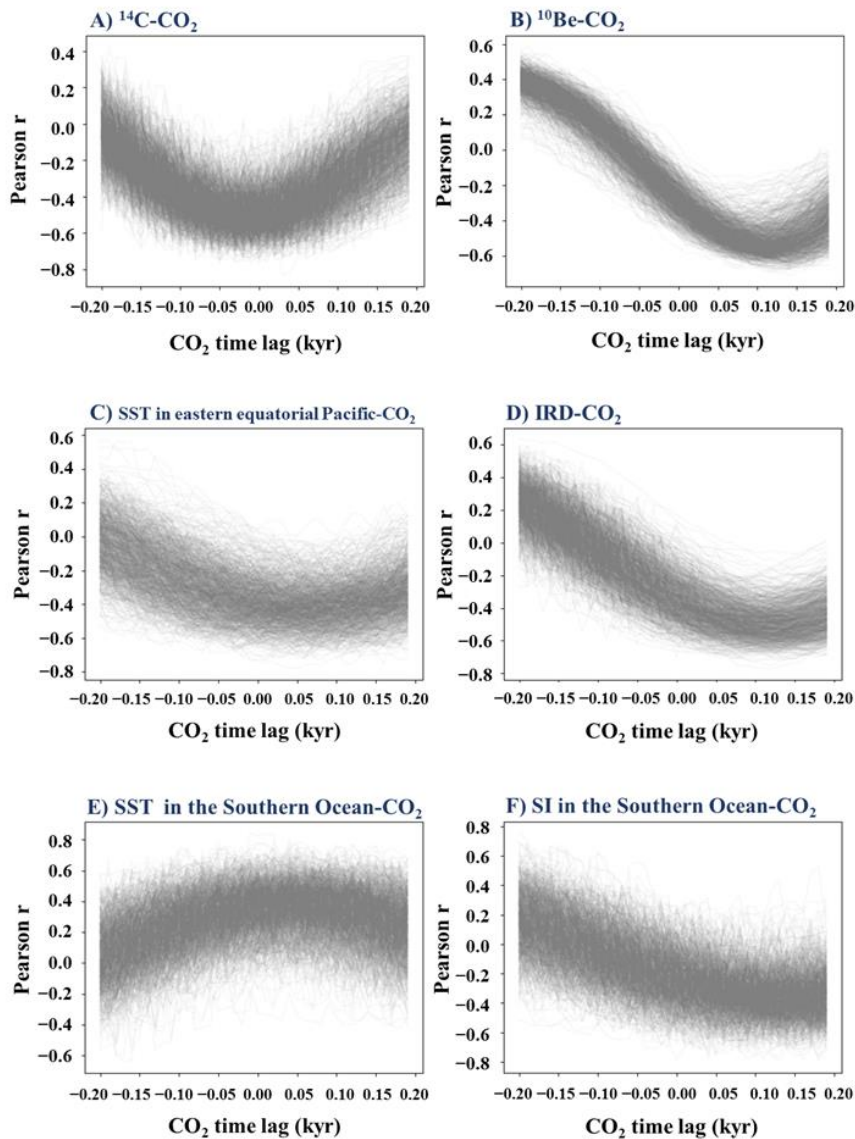


Figure 4. Correlation coefficients between CO₂ and proxies with CO₂ time lag calculated from Monte Carlo simulation. Vertical lines in black indicate zero time lag. Vertical lines in blue indicate maximum correlation coefficients between CO₂ and proxies with CO₂ time lag. A) ¹⁴C production rate and atmospheric CO₂. B) ¹⁰Be flux and atmospheric CO₂. C) SST in the eastern equatorial Pacific and atmospheric CO₂. D) IRD from the North Atlantic and atmospheric CO₂. E) SST in the East Equatorial Pacific indicating El Niño--like or La Niña--like conditions and atmospheric CO₂. F) SI in the East Equatorial Pacific and atmospheric CO₂.

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Figure 43. Correlation coefficients between CO₂ and proxies with CO₂ time lag calculated from Monte Carlo simulation. A) ¹⁴C production rate and atmospheric CO₂. B) ¹⁰Be flux and atmospheric CO₂. C) SST in the eastern equatorial Pacific and atmospheric CO₂. D) IRD from the North Atlantic and atmospheric CO₂. E) SST in the East Equatorial Pacific indicating El Niño-like or La Niña-like conditions and atmospheric CO₂. F) SI in the East Equatorial Pacific and atmospheric CO₂.

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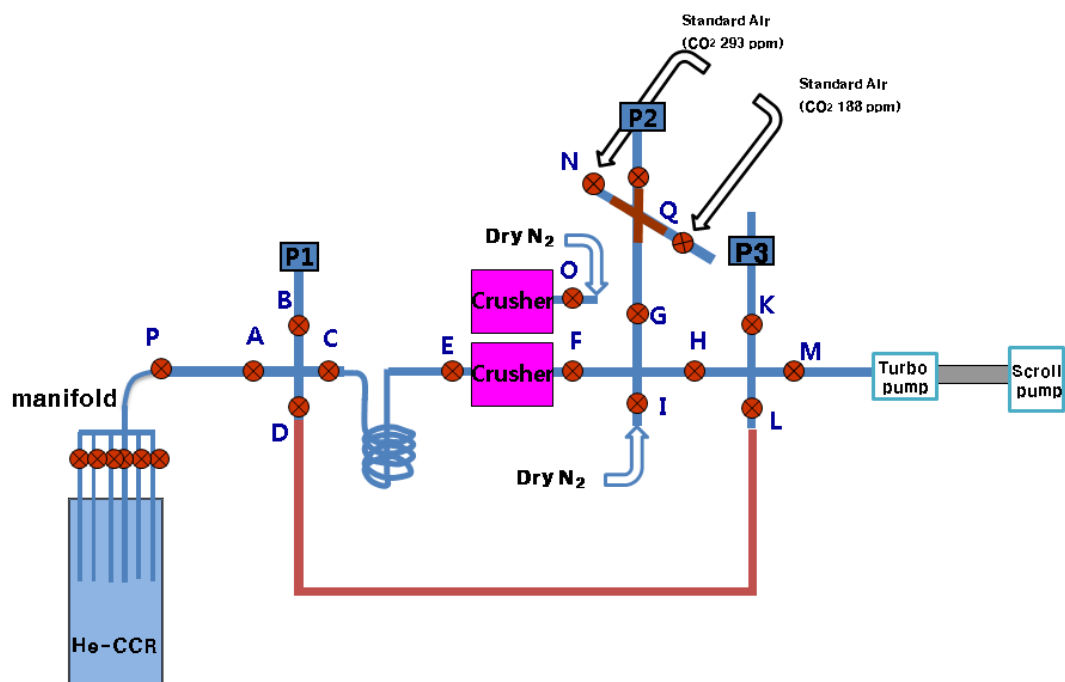


Figure S1. Schematic diagram of the dry extraction system used for this study. Details are described in Ahn et al. (2009). He-CCR stands for He-Closed Cycled Refrigerator.

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Table S1. Comparison of CO₂ analysis with two different standard airs (188.89 and 293.325 ppm) in order to check the linearity in the gas chromatograph.

<u>Siple</u> <u>Dome</u> <u>ice core</u>	<u>With 188.9 ppm Standard air</u>				<u>With 293.3 ppm Standard air</u>				<u>Difference</u>	
<u>Mean</u> <u>depth</u> <u>(m)</u>	<u>CO₂</u> <u>(ppm)</u>	<u>mean</u> <u>CO₂</u> <u>(ppm)</u>	<u>uncertainty</u> <u>(ppm)</u>	<u># of</u> <u>samples</u>	<u>CO₂</u> <u>(ppm)</u>	<u>mean</u> <u>CO₂</u> <u>(ppm)</u>	<u>uncertainty</u> <u>(ppm)</u>	<u># of</u> <u>samples</u>	<u>CO₂</u> <u>(ppm)</u>	<u>uncertainty</u> <u>(ppm)</u>
<u>640.53</u>	<u>248.7</u>	<u>249.6</u>	<u>0.9</u>	<u>2</u>	<u>250.7</u>	<u>249.6</u>	<u>0.9</u>	<u>3</u>	<u>0.0</u>	<u>1.3</u>
	<u>250.5</u>				<u>247.8</u>					
					<u>250.4</u>					
<u>644.22</u>	<u>246.0</u>	<u>246.0</u>	<u>0.1</u>	<u>2</u>	<u>247.8</u>	<u>247.2</u>	<u>0.6</u>	<u>2</u>	<u>1.2</u>	<u>0.6</u>
	<u>246.1</u>				<u>246.6</u>					
<u>668.09</u>	<u>241.7</u>	<u>242.2</u>	<u>0.5</u>	<u>2</u>	<u>242.3</u>	<u>243.1</u>	<u>0.8</u>	<u>2</u>	<u>0.9</u>	<u>0.9</u>
	<u>242.7</u>				<u>243.8</u>					
<u>671.51</u>	<u>243.7</u>	<u>243.0</u>	<u>0.7</u>	<u>2</u>	<u>242.8</u>	<u>243.5</u>	<u>0.4</u>	<u>3</u>	<u>0.5</u>	<u>0.8</u>
	<u>242.3</u>				<u>243.5</u>					
					<u>244.2</u>					
<u>673.47</u>	<u>241.1</u>	<u>240.4</u>	<u>0.8</u>	<u>2</u>	<u>239.1</u>	<u>239.8</u>	<u>0.7</u>	<u>2</u>	<u>-0.6</u>	<u>1.0</u>
	<u>239.6</u>				<u>240.5</u>					
									<u>Average</u>	<u>0.4</u>
										<u>0.9</u>

Table S2. Interlaboratory comparison between SNU (Seoul National University) and OSU (Oregon State University) using Siple Dome ice core.

<u>Depth range</u> <u>(m)</u>	<u>SNU CO₂</u> <u>(ppm)</u>	<u># of replicates</u>	<u>OSU CO₂</u> <u>(ppm)</u>	<u># of replicates</u>	<u>SNU-OSU</u> <u>(ppm)</u>
<u>490.17–490.22</u>	<u>266.8</u>	<u>2</u>	<u>265.7</u>	<u>3</u>	<u>1.1</u>
<u>500.40–500.45</u>	<u>263.8</u>	<u>2</u>	<u>264.1</u>	<u>4</u>	<u>-0.3</u>
<u>501.87–502.41</u>	<u>263.8</u>	<u>2</u>	<u>262.8</u>	<u>2</u>	<u>1.1</u>
<u>506.60–506.65</u>	<u>264.9</u>	<u>3</u>	<u>265.0</u>	<u>2</u>	<u>-0.1</u>
<u>522.90–523.10</u>	<u>266.7</u>	<u>2</u>	<u>266.3</u>	<u>2</u>	<u>0.5</u>
<u>523.28–523.33</u>	<u>265.2</u>	<u>2</u>	<u>265.8</u>	<u>2</u>	<u>-0.6</u>
<u>530.50–530.55</u>	<u>266.9</u>	<u>2</u>	<u>266.4</u>	<u>2</u>	<u>0.6</u>
				<u>Average</u>	<u>0.3</u>
				<u>Standard deviation</u>	<u>0.7</u>

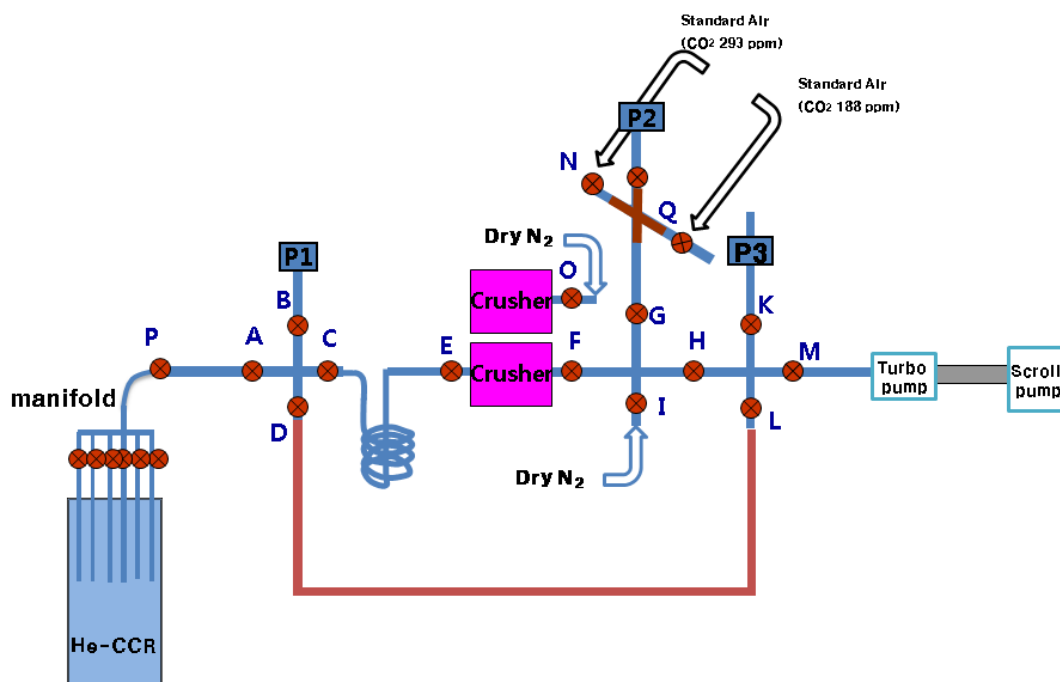


Figure S1. Schematic diagram of the dry extraction system used for this study. Details are described in Ahn et al. (2009). He-CCR stands for He Closed Cycled Refrigerator.

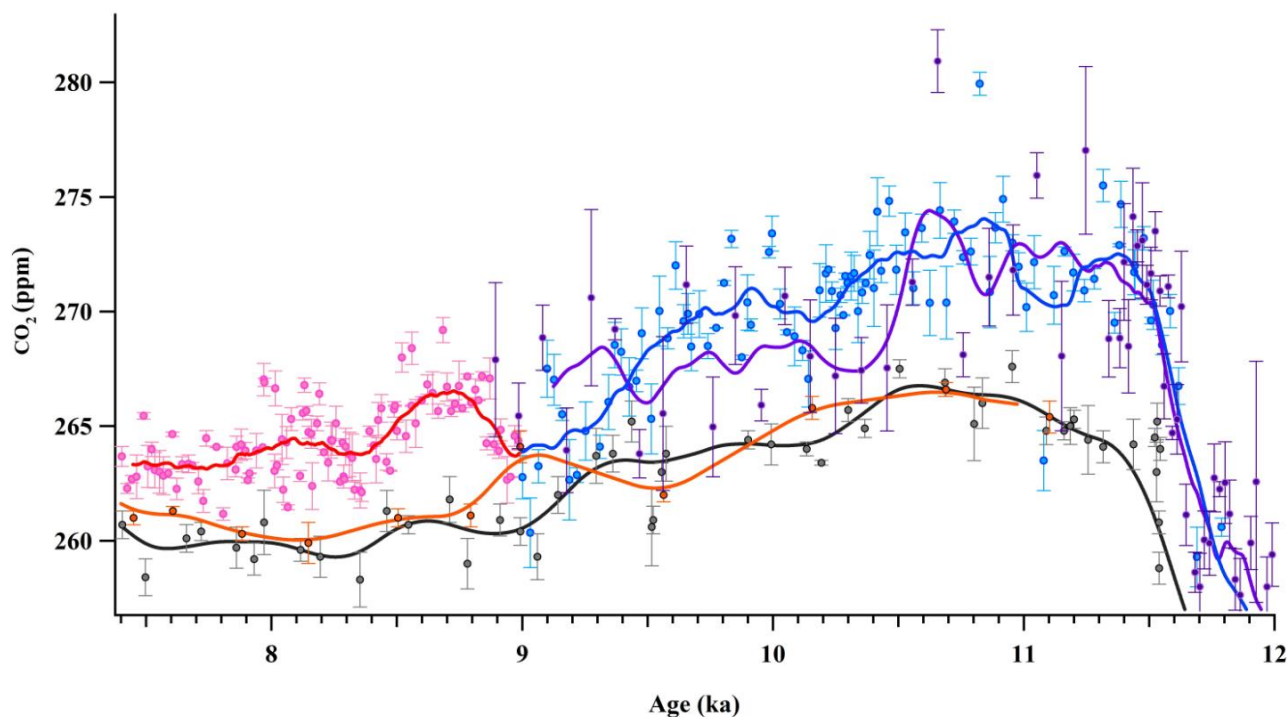


Figure S2. Comparison of Antarctic ice core CO₂ records. Pink and blue circles are Siple Dome ice core records obtained at Oregon State University (OSU) (Ahn et al., 2014) and Seoul National University (this study), respectively. Purple circles from WAIS Divide ice were analysed at OSU (Marcott et al., 2014). Black and orange circles are from the Dome C and Taylor Dome ice cores (Flückiger et al., 2002; Monnin et al., 2001; Monnin et al., 2004), respectively and they were analysed at University of Bern. Lines represent 250 year running means.

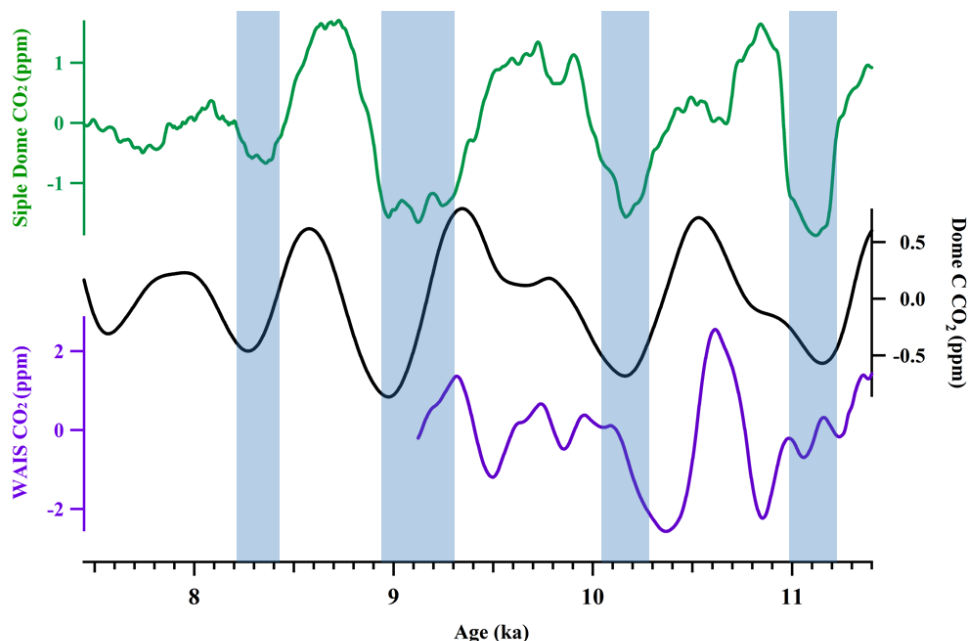


Figure S3. Comparison of Antarctic ice core CO₂ records. To see long-term trend and local minima of these data, the records were smoothed at ~250 years and high-pass filtered at $\frac{1}{1800}$ year⁻¹. Green line is Siple Dome ice core records obtained at OSU (Ahn et al., 2014) and SNU (this study). Black line is Dome C obtained from UB (Monnin et al., 2001; Monnin et al., 2004) and purple line indicates WAIS divide ice core from OSU (Marcott et al., 2014).

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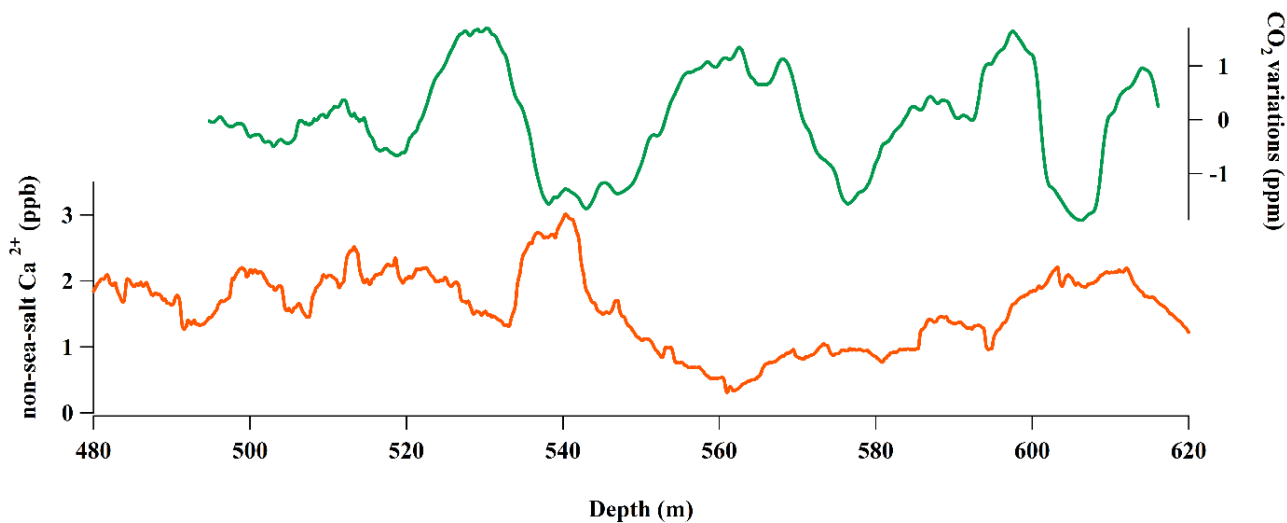


Figure S423. Comparison of Siple Dome millennial (filtered) CO₂ record (Ahn et al., 2014) and non-sea-salt Ca²⁺ (~~nssCa²⁺~~) on depth domain. The non-sea-salt Ca²⁺~~nssCa~~ data are 250-year running means.

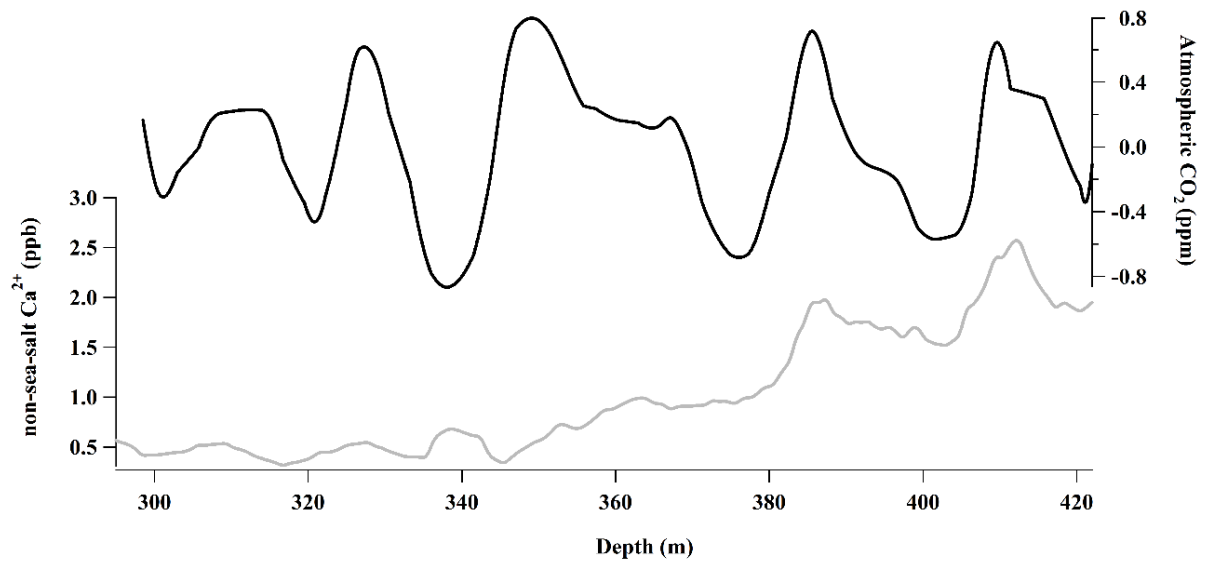
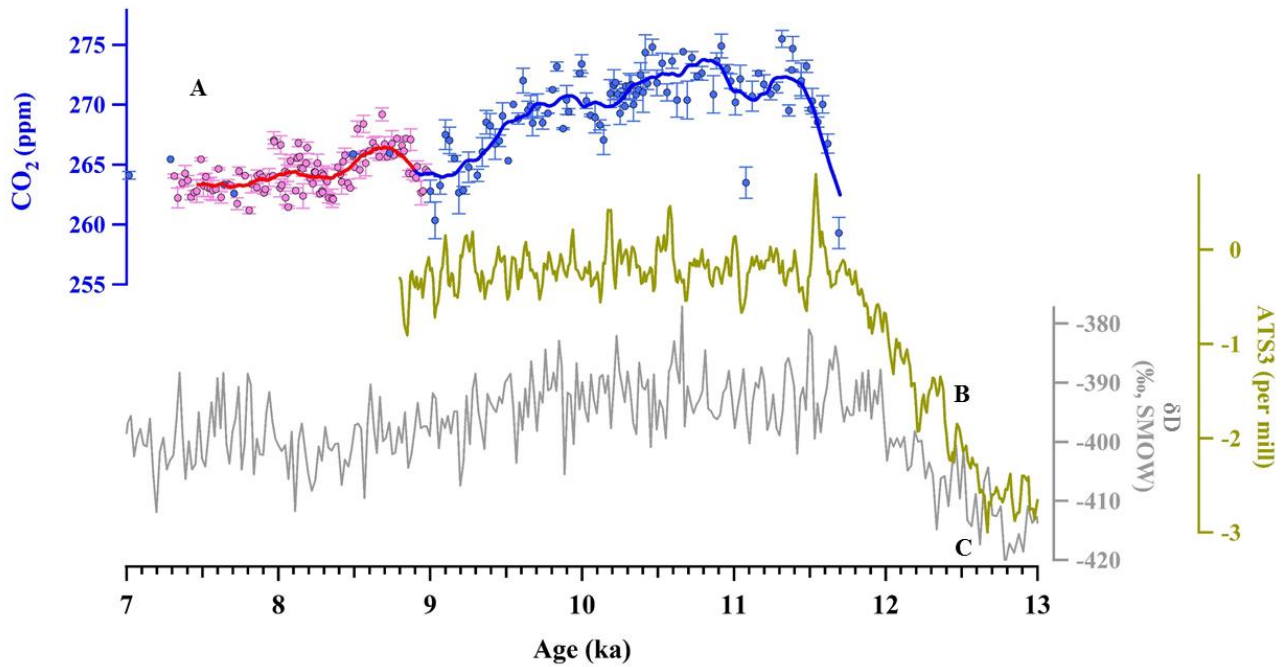


Figure S345. Comparison of Dome C millennial (filtered) CO₂ record (Monnin et al., 2001; Monnin et al., 2004) and non-sea-salt Ca²⁺Ca (nssCa) (Lambert et al., 2012) on depth domain. The non-sea-salt Ca²⁺nssCa data are 250-yr running means 250-year running mean.



- 5 **Figure S456.** High-resolution atmospheric CO₂ records obtained from Siple Dome ice core, Antarctica during the early Holocene. A) Pink and blue circles are Siple Dome ice core records obtained at OSU (Ahn et al., 2014) and SNU (this study), respectively. Lines in pink and blue represent 250-yr running means. B) Antarctic temperature Stack 3 (ATS3) developed by Buizert et al. (2018) using five records: Dome C, Dome Fuji, Talos Dome, EPICA Dronning Maud Land and WAIS Divide. C) Grey line indicates δD in the Dome C ice core, Antarctica (Jouzel et al., 2007).

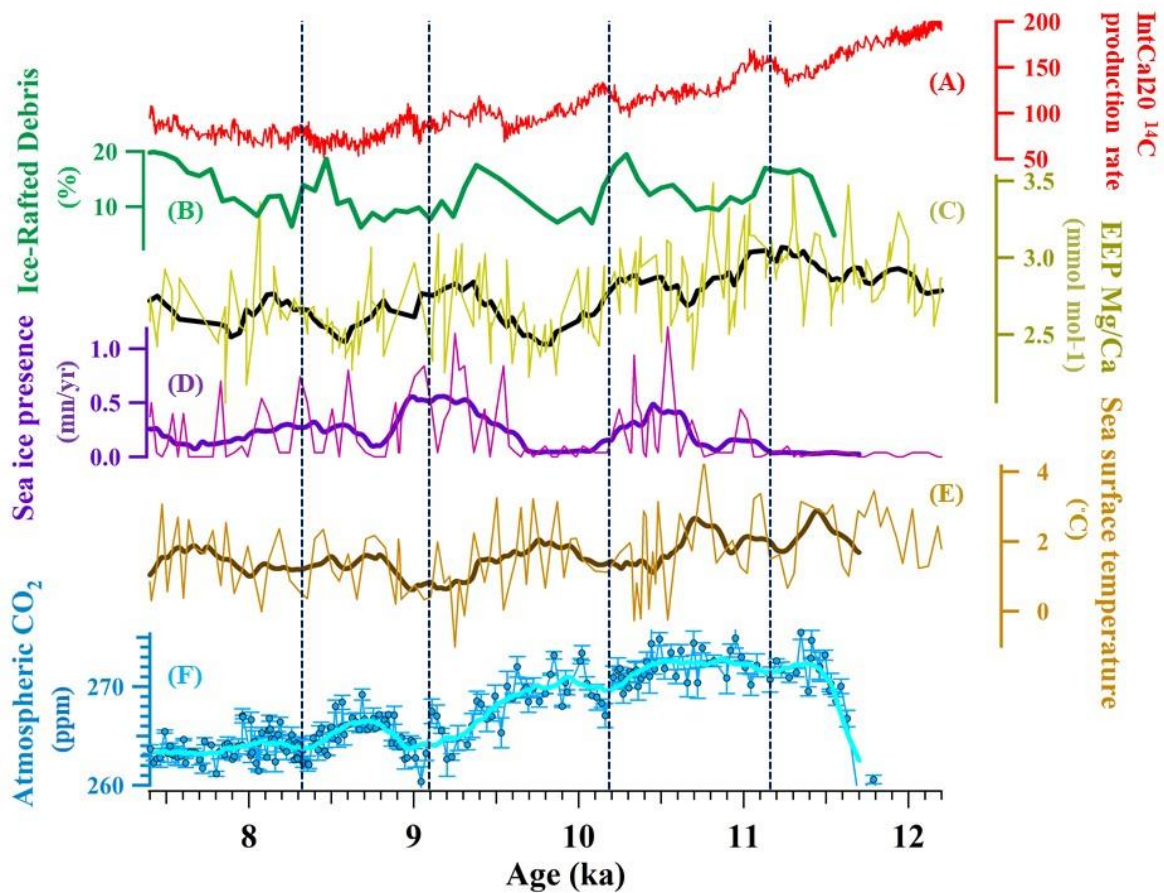


Figure S52. Comparison of atmospheric CO₂ with climatic proxy records over the early Holocene. Lines represent 250-yr running means. (A) IntCal20-¹⁴C production rate (Reimer et al., 2020). (B) Ice rafted debris stacked records from the North Atlantic regions on untuned calibrated ¹⁴C age model (Bond et al., 2001; Marchitto et al., 2010). (C) Sea surface temperature from the eastern equatorial Pacific indicating El Niño-like or La Niña-like conditions (Marchitto et al., 2010). (D) Sea ice presence from the Polar Front of the Southern Ocean on the chronology of Mortyn et al. (2003) (Nielsen et al., 2004). (E) Sea surface temperature from the Polar Front of the Southern Ocean on the chronology of Mortyn et al. (2003) (Nielsen et al., 2004). (F) Atmospheric CO₂ record from Siple Dome (in this study).

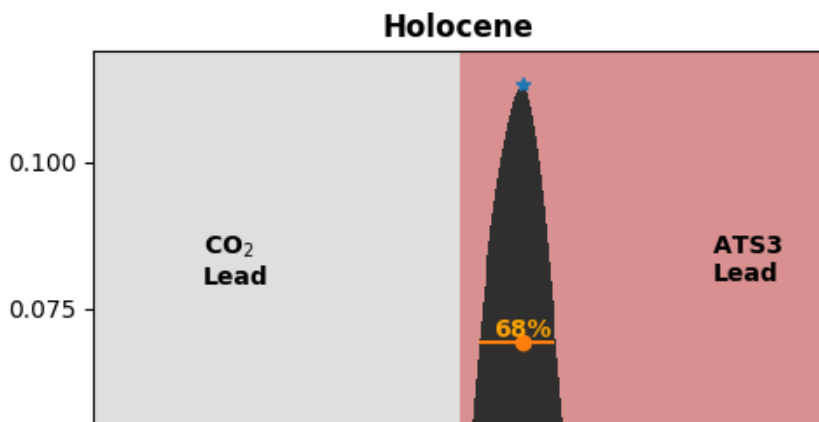


Figure S67. The relationship between CO₂ and Antarctic temperature Stack 3 (ATS3) at the onset of the Holocene. y axis indicates normalized probability density of an ATS3 lead. x axis indicate the time lag between temperature and CO₂ in years. Positive x axis indicate ATS3 leads CO₂. This graph shows ATS3 leads CO₂ by 190 yrs within a range of 60 to 285 yrs with 68% probability and within a range of -90 to 397 yrs with 95% probability.

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Table S1. Comparison of CO₂ analysis with two different standard airs (188.89 and 293.25 ppm) in order to check the linearity in the gas chromatograph.

Siple Dome ice core	With 188.9 ppm Standard air				With 293.3 ppm Standard air				Difference	
	CO ₂ (ppm)	mean CO ₂ (ppm)	uncertainty (ppm)	# of samples	CO ₂ (ppm)	mean CO ₂ (ppm)	uncertainty (ppm)	# of samples	CO ₂ (ppm)	uncertainty (ppm)
640.53	248.7	249.6	0.9	2	250.7	249.6	0.9	3	0.0	1.3
	250.5				247.8					
644.22	246.0	246.0	0.1	2	247.8	247.2	0.6	2	1.2	0.6
	246.1				246.6					
668.09	241.7	242.2	0.5	2	242.3	243.1	0.8	2	0.9	0.9
	242.7				243.8					
671.51	243.7	243.0	0.7	2	242.8	243.5	0.4	3	0.5	0.8
	242.3				243.5					
673.47	241.1	240.4	0.8	2	239.1	239.8	0.7	2	-0.6	1.0
	239.6				240.5					
Average									0.4	0.9

~~Table S2. Interlaboratory comparison between SNU and OSU using Siple Dome ice core.~~

Depth range	SNU CO₂	#of replicates	OSU CO₂	#of replicates	SNU-OSU
(m)	(ppm)		(ppm)		(ppm)
490.17-490.22	266.8	2	265.7	3	1.1
500.40-500.45	263.8	2	264.1	4	-0.3
501.87-502.41	263.8	2	262.8	2	1.1
506.60-506.65	264.9	3	265.0	2	-0.1
522.90-523.10	266.7	2	266.3	2	0.5
523.28-523.33	265.2	2	265.8	2	-0.6
530.50-530.55	266.9	2	266.4	2	0.6
				Average	0.3
				Standard deviation	0.7

5 Monte Carlo simulation

We calculate uncertainties of a smoothed CO₂ record by using Monte Carlo simulation. The uncertainty band is narrow, which is attributed to the removal of the high frequency signal by 250-yr running means. When the Monte Carlo simulation was conducted, we considered that each data follows a normal distribution. The width of the error band is affected by neighbouring data points. If the data points are close together, the error of neighbouring data points in the opposite direction can be canceled out, resulting in a narrow uncertainty band.

5 To assess variability on the millennial time scale, we evaluated 250-yr running means and their uncertainties by using a new Monte Carlo approach. Random sampling was made from a probability distribution for each measured value and its standard deviation. If the standard deviation was smaller than the average reproducibility of the measurement ($1\sigma = 0.87$ ppm), we used 0.87 ppm as the uncertainty of a measured value. Then, 1-yr interpolation and resampling were applied to generate an evenly-spaced time series and to calculate the 250-yr running means. We repeated this series of simulations 10,000 times and evaluated the mean of 250-yr running means and its uncertainty (shown as 2σ in Figure S6). We used a modified Akima method using the built-in makima function in Matlab for the interpolation. The different types of interpolation and smoothing methods resulted in insignificant differences in the 250-yr running means.

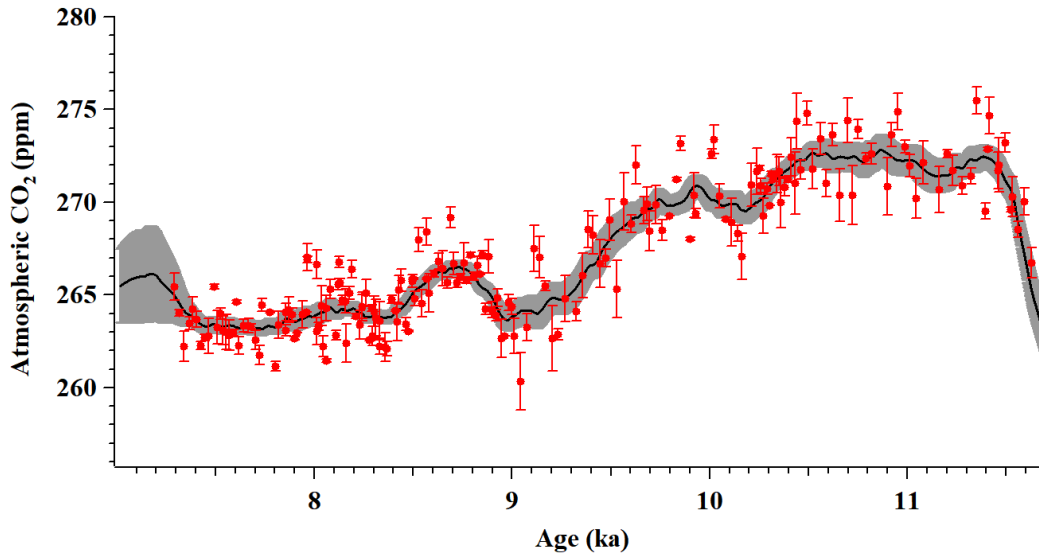


Figure S4S6. Red circles are Siple Dome ice core records during the early Holocene (11.7–7.4 ka). The black line indicates the average of 10,000 modified akima simulations showing an error-weighted average of the CO₂ record. The dark shaded indicates 2σ uncertainties calculated from modified akima simulations.

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Correlation coefficients between CO₂ and climate proxies

15 In order to assess the relationship between CO₂ and climate, we calculate correlations and estimate leads and lags between CO₂ and proxies of solar activity as well as climate proxies thought to be themselves related to solar activity. The Pearson correlation coefficient r is commonly used to verify relationships between variables. However, r does not take chronological uncertainty into account. As such, we apply a Monte Carlo procedure to estimate the correlations between CO₂ and climate proxies. In the procedure, we adjust the chronologies of the two series, within their chronological uncertainties, and re-calculate

r. We do so 1,000 times for each pair, allowing us to calculate a mean correlation coefficient that is more representative of the relationship between time-uncertain series.

We use a similar method to calculate the significance of this correlation against a random red-noise process. At each of the 1,000 steps, we use an AR(1) model (lag-1 auto regression) to fit data series. We then use these AR(1) characteristics to
5 randomly generate two synthetic series with the same red noise. Then, we calculate the percentage of correlations between the randomized synthetic series that are lower than the correlation coefficients of the real series to assess the significance of the correlation.

Finally, we can calculate the maximum-correlation lag between the two series. At each step in the 1,000 iteration Monte Carlo procedure, we calculate the lag which gives the maximum correlation by shifting one of the series by 10-yr increments, for
10 constant lags between -200 (a CO₂ lead) and 200 yrs (a CO₂ lag). Then, we make a histogram of the calculated maximum-correlation lags, from which the mode can be selected as an approximation of the phasing between the two series. We also report the maximum correlation between the two series, but note that this is not representative of the actual relationship (it simply gives an idea of how much the correlation can be improved by adding a lag between the two series).

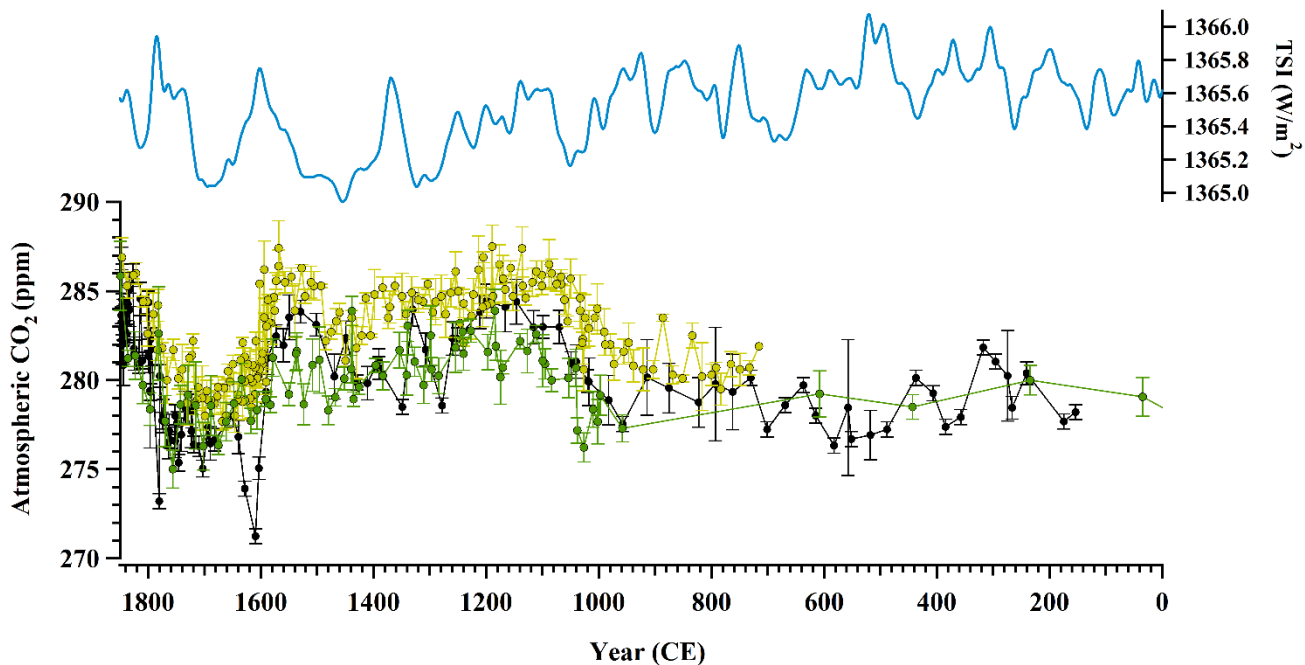


Figure S7. Atmospheric CO₂ from Antarctic ice cores during the last 2,000 yrs. Blue line: total solar irradiance (TSI) (Roth and Joos, 2013). Yellow dots: atmospheric CO₂ from WAIS Divide ice core (Ahn et al., 2012). Green dots: atmospheric CO₂ from EPICA Dronning Maud Land (EDML) (Monnin et al., 2004; Siegenthaler et al., 2005). Black dots: atmospheric CO₂ from Law Dome (Rubino et al., 2019).

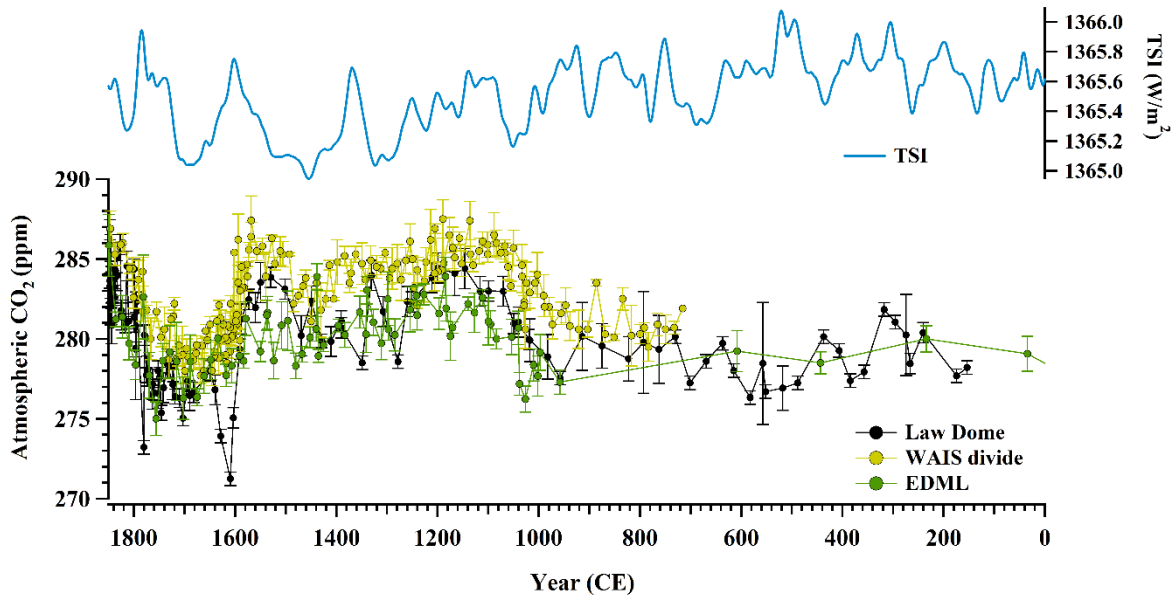


Figure 2. Atmospheric CO₂ from Antarctic ice cores during the last 2,000 years. Blue line: total solar irradiance (TSI) (Roth and Joos, 2013). Yellow dots: atmospheric CO₂ from WAIS Divide ice core (Ahn et al., 2012). Green dots: atmospheric CO₂ from EDML (Monnin et al., 2004; Siegenthaler et al., 2005). Black dots: atmospheric CO₂ from Law dome (Rubino et al., 2019).

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