Editor Comments

Many thanks for answering the comments raised by the different reviewers. This manuscript should be accepted but it would be nice to address the few points raised by the reviewer 3 on the coherency between the SP19 and WP2014 chronologies (and potential implications for a future WAIS chronology) in a revised manuscript. Thanks again for your efforts.

Thank you very much for your positive assessment of our article. Below are our responses to the reviewer comments as well as a revised manuscript. As you requested, we have addressed the points raised by reviewer 3. This includes the addition of a new figure showing the coherency between SP19 and WD2014 (Fig. S5). Our revisions also include clarification about the implications for the existing WD2014 timescale (lines 798-801 in this document).

Review #1

Thank you very much for the very useful and positive comments. We greatly appreciate your input. Below are our line-by-line responses to your review.

Winski et al. present a first chronology, named SP19, for the South Pole Ice Core (SPICECore), back to 54,302 yr BP. This chronology is based on a combination of 251 volcanic matching to the WAIS Divide ice core and annual layer counting (back to 11,341 yr BP for the latter). More precisely, the SP19 chronology is strictly tied to the WD2014 chronology at the depth of the volcanic matches, and then layer counting is used to interpolate in-between. Before 11,341 yr BP, a spline interpolation method is used instead of annual layer counting. The layer counting is based on CFA measurements of magnesium, sodium, sulfate, chloride and dust. It has been performed by 4 different operators and reconciliation is found a posteriori when there is a discrepancy. A comparison is also made with visual stratigraphy but chemical stratigraphy is preferred because it is found to be more accurate.

It is argued that the WD2014 is used because it is more precise (annual layer thickness is larger at WD) and to have a WD2014 compatible time scale. The relative uncertainty to WD2014 is small during the Holocene, generally less than 18 yr and always less than 25 yr. For older parts, the relative uncertainty to WD2014 is less than 124 yr. The accumulation rate which is found in the SPICEcore is found to be mainly due to the upstream spatial pattern of accumulation along the flow line. The accumulation reconstruction is also compatible with nitrate concentration, nitrate amplitude of seasonal variations and N-15 of N2 in air bubbles (through a dynamical Herron-Langway firn model for the latter). It is therefore argued that the SPICEcore is a good candidate to test the influence of surface accumulation rate on the Lock-In Depth.

This article is very well written and its content is consistent. I therefore recommend to accept it.
Thank you.

I only have a few technical corrections.
- l . 385: suppress "accumulation rate" since it is actually not plotted on this figure.
Done.

l. 489, l. 531 and l. 535: double space after dot.
Fixed.

Review #2

Thank you very much for the very useful and positive comments. We greatly appreciate your input. Below are our line-by-line responses to your review.

In the paper "The SP19 Chronology for the South Pole Ice Core - Part 1: Volcanic matching and annual-layer counting" by D. A. Winsky and co-authors a new timescale (called SP19) for the SPICEcore is presented. This new time scale was partly achieved by annual layer counting but the main guide to build the age scale is a robust volcanic match coming from a comparison with WAIS Divide ice core chronology. Given the best quality of annual layering in WAIS Divide (as shown in figure 4) I think that the authors choose the best methodological approach to build this time scale, considering WAIS as the most reliable annually counted scale. The discussion about the uncertainty of the age scale is really well done and confirms the goodness of this first SP19 age scale both for the Holocene period (with a maximum uncertainty of about 25 years) and for older ice (maximum 124 years in the longest time window without tie points). The paper is clear and well written and I recommend its publication after considering the following minor points.
Thank you.

Minor comments:
Figure 5: I would recommend to change the x-axis in kyr BP instead of using x10ˆ4. In my opinion it would be much readable.
We have changed the x-axis in Figure 5.

A list (table) of the volcanic horizons used to match WDC and SPICEcore would be valuable if inserted in the text or in supplementary material (even if archived at the NCDC or other repository).
The timescale and a full list of tie points will be archived in the supplementary material and will be available at the National Climate Data Center (www.ncdc.noaa.gov) and the U.S. Antarctic Program Data Center (http://www.usap-dc.org).

Line 427: I would change this sentence to "Below 798 m depth (start of the Holocene): : :."
Fixed.

Review #3

Thank you very much for the very useful and positive comments. We greatly appreciate your input. Below are our line-by-line responses to your review.

Review by Anders Svensson of manuscript entitled 'The SP19 Chronology for the South Pole Ice Core – Part 1: Volcanic matching and annual-layer counting' submitted to Climate of the Past by D. Winsky et al.

The manuscript (MS) introduces a stratigraphic chronology SP19 of the South Pole SPICE ice core based on 1) Holocene layer counting in high-resolution discrete chemistry samples and continuous records, and 2) a transfer of the WD2014 chronology based on identification of 251 common volcanic match points distributed over the last 54 ka. The layer counting is compared to a previously obtained independent layer counting from the same core based on visual stratigraphy alone. Furthermore, the authors are introducing an accumulation rate profile for the SPICE core based on a kink model. The model is compared to upstream accumulation patterns, to nitrate concentration profiles and to a d15N of N2 profile that all seem to support the obtained accumulation profile.

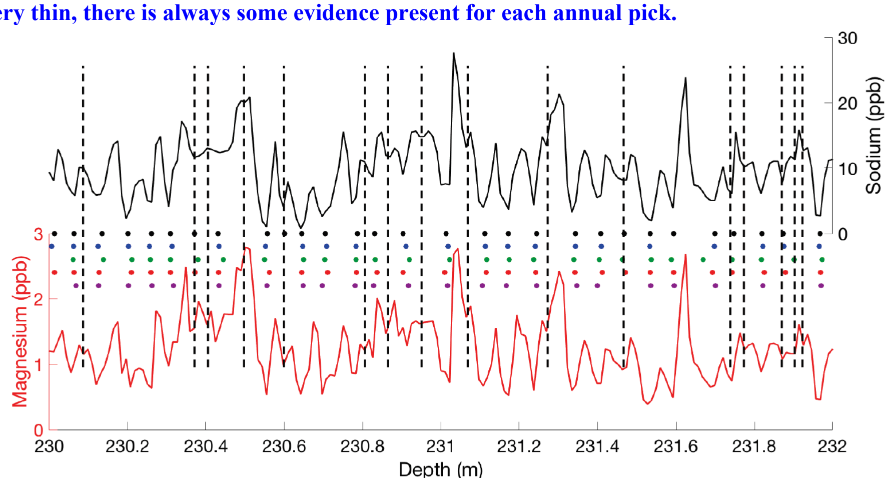
Overall, the MS is well written, well referenced and the figures are clear and illustrative of the study. The MS is well structured, the language is clear and the conclusions are well argued for.
Thank you.

I only have a few comments below for the authors to consider.
The authors perform multiple careful counting of annual layers of the Holocene using chemical parameters, continuous dust and conductivity. They then compare their resulting layer counting to an independent layer counting based on visual stratigraphy alone, and find an overall good agreement between the two approaches (Figure 9). Whereas this is a good test to see how well the two independent approaches are, it would probably have resulted in a better overall time scale, if all of the available high resolution records (chemistry + visual stratigraphy) had been combined in a common dating exercise from the beginning?
We agree that it would have been informative to combine the visual and chemical layer counting from the beginning. However, the visual layer counting was completed two years prior to the chemistry layer counts and had already been tied independently to WD2014 with electrical conductivity before the chemistry data were available or before we had agreed on an overall dating strategy. Since the timescale was ultimately linked with WD2014 using sulfate and electrical conductivity, any minor changes to the timescale resulting from an earlier reconciliation between methods would likely have little effect and would be within our uncertainty estimates.

Whereas I agree to the approach of transferring theWD2014 chronology to SPICE core rather than publishing a new independent time scale for SPICE core, it still seems like quite a large effort to do 4x independent layer counting of SPICE just to end up doing a transfer of time scale? Probably most of that time scale transfer could have been done based on a depth-depth matching alone (WDC – SPICE) similar to the approach taken in Figure 5?
Yes, this was a lot of effort! We initially hoped to produce an independent timescale based on layer counting, but we ultimately decided to synchronize the timescale with WAIS Divide for the reasons described in section 3.1. However, we believe the effort was useful in assessing the uncertainty in the timescale.

There are certain depth intervals (228-275m and 626-687m), where all of the independent layer counting plus the automated Straticounter dating approach consistently count significantly fewer annual layers than suggested by the transfer of the WD2014 time scale. I understand that this consistent undercounting is associated with periods of exceptionally low accumulation (upstream) at SPICE. Are the authors convinced, however, that their layer counting is wrong and that there is not a problem with the WD2014 counting in one or both of those periods? In other words, can the authors account for all of the 'missing' layers when they go back and recount the critical sections in SPICE core? I'm not suggesting that the authors should revise the WD2014 time scale, but independent checks are always useful, and considering the effort put into precise dating of SPICE, an outcome may be suggestions for future revisions of the WDC chronology?

This is a good question. We were able to account for all of the 'missing' layers needed during intervals where we had initially undercounted years. During these sections, the annual layering was less clear and many annual layers were missed by all five interpreters. However, with the knowledge that there should be extra years in a given interval, it was never difficult to find examples of less well-expressed annual layers needed to synchronize the timescale. Furthermore, the missed layers often fell in intervals with smaller layer thickness (lower accumulation), making it difficult to resolve the annual cycle with the sampling frequency. The figure below shows an example of such an area near 230 m, in a section that was initially very undercounted. This figure is equivalent to Figure 6 in the manuscript with sodium and magnesium plotted against the pick positions of the 5 interpreters. I have also added dotted lines to show where additional years were inserted for reconciliation. While many of the new years are poorly expressed in the chemistry or very thin, there is always some evidence present for each annual pick.



112,843 samples were collected and analyzed individually for this project! Did the authors consider doing fewer discrete and more continuous sample analysis? A CFA system optimized for depth/time resolution should be able to resolve the annual layering throughout the Holocene period. Of course, I'm not suggesting to do that now, but for future projects it may be an alternative?

Thank you for the suggestion! It was, indeed a lot of work to process such a high volume of samples. Our decision to analyze discrete samples was based on previous positive experiences with discrete sampling, the expertise of the team and the availability of personnel. We believe our efforts were worthwhile due to the high quality of the resulting

data.

A depth-difference relation figure between the two synchronized ice cores (SPICE, WDC) is very good for evaluating the synchronization and/or to identify regions where accumulation/thinning of the two records deviate. See Figure 2 of Seierstad et al., QSR, 2014: 'Consistently dated records from the Greenland GRIP, GISP2 and NGRIP ice cores for the past 104 ka reveal regional millennial-scale d18O gradients with possible Heinrich event imprint'.
Thank you for this reference. Since we have elected to tie the SP19 record exactly to the WD2014 record, a precisely equivalent figure would show a depth differences dominated by the very different accumulation rates (~7.4 cm/yr at SPICEcore and ~20 cm/yr at WAIS Divide) as well as very different flow patterns (SPICEcore is hundreds of kilometers from a divide – WAIS Divide is much closer). For instance, in WAIS Divide, 10,000 BP is located at 1800 meters depth – deeper than the bottom of SPICEcore. However, Fig. S1 shows some similar information with the age offset at a given depth between the initial, unreconciled layer counts with WD2014. A similar diagram to Fig. S1, or configured as done by Seierstad et al. 2014, in our case, would show a nearly flat line with no offset since we forced the timescale within +/- 1 year of each tie point.

After having nicely synchronized the SPICE and WDC records, it would be nice to see the two climate profiles (water isotopes) in the same figure on their common time scale. If for whatever reason that is not possible, maybe the Calcium profiles of the two ice cores can be shown together? It is difficult for the reader to evaluate the quality of the volcanic matching without seeing a comparison of some parameter of the two ice core records.
We have added this diagram to the supplemental material (Fig. S5) using calcium data.

Minor comments:
Accumulation mistakenly used in figure 7 caption, already mentioned by Frederic.
Fixed.

In Figure 3 is shown the seasonal variability of four impurities but not including nitrate. In section 4.5 and in Figure 11B the nitrate seasonality is discussed. Maybe it makes sense to include the seasonal variability of nitrate in Figure 3?
We have added a 5th panel in Figure 3 to show the nitrate data.

In Figure 7, the annual layer thickness appears to stay constant or even increase throughout the glacial part of the record. Wouldn't one normally expect a thinning of annual layering with depth?
You are correct that one would expect layer thickness to decrease with depth for a constant accumulation rate. However, if the accumulation was higher for older ages, the extra thinning experienced by those will not necessarily offset the greater initial thickness. Thus, the layer thickness minimum near 25 ka BP could be due to accumulation rates that were substantially lower between 20-30 ka BP than they were beforehand. There is another possibility as well: the thinning function at the South Pole may be complex (i.e. not monotonic) due to the location hundreds of kilometers from a divide with irregular bedrock topography and converging/diverging flow. We have discussed Holocene accumulation rates here because they relate to our timescale accuracy and because they are much less sensitive to thinning. However, we have deliberately left a detailed interpretation of the layer thickness record and accumulation reconstruction for ice older than the Holocene to future studies (for instance Fudge et al. 2019 – CPD).

The SP19 Chronology for the South Pole Ice Core - Part 1: Volcanic matching and annual-layer counting

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Abstract

The South Pole Ice Core (SPICEcore) was drilled in 2014-2016 to provide a detailed multi-proxy archive of paleoclimate conditions in East Antarctica during the Holocene and late Pleistocene. Interpretation of these records requires an accurate depth-age relationship. Here, we present the SP19 timescale for the age of the ice of SPICEcore. SP19 is synchronized to the WD2014 chronology from the West Antarctic Ice Sheet Divide (WAIS Divide) ice core using stratigraphic matching of 251 volcanic events. These events indicate an age of 54,302 +/- 519 years BP (before the year 1950) at the bottom of SPICEcore. Annual layers identified in sodium and magnesium ions to 11,341 BP were used to interpolate between stratigraphic volcanic tie points, yielding an annually-resolved chronology through the Holocene. Estimated timescale uncertainty during the Holocene is less than 18 years relative to WD2014, with the exception of the interval between 1800 to 3100 BP when uncertainty estimates reach +/- 25 years due to widely spaced volcanic tie points. Prior to the Holocene, uncertainties remain within 124 years relative to WD2014. Results show an average Holocene accumulation rate of 7.4 cm/yr (water equivalent). The time variability of accumulation rate is consistent with expectations for steady-state ice flow through the modern spatial pattern of accumulation rate. Time variations in nitrate concentration, nitrate seasonal amplitude, and $\delta^{15}N$ of N_2 in turn are as expected for the accumulation-rate variations. The highly variable yet well-constrained Holocene accumulation history at the site can help improve scientific understanding of deposition-sensitive climate proxies such as $\delta^{15}N$ of N_2 and photolyzed chemical compounds.

1. Introduction

Polar ice core records provide rich archives of paleoclimate information that have been used to advance understanding of the climate system. One of the great strengths of ice cores is the tightly constrained dating that permits interpretation of abrupt events and comparisons of phasing among records. Therefore, a critical phase in the development of any ice core record is the rigorous establishment of a depth-age relationship.

Several techniques are available to assign ages to each specific depth in an ice core. These include annual layer identification of chemical (e.g. Sigl et al. 2016; Andersen et al. 2006; Winstrup et al. 2012) and physical (e.g. Hogan and Gow 1997; Alley et al. 1997) ice properties, identification of stratigraphic horizons as relative age markers (e.g. Sigl et al. 2014; Bazin et al. 2013; Veres et al. 2013) and glaciological flow modeling (e.g. Parrenin et al. 2004). To establish a depth-age relationship for the South Pole Ice Core (hereafter SPICEcore), we use a combination of 1) annual layer counting of glaciochemical tracers and 2) stratigraphic matching of volcanic horizons to the West Antarctic Ice Sheet (WAIS) Divide ice core timescale "WD2014" (Sigl et al. 2016, Buizert et al. 2015).

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SPICEcore was drilled in 2014-2016 for the purpose of establishing proxy reconstructions of temperature, accumulation, atmospheric circulation and composition, and other earth system processes for the last 40,000 years (Casey et al. 2014). The SPICEcore record is the only ice core south of 80° S extending into the Pleistocene and is also located within one of the highest accumulation regions within interior East Antarctica (Casey et al. 2014). This provides the unique opportunity to develop the most highly resolved ice core record from interior East Antarctica. The South Pole is located at an elevation of 2835 m (Casey et al. 2014) and has a mean annual air temperature of -50°C (Lazzara et al. 2012). The high accumulation rate at South Pole (~8 cm yr^{-1} snow water equivalent, Mosley-Thompson et al. 1999; Lilien et al. 2018) relative to most of interior East Antarctica permits glaciochemical measurements at high temporal resolution. Occasional cyclonic events, particularly during winter months, bring seasonally variable amounts of sea salt, dust and other trace chemicals to the South Pole (Ferris et al. 2011; Mosley-Thompson and Thompson 1982; Parungo et al. 1981; Hogan 1997). Due to the favorable logistics and location at the geographic South Pole, the immediate area has been the site of several previous ice coring campaigns (e.g. Korotkikh et al. 2014; Budner and Cole-Dai 2003; Ferris et al. 2011; Meyerson et al. 2002; Mosley-Thompson and Thompson 1982). These ice cores contain records spanning the last two millennia, providing insight into seasonal chemistry variations and background values as well as recent snow accumulation trends.

In this paper, we focus on dating the ice itself; the dating of the gas record and the calculation of the gas-age/ice-age difference will be the subject of a future paper. The procedures used to generate the data necessary for ice-core dating and the dating techniques themselves are summarized in the remainder of the paper.

2. Measurements and Ice core data

2.1 Measurements

2.1.1 Fieldwork and Preparation Drilling began at the South Pole in the 2014/2015 austral summer season at a location 2.7 km from the Amundsen-Scott station, using the Intermediate Depth Drill designed and deployed by the U.S. Ice Drilling Program (Johnson et al. 2014). Drilling began at a depth of 5.10 m and reached a depth of 755 m in January 2015. Drilling continued during the 2015/2016 season, reaching a final depth of 1751 m. To extend the record to the surface, a 10 m core was hand-augered near the location of the main borehole. Ice core sections with a diameter of 98 mm and length of 1 m were packaged and shipped to the National Science Foundation Ice Core Facility (NSF-ICF) in Denver, Colorado. Each meter-long section of core was weighed and measured to calculate density and assign core depth. The cores were cut using bandsaws into CFA (continuous flow analysis) sticks with dimensions of 24 mm x 24 mm x 1 m and packaged in clean room grade, ultra-low outgassing polyethylene layflat tubing (Texas Technologies ULO) in preparation for the melter system at Dartmouth College. An additional 13 mm x 13 mm x 1 m stick was used for water-isotope analyses at the University of Colorado (see Jones et al., 2017 for water-isotope methods).

2.1.2 ECM measurements During core processing at the NSF-ICF, each core was cut and planed horizontally to produce a smooth, flat surface (Souney et al., 2014). Electrical conductivity measurements (ECM) were made with both direct current (DC) and alternating current (AC). We report only AC-ECM here, as it was the primary measurement for identifying volcanic peaks; further details are provided by Fudge et al. (2016a). Multiple tracks were made at different horizontal positions across the core (typically 3 tracks) and then averaged together. Measurements from each meter were normalized by the median to preserve the volcanic signal while providing a consistent baseline conductance to account for variations in electrode contact.

2.1.3 Visual Measurements Each core was examined by JF in a dark room with illumination from below. For some cores, particularly for depths greater than ~250 m, side-directed tray lighting using a scatter-diffuser was more effective at revealing features. All noteworthy internal features, stratigraphy, physical properties and seasonal indicators were documented by hand in paper log books.

Previous work at the South Pole shows that coarse-grained and/or depth-hoar layers form annually in late summer, often capped by a bubble-free wind-crust or iced crust up to ~1 mm thickness (Gow, 1965). We used these coarse-grained layers as the annual "picks" (noted as late-summers). The stratigraphy in the core was generally uniform and well-preserved, with the pattern identified by Gow (1965) continuing downward. The depths of all noted features were recorded to the nearest millimeter. Full details on visual layer counting are described in Fegyveresi et al. (2019).

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2.1.4. Ice Core Chemistry Analyses Ice sticks were melted and samples collected at Dartmouth College using a Continuous Flow Analysis – Discrete Sampling (CFA-DS) melt system (Osterberg et al. 2006). Stick ends were decontaminated by scraping with pre-cleaned ceramic (ZrO) knives. Cleaned sticks were then placed in pre-cleaned holders and melted on a melt head regulated by a temperature controller in a standup freezer. The melt head was made of 99.9995% pure chemical-vapor-deposited silicon carbide (CVD-SIC). CVD-SIC was chosen because of its ultra-high purity, high thermal conductivity, extreme hardness and excellent resistance to acids allowing for acid cleaning when not in use. The melt head design includes a 16x16x3 mm high tiered and rimmed inner section that was tapered with capillary slits to a center drain hole to minimize the risks of contamination from outer meltwater and wicking when melting porous firn (similar to Osterberg et al. 2006). This design provides a ≥4 mm buffer between the exterior of each ice stick and the edge of the center tiered section. Flexible plastic tines aligned on the four sides of the melt head keep the ice stick centered.

A peristaltic pump drew outer, contaminated meltwater away from the outer section through four waste lines. A second peristaltic pump drew clean meltwater from the center, tiered section of the melt head to a debubbler. The debubbler consisted of a short section of porous expanded PTFE tubing (Zeus Aeos 0000143895) and utilized pump pressure to force air through the tubing walls. The debubbled melt stream entered a splitter where it was separated into three fractions: one for major ion analyses, another for trace element analyses, and a third that passed through a particle counter and size analyzer (Klotz Abakus), an electrical conductivity meter (Amber Science 3084), and a flowmeter (Sensirion SLI-2000) before final collection in vials (Fig. 1). Samples were collected in cleaned vials using Gilson FC204 fraction collectors (cleaning procedures described in Osterberg et al. 2006). Samples were capped and kept frozen until additional analysis.

Core depths corresponding to each sample were tracked using custom software expanding on the concept of depth-point tracking developed by Breton et al. (2012). Simply, software tracks each depth point in the core as it progresses through the CFA-DS system until it reaches each collection vial. This is accomplished by using a combination of melt rate, flow rates, and system line volumes. Melt rates were measured with a weighted rotary encoder tracking displacement as the ice stick melts. Flow rates were measured by either an electronic flow meter or by calibrating the volume per revolution of each peristaltic pump tubing piece. Fraction collector advancements were made automatically based on melt rate, ice density (in firn), and the required sample volume and frequency. In addition, the software collected data from the inline particle counter and electronic conductivity meter. This system is capable of producing high-resolution, ultra-clean samples and has been used successfully in previous studies (e.g. Osterberg et al. 2017; Winski et al. 2017; Breton et al. 2012; Koffman et al. 2014). Samples corresponding to the top and bottom of each stick were assigned depths equal to the top and bottom depths measured at NSF-ICF, with intervening samples scaled linearly by the ratio of the NSF-ICF core lengths over the lengths measured by the depth encoder. This ensures that our data remain consistent with other SPICEcore datasets and there is no possibility of drift due to scraping core breaks, measurement or encoder errors.

Discrete ion chemistry samples were collected every 1.1 cm on average for the upper 800 m (Holocene) portion of the core and every 2.4 cm on average for older ice. In total, 112,843 samples were collected and analyzed using a Thermo Fisher Dionex ICS-5000 capillary ion chromatograph to determine the concentrations of the following major ions: nitrate, sulfate, chloride, sodium, potassium, magnesium and calcium. Liquid conductivity, particle concentration, and particle size distribution measurements were taken continuously with an effective resolution of 3 mm.

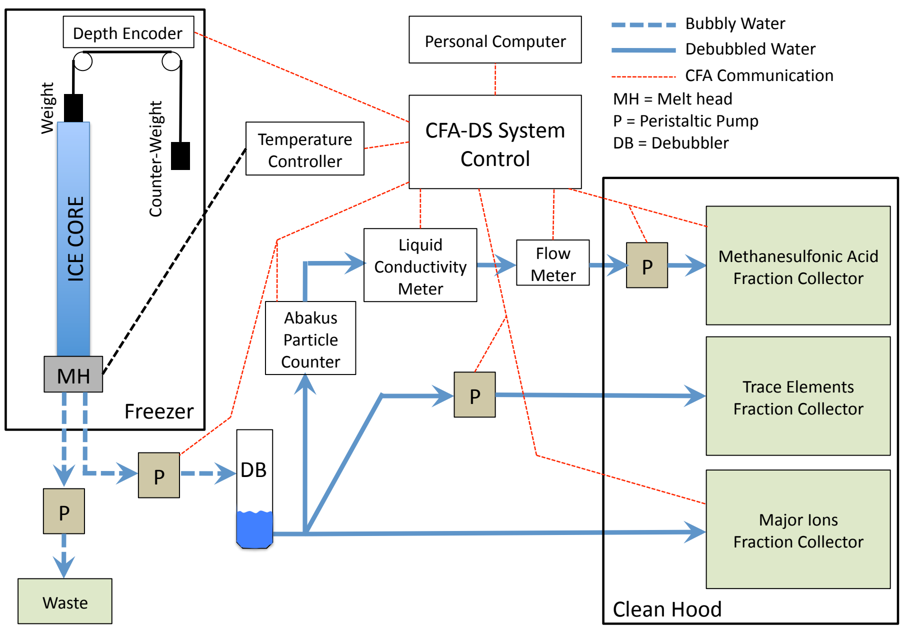

Figure 1: A schematic representation of the Dartmouth ice core melter system.

2.2 Chemistry Characteristics of SPICEcore

Previous research at the South Pole has shown that major sea salt ions (Cl^-, Na^+, Mg^{2+}) have winter maxima and summer minima when compared with the position of summer depth hoar layers (Cole-Dai and Mosley-Thompson 1999; Ferris et al. 2011). The same conclusion was reached through comparisons with seasonal isotopic fluctuations: sodium and magnesium peaks coincide with seasonal water-isotope minima (Legrand and Delmas 1984; Whitlow et al. 1992). These observations are consistent with sea salt aerosol measurements collected at the South Pole that demonstrate large sodium influx during winter months (Bodhaine et al. 1986; Bergin et al. 1998). The same seasonal pattern of sea salt deposition has been observed in Holocene strata of the WAIS Divide ice core (Sigl et al. 2016) and in other Antarctic ice cores (Kreutz et al. 1997; Curran et al. 1998; Wagenbach et al. 1998; Udisti et al. 2012). In the uppermost firn, seasonal chemistry is also influenced by the operation of South Pole station and its associated logistics (Casey et al. 2017).

In SPICEcore, sampling resolution is sufficiently high to consistently detect annual cyclicity in glaciochemistry throughout the Holocene. Clear annual signals are present in several glaciochemical species to a depth of 798 m (approximately 11341 BP), with the most prominent in sodium and magnesium (Figs. 2-3), which covary ($r = 0.95$; $p < 0.01$) and have coherent annual maxima and minima. Sulfate, chloride, AC-ECM, liquid conductivity, particle count and visual stratigraphy all exhibit discernable annual cyclicity.

The South Pole has long been recognized as a favorable location for identifying volcanic events, reflected by previous work on South Pole paleovolcanism (Ferris et al. 2011; Delmas et al. 1992; Budner and Cole-Dai 2003; Cole-Dai et al. 2009; Baroni et al. 2008; Cole-Dai and Thompson 1999; Palais et al. 1990). Volcanic events in SPICEcore are evident as peaks in sulfate and ECM rising well above background values. Within the Holocene, the median annual sulfate maximum is 60 ppb. This background level increases deeper in the core to values as high as 131 ppb between 18-26 ka BP, despite the lack of annual resolution during the Pleistocene. In contrast, sulfate concentration in volcanic events regularly exceeds 200 ppb with occasional concentrations as high as 1000 ppb for very large signals. For example, the pair of eruptions in 135 and 141 BP (1815 and 1809 CE), attributed to Tambora and Unknown in previous Antarctic studies (Delmas et al. 1992; Cole-Dai et al. 2000; Sigl et al. 2013) have peak sulfate concentrations of 518 and 281 ppb respectively, emerging well above seasonal background values of 60 ppb.

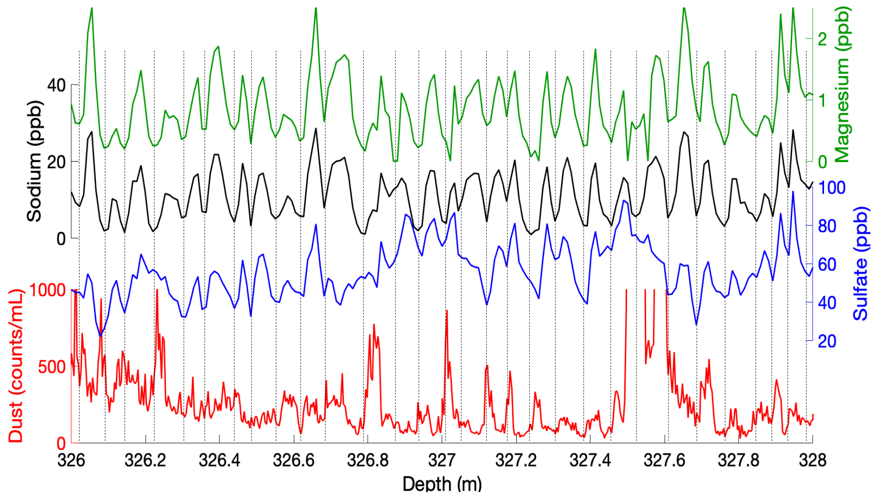

Figure 2: Example of annual layering in a representative segment of SPICEcore. Depicted are magnesium (green) and sodium (black) concentrations showing nearly identical variations and clear annual cyclicity. Sulfate (blue) has consistent but less pronounced layering, and dust (red; 1 micron size bin) has occasionally visible annual layering. Vertical dashed lines show annual pick positions based on the data shown.

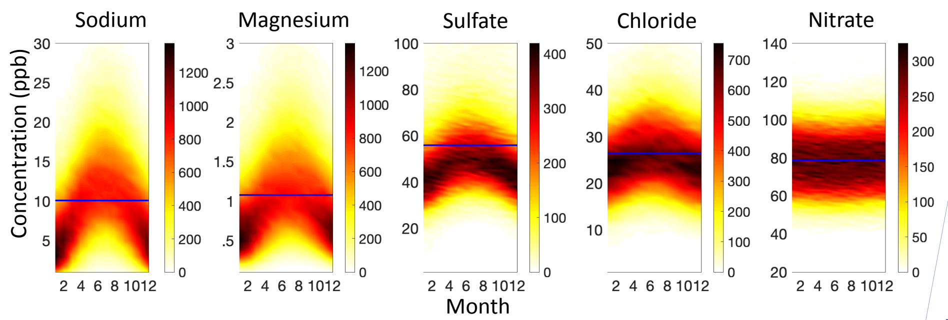

Figure 3: Seasonal variation in magnesium, sodium, sulfate, chloride and nitrate ion concentration in SPICEcore from -42 to 11341 BP (11383 total years). In each panel, the horizontal axis is month of the year (with 0 being Jan. 1^{st}) from linear interpolation between mean sample depth and the timescale. The vertical axis is concentration (ppb). The color scale indicates the density of measurements within gridded month and concentration bins. Concentration bin widths are 1 month (without claiming 1 month precision) and 1 ppb except for magnesium which is 0.1 ppb. The Holocene mean concentration of each ion is shown as a blue bar. Strong annual cyclicity is apparent in sodium and magnesium data. Annual cyclicity is weaker in sulfate, chloride and nitrate data.

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3. SPICEcore Dating Methods

3.1 Approach

The SPICEcore timescale (SP19) was developed by combining annual layer counting with volcanic event matching between SPICEcore and the WAIS Divide chronology. We identified 251 volcanic tie points that are clearly visible in both SPICEcore and WAIS Divide (Sigl et al. 2016). These tie points link SP19 with the WAIS Divide chronology, resulting in one of the most precisely dated interior East Antarctic records. Above 798 m, ages are interpolated between volcanic tie points using layer counts. Below 798 m, ages are interpolated between tie points by finding the smoothest annual layer thickness profile (minimizing the second derivative) that satisfies at least 95% of the tie points (following Fudge et al. 2014).

Although it is possible to create an independent, annually layer counted SPICEcore timescale during the Holocene, we linked the entire SP19 chronology to the WAIS Divide chronology for several reasons: (1) annual layers are insufficiently thick below 798 m (approximately 11341 BP) to consistently resolve individual years, requiring synchronization to another ice core to achieve the best possible dating accuracy. Tying the entire SP19 chronology to the WAIS Divide core ensures consistent temporal relationships between these two records; (2) although annual layers are remarkably well-preserved in SPICEcore chemistry, WAIS Divide has a higher accumulation rate (Banta et al., 2008; Fudge et al., 2016b; Koutnik et al. 2016) and stronger seasonality in chemical constituents (Sigl et al. 2016), producing more robust annual layering (Figure 4); (3) it is expected that some years at South Pole experience very low accumulation, resulting in a lack of an annually resolvable record during those years (Hamilton et al. 2004; Van der Veen et al. 1999; Mosley-Thompson et al. 1995, 1999); (4) an attempt to independently date the Holocene annual layers created drift of several percent at stratigraphic tie points. We therefore elected to anchor the SP19 timescale to WD2014, and use the annual layer counts as a means of interpolating between WD2014 tie points during the Holocene. The SP19 timescale spans -64 BP (2014 CE) to 54,302 +/- 519 BP, with the annually-dated Holocene section of the core extending to 11341 BP (798 m depth).

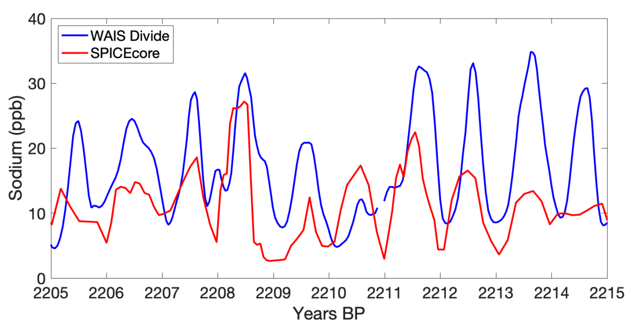

Figure 4: Annual layering of sodium in WAIS Divide (blue; Sigl et al. 2013) and SPICEcore (red). Annual layers in sodium are clear in both records but are more pronounced at WAIS Divide for most years.

3.2 Procedure for identifying matching events

The matching of volcanic events in sulfate and ECM records is commonly used to synchronize ice core timescales (e.g. Severi et al., 2007, 2012; Sigl et al. 2014; Fujita et al., 2015), including the recent extension of the annually-resolved WAIS Divide timescale to East Antarctic cores (Buizert et al., 2018). Volcanic matching is based on the depth pattern of events more than the magnitude of the events because the magnitude in individual ice cores can vary significantly across Antarctica depending on the location of the volcano and atmospheric transport to the ice core site. The volcanic matching between SPICEcore and WAIS Divide is based primarily on the sulfate record for SPICEcore and the combined sulfur and sulfate records for WAIS Divide (Buizert et al. 2018). AC-ECM from SPICEcore and WAIS Divide was used as a secondary data set and to fill small data gaps in the sulfate record. An example of the four data sets is shown in Figure 5.

The volcanic matches were performed independently by two interpreters (TJF and DF) and then reconciled by one (TJF) with concurrence from the other (DF). The position of each match was defined as the inception of the sulfate rise in order to most consistently reflect the timing of the volcanic event itself. Of the final 251 tie points, 229 were identified in the sulfate data by both interpreters. Of the remaining matches, 14 were made by one interpreter in the sulfate data, and at least one interpreter in the ECM data. One of the other matches was made only with ECM because of a gap in the sulfate data for SPICEcore. The last 7 matches were part of sequences not initially picked by one interpreter but deemed to be sufficiently distinct from the other events in the sequence to be included.

We note that the purpose of the volcanic matching was to develop a robust SPICEcore timescale, not to assess volcanic forcing. Thus, there are many potential volcanic matches that were not included either because they did not have the same level of certainty as the final 251 matches, or because they were in close proximity to the final matches and thus did not provide additional timescale constraints.

For the pre-Holocene section of the core, ages between the volcanic matches are interpolated by finding the smoothest annual layer thickness by minimizing the second derivative (Fudge et al., 2014). The goal of finding the smoothest annual layer thickness time series is to prevent sharp changes affecting the apparent duration of climate events on either side of a volcanic match point. The method allows the ages of the volcanic matches to vary within a threshold to produce a smoother annual layer thickness interpolation. The degree of smoothness was set such that 95% of the tie points are shifted by 1-year or less, which is a reasonable uncertainty on the precision of the volcanic matches.

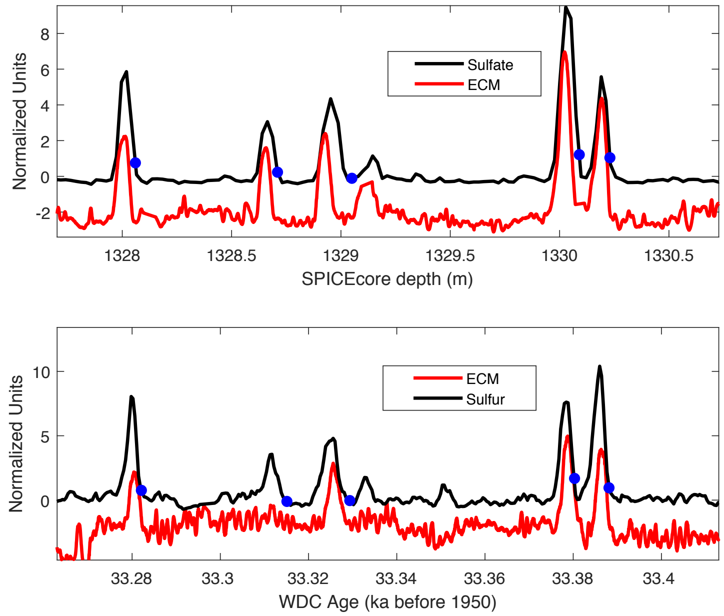

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Figure 5: An example of volcanic matching between SPICEcore (top) and WAIS Divide (bottom). Sulfate (black) and electrical conductivity (ECM; red) are shown for both ice cores. Here, five events are shown that link specific depths in SPICEcore to known ages in WAIS Divide. The position of the tie points is chosen at the beginning of the event (blue circles). The y-axis values are scaled for ease of visualization and do not indicate absolute measurement values.

3.3 Annual Layer Interpretation

Annual layer counting in SPICEcore was initially done independently of the volcanic matching with WAIS Divide. To minimize and quantify timescale uncertainty, five interpreters performed the layer counting independently: DW, DF, TJF, JF, and TC. Sodium and magnesium were the primary annual indicators, but electrical conductivity, dust concentration, sulfate, chloride and liquid conductivity were also helpful in delineating individual years. To remain consistent, each interpreter agreed to place the location of Jan. 1st for each year at the sodium/magnesium minimum, consistent with previous interpretation of South Pole sea salt seasonality (e.g. Ferris et al. 2011; Bergin et al. 1998). Two examples of annual layering including the Jan. 1st positions picked by each interpreter are shown in Figure 6. Shown here are sections of high (A) and low (B) agreement among the five interpreters.

This procedure resulted in five independent timescales to a depth of 540 m, containing between 6529 and 6807 years. The details of reconciling the five independent sets of layer counts are described in the Supplemental Information. Below 540 m, only one author (DW) continued with the layer counting once the decision to use the annual layers to interpolate between volcanic events had been made. The layer counting procedure resulted in an annually resolved timescale, fully independent of any external constraints, to a depth of 798.

Above 798 meters, 86 volcanic tie points were identified, producing 85 intervals within which a known number of years must be present. To make the layer-counted timescale consistent with these tie points, years were added or subtracted, as necessary, within each interval such that the layer-counted timescale passed through each tie point within +/- 1 year of its age, linking SPICEcore with the WAIS divide chronology. Procedural details for adding and subtracting layers by interval are discussed in the Supplemental Information. In most intervals, few years needed to be added or subtracted, with the average change in years equal to 5.6% of the interval length (Holocene intervals ranged from 6 to 747 years). In certain sections layer counting consistently differed from the WAIS-tied timescale. The most notable example is from 228 to 275 m depth where 105 years (14%) needed to be added.

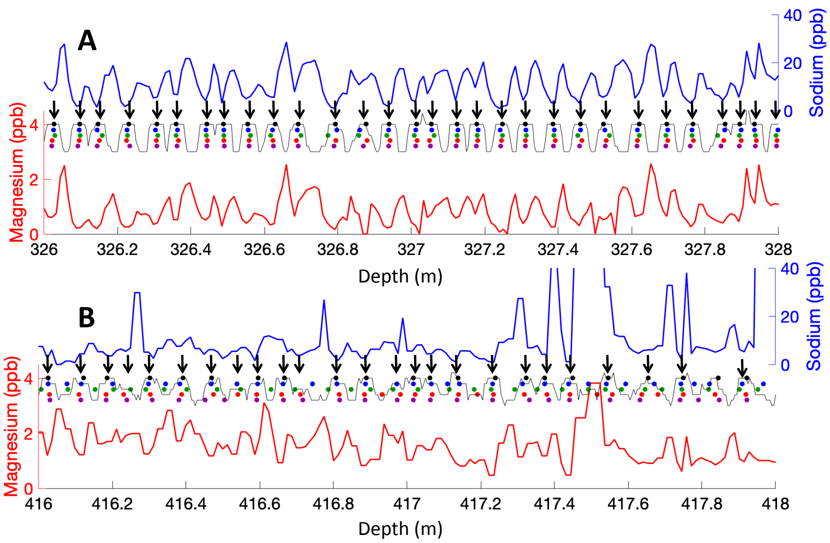

Figure 6: Representative sections of annual layer pick positions compared with magnesium (red) and sodium (blue) concentrations. Each interpreter is represented with a different color circle. Certain sections have excellent agreement among interpreters making reconciliation trivial (A), whereas other sections have poorly defined annual signals and associated disagreement among interpreters (B). The black line depicts the sum of all picks within +/- 2 cm; black arrows depict the final positions of the reconciled Jan. 1st annual layer picks.

4. Results and Discussion

4.1 Characteristics of the Timescale

The SP19 chronology extends from 2014 CE (-64 BP) at the surface to 54302 BP at 1751 m depth. The timescale and volcanic tie points are depicted in Figure 7 with volcanic tie points pinning the timescale also shown. Annual layer thicknesses near the surface are roughly 20 cm thick (owing to the low density of firn), decreasing rapidly to ~8 cm/yr by the firn-ice transition. The timescale is annually resolved between -64 and 11341 BP, below which resolution varies based on the distance between tie points. Using the methods in section 3.2 (Fudge et al. 2014), we report timescale values interpolated at 10-year resolution. The longest distance between tie points is 2476 years between 16348 and 19872 BP.

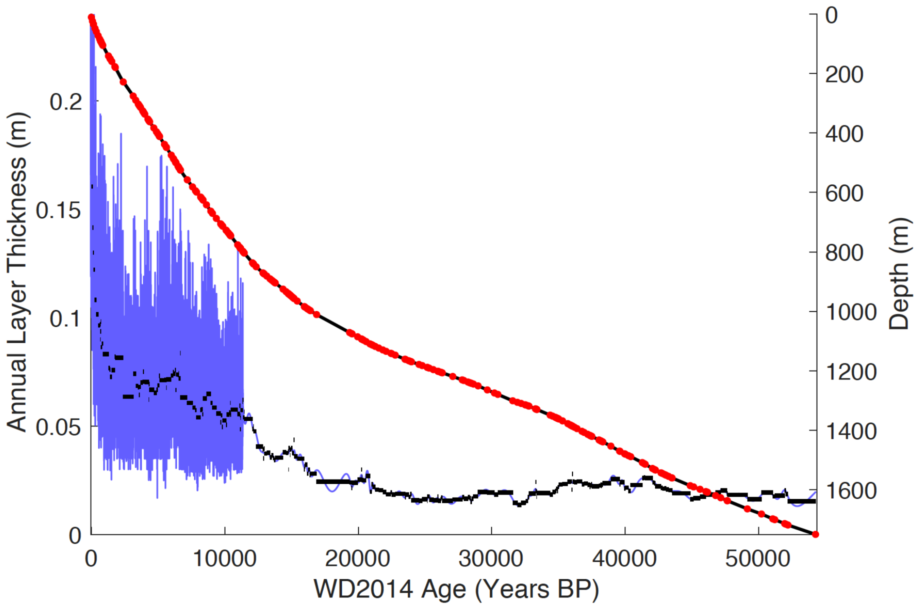

Figure 7: The SP19 timescale and layer thickness. The SP19 depth-age relationship (right y-axis, black line) is constrained by volcanic events (red dots) extending to 54302 BP. Annual layer thicknesses (left y-axis, blue) are shown at annual resolution during the Holocene and as decadally-interpolated thicknesses based on the smoothest annual layer thickness method (Fudge et al. 2014) during the Pleistocene. The average annual layer thickness during each volcanic interval is shown in black for comparison.

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4.2 Uncertainties

In discussing uncertainty values for SP19, the reported values are uncertainty *estimates* rather than rigorously quantified 1σ or 2σ values. There are several reasons for this: 1) the chemicals used to count annual layers have similar cyclicity and are not independent; 2) while each of the five interpreters counted layers independently, they were likely employing similar strategies; 3) certain years may not be well-represented in the data, providing insufficient information for accurate dating or quantifying uncertainty; 4) volcanic events were identified in clusters such that each event is not necessarily independent; 5) it is difficult to assign a numerical index of confidence to specific volcanic tie points. Instead, we discuss timescale uncertainties as uncertainty estimates, which are intended to approximate 2σ uncertainties but cannot be precisely defined as such. This approach follows that of Sigl et al. (2016).

We assess the SP19 timescale uncertainty with respect to the previously published WD2014 timescale (Sigl et al. 2016; Buizert et al. 2015). The absolute age uncertainty will always be equal to or greater than the uncertainty already associated with WD2014 (Buizert et al. 2015; Sigl et al. 2016; Fig. 8). In addition to the uncertainty in WD2014, there is also uncertainty in our ability to interpolate between stratigraphic tie points. During the Holocene, our layer-counting of sodium and magnesium concentration improves the timescale accuracy between tie points. Interpolation uncertainty can be estimated using the drift among the five different interpreters. We calculate the number of years picked by each interpreter in running intervals of 500 years in the final WD2014 synchronized timescale. Under ideal conditions, each interpreter would also pick 500 years within each interval, but on average the number of years picked by interpreters differs from the final timescale by 6.7%, usually by undercounting. This is similar to the metric described in section 3.3, wherein the average change in years needed to reconcile the layer counts and volcanic tie points was 5.6% of the interval length. Here, we report the larger and more conservative value of 6.7%. If our layer counting skill drifts by +/- 6.7% while unconstrained by volcanic tie points, then the interpolation uncertainties remain within +/- 18 years of WAIS Divide throughout the Holocene with the exception of a poorly-constrained interval between approximately 1800-3100 BP. The maximum uncertainty within the Holocene is +/- 25 years, occurring at roughly 2750 BP, where the nearest tie points are 373 years away at 2376 and 3123 BP. This relationship can be applied across the Holocene, with layers accumulating an uncertainty value equal to 6.7% of the distance to the nearest tie point (Fig. 8; blue).

Below 798 m depth (start of the Holocene), there were no annual layers to aid in our interpolation of the timescale, leading to larger uncertainties. Our assumption of the smoothest annual layer thickness (Fudge et al. 2014) satisfying tie points is the most accurate interpolation method in the absence of additional information, at least in Antarctic ice (Fudge et al. 2014). Using the WAIS Divide ice core as a test case, Fudge et al. (2014) estimated that the interpolation method accumulates uncertainties at a rate of 10% of the distance to the nearest tie-point, roughly 50% faster than the uncertainty of periods with identifiable annual layers. The longest interval with no volcanic constraints is between 16348 and 19872 BP. At 18110 BP, the center of the interval, the interpolation uncertainty reaches a maximum of 124 years, although uncertainties are proportionally lower in other intervals with closer volcanic tie points.

Figure 8 shows the total uncertainty estimates associated with the SP19 chronology, with interpolation uncertainties added to the published WAIS Divide uncertainties. The WD2014 and interpolation uncertainties are added in quadrature since the two sources of uncertainty are independent. The maximum estimated uncertainty in SP19 is 533 years at 34050 BP, the majority of which is attributed to uncertainties in WD2014. While it is not possible to rigorously quantify uncertainties throughout SP19, we believe these estimates provide reasonable and conservative values suitable for most paleoclimate applications. We acknowledge there is additional uncertainty related to the accuracy of our assigned stratigraphic tie points. Because of the conservative procedures discussed in section 3.1 wherein only unambiguous matches were used in linking the WAIS Divide and SPICEcore timescales, it is unlikely that any of these matches are in error. In previous work (Ruth et al. 2007), potential errors associated with tie points have been estimated by removing each tie point one at a time, and interpolating between the

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new series of tie points (with one point missing). If this procedure is repeated for each tie point and for each depth, the maximum error in age resulting from the erroneous inclusion of a tie point is approximately 83 years. However, because clusters of volcanic events were used to match the WAIS Divide and SPICEcore records, each tie point is not necessarily independent. Therefore, this method is more useful at sections of widely spaced tie points with greater potential uncertainties, but underestimates the uncertainties surrounding closely spaced events in SPICEcore and WAIS Divide. Examining calcium records from WAIS Divide (Markle et al. 2018) and SPICEcore shows concurrent timing in calcium variations between the two cores (Fig. S5), further supporting the choices of tie points.

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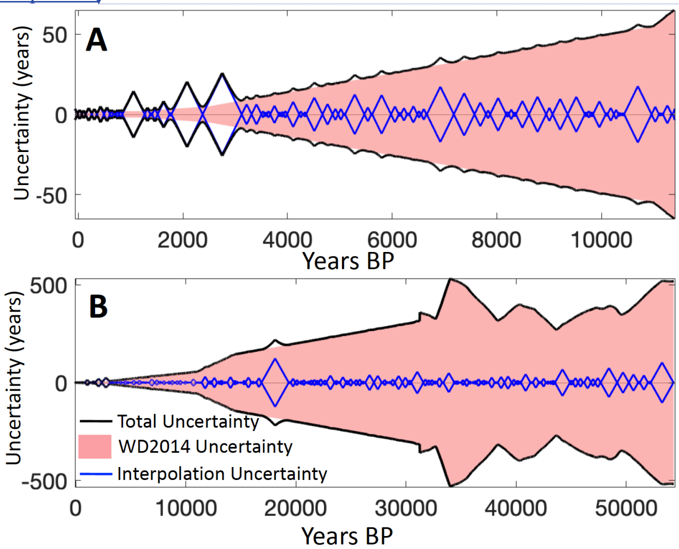

Figure 8: Uncertainty estimates in the SP19 timescale. The pink shading indicates the published uncertainty associated with the WAIS Divide timescale (Buizert et al., 2015; Sigl et al., 2016). The blue lines indicate the estimated uncertainty due to interpolation by layer counting (Holocene) and by finding the smoothest annual layer thickness history (Fudge et al. 2014; Pleistocene). Total uncertainty (black) is defined here as the root sum of the squares of the interpolation and WD2014 uncertainties. Total uncertainty estimates remain within +/- 50 years for most of the Holocene (A), but are as high as 533 years in the Pleistocene (B).

4.3 Comparison with Visual Stratigraphy

Visual stratigraphy in SPICEcore provides an independent check on the glaciochemical layer counting we used to interpolate the Holocene depth-age scale between tie points. Visual layer counting was conducted to a depth of 735 m (~10,250 years BP; Fegyveresi et al. 2017). We calculate the offset between the visual stratigraphic timescale and a linear interpolation between tie points and do the same for the chemistry layer counts (Fig. 9). If both the chemical and visual layer counting methods are capturing the true variability in layer thickness within intervals, then both would show the same structure within each interval.

There is broad correspondence between visual and chemical stratigraphy at all depths, which, with their almost completely independent origin and measurements techniques, is highly reassuring. In detail, though, there is little high-frequency correspondence between visual and chemical layer counts below 1400 BP (150 m depth), although a direct comparison is not possible since visible layer counts were not linked to stratigraphic tie points between 1400-2400 BP and 8400-9500 BP. Furthermore, visible layer counts were matched to the tie points within error of the WAIS Divide timescale, whereas the chemistry layer counts were forced to match within +/- 1 year of each tie point. In counting visible layers, occasional under- and overcounting of depth hoar layers within annual strata is likely, especially in deeper ice where thinning will make adjacent layers appear even closer. There were some intervals (e.g. 2000 – 2500 BP) in the core that appeared more homogeneous during viewing, and therefore annual layer choices have a higher level of uncertainty. Because of the differences between methodologies in matching to tie points and because of the uncertainties in visual counting below 2000 BP (200 m), we did not attempt to reconcile the visible and chemical layer counts, but instead rely only on the annual layers in the chemistry data.

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Between 100 and 1400 BP, both visible and glaciochemical timescales remain remarkably coherent and do not indicate drift of more than +/- 2 years. Over this interval, the correlation between the visible and chemical layer offsets from constant annual layer thickness (red and blue curves in Figure 9) is 0.74. The correlation between the two layer counting methods is as high as r = 0.85 between the tie points at 841 and 1268 BP. The discrepancy within the top 100 years is due to the tie point at 10.58 m, which was not included at the time of visible layer counting, as well as low layer chemical counting confidence within the firn column. There is no obvious relation between the accumulation rate and statistical agreement among methods.

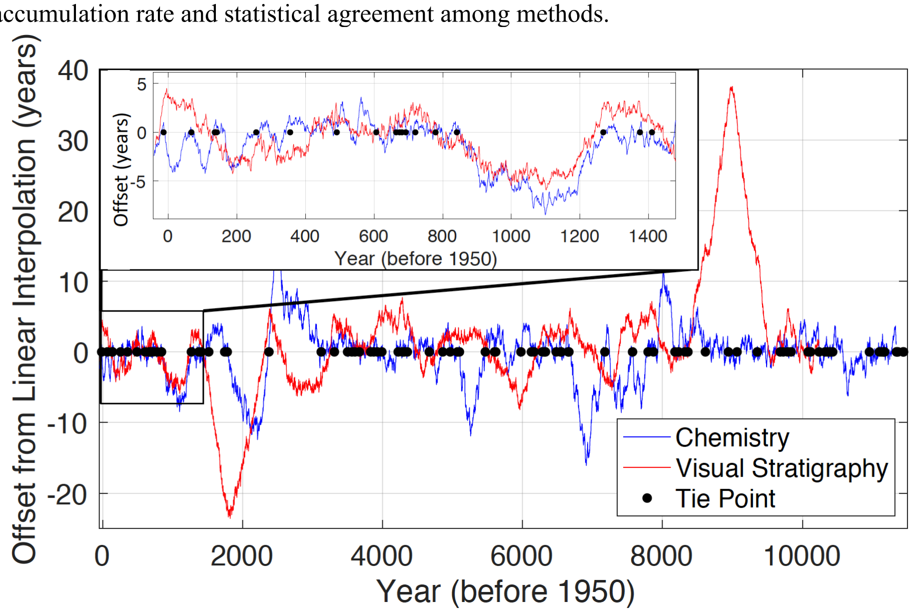

Figure 9: Comparison between visible layer (red) and chemistry-based (blue) Holocene annual timescales. Both curves are shown as residual values with respect to a linear

interpolation between tie points (black circles). When the shape of the red and blue curves is similar between tie points, we infer relatively high accuracy in both methods. The region showing the closest agreement between methods is shown in the inset with both curves remaining within 2 years of each other despite a long section with no tie points (841 to 1286 BP).

4.4 Accumulation Rate History

The SP19 timescale allows us to produce annually-resolved estimates of past snow accumulation to 11341 BP (Fig. 10). We apply a Dansgaard-Johnsen model (Dansgaard et al. 1969) to estimate the amount of thinning undergone by each layer of ice. Since the entirety of the Holocene in SPICEcore is located within the top third of the core (over 1900 m above the bed), the challenges associated with reconstructing surface accumulation are smaller than at sites with records closer to the bed (e.g. Kaspari et al. 2008, Thompson et al. 1998, Winski et al. 2017). Radar measurements indicate a bed depth at the South Pole of 2812 m, giving an ice-equivalent thickness of 2774 m, using the South Pole density function developed by Kuivinen et al. (1982). We used a kink height of 20% of the ice thickness and an input surface accumulation rate of 8 cm/yr (water equivalent), consistent with the parameters used by Lilien et al. (2018). The average Holocene accumulation rate is 7.4 cm/yr (water equivalent), in excellent agreement with results of previous studies (Hogan and Gow 1997; 7.5 cm/yr to 2000 BP; Mosley-Thompson et al. 1999 – 6.5-8.5 cm/yr for late 20th century). The upstream flow dynamics are too complicated for a static 1-D model to accurately determine the thinning function before the Holocene.

As discussed in Lilien et al. (2018), Koutnik et al. (2016), and Waddington et al. (2007), South Pole layer thicknesses are affected by 1) spatial variability in surface accumulation being advected to South Pole; 2) past climate-related changes in snow accumulation; and 3) post-depositional thinning due to ice flow. Thinning models can account for only the third factor. Understanding of Holocene climate history as recorded at other sites and in other indicators in SPICEcore, combined with knowledge of the modern upglacier variation in accumulation (Lilien et al., 2018), make it clear that the Holocene SPICEcore time-variations in accumulation are primarily from advection of spatial variations. Figure 10 shows Holocene accumulation rate in SPICEcore (black) compared with geophysically derived accumulation estimates over space using ice-penetrating radar (blue, details in Lilien et al. 2018). Using the present-day surface velocity field and the inferred 15% increase in flow rate, present day upstream surface accumulation rates were matched with corresponding ages at the SPICEcore borehole (Lilien et al. 2018). The close match between present-day near-surface accumulation rates upstream and the annual accumulation rate in SPICEcore shows that the millennial-scale signal of accumulation rate in SPICEcore is related to spatial patterns of snow accumulation upstream of South Pole.

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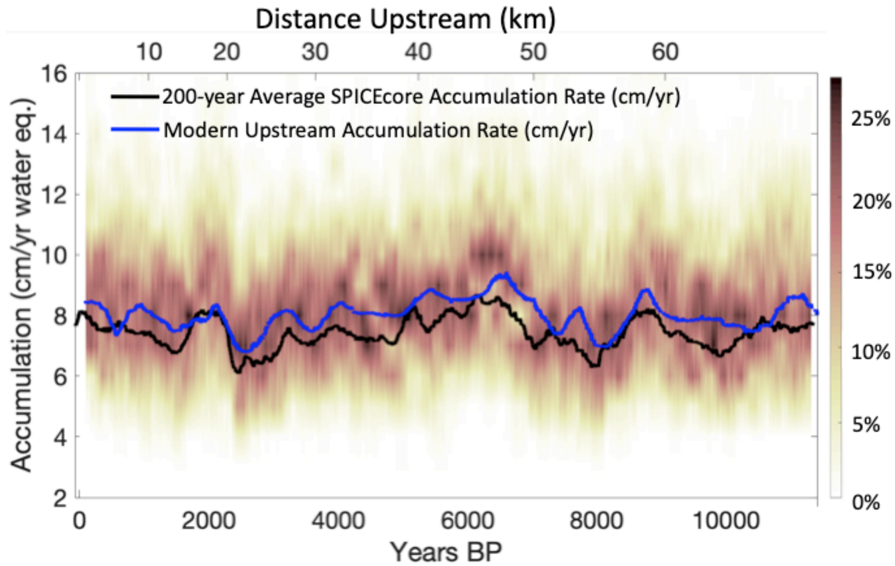

Figure 10: The Holocene accumulation rate history in SPICEcore. Shading indicates a running histogram of accumulation rate with darker colors indicative of more years at a given accumulation rate. The color axis (left) indicates percentage of years with a given accumulation rate within 1 cm accumulation bins across 200-year sliding intervals. The solid black line is the 200-year running mean of accumulation rate. These data are compared with modern spatial accumulation rates upstream of SPICEcore (blue; upper x-axis; Lilien et al. 2018).

A striking feature in the Holocene accumulation record in SPICEcore is the sharp dip centered on 2400 BP. Annual layers were notably less clear in that portion of SPICEcore because low accumulation rates led to low sampling resolution (5-6 samples/year). For instance, in the interval between 228-275 m, the interpreters picked between 511 and 670 years, when 747 years are present based on the volcanic tie points. Because the undercounting of layers in the development of SP19 is coincident with low accumulation rates, we are confident that this undercounting is due to poorly resolved layers in SPICEcore rather than to erroneous tie points or errors in the WD2014 chronology.

The cause of the sharp drop in accumulation is not clear. Modern accumulation rates upstream of SPICEcore were measured using a 20 m-deep isochron imaged with ice penetrating radar (Lilien et al. 2018). These results show lower accumulation in the location where the 2400 BP ice originated (Fig. 10). However, the modern upstream spatial pattern of accumulation shows a decline that is both more gradual and less than half the magnitude of the 2400 BP change in SPICEcore. It is possible that this represents a climatic signal, but we note sharp accumulation variations at this time that are not observed in the WAIS Divide core (Fudge et al. 2016b; Koutnik et al. 2016). Instead, we hypothesize that this event was most likely a transient local accumulation anomaly. Farther upstream at ~75km from South Pole, there is an accumulation low where the rate of change is approximately 3 cm/yr in 2 km. With the current South Pole ice flow velocity of 10 m/yr, this could explain a 3 cm/yr decrease in 200 years, similar to what is observed at 2400 BP. If a climate-driven accumulation anomaly did contribute to this sharp change, these anomalies do not appear to be common, as we see no other large and sustained change in the annual timescale.

On sub-centennial timescales, the effects of upstream advection of spatial accumulation patterns are likely smaller, such that annual-to-decadal patterns in snow accumulation in SPICEcore may be indicative of climate conditions. Previous studies have used a snow stake field 400 m to the east (upwind) of South Pole station to assess recent trends in accumulation rate with differing results. Mosley-Thompson et al. (1995, 1999) found a trend of increasing snow accumulation during the late 20th century, while Monaghan et al. (2006) and Lazzara et al. (2012) found decreasing snow accumulation trends between 1985-2005 and 1983-2010, respectively. No significant trends exist in the SPICEcore accumulation record within the last 50 years, although there is a significant ($p = 0.046$) increasing trend in snow accumulation in SPICEcore since 1900. Note that errors in measured firn density would influence this accumulation trend.

4.5 Nitrate Variability, $\delta^{15}N$ of N_2, and Accumulation

SPICEcore nitrate concentrations provide independent support for the Holocene accumulation rate history implied by the SP19 timescale. Previous studies have recognized an association between accumulation rate and nitrate concentration among ice core sites (Rothlisberger et al. 2002). Nitrate in surface snow, exposed to sunlight, results in photolytic reactions that volatilize nitrate and release it to the atmosphere (Erbland et al. 2013, Grannas et al. 2007; Rothlisberger et al. 2000). Evaporation of HNO_3 may also significantly contribute to nitrate loss in the surface snow (Munger et al. 1999; Grannas et al. 2007). Under low-accumulation conditions such as in East Antarctica, the amount of time snow is exposed at the surface is the dominant control on nitrate concentration, such that with more accumulation, snow is more rapidly buried and retains higher nitrate concentrations (Rothlisberger et al. 2000).

There is close correspondence between accumulation rate and nitrate concentration in SPICEcore (Fig. 11A). This association is strongest on multidecadal to multicentennial timescales with correlation coefficients between accumulation rate and nitrate reaching peak values after 512-year smoothing ($r = 0.60$; Fig. 11 inset). Although the smoothing makes standard metrics of statistical significance inapplicable, the similarity between time series is expected given the previous work described above. Among sites, an inverse relationship exists between seasonal amplitude of nitrate concentration and accumulation rate. High-accumulation sites such as Summit, Greenland exhibit strong annual nitrate layering, whereas low-accumulation sites such as Vostok (~2 cm w.e./yr; Ekaykin et al. 2004) and Dome C (~3.6 cm w.e./yr; Petit et al. 1982) do not show annual nitrate layers at all (Rothlisberger et al. 2000). SPICEcore has much higher accumulation rates than Vostok or Dome C, and retains weak intra-annual variability in nitrate. While minor compared with multi-annual and longer variability, nitrate seasonal cyclicity, wherein nitrate often peaks in the summer months (described in Grannas et al. 2007; Davis et al. 2004) is discernable in the SPICEcore nitrate record. As expected, the seasonal amplitude of nitrate over the Holocene closely follows nitrate concentration and accumulation rate (Figure 11B) and is even more highly correlated

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with accumulation than nitrate concentration itself, especially on multicentennial to millennial timescales ($r = 0.80$ at 512-year smoothing). Nitrate and accumulation rate are entirely independent variables in terms of their measurement, adding confidence to the annual layer counting and tie points underlying the SP19 chronology.

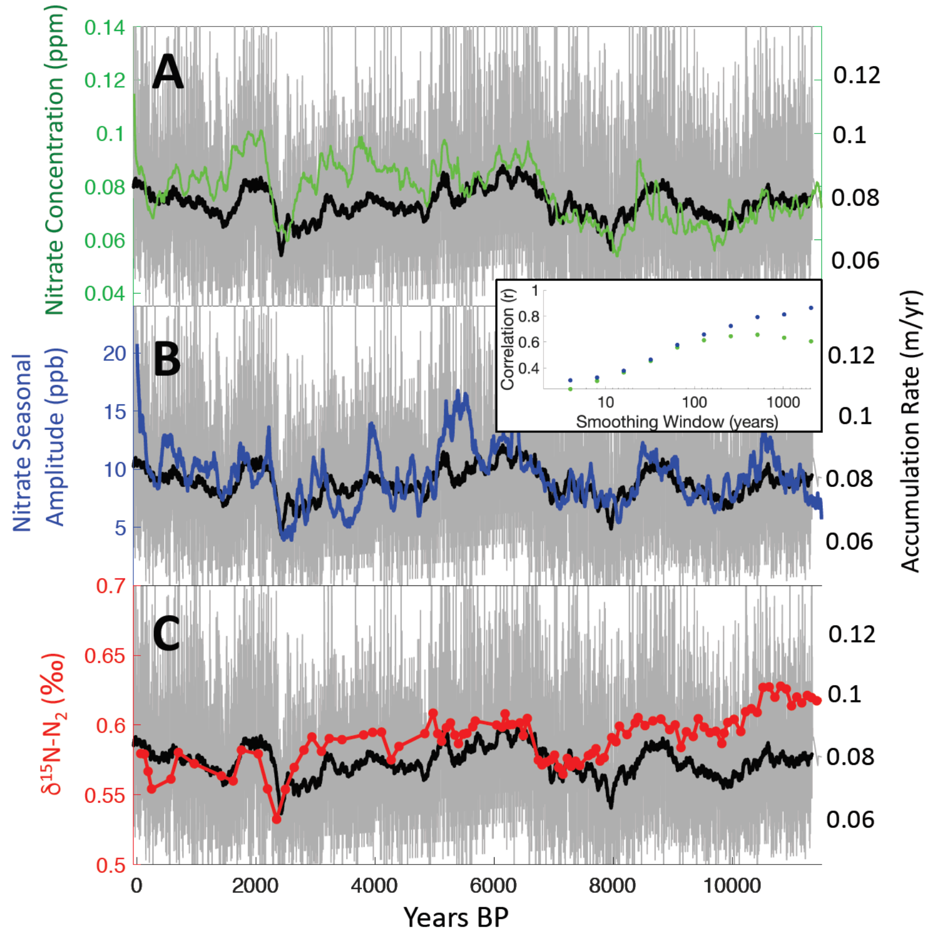

Figure 11: The Holocene accumulation rate at the South Pole compared with nitrate and δ^{15}N-N_2. In each panel, annual accumulation rates are depicted in gray, with the running 100-year mean shown in black. These results are compared with 100-year median annual values of nitrate concentration (A) and seasonal amplitude in nitrate concentration (B) as well as δ^{15}N-N_2 values (C). All three metrics exhibit shared variability on multicentennial to millennial timescales. The inset shows the correlation between accumulation rate and nitrate concentration (green) from panel A, and between accumulation rate and nitrate seasonal amplitude (blue) from panel B, against length of the smoothing window, with both exhibiting high correlations, especially at lower frequencies.

The relationship between inferred variations in accumulation rate and nitrate concentration breaks down prior to the Holocene, but a relationship between nitrate and calcium concentrations emerges. During the Pleistocene, the correlation between centennial median of calcium and nitrate is $r = 0.80$ ($p < 0.01$; Figure 12), compared with $r = 0.26$ ($p < 0.01$) during the Holocene. Rothlisberger et al. (2000, 2002) observed the same pattern at Dome C, and attributed it to the stabilization of nitrate through interaction with calcium and dust. They proposed that $CaCO_3$ and HNO_3 react to form $Ca(NO_3)_2$, which is more resistant to photolysis and consequently leads to higher concentrations of nitrate in the glacial age snowpack despite lower accumulation rates. The stabilization effect of calcium apparently overtakes photolysis and evaporation of nitrate in terms of importance only at the very high calcium concentrations as seen in the pre-Holocene ice.

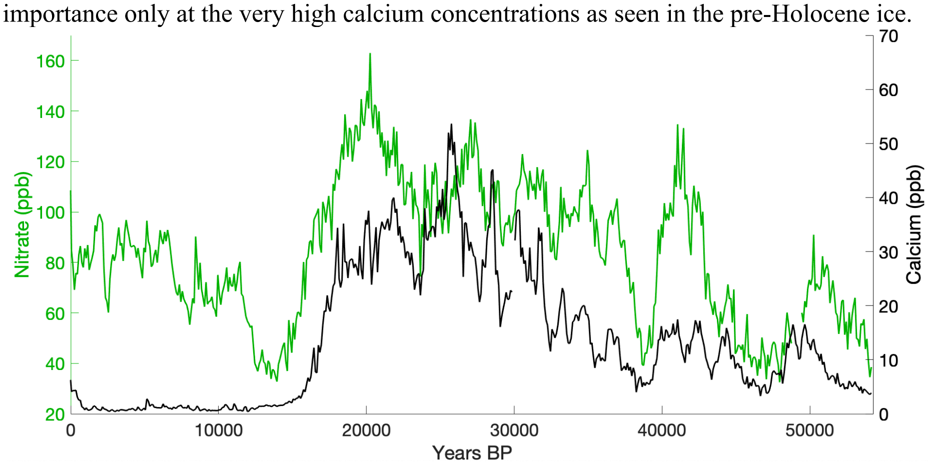

Figure 12: Nitrate and calcium concentrations in SPICEcore. There is low centennial-scale correlation ($r = 0.26$; $p < 0.01$) between calcium and nitrate ions during the Holocene, when accumulation is the dominant control on nitrate concentration (Fig. 11). During the Pleistocene, centennial median nitrate and calcium are positively correlated ($r = 0.80$; $p < 0.01$).

Stable isotope ratios of atmospheric diatomic nitrogen ($\delta^{15}N-N_2$) in trapped air in SPICEcore show a pattern similar to accumulation rate within the Holocene (Fig. 11C). $\delta^{15}N-N_2$ values were measured using the procedures described by Petrenko et al. (2006). The $\delta^{15}N-N_2$ in ice cores is driven by gravitational enrichment and is a proxy for past thickness of the firn column (Sowers et al 1992). Firn densification rates depend primarily on temperature and overburden pressure, with the second parameter closely linked to the accumulation rate at the site. Low temperatures and high accumulation rates both act to thicken the firn, thereby increasing $\delta^{15}N-N_2$ (Herron and Langway 1980, Goujon 2003).

We perform a simple attribution study to see whether $\delta^{15}N-N_2$ variations can be explained by reconstructed accumulation history or variable temperature. We compare three climatic scenarios in a dynamical version of the Herron-Langway densification model (Buizert et al. 2014). The first uses variable temperature (from $\delta^{18}O$ using a scaling ratio of 0.8‰/°C) and variable accumulation (from annual layer thickness) forcing; a second uses constant temperature (-51.5 ^{o}C) and the variable accumulation forcing; a third uses variable temperature and constant accumulation (7.8 cm/yr) forcing.

The correlations between the $\delta^{15}N\text{-}N_2$ data and each model run are displayed in Fig. 13 for both raw and detrended time series. The model scenario forced by both temperature and accumulation has the best correspondence with the $\delta^{15}N\text{-}N_2$ data ($r = 0.65$; $p < 0.01$). While secular changes in temperature appear to be driving the decreasing trend in $\delta^{15}N\text{-}N_2$, millennial-scale fluctuations in $\delta^{15}N\text{-}N_2$ appear to be driven by accumulation, supported by the high correlation ($r = 0.64$; $p < 0.01$) with the accumulation-only model run using detrended time series. In particular, a sharp drop in $\delta^{15}N\text{-}N_2$ is present at approximately 2400 BP, coincident with (and driven by) the local minimum in accumulation. These experiments provide additional confidence in the reconstructed accumulation history. To our knowledge, these data represent the best observation of accumulation-driven $\delta^{15}N\text{-}N_2$ variation, making it a valuable target for benchmarking firn densification model performance (Lundin et al. 2017).

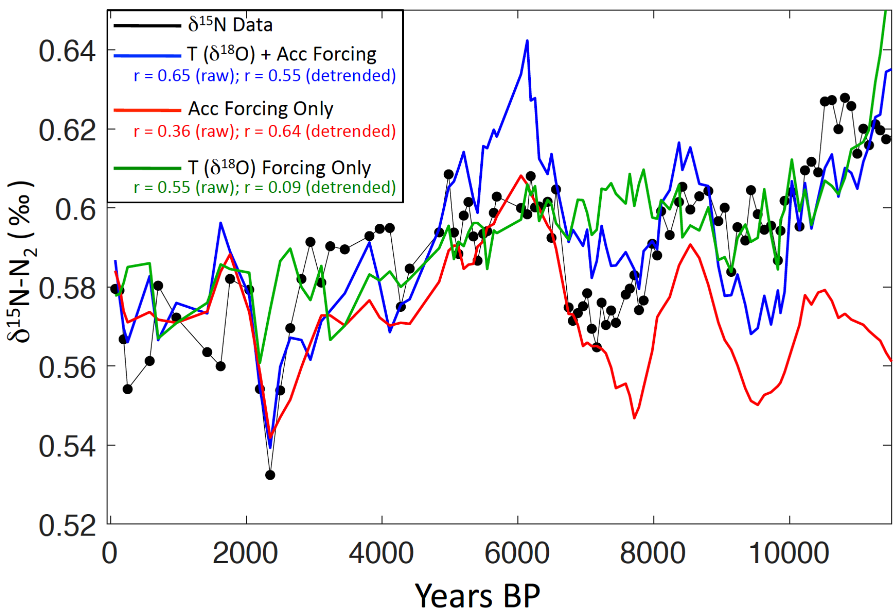

Figure 13: Results from three firn models compared with $\delta^{15}N\text{-}N_2$ variations in SPICEcore (black). The model run incorporating only $\delta^{18}O$-based temperature (green) does not capture the millennial-scale variations in $\delta^{15}N\text{-}N_2$, whereas the models using only accumulation (red) and both accumulation and $\delta^{18}O$-based temperature (blue) are able to reproduce the observed millennial-scale $\delta^{15}N\text{-}N_2$ changes. Correlations between the $\delta^{15}N\text{-}N_2$ data and the three model runs are reported in the legend with correlation coefficients calculated for both raw and linearly detrended time series.

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5. Summary

The SP19 includes the last 54,366 (-64 to 54,302 BP) years, and is the oldest and most well-constrained ice core timescale from the South Pole. SP19 was developed using

251 volcanic events that link the SPICEcore timescale with the WAIS Divide chronology WD2014 (Sigl et al. 2016; Buizert et al. 2015). High-resolution chemical records in SPICEcore during the Holocene provide the only annually resolved full-Holocene paleoclimate record in interior East Antarctica. Within the Holocene, SP19 uncertainties are in the range of +/- 18 years with respect to WAIS Divide, with the exception of the interval between 1800-3100 BP when low accumulation and sparse volcanic controls lead to uncertainties as high as +/- 25 years. During the Pleistocene, SP19 uncertainties are inversely related to the density of tie points, with maximum uncertainties reaching +/- 124 years relative to WD2014. Results show an average Holocene accumulation rate of 7.4 cm/yr with millennial-scale variations that are closely linked with advection of spatial surface-accumulation patterns upstream of the drill site. Nitrate concentrations, nitrate seasonal amplitude, and $\delta^{15}N\text{-}N_2$ variability are positively correlated with accumulation rate during the Holocene, providing independent confirmation of the SP19 chronology.

Competing Interests

The authors declare that they have no conflict of interest.

Data Availability

The SP19 chronology, associated tie points, uncertainty estimates and supporting data sets will be archived at the National Climate Data Center (www.ncdc.noaa.gov) and the U.S. Antarctic Program Data Center (http://www.usap-dc.org) with the publication of this paper.

Author Roles

All authors contributed data to this study. DW, DF, EO, JCD, ZT, KK, and NO measured the ice core chemistry. TJF and EDW collected the ECM data. JF and RA performed the visual analysis. CB, JE, EB, RB, JS, JF and TS made the gas measurements. ES, EK, TJ, and VM made the isotope measurements. DW, TJF, DF, JF and TC performed the annual layer counting. TJF and DF performed the volcanic matching. DW, TJF, DF, EO, JF and CB wrote the paper with contributions from all authors.

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Supplemental Information

S1. Annual Layer Counting Methods

Layer counting was done in two phases. The first phase spans 5 to 540 m depth and the second phase includes 540 to 798 m depth. The specific annual-layer counting procedure is described for each phase.

S1.1 Layer Counting from 5-540 meters

Above 540 m, all five interpreters independently picked annual layers using the glaciochemical time series described above. These efforts resulted in five separate timescales. Over the top 540 m of SPICEcore, the different interpreters identified between 6529 and 6807 years, with sometimes inconsistent offsets between the interpreters. These five timescales are plotted in Figure S1 with respect to a timescale passing through each of the tie points developed in section 3.2, with positive values indicating that layer counts are missing in order to synchronize ages with WAIS Divide (blue lines). To combine the 5 sets of timescales into a single unified timescale, we first identified clusters of individual layer picks. For depths above 540 m, we calculated the sum of the number of individual layer picks within a moving window +/- 2 cm wide at 0.5 cm increments. In an ideal scenario, with all 5 researchers picking a layer in the same position, reconciliation among the sets of picks is simple (Fig. 6A). However, in some areas choosing between the 5 sets of picks is non-trivial (Fig. 6B). DW and DF individually and independently reviewed the five sets of picks using independently established criteria to decide whether each cluster of picks represented a year for inclusion within the timescale or not. These decisions generated two new sets of timescales containing 6791 and 6856 years within the top 540 m (Figure S1, black dashed). While eliminating most of the discrepancies among the five interpreters, there remained a difference of 65 years (1%) and 481 specific locations in the core where DW and DF made different choices about the presence of an annual demarcation. DW made one final round of choices after investigating the chemical stratigraphy surrounding each of these years to reconcile the remaining differences into a single timescale containing 6826 years above 540 m (Figure S1, red). For comparison, roughly 6932 years would be expected between 5 and 540 m based on volcanic synchronization, indicating that our layer counting missed at least 106 years out 6932 (1.5%).

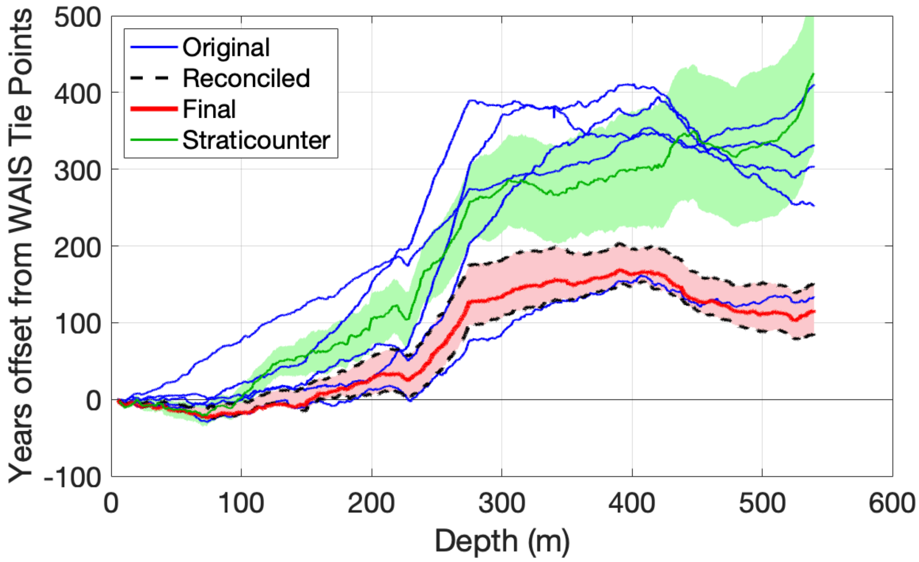

Figure S1: Offset between annual layer counting and SPICEcore-WAIS Divide tie points. Positive values on the y-axis indicate younger SPICEcore layer count ages relative to the synchronization with WAIS Divide (i.e. interpreters were missing years). Each blue line represents one of the 5 original independent sets of layer counts. The dashed black lines show the reconciled layer counted timescales after independent merging of the five original sets of picks by DW and DF. The red line indicates the final independent layer counted timescale. The green line shows the offset from SPICEcore-WAIS Divide tie points of automated layer counting using Straticounter (Winstrup et al. 2012) with the green shading inclusive of the 5th to 95th percentile ages. All of the layer counting efforts depicted were done independently of the stratigraphic tie points. All interpreters undercounted years, particularly during the interval from 228 to 275 m depth.

S1.2 540-798 meters

Between 540 and 798 m, sampling resolution permitted further annual layer counting. One interpreter (DW) continued the counting to 798 m, below which point annual layers were not consistently detectable. DW counted 4597 layers between 540 and 798 m, leading to an initial age of 11321 BP at the bottom of the annually dated section of the core.

S1.3 Sections with Missing Data

S1.3.1 Gaps and Damaged Core Within the top 798 m, where layer counting took place, there is a total of 2.74 m with missing data due to poor core quality or melter system errors. To fill these gaps in the timescale, we assigned an annual layer spacing equal to the average annual layer thickness of the 10 years above and below the gap. Layer thicknesses within the gap were rounded to make an integer number of equally spaced years. In total, 31 years were interpolated using this procedure.

S1.3.2 Dating the Top 5 meters The SPICEcore chemistry dataset begins at 5.15 m below the surface. We used chemistry from a hand-augered (HA) core drilled near the SPICEcore drill site to date the uppermost firn. The HA core was recovered from the surface to a depth of 9.86 m during the 2014/2015 field season; the same time as the beginning of SPICEcore drilling. The short length allows us to date the HA core with extra care using the following measurements: chloride, sulfate, sodium, magnesium, liquid conductivity, particle counts (small channel), Cl^-/Na^+ and SO_4^{2-}/Na^+. Layer counting of the HA core indicates a date at 5.15 m between 1992 and 1993 (5.03=1993, 5.29=1992). The nearest volcanic tie point to the top of SPICEcore is at 10.58 m depth with an age of -14 years before 1950 (1964 CE).

S1.4 Validation with Straticounter

We performed a semi-independent check on our manual layer counting ability using Straticounter, an automated layer counting software package described in (Winstrup et al. 2012). This software has been used to aid in the dating of previous ice cores in Greenland (Winstrup et al. 2012), Antarctica (Sigl et al. 2016), and Alaska (Winski et al. 2017). We used the Straticounter program to identify annual layers within the top 540 m of SPICEcore given sodium, magnesium, sulfate and microparticle data, as well as the reconciled version of our layer counts. Results produce ages ranging from 6408 to 6615 years between 5-540 m, agreeing closely with 4 of the 5 interpreters, but differing from the stratigraphically matched timescale by approximately 250-500 years (Figure S1, green). Because of the scrutiny applied to the manual layer counting efforts, and because the reconciled version of the hand-picked annual layer chronology is closer to the stratigraphically coordinated timescale, we use the manual layer counts to interpolate between tie points.

S2. Reconciling Layer Counts with Stratigraphic Ties

To a depth of 798 m, 86 volcanic tie points to WAIS Divide were identified, bracketing 85 depth intervals within which a known number of years must be present. To make the layer-counted timescale consistent with these tie points, years were added or subtracted, as necessary, within each interval such that the layer-counted timescale passes through each tie point within +/- 1 year of its age, linking SPICEcore with the WAIS Divide WD2014 chronology. In most intervals, very few years (1-5) needed to be added or subtracted, although in certain sections layer counting consistently differed from the WAIS-tied timescale. For instance, 105 years are missing between 228 and 275 m while 78 extra years were counted between 626 and 687 m. Because of the different counting methods, procedures for adding and subtracting years differ above and below 540 m and are described separately.

Above 540 m, years were preferentially added or subtracted where DF and DW disagreed in their final reconciliation (see Section S1.1). If an interval required the addition or subtraction of a number of years that exceeded the number of disagreements between DF and DW within the same interval, we first added years with 5 picks, then 4, then 3, then 2, then 1 until we met the required number of years within each interval (the opposite order applying to the subtraction of years). Between 540 and 798 m only one interpreter (DW) picked layers, so positions where years were added or subtracted were selected manually. After adding or subtracting the appropriate number of years within each interval, the layer-counted timescale passes within +/- 1 year of each stratigraphic tie point. The Holocene layer-counted timescale is then merged with the Pleistocene smoothest annual layer thickness interpolated timescale (Fudge et al. 2014) to form the final SP19 timescale.

S3. Layer Counting Performance

Using multiple interpreters to develop the timescale provides the ability to assess which areas contain better agreement among the different sets of picks. We assume that our layer count chronology is more robust in regions where all 5 sets of picks agree (e.g. Fig. 6A) than in regions with high discrepancy among picks (e.g. Fig. 6B). We create the following index of layer count quality using the following rules: For picks where all interpreters assigned a year within +/- 2 cm, we assign a value of 1. For picks where there was disagreement among the five interpreters, but agreement between DW and DF while reconciling, we assign a value of 0. For picks where there was disagreement between DW and DF while reconciling, we assign a value of -1. By calculating smoothed values of the layer count quality index over the top 540 m, patterns emerge showing areas of higher and lower layer counting confidence (Fig. S2). Most notable is the section between 412 and 456 m (5100 to 5700 BP) where analytical issues obscured robust annual signals. Fortunately, this section is well constrained by volcanic events. The very low accumulation values centered on 2400 BP are associated with another interval of slightly lower certainty in our layer counting ability. This is partly due to a lack of stratigraphic tie points, but the low accumulation here also caused all interpreters to consistently undercount years (between 228 and 275 m) leading to greater potential uncertainties.

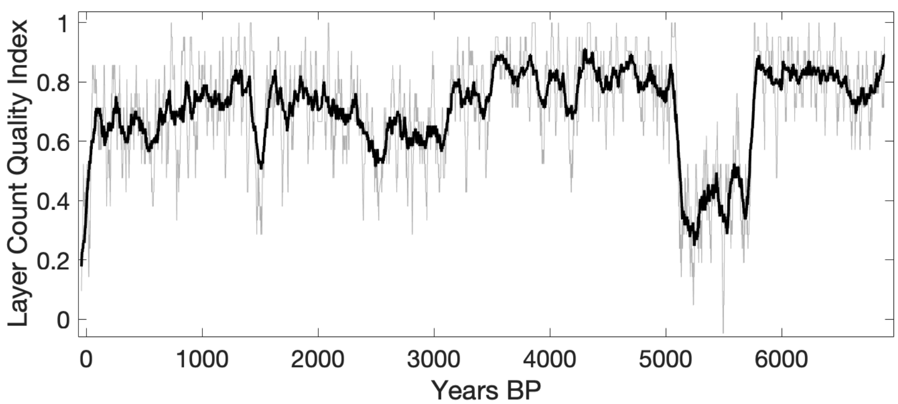

Figure S2: An index of layer counting quality over time. Higher values indicate greater confidence in layer counting ability. The index reflects the level of agreement among the five interpreters (see text). The 20-year (light gray) and 100-year (bold) running means are shown.

S4. Interval from 180 – 275 meters

We have identified only one tie point in the interval from 180m to 275m. Both interpreters (TJF, DF) have spent considerable time examining this interval but have not consistently identified the same events. The challenge of matching volcanic events in this period has two primary causes: 1) a sharp change in accumulation rate and 2) numerous small volcanic events. These two factors combine to make it difficult to distinguish sequences of volcanic eruptions because the depth between events may be varying due to the change in accumulation rather than a change in duration, and because there is no distinct pattern to relative amplitudes of the events. One interpreter revisited this interval to make another round of volcanic matches, which were made without any direct information from the annual interpolation. These matches indicate a similar timescale to the annual interpretation, which can be seen in Figure S3 of the average annual layer thickness. However, the second interpreter did not find the same matches when revisiting the interval and thus we exclude them from the underlying timescale.

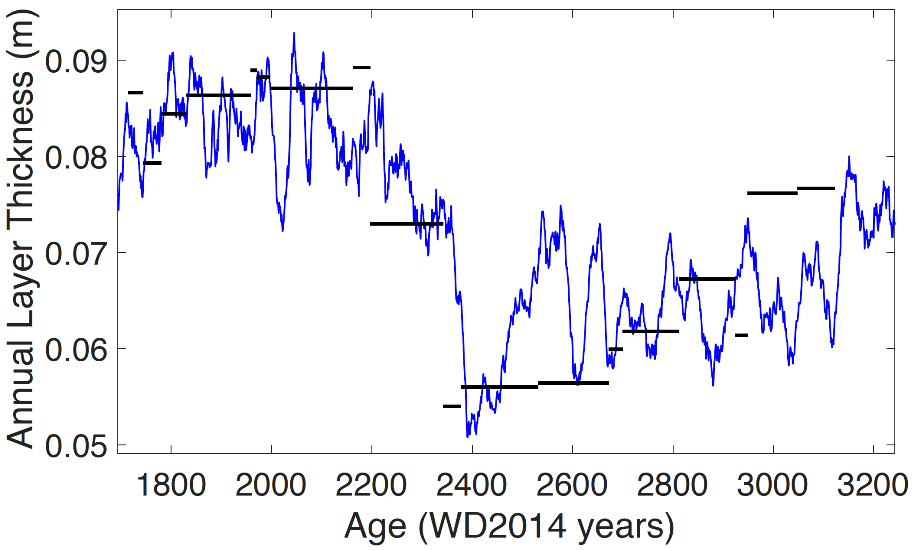

Figure S3: Annual layer thickness comparison between the SP19 timescale (blue – smoothed) and independent volcanic matches to WAIS-Divide (black). The volcanic matches were made without reference to the annual layer counts yet show a very similar pattern of annual layer thickness, further improving confidence in an accumulation anomaly at this time.

S5. Time Series Comparisons Between WAIS-Divide and SPICEcore

Below, we show the calcium time series presented in Figure 12 of this manuscript compared with an equivalent 100-year running median calcium record from WAIS-Divide. Given the broad similarity in millennial-scale calcium variability among Antarctic ice core records (e.g. Markle et al. 2018), there should be visible synchroneity in the timing of events between the two records. Figure S5, shows that both the WAIS-

Divide and SPICEcore calcium records are closely matched (r=0.96, p<0.001) with a maximum correlation with no offset between the two datasets. This provides further support that none of the tie points selected to link the two ice cores are erroneous.

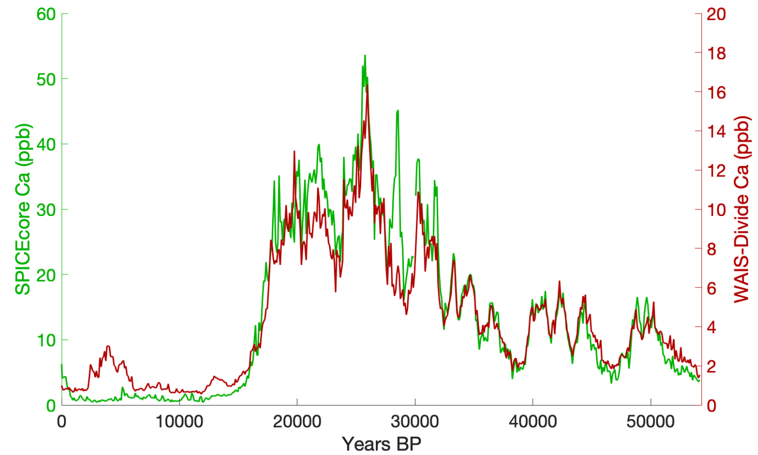

Figure S5: Calcium concentrations in SPICEcore (green) compared with those in WAIS-Divide (red). Both datasets are shown as 100-year medians. Shared events in calcium concentration among the two cores are closely synchronized, supporting choices of tie points between SP19 and WD2014. SPICEcore calcium data are preliminary.