

# Deep Ocean Temperatures through Time

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## Abstract

Benthic oxygen isotope records are commonly used as a proxy for global mean surface temperatures during the late Cretaceous and Cenozoic, and the resulting estimates have been extensively used in characterising major trends and transitions in the climate system, and for analysing past climate sensitivity. However, some fundamental assumptions governing this proxy have rarely been tested. Two key assumptions are: (a) benthic foraminiferal temperatures are geographically well-mixed and are linked to surface high latitude temperatures, and (b) surface high latitude temperatures are well correlated with global mean temperatures. To investigate the robustness of these assumptions through geological time, we performed a series of 109 climate model simulations using a unique set of paleogeographical reconstructions covering the entire Phanerozoic at the stage-level. The simulations have been run for at least 5000 model years to ensure that the deep ocean is in dynamic equilibrium. We find that the correlation between deep ocean temperatures and global mean surface temperatures is good for the Cenozoic and thus the proxy data are reliable indicators for this time period, albeit with a standard error of 2K. This uncertainty has not normally been assessed and needs to be combined with other sources of uncertainty when, for instance, estimating climate sensitivity based on using  $\delta^{18}\text{O}$  measurements from benthic foraminifera. The correlation between deep and global mean surface temperature becomes weaker for pre-Cenozoic time periods (when the paleogeography is significantly different than the present-day). The reasons for the weaker correlation include variability in the source region of the deep water (varying hemispheres but also varying latitudes of sinking), the depth of ocean overturning (some extreme warm climates have relatively shallow and sluggish circulations weakening the link between surface and deep ocean), and the extent of polar amplification (e.g. ice albedo feedbacks). Deep ocean sediments prior to the Cretaceous are rare, so extending the benthic foram proxy further into deeper time is problematic, but the model results presented here would suggest that the deep ocean temperatures from such time periods would probably be an unreliable indicator of global mean surface conditions.

32

## 33 1. Introduction

34

35 One of the most widely used proxies for estimating global mean surface temperature through the  
36 last 100 million years is benthic  $\delta^{18}\text{O}$  measurements from deep sea foraminifera (Zachos et al.,  
37 2001), (Zachos et al., 2008), (Cramer et al., 2009), (Friedrich et al., 2012), (Westerhold et al., 2020).  
38 Two key underlying assumptions are that  $\delta^{18}\text{O}$  from benthic foraminifera represents deep ocean  
39 temperature (with a correction for ice volume and any vital effects), and that the deep ocean water  
40 masses originate from surface water in polar regions. By further assuming that polar surface  
41 temperatures are well correlated with global mean surface temperatures, then deep ocean isotopes  
42 can be assumed to track global mean surface temperatures. More specifically, (Hansen et al., 2008),  
43 and (Hansen and Sato, 2012) argue that changes in high latitude sea surface temperatures are  
44 approximately proportional to global mean surface temperatures because changes are generally  
45 amplified at high latitudes but that this is offset because temperature change is amplified over land  
46 areas. They therefore directly equate changes in benthic ocean temperatures with global mean  
47 surface temperature.

48 The resulting estimates of global mean surface air temperature have been used to understand past  
49 climates (e.g. (Zachos et al., 2008), (Westerhold et al., 2020)). Combined with estimates of  
50 atmospheric  $\text{CO}_2$  they have also been used to estimate climate sensitivity (e.g. (Hansen et al., 2013))  
51 and hence contribute to the important ongoing debate about the likely magnitude of future climate  
52 change.

53 However, some of the underlying assumptions behind the method remains largely untested, even  
54 though we know that there are major changes to paleogeography and consequent changes in ocean  
55 circulation and location of deep-water formation in the deep past (e.g. (Lunt et al., 2010; Nunes and  
56 Norris, 2006); (Donnadieu et al., 2016); (Farnsworth et al., 2019a); (Ladant et al., 2020)). Moreover,  
57 the magnitude of polar amplification is likely to vary depending on the extent of polar ice caps, and  
58 changes in cloud cover (Sagoo et al., 2013), (Zhu et al., 2019). These issues are likely to modify the  
59 correlation between deep ocean temperatures and global mean surface temperature or, at the very  
60 least, increase the uncertainty in reconstructing past global mean surface temperatures.

61 The aim of this paper is two-fold, (1) we wish to document the setup and initial results from a unique  
62 set of 109 climate model simulations of the whole Phanerozoic era (last 540 million years) at the  
63 stage level (approximately every 5 million years), and (2) we will use these simulations to investigate

64 the accuracy of the deep ocean temperature proxy in representing global mean surface  
65 temperature.

66 The focus of the work is to examine the link between benthic ocean temperatures and surface  
67 conditions. However, we evaluate the fidelity of the model by comparing the model predicted ocean  
68 temperatures to estimates of the isotopic temperature of the deep ocean during the past 110  
69 million years ((Zachos et al., 2008), (Cramer et al., 2009), (Friedrich et al., 2012)), and model  
70 predicted surface temperatures to the sea surface temperatures estimates of (O'Brien et al., 2017)  
71 and (Cramwinckel et al., 2018). This gives us confidence that the model is behaving plausibly but we  
72 emphasise that the fidelity of the simulations is strongly influenced by the accuracy of CO<sub>2</sub> estimates  
73 through time. We then use the complete suite of climate simulations to examine changes in ocean  
74 circulation, ice formation, and the impact on ocean and surface temperature. Our paper will not  
75 consider any issues associated with assumptions regarding the relationship between deep-sea  
76 foraminifera  $\delta^{18}\text{O}$  and various temperature calibrations because our model does not simulate the  
77  $\delta^{18}\text{O}$  of sea water (or vital effects).

78

## 79 2. Simulation Methodology

### 80 2.1 Model Description

81 We use a variant of the Hadley Centre model, HadCM3 ((Pope et al., 2000), (Gordon et al., 2000))  
82 which is a coupled atmosphere-ocean-vegetation model. The specific version, HadCM3BL-M2.1aD, is  
83 described in detail in (Valdes et al., 2017). The model has a horizontal resolution of 3.75° x 2.5° in  
84 longitude/latitude (roughly corresponding to an average grid box size of ~300km) in both the  
85 atmosphere and the ocean. The atmosphere has 19 unequally spaced vertical levels, and the ocean  
86 has 20 unequally spaced vertical levels. To avoid singularity at the poles, the ocean model always has  
87 to have land at the poles (90N and 90S), but the atmosphere model can represent the poles  
88 correctly (i.e. in the pre-industrial geography, the atmosphere considers there is sea ice covered  
89 ocean at the N. Pole but the ocean model has land and hence there is no ocean flow across the  
90 pole). Though HadCM3 is a relatively low resolution and low complexity model compared to the  
91 current CMIP5/CMIP6 state-of-the-art model, its performance at simulating modern climate is  
92 comparable to many CMIP5 models (Valdes et al., 2017). The performance of the dynamic  
93 vegetation model compared to modern observations is also described in (Valdes et al., 2017) but the  
94 modern deep ocean temperatures are not described in that paper. We therefore include a  
95 comparison to present day observed deep ocean temperatures in section 3.1.

96 To perform paleo-simulations, several important modifications to the standard model described in  
97 (Valdes et al., 2017) must be incorporated:

98 (a) The standard pre-industrial model uses a prescribed climatological pre-industrial ozone  
99 concentration (i.e. prior to the development of the “ozone” hole) which is a function of  
100 latitude, atmospheric height, and month of the year. However, we do not know what the  
101 distribution of ozone should be in these past climates. (Beerling et al., 2011) modelled small  
102 changes in tropospheric ozone for the early Eocene and Cretaceous but no comprehensive  
103 stratospheric estimates are available. Hence most paleoclimate model simulations assume  
104 unchanging concentrations. However, there is a problem with using a prescribed ozone  
105 distribution for paleo-simulations because it does not incorporate ozone feedbacks  
106 associated with changes in tropospheric height. During warm climates, the model predicts  
107 that the tropopause would rise. In the real world, ozone would track the tropopause rise.  
108 However, this rising ozone feedback is not included in our standard model. This leads to  
109 substantial extra warming and artificially increases the apparent climate sensitivity.  
110 Simulations of future climate change have shown that ozone feedbacks can lead to an over-  
111 estimate of climate sensitivity by up to 20% ((Dietmuller et al., 2014), (Nowack et al., 2015))  
112 (Hardiman et al., 2019). Therefore, to incorporate some aspects of this feedback, we have  
113 changed the ozone scheme in the model. Ozone is coupled to the model predicted  
114 tropopause height every model timestep in the following simple way:

- 115 •  $2.0 \times 10^{-8}$  kg/kg in the troposphere
- 116 •  $2.0 \times 10^{-7}$  kg/kg at the tropopause
- 117 •  $5.5 \times 10^{-6}$  kg/kg above the tropopause
- 118 •  $5.5 \times 10^{-6}$  kg/kg at the top model level.

119 These values are approximate averages of present-day values and were chosen so that the  
120 tropospheric climate of the resulting pre-industrial simulation was little altered compared  
121 with the standard preindustrial simulations; the resulting global mean surface air  
122 temperatures differed by only 0.05 °C. These modifications are similar to those used in the  
123 FAMOUS model (Smith et al., 2008) except that the values in the stratosphere are greater in  
124 our simulation, largely because our model vertical resolution is higher than in FAMOUS.

125 Note that these changes improve upon the scheme used by (Lunt et al., 2016) and  
126 (Farnsworth et al., 2019a). They used much lower values of stratospheric ozone and had no  
127 specified value at the top of the model. This resulted in their model having ~ 1°C cold bias

128 for pre-industrial temperatures and may have also affected their estimates of climate  
129 sensitivity.

130 (b) The standard version of HadCM3 conserves the total volume of water throughout the  
131 atmosphere and ocean (including in the numerical scheme) but several processes in the  
132 model “lose or gain” water:

- 133 1. Snow accumulates over ice sheets but there is no interactive loss through iceberg  
134 calving resulting in an excess loss of fresh water from the ocean.
- 135 2. The model caps salinity at a maximum of 45 PSU (and a minimum of 0 PSU), by  
136 artificially adding/subtracting fresh water to the ocean. This mostly affects small,  
137 enclosed seas (such as the Red Sea or enclosed Arctic) where the model does not  
138 represent the exchanges with other ocean basins.
- 139 3. Modelled river runoff includes some river basins which drain internally. These often  
140 correspond to relatively dry regions, but any internal drainage simply disappears  
141 from the model.
- 142 4. The land surface scheme includes evaporation from sub-grid scale lakes (which are  
143 prescribed as a lake fraction in each grid box, at the start of the run). The model  
144 does not represent the hydrological balance of these lakes, consequently the  
145 volume of the lakes does not change. This effectively means that there is a net  
146 source/sink of water in the model in these regions.

147 In the standard model, these water sources/sinks are approximately balanced by a flux of  
148 water into the surface ocean. This is prescribed at the start of the run and does not vary  
149 during the simulations. It is normally set to a pre-calculated estimate based on an old  
150 HadCM3-M1 simulation. The flux is strongest around Greenland and Antarctica and is  
151 chosen such that it approximately balances the water loss described in (1) i.e. the net snow  
152 accumulation over these ice sheets. There is an additional flux covering the rest of the  
153 surface ocean which approximately balances the water loss from the remaining three terms  
154 (2-4). The addition of this water flux keeps the global mean ocean salinity approximately  
155 constant on century time scales. However, depending on the simulation, the drift in average  
156 oceanic salinity can be as much as 1PSU per thousand years and thus can have a major  
157 impact on ultra-long runs of >5000 years (Farnsworth et al., 2019a).

158 For the paleo-simulations in this paper, we therefore take a slightly different approach.  
159 When ice sheets are present in the Cenozoic, we include the water flux (for the relevant

160 hemisphere) described in (1) above, based on modern values of iceberg calving fluxes for  
161 each hemisphere. However, to ensure that salinity is conserved, we also interactively  
162 calculate an additional globally uniform surface water flux based on relaxing the volume  
163 mean ocean salinity to a prescribed value on a 20-year timescale. This ensures that there is  
164 no long-term trend in ocean salinity. Tests of this update on the pre-industrial simulations  
165 revealed no appreciable impact on the skill of the model relative to the observations. We  
166 have not directly compared our simulations to the previous runs of the (Farnsworth et al.,  
167 2019a) because they use different CO<sub>2</sub> and different paleogeographies. However in practice,  
168 the increase of salinity in their simulations is well mixed and seems to have relatively little  
169 impact on the overall climate and ocean circulation.

170 We have little knowledge of whether ocean salinity has changed through time, and so keep  
171 the prescribed mean ocean salinity constant across all simulations.

172

## 173 2.2 Model Boundary Conditions

174 There are several boundary conditions that require modification through time. In this sequence of  
175 simulations, we only modify three key time-dependent boundary conditions: 1) the solar constant, 2)  
176 atmospheric CO<sub>2</sub> concentrations and, 3) paleogeographic reconstructions. We set the surface soil  
177 conditions to a uniform medium loam everywhere. All other boundary conditions (such as orbital  
178 parameters, volcanic aerosol concentrations etc.) are held constant at pre-industrial values.

179 The solar constant is based on (Gough, 1981) and increases linearly at an approximate rate of 11.1  
180 Wm<sup>-2</sup> per 100 Ma (0.8% per 100Ma), to 1365Wm<sup>-2</sup> currently. If we assume a planetary albedo of 0.3,  
181 and a climate sensitivity of 0.8 °C /Wm<sup>-2</sup> (approximately equivalent to 3°C per doubling of CO<sub>2</sub>), then  
182 this is equivalent to a temperature increase of ~.015°C per million years (~8°C over the whole of the  
183 Phanerozoic).

184 Estimates of atmospheric CO<sub>2</sub> concentrations have considerable uncertainty. We, therefore, use two  
185 alternative estimates (fig. 1a). The first uses the best fit Loess curve from (Foster et al., 2017), which  
186 is also very similar to the newer data from (Witkowski et al., 2018). The CO<sub>2</sub> levels have considerable  
187 short and long-term variability throughout the time period. Our second estimate removes much of  
188 the shorter-term variability in the Foster (2017) curve. It was developed for two reasons. Firstly, a  
189 lot of the finer temporal structure in the Loess curve is a product of differing data density of the raw  
190 data and does not necessarily correspond to real features. Secondly, the smoother curve was heavily  
191 influenced by a previous (commercially confidential) sparser sequence of simulations using non-  
192 public paleogeographic reconstructions. The resulting simulations were generally in good agreement

193 with terrestrial proxy datasets (Harris et al., 2017). Specifically, using commercial in confidence  
194 paleogeographies, we have performed multiple simulations at different CO<sub>2</sub> values for several stages  
195 across the last 440 million years and tested the resulting climate against commercial-in-confidence  
196 proxy data (Harris et al., 2017). We then selected the CO<sub>2</sub> that best matched the data. For the  
197 current simulations, we linearly interpolated these CO<sub>2</sub> values to every stage. The resulting CO<sub>2</sub>  
198 curve looks like a heavily smoothed version of the Foster curve and is within the (large) envelope of  
199 CO<sub>2</sub> reconstructions. The first-order shapes of the two curves are similar, though they are very  
200 different for some time periods (e.g. Triassic and Jurassic). In practice, both curves should be  
201 considered an approximation to the actual evolution of CO<sub>2</sub> through time which remains uncertain.

202 We refer to the simulation using the second set of CO<sub>2</sub> reconstructions as the “smooth” CO<sub>2</sub>  
203 simulations, though it should be recognised that the Foster CO<sub>2</sub> curve has also been smoothed. The  
204 Foster CO<sub>2</sub> curve extends back to only 420 Ma, so we have proposed two alternative extensions back  
205 to 540 Ma. Both curves increase sharply so that the combined forcing of CO<sub>2</sub> and solar constant are  
206 approximately constant over this time period (Foster et al., 2017). The higher CO<sub>2</sub> in the Foster curve  
207 relative to the “smooth” curve is because the initial set of simulations showed that the Cambrian  
208 simulations were relatively cool compared to data estimates for the period (Henkes et al., 2018).

### 209 2.3 Paleogeographic Reconstructions

210 The 109 paleogeographic maps used in the HadleyCM3 simulations are digital representations of the  
211 maps in the PALEOMAP Paleogeographic Atlas (Scotese, 2016); (Scotese and Wright, 2018). Table 1  
212 lists all the time intervals that comprise the PALEOMAP Paleogeographic Atlas. The Paleo Atlas  
213 contains one map for nearly every stage in the Phanerozoic. A paleogeographic map is defined as a  
214 map that shows the ancient configuration of the ocean basins and continents, as well as important  
215 topographic and bathymetric features such as mountains, lowlands, shallow sea, continental  
216 shelves, and deep oceans. Paleogeographic reconstructions older than the oldest ocean floor (~Late-  
217 Jurassic) have uniform deep ocean floor depth.

218 Once the paleogeography for each time interval has been mapped, this information is then  
219 converted into a digital representation of the paleotopography and paleobathymetry. Each digital  
220 paleogeographic model is composed of over 6 million grid cells that capture digital elevation  
221 information at a 10 km x 10 km horizontal resolution and 40-meter vertical resolution. This  
222 quantitative, paleo-digital elevation model, or “paleoDEM”, allows us to visualize and analyse the  
223 changing surface of the Earth through time using GIS software and other computer modelling  
224 techniques. For use with the HadCM3L climate model, the original high-resolution elevation grid was  
225 reduced to a ~111 km x ~111 km (1° x 1°) grid.

226 For a detailed description of how the paleogeographic maps and paleoDEMs were produced the  
227 reader is referred to (Scotese, 2016); (Scotese and Schettino, 2017); (Scotese and Wright, 2018).  
228 (Scotese and Schettino, 2017) includes an annotated bibliography of the more than 100 key sources  
229 of paleogeographic information. Similar paleogeographic paleoDEMs have been produced by  
230 (Baatsen et al., 2016) and (Verard et al., 2015).

231 The raw paleogeographic data reconstructs paleo-elevations and paleo-bathymetry at a resolution of  
232  $1^\circ \times 1^\circ$ . These data were re-gridded to  $3.75^\circ \times 2.5^\circ$  resolution that matched the GCM using a simple  
233 area (for land sea mask) or volume (for orography and bathymetry) conserving algorithm. The  
234 bathymetry was lightly smoothed (using a binomial filter) to ensure that the ocean properties in the  
235 resulting model simulations were numerically stable. The high latitudes had this filter applied  
236 multiple times. The gridding sometimes produced single grid point enclosed ocean basins,  
237 particularly along complicated coastlines, and these were manually removed. Similarly, important  
238 ocean gateways were reviewed to ensure that the re-gridded coastlines preserved these structures.  
239 The resulting global fraction of land is summarized in fig.1b and examples are shown in figure 2. The  
240 original reconstructions can be found at [https://www.earthbyte.org/paleodem-resource-scotese-](https://www.earthbyte.org/paleodem-resource-scotese-and-wright-2018/)  
241 [and-wright-2018/](https://www.earthbyte.org/paleodem-resource-scotese-and-wright-2018/). Maps of each HadCM3L paleogeography are included in the supplementary  
242 figures.

243 The paleogeographic reconstructions also include an estimate of land ice area ((Scotese and Wright,  
244 2018); fig.1c). These were converted to GCM boundary conditions assuming a simple parabolic  
245 shape to estimate the ice sheet height. These ice reconstructions suggest small amounts of land ice  
246 were present during the early Cretaceous, unlike (Lunt et al., 2016) who used ice-free Cretaceous  
247 paleogeographies.

## 248 2.4 Spin up Methodology

249 The oceans are the slowest evolving part of the modelled climate system and can take multiple  
250 millennia to reach equilibrium, depending on the initial condition and climate state. To speed up the  
251 convergence of the model, we initialized the ocean temperatures and salinity with the values from  
252 previous model simulations from similar time periods using the commercial in confidence  
253 paleogeographies. Specifically, we had a set of 17 simulations covering the last 440Ma. We selected  
254 the nearest simulation to the time period. For instance, the 10.5 Ma, 14.9 Ma, and the 19.5Ma  
255 simulations were initialised from the 13Ma simulation performed using the alternative  
256 paleogeographies. Table 2 summarises the simulations performed in this study and shows the  
257 initialisation of the model. The Foster CO<sub>2</sub> simulations were initialised from the end point of the  
258 smooth CO<sub>2</sub> simulations. In the first set of simulations (smooth CO<sub>2</sub>) we also attempted to accelerate



259 the spin up by using the ocean temperature trends at year 500 to linearly extrapolate the bottom 10  
260 level temperatures for a further 1000 years. This had limited success and was not repeated. The  
261 atmosphere variables were also initialized from the previous model simulations but the spin-up of  
262 the atmosphere is much more rapid and did not require further intervention.

263 Simulations were run in parallel so were not initialised from the previous stage results using these  
264 paleogeographies. In total, we performed almost 1 million years of model simulation and if we ran  
265 simulations in sequence, it would have taken 30 years to complete the simulations. By running these  
266 in parallel, initialised from previous modelling studies, we reduced the total run time to 3 months,  
267 albeit using a substantial amount of our high-performance computer resources.

268 Although it is always possible that a different initialization procedure may produce different final  
269 states, it is impossible to explore the possibility of hysteresis/bistability without performing many  
270 simulations for each period, which is currently beyond our computing resources. Previous studies  
271 using HadCM3L (not published) with alternative ocean initial states (isothermal at 0C, 8C, and 16C)  
272 have not revealed multiple equilibria but this might have been because we did not locate the  
273 appropriate part of parameter space that exhibits hysteresis. However, other studies have shown  
274 such behaviour (e.g. (Baatsen et al., 2018)). This remains a caveat of our current work and which we  
275 wish to investigate when we have sufficient computing resource.

276 The simulations were then run until they reached equilibrium, as defined by:

- 277 1. The globally and volume integrated annual mean ocean temperature trend is less than  
278  $1^{\circ}\text{C}/1000$  year, in most cases considerably smaller than this. We consider the volume  
279 integrated temperature because it includes all aspects of the ocean. However, it is  
280 dominated by the deep ocean trends and is near identical to the trends at a depth of 2731m  
281 (the lowest level that we have archived for the whole simulation).
- 282 2. The trends in surface air temperature are less than  $0.3^{\circ}\text{C}/1000$  year.
- 283 3. The net energy balance at the top of the atmosphere, averaged over 100-year period at the  
284 end of the simulation, is less than  $0.25 \text{ Wm}^{-2}$  (in more than 80% of the simulations, the  
285 imbalance is less than  $0.1 \text{ Wm}^{-2}$ ). The Gregory plot (Gregory et al., 2004) implies surface  
286 temperatures are within  $0.3^{\circ}\text{C}$  of the equilibrium state.

287 These target trends were chosen somewhat arbitrarily but are all less than typical orbital time scale  
288 variability (e.g. temperature changes since the last deglaciation were approximately  $5^{\circ}\text{C}$  over 10,000  
289 years). Most simulations were well within these criteria. 70% of simulations had residual net energy  
290 balances at the top of the atmosphere of less than  $0.1 \text{ Wm}^{-2}$ , but a few simulations were slower to

291 reach full equilibrium. The strength of using multiple constraints is that a simulation may, by chance,  
292 pass one or two of these criteria but were unlikely to pass all three tests. For example, all the models  
293 that we extended failed at least two of the criteria. The resulting time series of volume integrated  
294 global, annual mean ocean temperatures are shown in fig. 3. The supplementary figures also include  
295 this for each simulation, as well as the trends at 2731m.

296 The “smooth” CO<sub>2</sub> simulations were all run for 5050 model years and satisfied the criteria. The  
297 Foster-CO<sub>2</sub> simulations were initially run for a minimum of 2000 years (starting from the end of the  
298 5000-year runs), at which point we reviewed the simulations relative to the convergence criteria. If  
299 the simulations had not converged, we extended the runs for an additional 3000 years. If they had  
300 not converged at the end of 5000 years, we extended them again for an additional 3000 years. After  
301 8000 years, all simulations had converged based on the convergence criteria. In general, the slowest  
302 converging simulations corresponded to some of the warmest climates (final temperatures in figure  
303 3b and 3c were generally warmer than in figure 3a). It cannot be guaranteed that further changes  
304 will not occur; however, we note that the criteria and length of the simulations greatly exceed PMIP-  
305 LGM (Kageyama et al., 2017) and PMIP-DeepMIP (Lunt et al., 2017) protocols.

306

### 307 3. Results

#### 308 3.1 Comparison of Deep Ocean Temperatures to Benthic Ocean Data

309 Before using the model to investigate the linkage of deep ocean temperatures to global mean  
310 surface temperatures, it is interesting to evaluate whether the modelled deep ocean temperatures  
311 agree with the deep ocean temperatures obtained from the isotopic studies of benthic foraminifera  
312 (Friedrich et al., 2012; Zachos et al., 2008). It is important to note that the temperatures are strongly  
313 influenced by the choice of CO<sub>2</sub>, so we are not expecting complete agreement, but we simply wish to  
314 evaluate whether the model is within plausible ranges. If the modelled temperatures were in  
315 complete disagreement with data, then it might suggest that the model was too far away from  
316 reality to allow us to adequately discuss deep ocean/surface ocean linkages. If the modelled  
317 temperatures are plausible, then it shows that we are operating within the correct climate space. A  
318 detailed comparison of modelled surface and benthic temperatures to data throughout the  
319 Phanerozoic, using multiple CO<sub>2</sub> scenarios, is the subject of a separate ongoing project.

320 Figure 4a compares the modelled deep ocean temperature to the foraminifera data from the  
321 Cenozoic and Cretaceous (115 Ma). The observed isotope data are converted to deep ocean  
322 temperature using the procedures described by (Hansen et al., 2013). The modelled deep  
323 temperature shown in fig.4a (solid line) is the average temperature at the bottom level of the model,  
324 excluding depths less than 1000m (to avoid continental shelf locations which are typically not  
325 included in benthic data compilations). The observed benthic data are collected from a range of  
326 depths and are rarely at the very deepest levels (e.g. the new cores in (Friedrich et al., 2011) are  
327 from current water depths ranging from 1899m to 3192m). Furthermore, large data compilations  
328 rarely include how the depth of a particular site changed with time, and thus effectively assume that  
329 any differences between basins and through time are entirely due to climate change and not to  
330 changes in depth. Hence throughout the rest of the paper we frequently use the modelled 2731m  
331 temperatures as a surrogate for the true benthic temperature. This is a pragmatic definition because  
332 the area of deep ocean reduces rapidly (e.g. there is typically only 50% of the globe deeper than  
333 3300m). To evaluate whether this procedure gave a reasonable result, we also calculated the global  
334 average temperature at the model bottom, and at the model level at a depth of 2731m. The latter is  
335 shown by the dashed line in figure 4a. In general, the agreement between model bottom water  
336 temperatures and 2731m temperatures is very good. The standard deviation between model  
337 bottom water and constant depth of 2731m is 0.7°C, and the maximum difference is 1.4°C.  
338 Compared to the overall variability, this is a relatively small difference and shows that it is  
339 reasonable to assume that the deep ocean has weak vertical gradients.

340 The total change in benthic temperatures over the late Cretaceous and Cenozoic is well reproduced  
341 by the model, with the temperatures associated with the “smooth” CO<sub>2</sub> record being particularly  
342 good. We do not expect the model to represent sub-stage changes (100,000’s of years) such as the  
343 PETM excursion or OAEs, but we do expect that the broader temperature patterns should be  
344 simulated.

345 Comparison of the two simulations illustrates how strongly CO<sub>2</sub> controls global mean temperature.  
346 The Foster-CO<sub>2</sub> driven simulation substantially differs from the estimates of deep-sea temperature  
347 obtained from benthic forams and is generally a poorer fit to data. The greatest mismatch between  
348 the Foster curve and the benthic temperature curve is during the late Cretaceous and early  
349 Paleogene. Both dips in the Foster-CO<sub>2</sub> simulations correspond to relatively low estimates of CO<sub>2</sub>  
350 concentrations. For these periods, the dominant source of CO<sub>2</sub> values is from paleosols (fig.1) and  
351 thus we are reliant on one proxy methodology. Unfortunately, the alternative CO<sub>2</sub> reconstructions of  
352 (Witkowski et al., 2018) have a data gap during these periods.

353 A second big difference between the Foster curve and the benthic temperature curve occurs during  
354 the Cenomanian-Turonian. This difference is similarly driven by a low estimate of CO<sub>2</sub> in the Foster-  
355 CO<sub>2</sub> curve. These low CO<sub>2</sub> values are primarily based on stomatal density indices. As can be seen in  
356 figure 1, stomatal indices frequently suggest CO<sub>2</sub> levels lower than estimates obtained by other  
357 methods. The CO<sub>2</sub> estimates by (Witkowski et al., 2018) generally supports the higher levels of CO<sub>2</sub>  
358 (near to 1000 ppmv) that are suggested by the “smooth” CO<sub>2</sub> curve.

359 Both sets of simulations underestimate the warming during the middle Miocene. This issue has been  
360 seen before in other models e.g. (You et al., 2009), (Knorr et al., 2011), (Krapp and Jungclaus, 2011)  
361 (Goldner et al., 2014) (Steinhorsdottir, 2021). In order to simulate the surface warmth of the middle  
362 Miocene (15 Ma), CO<sub>2</sub> concentrations in the range 460–580 ppmv were required, whereas the CO<sub>2</sub>  
363 reconstructions for this period (Foster et al., 2017) are generally quite low (250-400ppmv). This  
364 problem may be either due to the climate models having too low a climate sensitivity or that the  
365 estimates of CO<sub>2</sub> are too low (Stoll et al., 2019).

366 The original compilation of (Zachos et al., 2008) represented a relatively small portion of the global  
367 ocean and the implicit assumption was made that these results represented the entire ocean basin.  
368 (Cramer et al., 2009) examined the data from an ocean basin perspective and suggested that these  
369 inter-basin differences were generally small during the Late Cretaceous and early Paleogene (90Ma –  
370 35 Ma) and the differences between ocean basins were larger during the late Paleogene and early  
371 Neogene. Our model largely also reproduces this pattern. Figure 5 shows the ocean temperature at  
372 2731 m during the late Cretaceous (69 Ma), the late Eocene (39 Ma) and the Oligocene (31 Ma) for

373 the “smooth”-CO<sub>2</sub> simulations. In the late Cretaceous, the model temperatures are almost identical  
374 in the North Atlantic and Pacific (8°C – 10°C). There is warmer deep water forming in the Indian  
375 Ocean (deep mixed layer depths, not shown), off the West coast of Australia (10°C – 12°C), but  
376 otherwise the pattern is very homogeneous. This is in agreement with some paleoreconstructions  
377 for the Cretaceous (e.g. (Murphy and Thomas, 2012)).

378 By the time we reach the late Eocene (39 Ma), the North Atlantic and Pacific remain very similar but  
379 cooler deep water (6°C – 8°C) is now originating in the South Atlantic. The South Atlantic cool  
380 bottom water source remains in the Oligocene, but we see a strong transition in the North Atlantic  
381 to an essentially modern circulation with the major source of deep, cold water occurring in the high  
382 southerly latitudes (3°C – 5°C) and strong gradient between the North Atlantic and Pacific.

383 Figure 5 also shows the modelled deep ocean temperatures for present day (Fig. 5d) compared to  
384 the World Ocean Atlas Data (fig. 5e). It can be seen that the broad patterns are well reproduced in  
385 the model, with good predictions of the mean temperature of the Pacific. The model is somewhat  
386 too warm in the Atlantic itself and has a stronger plume from the Mediterranean than is shown in  
387 the observations.

388

### 389 3.2 Comparison of Model Sea Surface Temperature to Proxy Data

390 The previous section focused on benthic temperatures, but it is also important to evaluate whether  
391 the modelled sea surface temperatures are plausible (within the uncertainties of the CO<sub>2</sub>  
392 reconstructions). Figure 4b shows a comparison between the model simulations of sea surface  
393 temperature and two published synthesis of proxy SST data. (O'Brien et al., 2017) compiled TEX<sub>86</sub>  
394 and δ<sup>18</sup>O for the Cretaceous, separated into tropical and high-latitude (polewards of 48°) regions.  
395 (Cramwinckel et al., 2018) compiled early Cenozoic tropical SST data, using Tex<sub>86</sub>, δ<sup>18</sup>O, Mg/Ca and  
396 clumped isotopes. We compare these to modelled SST for the region 15°S to 15°N, and for the  
397 average of Northern and Southern hemispheres between 47.5° and 60°. The proxy data includes  
398 sites from all ocean basins and so we also examined the spatial variability within the model. This  
399 spatial variability consists of changes along longitude (effectively different ocean basins) and  
400 changes with latitude (related to the gradient between equator and pole). We therefore calculated  
401 the average standard deviation of SST relative to the zonal mean at each latitude (this is shown by  
402 the smaller tick marks) and the total standard deviation of SST relative to the regional average. In  
403 practice, the equatorial values are dominated by inter-basin variations and hence the two measures  
404 of spatial variability are almost identical. The high latitude variability has a bigger difference  
405 between the longitudinal variations and the total variability, because the equator-to-pole

406 temperature gradient (i.e. the temperatures at the latitude limits of the region are a few degrees  
407 warmer/colder than the average). The spatial variability was very similar for the “smooth”-CO<sub>2</sub> and  
408 Foster-CO<sub>2</sub> simulations so, for clarity, on figure 4b we only show the results as error bars on the  
409 model Foster-CO<sub>2</sub> simulations.

410 Overall, the comparison between model and data is generally reasonable. The modelled equatorial  
411 temperatures largely follow the data, albeit with considerable scatter in the data. Both simulations  
412 tend to be towards the warmest equatorial data in the early Cretaceous (Albian). These  
413 temperatures largely come from Tex<sub>86</sub> data. There are many  $\delta^{18}\text{O}$  based SST which are significantly  
414 colder during this period. This data almost exclusively comes from cores 1050/1052 which are in the  
415 Gulf of Mexico. It is possible that these data are offset due to a bias in the  $\delta^{18}\text{O}$  of sea water because  
416 of the relatively enclosed region. The Foster-CO<sub>2</sub> simulations are noticeably colder than the data at  
417 the Cenomanian peak warmth, which is presumably related to the relatively low CO<sub>2</sub> as discussed for  
418 the benthic temperatures. The benthic record also showed a cool (low CO<sub>2</sub>) bias in the late  
419 Cretaceous. This is not such an obvious feature of the surface temperatures. The Foster simulations  
420 are colder than the “smooth”-CO<sub>2</sub> simulations during the late Cretaceous but there is not a strong  
421 mismatch between model and data. Both simulations are close to the observations, though the  
422 “smooth”-CO<sub>2</sub> simulations better matches the high-latitude data (but is slightly poorer with the  
423 tropical data).

424 The biggest area of disagreement between model and data is at high latitudes in the mid-Cretaceous  
425 warm period. In common with previous work with this model in the context of the Eocene (Lunt et  
426 al., 2021) the model is considerably cooler than the data, with a 10-15°C mismatch between models  
427 and data. The polar sea surface temperature estimates may have a seasonal bias because  
428 productivity is likely to be higher during the warmer summer months and, if we select the summer  
429 season temperatures from the model, then the mismatch is slightly reduced by about 4°C. The  
430 problem of a cool high latitudes in models is seen in many model studies and there is increasing  
431 evidence that this is related to the way that the models simulate clouds ((Kiehl and Shields, 2013);  
432 (Sagoo et al., 2013); (Upchurch et al., 2015; Zhu et al., 2019)). Of course, in practice deep water is  
433 formed during winter so the benthic temperatures do not suffer from a summer bias.

### 434 3.3 Correlation of Deep Ocean Temperatures to Polar Sea Surface Temperatures

435 The previous sections showed that that the climate model was producing a plausible reconstruction  
436 of past ocean temperature changes, at least within the uncertainties of the CO<sub>2</sub> estimates. We now  
437 use the HadCM3L model to investigate the links between deep ocean temperature and global mean  
438 surface temperature.

439 In theory, the deep ocean temperature should be correlated with the sea surface temperature at the  
440 location of deep-water formation which is normally assumed to be high latitude surface waters in  
441 winter. We therefore compare deep ocean temperatures (defined as the average temperature at the  
442 bottom of the model ocean, where the bottom must be deeper than 1000 m) with the average  
443 winter sea surface temperature polewards of 60° (fig. 6). Winter is defined as December, January,  
444 and February in the northern hemisphere and June, July, and August in the southern hemisphere.  
445 Also shown in Figure 6 is the best fit line, which has a slope of 0.40 (+/-0.05 at the 97.5% level), an  $r^2$   
446 of 0.59, and a standard error of 1.2°C. We obtained very similar results when we compared the polar  
447 sea surface temperatures with the average temperature at 2731m instead of the true benthic  
448 temperatures. We also compared the deep ocean temperatures to the mean polar sea surface  
449 temperatures when the mixed layer depth exceeded 250 m (poleward of 50°). The results were  
450 similar although the scatter was somewhat larger ( $r^2=0.48$ ).

451 Overall, the relationship between deep ocean temperatures and polar sea surface temperatures is  
452 clear (Figure 6) but there is considerable scatter around the best fit line, especially at the high end,  
453 and the slope is less steep than perhaps would be expected (Hansen and Sato, 2012). The scatter is  
454 less for the Cenozoic and late Cretaceous (up to 100 Ma: green and orange dots and triangles). If we  
455 used only Cenozoic and late Cretaceous simulations, then the slope is similar (0.43) but  $r^2=0.92$  and  
456 standard error=0.47°C. This provides strong confirmation that benthic data is a robust  
457 approximation to polar surface temperatures when the continental configuration is similar to the  
458 present.

459 However, the scatter is greater for older time periods, with the largest divergence observed for the  
460 warm periods of the Triassic and early Jurassic, particularly for the Foster CO<sub>2</sub> simulations (purple  
461 and blue dots). Examination of climate models for these time periods reveals relatively sluggish and  
462 shallow ocean circulation, with weak horizontal temperature gradients at depth (though salinity  
463 gradients can still be important, (Zhou et al., 2008)). For instance, in the Ladinian stage, mid-Triassic  
464 (~240Ma) the overturning circulation is extremely weak (Fig. 7). The maximum strength of the  
465 northern hemisphere overturning cell is less than 10 Sv and the southern cell is less than 5 Sv. Under  
466 these conditions, deep ocean water does not always form at polar latitudes. Examination of the  
467 mixed layer depth (not shown) shows that during these time periods, the deepest mixed layer  
468 depths are in the sub-tropics. In subtropics, there is very high evaporation relative to precipitation  
469 (due to the low precipitation and high temperatures). This produces highly saline waters that sinks  
470 and spread out into the global ocean.

471 The idea that deep water may form in the tropics is in disagreement with early hypothesis (e.g.  
472 (Emiliani, 1954)) but they were only considering the Tertiary and our model does not simulate any

473 low latitude deep water formation during this period. We only see significant tropical deep water  
474 formation for earlier periods, and this has previously been suggested as a mechanism for warm  
475 Cretaceous deep water formation (Brass et al., 1982), (Kennett and Stott, 1991). Deep water  
476 typically forms in convective plumes. (Brass et al., 1982) showed that the depth and spreading of  
477 these plumes is related to the buoyancy flux with the greatest flux leading to bottom water and  
478 plumes of lesser flux leading to intermediate water. (Brass et al., 1982) suggested that this could  
479 occur in warm conditions in the tropics, particularly if there was significant epicontinental seaways  
480 and hypothesised that it “has been a dominant mechanism of deep-water formation in historical  
481 times”. It is caused by a strong buoyancy flux linked to strong evaporation at high temperatures.

482

483 Our computer model simulations are partly consistent with this hypothesis. The key aspect for the  
484 model is a relatively enclosed seaway in the tropics and warm conditions. The paleogeographic  
485 reconstructions (see supplementary figures) suggest an enclosed Tethyan-like seaway starting in the  
486 Carboniferous and extending through to the Jurassic and early Cretaceous. However, the colder  
487 condition of the Carboniferous prevents strong tropical buoyancy fluxes. When we get into the  
488 Triassic and Jurassic, the warmer conditions lead to strong evaporation at low latitudes and bottom  
489 water formation in the tropics. This also explains why we see more tropical deep water (and hence  
490 poorer correlations between deep and polar surface temperatures in figure 6) when using the Foster  
491 CO<sub>2</sub> since this is generally higher (and hence warmer) than the smoothed CO<sub>2</sub> record.

492

493 An example of the formation of tropical deep water is shown in fig. 8. This shows a vertical cross-  
494 section of temperature and salinity near the equator for the Ladinian stage, mid-Triassic (240Ma).  
495 The salinity and temperature cross-section clearly shows high salinity warm waters sinking to the  
496 bottom of the ocean and spreading out. This is further confirmed by the water age tracer, fig. 9. This  
497 shows the water age (measured as time since it experienced surface conditions, see (England, 1995))  
498 at 2731m in the model for the Permian, Triassic, Cretaceous and present day. The present-day  
499 simulation shows that the youngest water is in the N. Atlantic and off the coast of Antarctica,  
500 indicating that this is where the deep water is forming. By contrast, the Triassic period shows that  
501 the youngest water is in the tropical Tethyan region and that it spreads out from there to fill the rest  
502 of the ocean basin. There is no young water at high latitudes, confirming that the source of bottom  
503 water is tropical only. For the Permian, although there continues to be a Tethyan-like tropical  
504 seaway, the colder conditions mean that deep water is again forming at high latitudes only. The  
505 Cretaceous is more complicated. It shows younger water in the high latitudes, but also shows some



506 young water in the Tethys which merges with the high latitude waters. Additional indicators of the  
507 transitional nature of the Cretaceous are the mixed layer depth (see supplementary figures). This is  
508 a measure of where water is mixing to deeper levels. For this time period, there are regions of deep  
509 mixed layer in both the tropics and high latitudes, whereas it is only deep in the tropics for the  
510 Triassic and at high latitudes for present day.

511

512 This mechanism for warm deep water formation has also been seen in other climate models (e.g.  
513 (Barron and Peterson, 1990)). However, (Poulsen et al., 2001) conclude that in his model of the  
514 Cretaceous high-latitudes sources of deep water diminish with elevated CO<sub>2</sub> concentrations but did  
515 not see the dominance of tropical sources. Other models (e.g. (Ladant et al., 2020)) do not show any  
516 significant tropical deep-water formation, suggesting that this feature is potentially a model-  
517 dependent result.

518 The correlation between deep ocean temperatures and the temperature of polar surface waters  
519 differs between the “smooth” CO<sub>2</sub> simulations and the Foster CO<sub>2</sub> simulations. The slope is only 0.30  
520 ( $r^2=0.57$ ) for the “smooth” CO<sub>2</sub> simulations whereas the slope is 0.48 ( $r^2=0.65$ ) for the Foster  
521 simulations. This is because CO<sub>2</sub> is a strong forcing agent that influences both the surface and deep  
522 ocean temperatures. By contrast, if the CO<sub>2</sub> does not vary as much, then the temperature does not  
523 vary as much, and the influence of paleogeography becomes more important. These  
524 paleogeographic changes generally cause subtle and complicated changes in ocean circulation that  
525 affect the location and latitude of deep-water formation.

526 In contrast, the mid-Cretaceous is also very warm but the continental configuration (specifically, land  
527 at high southern latitudes) favours the formation of cool, high latitude deep water. Throughout the  
528 Cretaceous there is significant southern high latitude source of deep water and hence deep-water  
529 temperatures are well correlated with surface high latitude temperatures. The strength of this  
530 connection, however, may be over exaggerated in the model. Like many climate models, HadCM3  
531 underestimates the reduction in the pole-to-Equator sea surface temperature (Lunt et al., 2012),  
532 (Lunt et al., 2021). This means that during the Cretaceous the high latitudes are probably too cold.  
533 Consequently, some seasonal sea ice does form which encourages the formation of cold deep-water,  
534 via brine rejection.

535 In the late Eocene (~40 Ma), the ocean circulation is similar to the Cretaceous, but the strong  
536 southern overturning cell is closer to the South Pole, indicating that the main source of deep water  
537 has moved further polewards. The poleward movement of the region of downwelling waters

538 explains some of the variability between deep ocean temperatures and temperature of polar surface  
539 waters.

540 For reference, we also include the present-day meridional circulation. The modern southern  
541 hemisphere circulation is essentially a strengthening of late Eocene meridional circulation. The  
542 Northern hemisphere is dominated by the Atlantic meridional overturning circulation. The Atlantic  
543 circulation pattern does not resemble the modern pattern of circulation until the Miocene.

#### 544 3.4 Surface Polar Amplification 545

546 The conceptual model used to connect benthic ocean temperatures to global mean surface  
547 temperatures assumes that there is a constant relationship between high latitude sea surface  
548 temperatures and global mean annual mean surface air temperature. (Hansen and Sato, 2012)  
549 argue that this amplification is partly related to ice-albedo feedback but also includes a factor  
550 related to the contrasting amplification of temperatures on land compared to the ocean. To  
551 investigate the stability of this relationship, fig. 10 shows the correlation between polar winter sea  
552 surface temperatures ( $60^{\circ}$  -  $90^{\circ}$ ) and global mean surface air temperature. The polar temperatures  
553 are the average of the two winter hemispheres (i.e. average of DJF polar SSTs in the Northern  
554 hemisphere and JJA polar SSTs in the Southern hemisphere). Also shown is a simple linear  
555 regression, with an average slope of 1.3 and with an  $r^2 = 0.79$ . If we only use Northern polar winter  
556 temperatures, the slope is 1.1; if we only use Southern polar winter temperatures, then the slope is  
557 0.7. Taken separately, the scatter about the mean is considerably larger ( $r^2$  of 0.5 and 0.6  
558 respectively) than the scatter if both data sets are combined ( $r^2 = 0.79$ ). The difference between the  
559 southern and northern hemisphere response complicates the interpretation of the proxies and leads  
560 to potentially substantial uncertainties.

561 As expected, there appears to be a strong non-linear component to the correlation. There are two  
562 separate regimes: 1) one with a steeper slope during colder periods (average polar winter  
563 temperature less than about  $1^{\circ}\text{C}$ ), and 2) a shallower slope for warmer conditions. This is strongly  
564 linked to the extent of sea-ice cover. Cooler periods promote the growth of sea-ice which  
565 strengthens the ice-albedo feedback mechanism resulting in a steeper overall temperature gradient  
566 (strong polar amplification). Of course, the ocean sea surface temperatures are constrained to be -  
567  $2^{\circ}\text{C}$  but an expansion of sea ice moves this further equatorward. Conversely, the warmer conditions  
568 result in less sea ice and hence a weaker sea ice-albedo feedback resulting in a weaker temperature  
569 gradient (reduced polar amplification). This suggests that using a simple linear relationship (as in  
570 (Hansen et al., 2008)) could be improved upon.

571 Examining the Foster CO<sub>2</sub> and “smooth” CO<sub>2</sub> simulations reveals an additional factor. If we examine  
572 the “smooth” CO<sub>2</sub> simulations only, then the best fit linear slope is slightly less than the average  
573 slope (1.1 vs 1.3). This can be explained by the fact that we have fewer very cold climates  
574 (particularly in the Carboniferous) due to the relatively elevated levels of CO<sub>2</sub>. However, the scatter  
575 in the “smooth” CO<sub>2</sub> correlation is much larger, with an  $r^2$  of only 0.66. By comparison, correlation  
576 between global mean surface temperature and polar sea surface temperature using the Foster CO<sub>2</sub>  
577 has a similar overall slope to the combined set and a smaller amount of scatter. This suggests that  
578 CO<sub>2</sub> forcing and polar amplification response have an important impact on the relationship between  
579 global and polar temperatures. The variations of carbon dioxide in the Foster set of simulations are  
580 large and they drive large changes in global mean temperature. Conversely significant sea-ice albedo  
581 feedbacks characterize times when the polar amplification is important. There are several well  
582 studied processes that lead to such changes, including albedo effects from changing ice but also  
583 from poleward heat transport changes, cloud cover, and latent heat effects ((Alexeev et al., 2005;  
584 Holland and Bitz, 2003; Sutton et al., 2007)). By contrast, the “smooth” CO<sub>2</sub> simulations have  
585 considerably less forcing due to CO<sub>2</sub> variability which leads to a larger paleogeographic effect. For  
586 instance, when there is more land at the poles, there will be more evaporation over the land areas  
587 and hence simple surface energy balance arguments would suggest different temperatures ((Sutton  
588 et al., 2007)) .

589 In figure 10, there are a few data points which are complete outliers. These correspond to  
590 simulations in the Ordovician; the outliers happen irrespective of the CO<sub>2</sub> model that is used.  
591 Inspection of these simulations shows that the cause for this discrepancy is related to two factors:  
592 1) a continental configuration with almost no land in the Northern hemisphere and, 2) a  
593 reconstruction which includes significant southern hemisphere ice cover (see fig.1 and fig 2).  
594 Combined, these factors produced a temperature structure which is highly non-symmetric, with the  
595 Southern high latitudes being more than 20°C colder than the Northern high latitudes. This anomaly  
596 biases the average polar temperatures shown in figure 10.

### 597 3.5 Deep Ocean Temperature versus Global Mean Temperature 598

599 The relationships described above help to understand the overall relationship between deep ocean  
600 temperatures and global mean temperature. Figure 11 shows the correlation between modelled  
601 deep ocean temperatures (> 1000 m) and global mean surface air temperature, and figure 12 shows  
602 a comparison of changes in modelled deep ocean temperature compared to model global mean  
603 temperature throughout the Phanerozoic.

604 The overall slope is 0.64 (0.59 to 0.69) with an  $r^2 = 0.74$ . If we consider the last 115 Ma (for which  
605 exists compiled benthic temperatures), then the slope is slightly steeper (0.67 with an  $r^2 = 0.90$ ).  
606 Similarly, the “smooth”-CO<sub>2</sub> and the Foster-CO<sub>2</sub> simulation results have very different slopes. The  
607 “smooth”-CO<sub>2</sub> simulations have a slope of 0.47, whereas the Foster-CO<sub>2</sub> simulations have a slope of  
608 0.76. The root mean square departure from the regression line in figure 11 is 1.3°C. Although we  
609 could have used a non-linear fit as we might expect such a relationship if the pole-to-equator  
610 temperature gradient changes, all use of benthic temperatures as a global mean surface  
611 temperature proxy are based on linear relationship.

612

613 The relatively good correlations in the fig.11 are confirmed when examining fig.12a and 12b. On  
614 average, the deep ocean temperatures tend to underestimate the global mean change (fig.12b)  
615 which is consistent with the regression slope being less than 1. However, the errors are substantial  
616 with largest errors occurring during the pre-Cretaceous and can be 4-6 °C. This is an appreciable  
617 error that would have a substantial impact on estimates of climate sensitivity. Even within the late  
618 Cretaceous and Cenozoic, the errors can exceed 2°C which can exceed 40% of the total change.

619 The characteristics of the plots can best be understood in terms of figures (6 and 10). For instance,  
620 most of the Carboniferous simulations plot below the regression line because the polar SSTs are not  
621 well-correlated with the global mean temperature (figure 10). By contrast, the Triassic and Jurassic  
622 Foster CO<sub>2</sub> simulations plot above the regression line because the deep ocean temperature is not  
623 well-correlated with the polar temperatures (figure 6).

#### 624 4. Discussion and Conclusion

625 The paper has presented the results from two unique sets of paleoclimate simulations covering the  
626 Phanerozoic. The focus of the paper has been to use the HadCM3L climate model to evaluate how  
627 well we can predict global mean surface temperatures from benthic foram data. This is an important  
628 consideration because benthic microfossil data are one of the few datasets used to directly estimate  
629 past global mean temperatures. Other methods, such as using planktonic foraminiferal estimates,  
630 are more challenging because the sample sites are geographically sparse, so it is difficult to  
631 accurately estimate the global mean temperature from highly variable and widely dispersed data.  
632 This is particularly an issue for older time periods when fewer isotopic measurements from  
633 planktonic microfossils are available and can result in a bias because most of the isotopic  
634 temperature sample localities are from tropical latitudes (30°S – 30°N) (Song et al., 2019).

635 By contrast, deep ocean temperatures are more spatially uniform. Hence. benthic foram data has  
636 frequently been used to estimate past global mean temperatures and climate sensitivity (Hansen et

637 al., 2013). Estimates of uncertainty for deep ocean temperatures incorporate uncertainties from CO<sub>2</sub>  
638 and from the conversion of  $\delta^{18}\text{O}$  measurements to temperature but have not been able to assess  
639 assumptions about the source regions for deep ocean waters and the importance polar  
640 amplification. Of course, in practice, lack of ocean sea floor means that benthic compilations exist  
641 only for the last 110Ma.

642 Changes in heat transport also play a potentially important role in polar amplification. In the  
643 supplementary figure, we show the change in atmosphere and ocean poleward heat fluxes for each  
644 time period. Examination of the modelled poleward heat transport by the atmosphere and ocean  
645 shows a very complicated pattern, with all time periods showing the presence of some Bjerknes  
646 compensation (Bjerknes, 1964) (see (Outten et al., 2018) for example in CMIP5 models). Bjerknes  
647 compensation is where the change in ocean transport is largely balanced by an equal but opposite  
648 change in atmospheric transport. For instance, compared to present day, the mid-Cretaceous and  
649 Early Eocene warm simulations shows a large increase in northward atmospheric heat transport,  
650 linked with enhanced latent heat transport associated with the warmer, moister atmosphere.  
651 However, this is partly cancelled by an equal but opposite change in the ocean transport. E.g.  
652 compared to present day, the early Eocene northern hemisphere atmospheric heat transport  
653 increases by up to 0.5PW, but the ocean transport is reduced by an equal amount. The net  
654 transport from equator to the N. Pole changes by less than 0.1PW (i.e. less than 2% of total). Further  
655 back in time, the compensation is still apparent, but the changes are more complicated, especially  
656 when the continents are largely in the Southern hemisphere. Understanding the causes of these  
657 transport changes will be the subject of another paper.

658 We have shown that although the expected correlation between benthic temperatures and high-  
659 latitude surface temperatures exists, the correlation has considerable scatter. This is caused by  
660 several factors. Changing paleogeographies results in changing locations for deep water formation.  
661 Some paleogeographies result in significant deep-water formation in the Northern hemisphere (e.g.  
662 our present-day configuration) although for most of the Phanerozoic, the dominant source of deep-  
663 water formation has been southern hemisphere. Similarly, even when deep water is formed in just  
664 one hemisphere, there can be substantial regional and latitudinal variations in its location and the  
665 corresponding temperatures. Finally, during times of very warm climates (e.g. mid-Cretaceous) the  
666 overturning circulation can be very weak and there is a marked decoupling between the surface  
667 waters and deep ocean. In the HadCM3 model during hothouse time periods, high temperatures and  
668 high rates of evaporation produce hot and saline surface waters which sink to become intermediate  
669 and deep waters at low latitudes.

670 Similar arguments can be made regarding the link between global mean temperature and the  
671 temperature at high latitudes. Particularly important is the area of land at the poles and the extent  
672 of sea ice/land ice. Colder climates and paleogeographic configurations with more land at the pole  
673 will result in a steeper latitudinal temperature gradient and hence exhibit a changing relationship  
674 between polar and global temperatures. But the fraction of land versus ocean is also important.

675 Finally, the overall relationship between deep ocean temperatures and global mean temperature is  
676 shown to be relatively linear, but the slope is quite variable. In the model simulations using the  
677 “smooth” CO<sub>2</sub> curve, the slope is substantially shallower (0.48) than slope obtained using the Foster  
678 CO<sub>2</sub> curve (0.76). This is related to the different controls that CO<sub>2</sub> and paleogeography exert (as  
679 discussed above). In the simulation that uses the “smooth” CO<sub>2</sub> data set, the levels of CO<sub>2</sub> do not  
680 vary much, so the paleogeographic controls are more pronounced.

681 This raises the interesting conundrum that when trying to use reconstructed deep ocean  
682 temperatures and CO<sub>2</sub> to estimate climate sensitivity, the interpreted global mean temperature also  
683 depends, in part, on the CO<sub>2</sub> concentrations. However, if we simply use the combined slope, then  
684 the root mean square error is approximately 1.4°C, and the maximum error is over 4°C. The root  
685 mean square error is a relatively small compared to the overall changes and hence the resulting  
686 uncertainty in climate sensitivity associated with this error is relatively small (~15%) and the CO<sub>2</sub>  
687 uncertainty dominates. However, the maximum error is potentially more significant.

688 Our work has not addressed other sources of uncertainty. In particular, it would be valuable to use a  
689 water isotope-enabled climate model to better address the uncertainties associated with the  
690 conversion of the observed benthic δ<sup>18</sup>O to temperature. This requires assumptions about the δ<sup>18</sup>O  
691 of sea water. We hope to perform such simulation in future work, though this is a particularly  
692 challenging computational problem because the isotope enabled model is significantly slower and  
693 the completion of the multi-millennial simulations required for deep ocean estimates would take  
694 more than 18 months to complete.

695 Our simulations extend and develop those published by (Lunt et al., 2016), and (Farnsworth et al.,  
696 2019a; Farnsworth et al., 2019b). The simulations reported in this paper used the same climate  
697 model (HadCM3L) but used an improved ozone concentration and corrected a salinity drift that can  
698 lead to substantial changes over the duration of the simulation. Our simulations also use an  
699 alternative set of geographic reconstructions that cover a larger time period (540 Ma – Modern).  
700 They also include realistic land ice cover estimates, which were not included in the original  
701 simulations (except for the late Cenozoic) but generally have a small impact in the Mesozoic.

702 Similarly, the new simulations use two alternative models for past atmospheric CO<sub>2</sub> use more  
703 realistic variations in CO<sub>2</sub> through time (compared with idealised constant values in Farnsworth et al  
704 and Lunt et al), while at the same time recognizing the levels of uncertainty. Although the Foster CO<sub>2</sub>  
705 curve is more directly constrained by CO<sub>2</sub> data, it should be noted that this data come from multiple  
706 proxies and there are large gaps in the data set. There is evidence that the different proxies have  
707 different biases, and it is not obvious that the correct approach is to simply fit a Loess-type curve to  
708 the CO<sub>2</sub> data. This is exemplified by the Maastrichtian. The Foster Loess curve shows a minimum in  
709 CO<sub>2</sub> during the Maastrichtian which results in the modelled deep ocean temperatures being much  
710 too cold. However, detailed examination of the CO<sub>2</sub> data shows most of the Maastrichtian data is  
711 based on stomatal index reconstructions which often are lower than other proxies. Thus, the  
712 Maastrichtian low CO<sub>2</sub>, relative to other periods, is potentially driven by changing the proxy rather  
713 than by real temporal changes.

714 Though the alternative, “smooth” CO<sub>2</sub> curve is not the optimum fit to the data, it does pass through  
715 the cloud of individual CO<sub>2</sub> reconstructions and hence represents one possible “reality”. For the Late  
716 Cretaceous and Cenozoic, the “smooth” CO<sub>2</sub> simulation set does a significantly better job simulating  
717 the deep ocean temperatures of the Friedrich/Cramer/Zachos curve.

718 Although the focus of the paper has been the evaluation of the modelled relationship between  
719 benthic and surface temperatures, the simulations are a potentially valuable resource for future  
720 studies. This includes using the simulations for paleoclimate/climate dynamic studies and for climate  
721 impact studies, such as ecological niche modelling. We have therefore made available on our  
722 website the results from our simulations

723 ([https://www.paleo.bristol.ac.uk/ummodel/scripts/papers/Valdes\\_et\\_al\\_2021.html](https://www.paleo.bristol.ac.uk/ummodel/scripts/papers/Valdes_et_al_2021.html))

724 Data Availability

725 All simulation data is available from:

726 [https://www.paleo.bristol.ac.uk/ummodel/scripts/papers/Valdes\\_et\\_al\\_2021.html](https://www.paleo.bristol.ac.uk/ummodel/scripts/papers/Valdes_et_al_2021.html)

727 Author contributions

728 Study was developed by all authors. All model simulations were performed by PJV who also  
729 prepared the manuscript with contributions from all co-authors.

730 Competing interests

731 The authors declare that they have no conflict of interest.

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733

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738 Institution's National Museum of Natural History and was initiated at a workshop supported by  
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740 Advanced Computing Research Centre, University of Bristol (<http://www.bris.ac.uk/acrc/>). The  
741 authors declare that they have no competing interests. Data and materials availability: All data  
742 needed to evaluate the conclusions in the paper are present in the paper. Model data can be  
743 accessed at [www.bridge.bris.ac.uk/resources/simulations](http://www.bridge.bris.ac.uk/resources/simulations).

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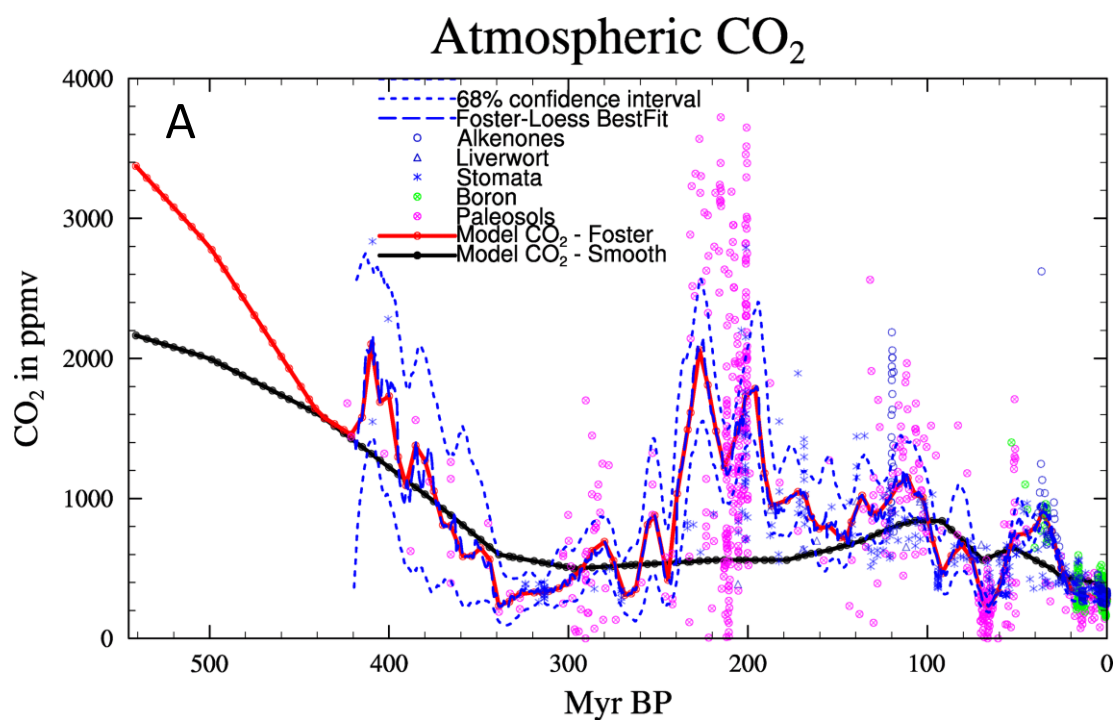
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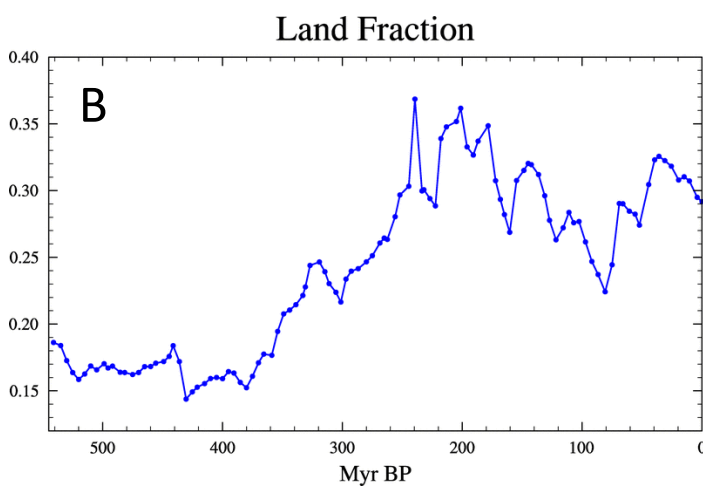


750 Figures

751 **Figure 1.** Summary of boundary condition changes to model of the Phanerozoic, (a) CO<sub>2</sub>  
752 reconstructions (from Foster et al. 2017) and the two scenarios used in the models, (b) Land-sea  
753 fraction from the paleogeographic reconstructions, and (c) land ice area input into model. The  
754 paleogeographic reconstructions can be accessed at [https://www.earthbyte.org/paleodem-](https://www.earthbyte.org/paleodem-resource-scotese-and-wright-2018/)  
755 [resource-scotese-and-wright-2018/](https://www.earthbyte.org/paleodem-resource-scotese-and-wright-2018/). An animation of the high-resolution (1° x 1°) and model  
756 resolution (3.75° longitude x 2.5° latitude) maps can be found here:  
757 [https://www.paleo.bristol.ac.uk/~ggpjv/scotese/scotese\\_raw\\_moll.normal\\_scotese\\_moll.normal.ht](https://www.paleo.bristol.ac.uk/~ggpjv/scotese/scotese_raw_moll.normal_scotese_moll.normal.ht)  
758 [ml](https://www.paleo.bristol.ac.uk/~ggpjv/scotese/scotese_raw_moll.normal_scotese_moll.normal.ht)

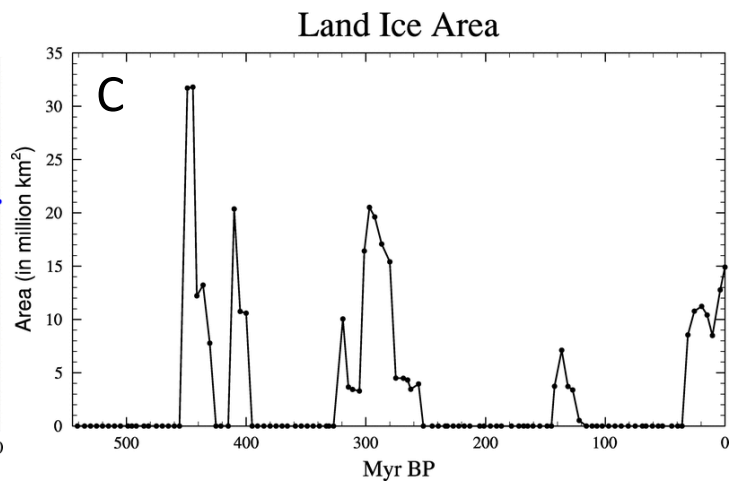


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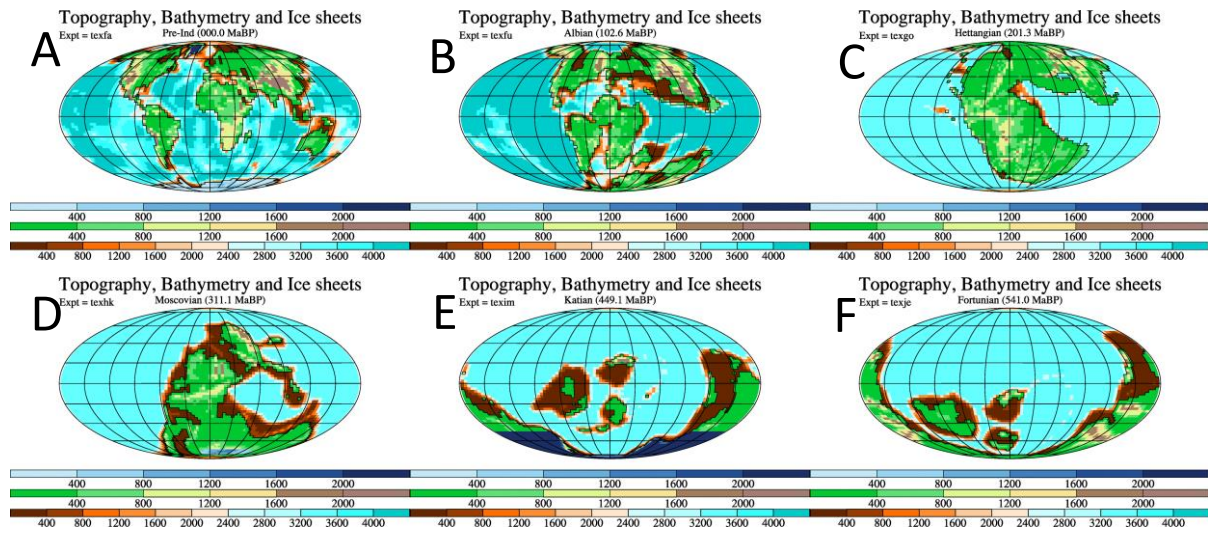
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762 **Figure 2.** A few example paleogeographies, once they have been re-gridded onto the HadCM3L grid.  
 763 The examples are for (a) present day, (b) Albian, 102.6Ma (Lower Cretaceous), (c) Hettangian,  
 764 201.3Ma (lower Jurassic), (d) Moscovian, 311.1Ma (Pennsylvanian, Carboniferous), (e) Katian,  
 765 449.1Ma (Upper Ordovician), and (f) Fortunian, 541.0Ma (Cambrian). The top color legend refers to  
 766 the height of the ice sheets (if they exist), the middle color legend refers to heights on land (except  
 767 ice), and the lower color legend refers to the ocean bathymetry. All units are meters.

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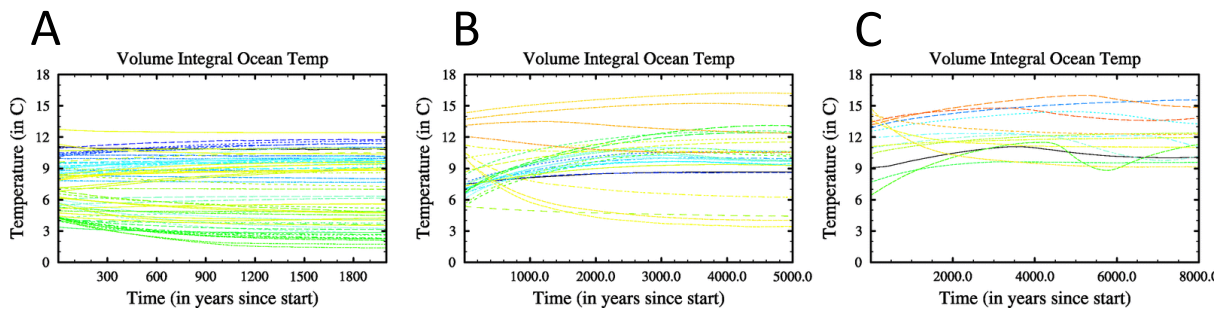
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780 **Figure 3.** Time series of the annual, volume mean ocean temperature for all 109 simulations. (a)  
 781 shows those simulations for which 2000 years was sufficient to satisfy the convergence criteria  
 782 described in text (these were for all simulations listed in table 1 except those listed in (b) and (c)), (b)  
 783 those simulation which required 5000 years (these were for all the simulations for 31.0, 35.9, 39.5,  
 784 55.8, 60.6, 66.0, 69.0, 102.6, 107.0, 121.8, 127.2, 154.7, 160.4, 168.2, 172.2, 178.4, 186.8, 190.8,  
 785 196.0, 201.3, 204.9, 213.2, 217.8, 222.4, 227.0, 232.0, and 233.6 Ma BP), and (c) those simulation  
 786 which required 8000 years (these were simulations for 44.5, 52.2, 86.7, 91.9, 97.2, 111.0, 115.8,  
 787 131.2, 136.4, 142.4, 145.0, 148.6, 164.8, and 239.5 Ma BP). The different coloured lines show the  
 788 different runs. The plot simply show the extent to which all runs have reached steady state. For  
 789 more details about specific simulations, please see the supplementary figures.

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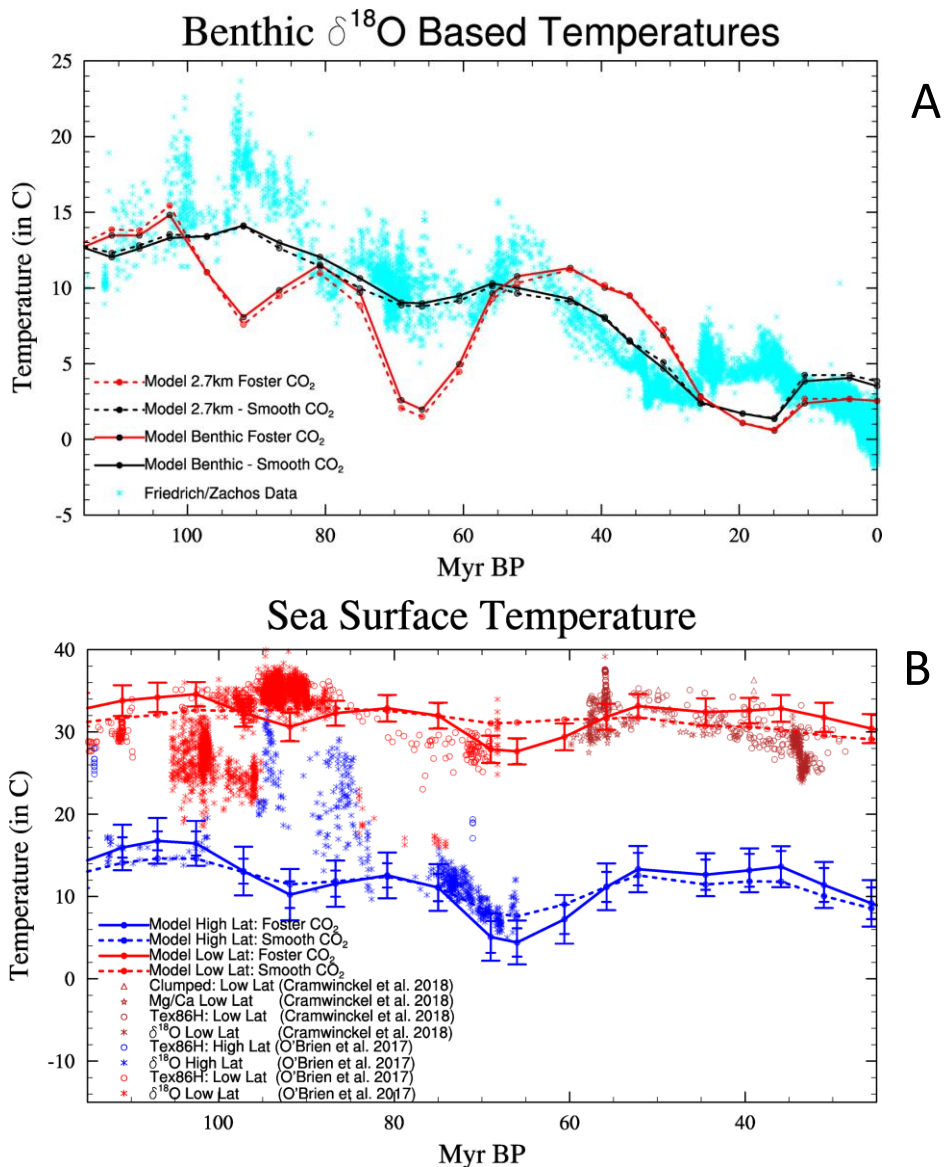
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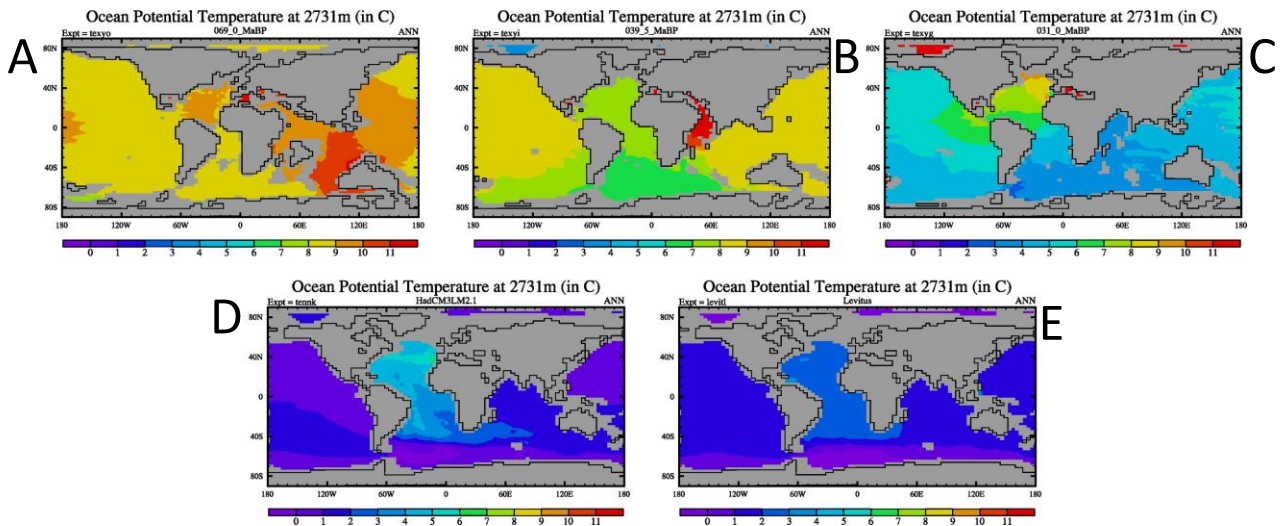
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794 **Figure 4.** (a) Comparison of modelled deep ocean temperatures versus those from (Zachos et al.,  
 795 2008) and (Friedrich et al., 2012) converted to temperature using the formulation in (Hansen et al.,  
 796 2013). The model temperatures are global averages over the bottom layer of the model but excludes  
 797 shallow marine settings (less than 1000m). The dashed lines show the modelled global average  
 798 ocean temperatures at the model layer centered at 2731m, and (b) Comparison of modelled sea  
 799 surface temperatures with the compilations of (O'Brien et al., 2017) and (Cramwinckel et al., 2018).  
 800 The data is a combination of Tex<sub>86</sub>, δ<sup>18</sup>O, Mg/Ca, and clumped Isotope data. The model data shows  
 801 low latitude temperatures (averaged from 10S to 10N) and high latitude temperatures (averaged  
 802 over 47.5N to 65N and 47.5S to 65S). The Foster-CO<sub>2</sub> simulations also show a measure of the spatial  
 803 variability. The large bars show the spatial standard deviation across the whole region, and the  
 804 smaller bars shows the average spatial standard deviation along longitudes within the region. Note  
 805 that the ranges of both the x and y-axis differ between (a) and (b).

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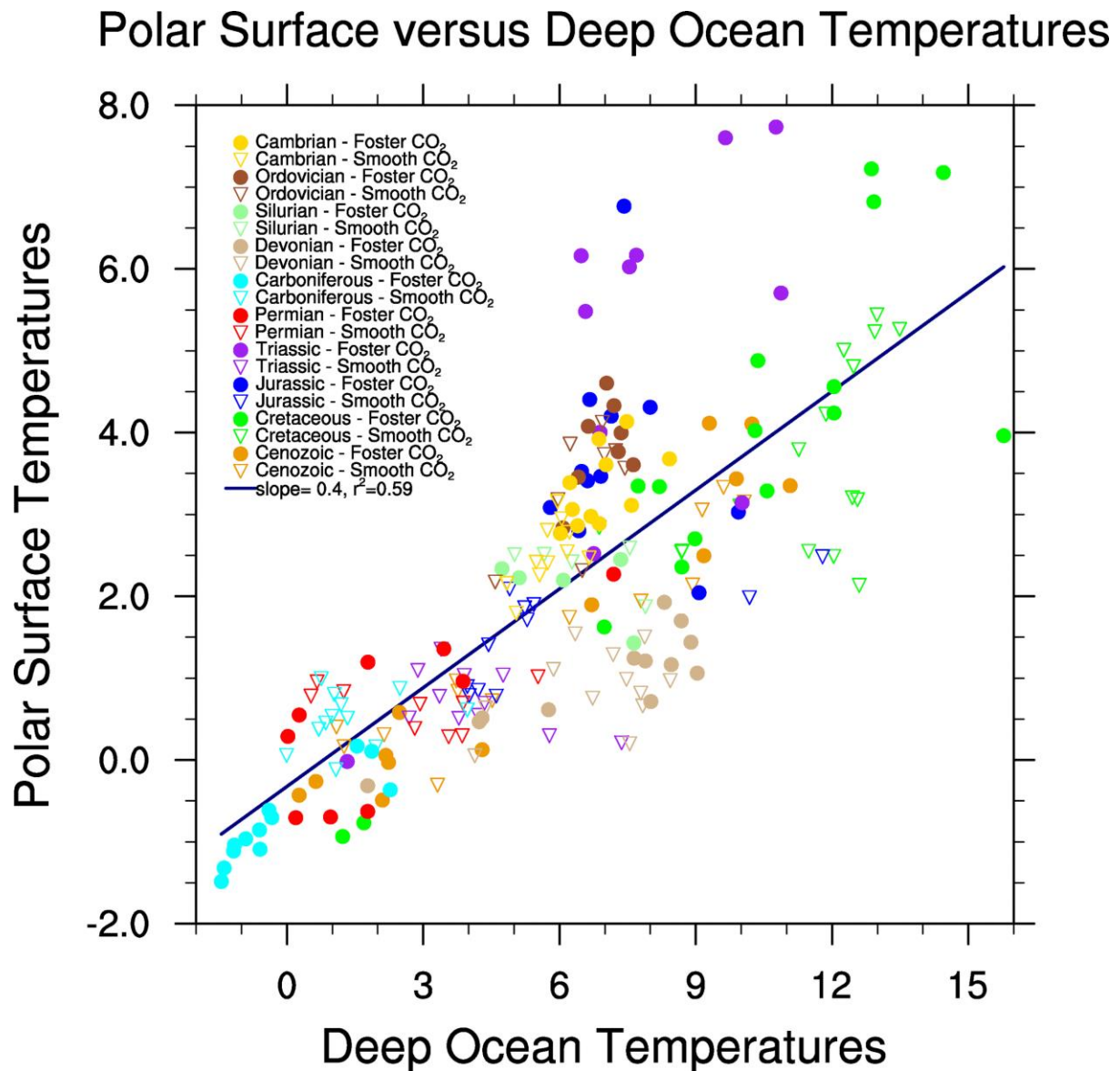
807 **Figure 5.** Modelled annual mean ocean temperatures are 2731m depth for three example past time  
 808 periods. The left figure is for the late Cretaceous, the center for the late Eocene (39.5Ma), and the  
 809 right for the Oligocene (31Ma). These are results from the smooth-CO<sub>2</sub> set of simulations which  
 810 agree better with the observed benthic temperature data. Also included are the pre-industrial  
 811 simulation and World Ocean Atlas 1994 observational data, provided by the NOAA-ESRL Physical  
 812 Sciences Laboratory, Boulder Colorado from their web site at <https://psl.noaa.gov/>. The thin black  
 813 lines show the coastlines, and the grey areas are showing where the ocean is shallower than 2731m.



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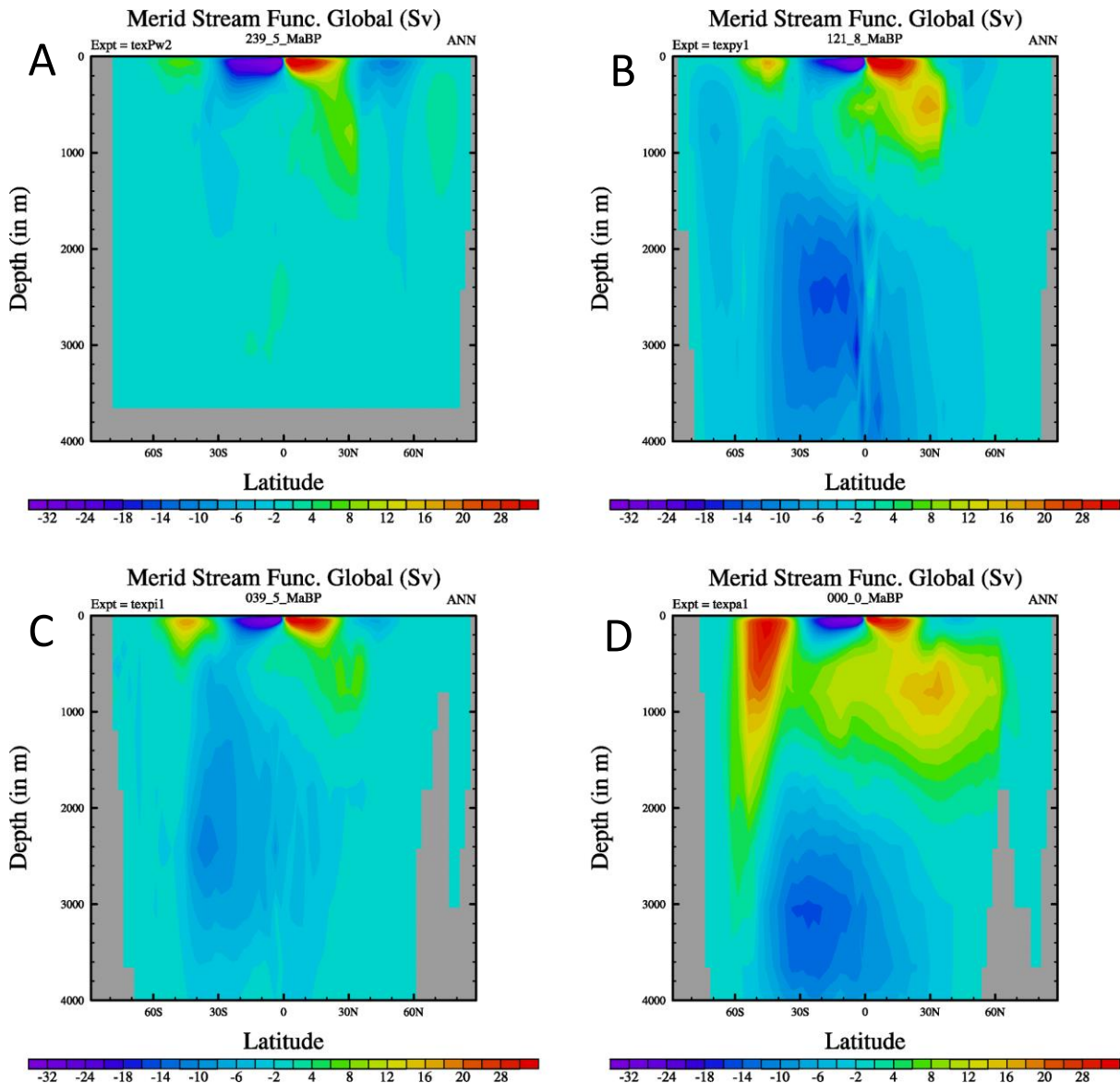
816 **Figure 6.** Correlations between deep ocean temperatures and surface polar sea surface  
 817 temperatures. The deep ocean temperatures are defined as the average temperature at the bottom  
 818 of the model ocean, where the bottom must be deeper than 1000m. The polar sea surface  
 819 temperatures are the average winter (i.e. northern polar in DJF and southern polar in JJA) sea  
 820 surface temperature polewards of 60°. The inverted triangles show the results from the smooth CO<sub>2</sub>  
 821 simulations and the dots refer to the Foster CO<sub>2</sub> simulations. The colors refer to different geological  
 822 era.



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824 **Figure 7.** Global Ocean overturning circulation (in Sverdrup) for four different time periods for the  
 825 Foster-CO<sub>2</sub> simulations. Positive (yellow/red) values correspond to a clockwise circulation, negative  
 826 (dark blue/purple) values represent an anti-clockwise circulation. (a) Middle Triassic, Ladinian,  
 827 239.5Ma, (b) Lower Cretaceous, Aptian, 121.8 Ma, (c) Late Eocene, Bartonian, 39.5Ma, and (d)  
 828 Present Day. Paleogeographic reconstructions older than the oldest ocean floor (~Late-Jurassic) have  
 829 uniform deep ocean floor depth.

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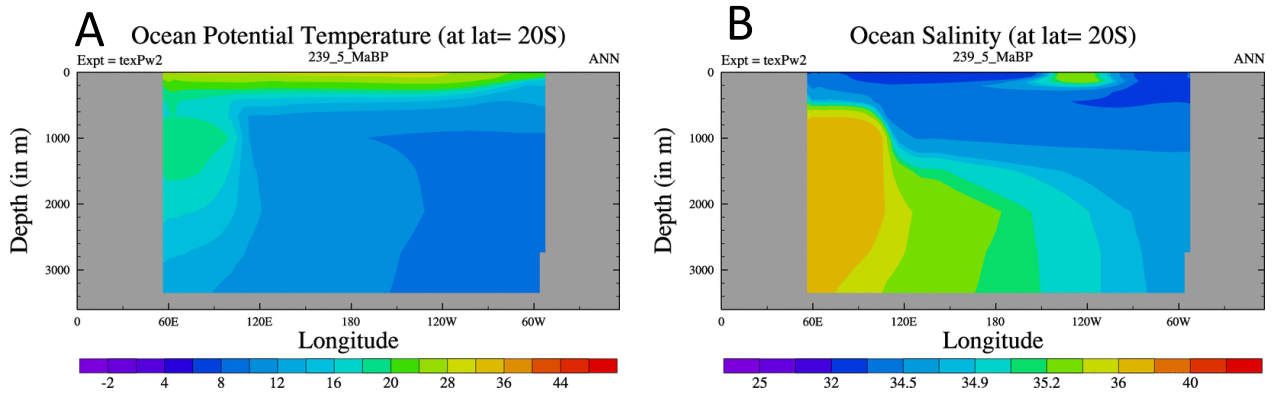


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833 **Figure 8.** Longitudinal cross section at 20S of (a) ocean potential temperature and (b) salinity for the  
834 Ladanian (240Ma). Temperature is in C and salinity is in PSU.  
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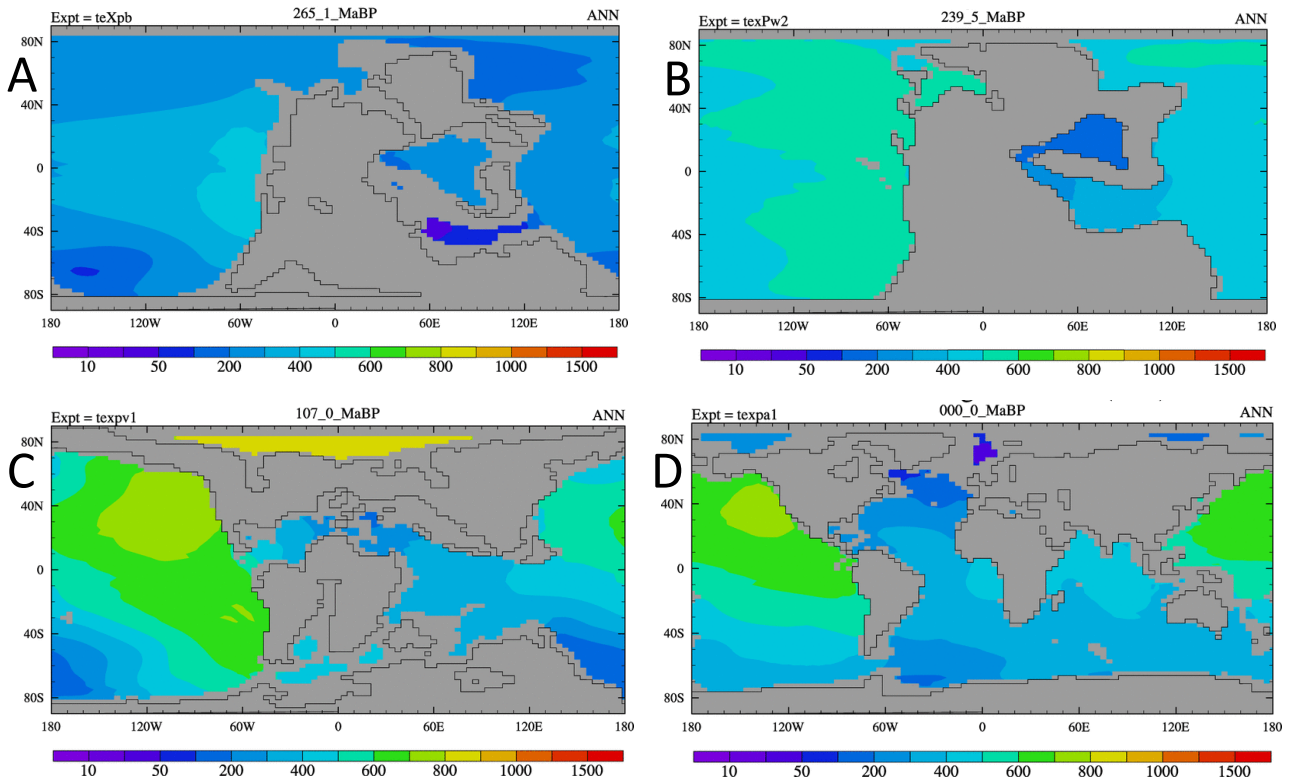


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840 **Figure 9.** Modelled age of water tracer at 2731m for 4 different time periods (a) 265Ma, (b) 240Ma,  
841 (c) 107Ma, and (d) 0Ma. Units are years.

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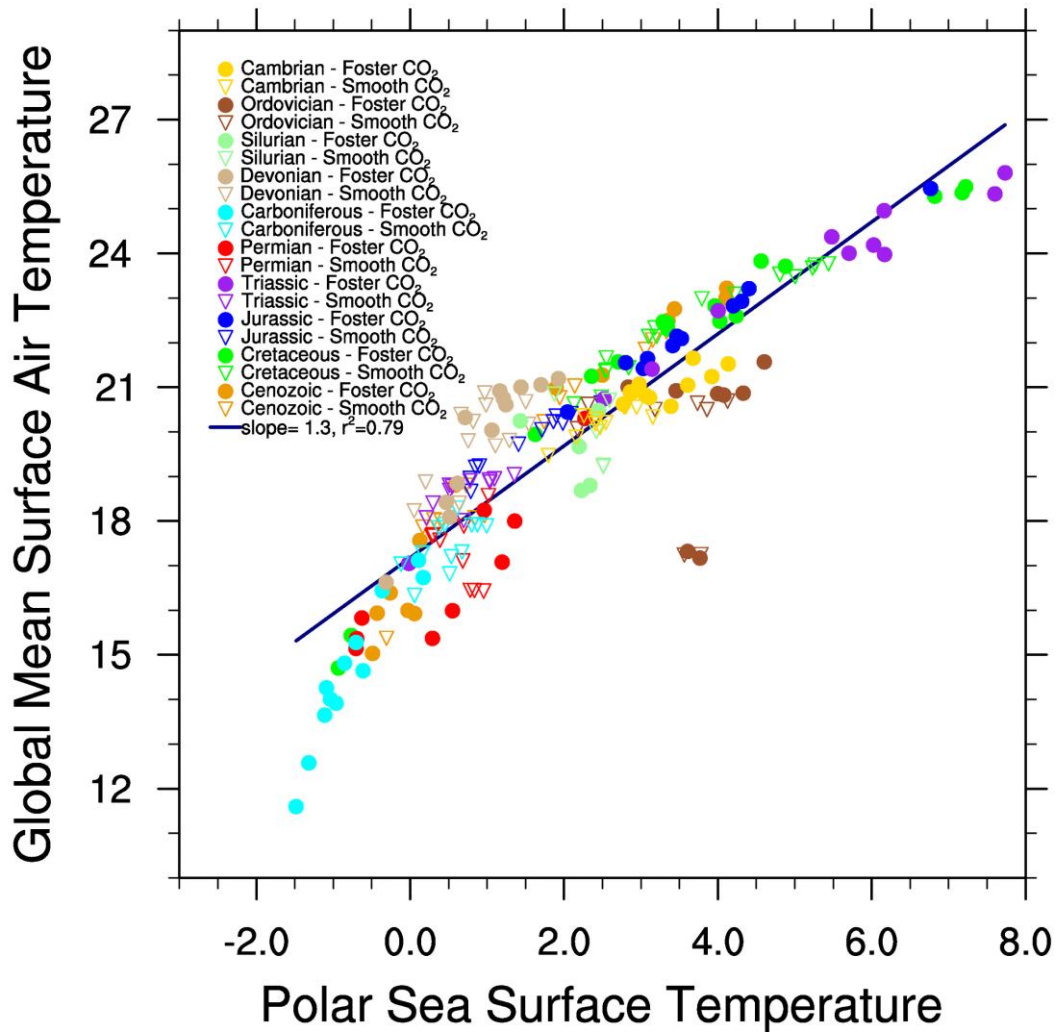
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849 **Figure 10.** Correlation between high latitude ocean temperatures (polewards of 60°) and the annual  
850 mean, global mean surface air temperature. The polar temperatures are the average of the two  
851 winter hemispheres (i.e. northern DJF and southern JJA). Other details as in figure 6.

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## Global Mean Surface Air Temperature versus Polar SST

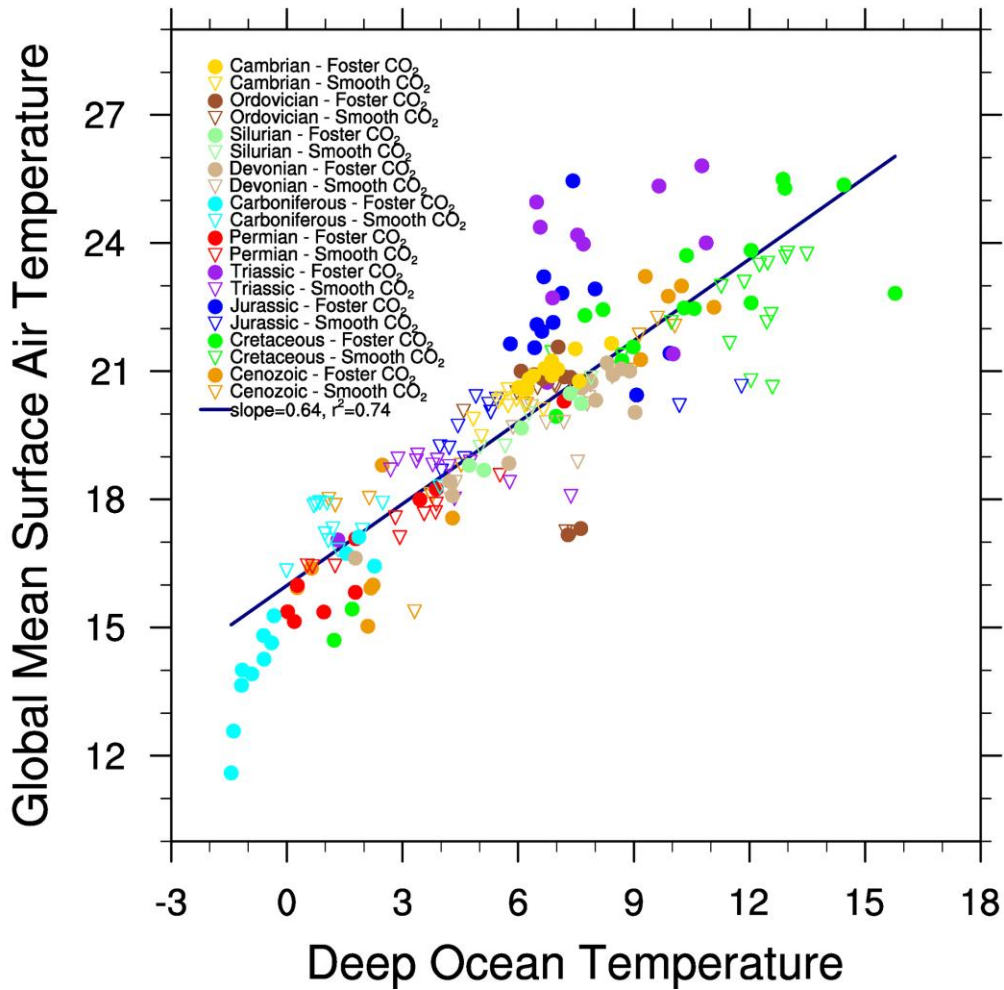


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855 **Figure 11.** Correlation between the global mean, annual mean surface air temperature and the deep  
856 ocean temperature. The deep ocean temperatures are defined as the average temperature at the  
857 bottom of the model ocean, where the bottom must be deeper than 1000m. Other details as in  
858 figure 6.

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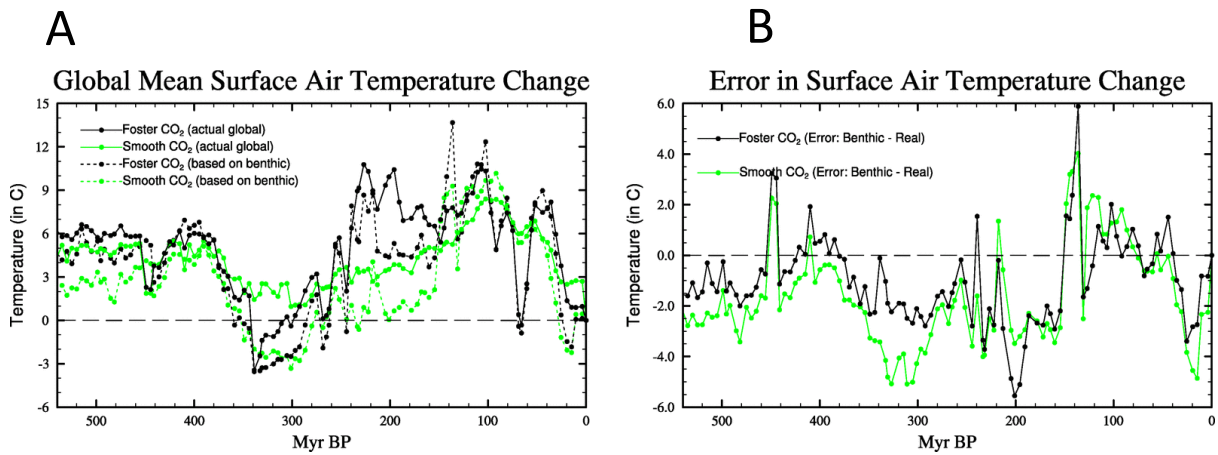
## Surface Air Temperature versus Deep Ocean Temperature



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862 **Figure 12.** Phanerozoic Time series of modelled temperature change (relative to pre-Industrial) for  
863 the smooth (green lines) and Foster-CO<sub>2</sub> (black) simulations (a) shows the actual modelled global  
864 mean surface air temperature (solid lines) whereas the dashed line shows the estimate based on  
865 deep ocean temperatures, and (b) error in the estimate of global mean temperature change if based  
866 on deep ocean temperatures (i.e. deep ocean – global mean surface temperatures).  
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Table I. List of Paleogeographic Maps and PaleoDEMs

Map Number	Stratigraphic Age Description	Plate Model Age
1	Present-day (Holocene, 0 Ma)	0
2	<i>Last Glacial Maximum (Pleistocene, 21 ky) *</i>	0
3	<i>Late Pleistocene (122 ky) *</i>	0
4	<i>Middle Pleistocene (454 ky) *</i>	0
5	<i>Early Pleistocene (Calabrian, 1.29 Ma) *</i>	0
6	<i>Early Pleistocene (Gelasian, 2.19) *</i>	0
7	Late Pliocene (Piacenzian, 3.09)	5
8	<i>Early Pliocene (Zanclean, 4.47 Ma) *</i>	5
9	<i>latest Miocene (Messinian, 6.3 Ma) *</i>	5
10	Middle/Late Miocene (Serravallian&Tortonian, 10.5 Ma)	10
11	Middle Miocene (Langhian, 14.9 Ma)	15
12	Early Miocene (Aquitanian&Burdigalian, 19.5 Ma)	20
13	Late Oligocene (Chattian, 25.6 Ma)	25
14	Early Oligocene (Rupelian, 31 Ma)	30
15	Late Eocene (Priabonian, 35.9 Ma)	35
16	late Middle Eocene (Bartonian, 39.5 Ma)	40
17	early Middle Eocene (Lutetian, 44.5 Ma)	45
18	Early Eocene (Ypresian, 51.9 Ma)	50
19	Paleocene/Eocene Boundary (PETM, 56 Ma)	55
20	Paleocene (Danian&Thanetian, 61 Ma)	60
21	KT Boundary (latest Maastrichtian, 66 Ma)	65
22	Late Cretaceous (Maastrichtian, 69 Ma)	70
23	Late Cretaceous (Late Campanian, 75 Ma)	75
24	Late Cretaceous (Early Campanian, 80.8 Ma)	80
25	Late Cretaceous (Santonian&Coniacian, 86.7 Ma)	85
26	Mid-Cretaceous (Turonian, 91.9 Ma)	90

27	Mid-Cretaceous (Cenomanian, 97.2 Ma)	95
28	Early Cretaceous (late Albian, 102.6 Ma)	100
29	Early Cretaceous (middle Albian, 107 Ma)	105
30	Early Cretaceous (early Albian, 111 Ma)	110
31	Early Cretaceous (late Aptian, 115.8 Ma)	115
32	Early Cretaceous (early Aptian, 121.8 Ma)	120
33	Early Cretaceous (Barremian, 127.2 Ma)	125
34	Early Cretaceous (Hauterivian, 131.2 Ma)	130
35	Early Cretaceous (Valanginian, 136.4 Ma)	135
36	Early Cretaceous (Berriasian, 142.4 Ma)	140
37	Jurassic/Cretaceous Boundary (145 Ma)	145
38	Late Jurassic (Tithonian, 148.6 Ma)	150
39	Late Jurassic (Kimmeridgian, 154.7 Ma)	155
40	Late Jurassic (Oxfordian, 160.4 Ma)	160
41	Middle Jurassic (Callovian, 164.8 Ma)	165
42	Middle Jurassic (Bajocian&Bathonian, 168.2)	170
43	Middle Jurassic (Aalenian, 172.2 Ma)	175
44	Early Jurassic (Toarcian, 178.4 Ma)	180
45	Early Jurassic (Pliensbachian, 186.8 Ma)	185
46	Early Jurassic (Sinemurian/Pliensbachian, 190.8 Ma)	190
47	Early Jurassic (Hettangian&Sinemurian, 196 Ma)	195
48	Late Triassic (Rhaetian/Hettangian, 201.3 Ma)	200
49	Late Triassic (Rhaetian, 204.9 Ma)	205
50	Late Triassic (late Norian, 213.2 Ma)	210
51	Late Triassic (mid Norian, 217.8 Ma)	215
52	Late Triassic (early Norian, 222.4 Ma)	220
53	Late Triassic (Carnian/Norian 227 Ma)	225
54	Late Triassic (Carnian, 232 Ma)	230

55	Late Triassic (early Carnian, 233.6)	235
56	Middle Triassic (Ladinian, 239.5 Ma)	240
57	Middle Triassic (Anisian, 244.6 Ma)	245
58	Permo-Triassic Boundary (252 Ma)	250
59	Late Permian (Lopingian, 256 Ma)	255
60	late Middle Permian (Capitanian, 262.5 Ma)	260
61	Middle Permian (Wordian/Capitanian Boundary 265.1 Ma)	265
62	Middle Permian (Roadian&Wordian, 268.7 Ma)	270
63	Early Permian (late Kungurian, 275 Ma)	275
64	Early Permian (early Kungurian, 280 Ma)	280
65	Early Permian (Artinskian, 286.8 Ma)	285
66	Early Permian (Sakmarian, 292.6 Ma)	290
67	Early Permian (Asselian, 297 Ma)	295
68	Late Pennsylvanian (Gzhelian, 301.3 Ma)	300
69	Late Pennsylvanian (Kasimovian, 305.4 Ma)	305
70	Middle Pennsylvanian (Moscovian, 311.1 Ma)	310
71	Early/Middle Carboniferous (Baskirian/Moscovian boundary, 314.6 Ma)	315
72	Early Pennsylvanian (Bashkirian, 319.2 Ma)	320
73	Late Mississippian (Serpukhovian, 327 Ma)	325
74	Late Mississippian (Visean/Serpukhovian boundary, 330.9 Ma)	330
75	Middle Mississippian (late Visean, 333 Ma)	335
76	Middle Mississippian (middle Visean, 338.8Ma)	340
77	Middle Mississippian (early Visean, 344 Ma)	345
78	Early Mississippian (late Tournaisian, 349 Ma)	350
79	Early Mississippian (early Tournaisian, 354Ma)	355
80	Devono-Carboniferous Boundary (358.9 Ma)	360
81	Late Devonian (middle Famennian, 365.6 Ma)	365

82	Late Devonian (early Famennian, 370 Ma)	370
83	Late Devonian (late Frasnian, 375 Ma)	375
84	Late Devonian (early Frasnian, 380 Ma)	380
85	Middle Devonian (Givetian, 385.2 Ma)	385
86	Middle Devonian (Eifelian, 390.5 Ma)	390
87	Early Devonian (late Emsian, 395 Ma)	395
88	Early Devonian (middle Emsian, 400 Ma)	400
89	Early Devonian (early Emsian, 405 Ma)	405
90	Early Devonian (Pragian, 409.2 Ma)	410
91	Early Devonian (Lochkovian, 415 Ma)	415
92	Late Silurian (Pridoli, 421.1 Ma)	420
93	Late Silurian (Ludlow, 425.2 Ma)	425
94	Middle Silurian (Wenlock, 430.4 Ma)	430
95	Early Silurian (late Llandovery, 436 Ma)	435
96	Early Silurian (early Llandovery, 441.2 Ma)	440
97	Late Ordovician (Hirnantian, 444.5 Ma)	445
98	Late Ordovician (Katian, 449.1 Ma)	450
99	Late Ordovician (Sandbian, 455.7 Ma)	455
100	Middle Ordovician (late Darwillian, 460 Ma)	460
101	Middle Ordovician (early Darwillian, 465 Ma)	465
102	Early Ordovician (Floian/Dapingian boundary, 470 Ma)	470
103	Early Ordovician (late Early Floian, 475 Ma)	475
104	Early Ordovician (Tremadoc, 481.6 Ma)	480
105	Cambro-Ordovician Boundary (485.4 Ma)	485
106	Late Cambrian (Jiangshanian, 491.8 Ma)	490
107	Late Cambrian (Pabian, 495.5 Ma)	495
108	late Middle Cambrian (Guzhangian, 498.8 Ma)	500
109	late Middle Cambrian (early Epoch 3, 505 Ma)	505

110	early Middle Cambrian (late Epoch 2, 510 Ma)	510
111	early Middle Cambrian (middle Epoch 2, 515 Ma)	515
112	Early/Middle Cambrian boundary (520 Ma)	520
113	Early Cambrian (late Terreneuvian, 525 Ma)	525
114	Early Cambrian (middle Terreneuvian, 530 Ma)	530
115	Early Cambrian (early Terreneuvian, 535 Ma)	535
116	Cambrian/Precambrian boundary (541 Ma)	540

1053 \* *Simulations were not run for the time intervals highlighted in italics.*

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1056 Table 2. Summary of Model Simulations

1057 The table summarises the simulations performed in this study. The 6<sup>th</sup> and 7<sup>th</sup> column refers to how  
 1058 the model was initialised. The smooth CO<sub>2</sub> simulations were initialised from existing simulations  
 1059 using paleogeographies which were provided commercial in confidence. The time period for which  
 1060 these paleogeographies correspond to are listed in column 6, and the CO<sub>2</sub> value used in the runs is in  
 1061 column 7. The Foster CO<sub>2</sub> simulations were initialised from the end point of the smooth CO<sub>2</sub>  
 1062 simulations.

Time Period (in Ma)	CO <sub>2</sub> (in ppmv) for the Smooth CO <sub>2</sub>	CO <sub>2</sub> (in ppmv) for the Foster CO <sub>2</sub>	Length of Run (in years) (Smooth CO <sub>2</sub> )	Length of Run (in years) (Foster CO <sub>2</sub> )	Time period (in Ma) from existing simulations used for the initial condition for smooth CO <sub>2</sub> simulation	CO <sub>2</sub> (in ppmv) used for the initial conditions
0.0	280	276	5000	5000	0	280
3.1	384	298	5000	2000	3	401
10.5	410	299	5000	2000	13	280
14.9	423	310	5000	2000	13	280
19.5	430	338	5000	2000	13	280
25.6	439	502	5000	2000	26	560
31.0	500	764	5000	5000	26	560
35.9	533	901	5000	5000	26	560
39.5	557	796	5000	5000	26	560
44.5	594	751	5000	8000	26	560
51.9	649	736	5000	8000	52	560

56.0	630	570	5000	5000	52	560
61.0	604	335	5000	5000	52	560
66.0	576	229	5000	5000	69	560
69.0	560	262	5000	5000	69	560
75.0	633	559	5000	2000	69	560
80.8	704	667	5000	2000	69	560
86.7	775	590	5000	8000	69	560
91.9	839	466	5000	8000	92	560
97.2	840	707	5000	8000	92	560
102.6	840	1008	5000	5000	92	560
107.0	840	1028	5000	5000	92	560
111.0	827	1148	5000	8000	92	560
115.8	811	1103	5000	8000	92	560
121.8	784	986	5000	5000	92	560
127.2	752	898	5000	5000	92	560
131.2	728	896	5000	5000	92	560
136.4	699	1020	5000	8000	136	840
142.4	677	832	5000	8000	136	840
145.0	667	713	5000	8000	136	840
148.6	654	721	5000	8000	136	840
154.7	631	802	5000	5000	155	560
160.4	617	785	5000	5000	155	560
164.8	606	868	5000	8000	155	560
168.2	596	1019	5000	5000	167	840
172.2	581	1046	5000	5000	167	840
178.4	560	986	5000	5000	178	1120

186.8	560	949	5000	5000	178	1120
190.8	560	1181	5000	5000	178	1120
196.0	560	1784	5000	5000	178	1120
201.3	560	1729	5000	5000	178	1120
204.9	560	1503	5000	5000	218	560
213.2	560	1223	5000	5000	218	560
217.8	560	1481	5000	5000	218	560
222.4	557	1810	5000	5000	218	560
227.0	553	2059	5000	5000	218	560
232.0	549	1614	5000	5000	218	560
233.6	548	1492	5000	5000	218	560
239.5	543	1034	5000	8000	218	560
244.6	540	419	5000	2000	218	560
252.0	534	879	5000	2000	257	1120
256.0	531	811	5000	2000	257	1120
262.5	526	352	5000	2000	257	1120
265.1	524	321	5000	2000	257	1120
268.7	521	311	5000	2000	257	1120
275.0	517	556	5000	2000	257	1120
280.0	513	690	5000	2000	257	1120
286.8	508	626	5000	2000	257	1120
292.6	503	495	5000	2000	297	280
297.0	500	445	5000	2000	297	280
301.3	510	393	5000	2000	297	280
305.4	520	358	5000	2000	297	280
311.1	534	338	5000	2000	297	280



314.6	542	327	5000	2000	297	280
319.2	553	328	5000	2000	297	280
327.0	571	317	5000	2000	297	280
330.9	581	296	5000	2000	339	420
333.0	586	263	5000	2000	339	420
338.8	600	233	5000	2000	339	420
344.0	653	565	5000	2000	339	420
349.0	705	645	5000	2000	339	420
354.0	758	589	5000	2000	339	420
358.9	809	587	5000	2000	339	420
365.6	880	806	5000	2000	339	420
370.0	926	811	5000	2000	339	420
375.0	979	1052	5000	2000	339	420
380.0	1029	1269	5000	2000	339	420
385.2	1079	1377	5000	2000	377	1680
390.5	1131	1093	5000	2000	377	1680
395.0	1174	1297	5000	2000	377	1680
400.0	1223	1731	5000	2000	377	1680
405.0	1271	1689	5000	2000	377	1680
409.2	1319	2102	5000	2000	377	1680
415.0	1368	1579	5000	2000	377	1680
421.1	1427	1457	5000	2000	377	1680
425.2	1466	1490	5000	2000	377	1680
430.4	1517	1531	5000	2000	377	1680
436.0	1571	1576	5000	2000	377	1680
441.2	1614	1643	5000	2000	439	1877

444.5	1636	1708	5000	2000	439	1877
449.1	1666	1799	5000	2000	439	1877
455.7	1710	1929	5000	2000	439	1877
460.0	1738	2013	5000	2000	439	1877
465.0	1770	2111	5000	2000	439	1877
470.0	1803	2210	5000	2000	439	1877
475.0	1836	2308	5000	2000	439	1877
481.6	1879	2438	5000	2000	439	1877
485.4	1904	2513	5000	2000	439	1877
491.8	1946	2639	5000	2000	439	1877
495.5	1970	2711	5000	2000	439	1877
498.8	1992	2776	5000	2000	439	1877
505.0	2020	2870	5000	2000	439	1877
510.0	2040	2940	5000	2000	439	1877
515.0	2060	3010	5000	2000	439	1877
520.0	2080	3080	5000	2000	439	1877
525.0	2100	3150	5000	2000	439	1877
530.0	2120	3220	5000	2000	439	1877
535.0	2140	3290	5000	2000	439	1877
541.0	2164	3374	5000	2000	439	1877

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