



## Early Eocene vigorous ocean overturning and its contribution to a warm Southern Ocean

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1 **Abstract.** The early Eocene (~55 Ma) is the warmest period, and most likely characterized by the highest atmospheric  
2 CO<sub>2</sub> concentrations, of the Cenozoic era. Here, we analyze simulations of the early Eocene performed with the IPSL-  
3 CM5A2 coupled climate model set up with paleogeographic reconstructions of this period from the DeepMIP project,  
4 with different levels of atmospheric CO<sub>2</sub>, and compare them with simulations of the modern conditions. This allows  
5 us to explore the changes of the ocean circulation and the resulting ocean meridional heat transport. At a CO<sub>2</sub> level of  
6 840 ppm, the Early Eocene simulation is characterized by a strong abyssal overturning circulation in the Southern  
7 Hemisphere (40 Sv at 60°S), fed by deep water formation in the three sectors of the Southern Ocean. Deep convection  
8 in the Southern Ocean is favored by the closed Drake and Tasmanian passages, which provide western boundaries for  
9 the build-up of strong subpolar gyres in the Weddell and Ross seas, in the middle of which convection develops. The  
10 strong overturning circulation, associated with the subpolar gyres, sustains the poleward advection of saline  
11 subtropical water to the convective region in the Southern Ocean, maintaining deep-water formation. This salt-  
12 advection feedback mechanism works similarly in the present-day North Atlantic overturning circulation. The strong  
13 abyssal overturning circulation in the 55 Ma simulations primarily results in an enhanced poleward ocean heat  
14 transport by 0.3–0.7 PW in the Southern Hemisphere compared to modern conditions, reaching 1.7 PW southward at  
15 20°S, and contributing to maintain the Southern Ocean and Antarctica warm in the Eocene. Simulations with different  
16 atmospheric CO<sub>2</sub> levels show that the ocean circulation and heat transport are relatively insensitive to CO<sub>2</sub>-doubling.  
17 **Keywords:** Early Eocene, overturning circulation, deep water formation, oceanic heat transport.



## 18 1. Introduction

19 Proxy-based temperature reconstructions suggest that the early Eocene (55–50 Ma) was one of the warmest intervals  
20 in the geological history and the warmest of the Cenozoic (Zachos et al., 2001; Cramer et al., 2011; Dunkley Jones et  
21 al., 2013). More specifically, the EECO (Early Eocene Climatic Optimum) occurred 51–53 Ma, but shorter (less than  
22 tens of thousands of years) hyperthermal events such as the PETM (Paleocene-Eocene Thermal Maximum) about 55  
23 Ma ago (Zachos et al., 2008) also occurred. The Southern Hemisphere was particularly warm at that time, as shown by  
24 inferred surface ocean temperatures exceeding 20°C at high-latitudes (e.g. Evans et al., 2018 and references therein),  
25 and by the absence of a perennial ice over Antarctica until the onset of the Antarctic ice sheet at the Eocene-Oligocene  
26 Boundary, ~34 Ma when CO<sub>2</sub> abruptly declined below a certain threshold (Galeotti et al., 2016; Gasson et al., 2014;  
27 Ladant et al., 2014). In the early Eocene, high levels of CO<sub>2</sub> in the atmosphere are undoubtedly a critical contributor  
28 to the extremely warm climate, with the global temperature increasing by more than 5°C in less than 10 000 years  
29 (Zachos et al., 2001, 2008; Huber and Caballero, 2011; Anagnostou et al., 2016), but they do not fully explain the  
30 extreme warmth at high-latitudes and the reduced equator-to-pole temperature gradient (Huber and Caballero, 2011).

31 In addition to the much higher levels of CO<sub>2</sub> in the atmosphere, one of the main differences between the early Eocene  
32 and our modern climate lies in the very contrasted bathymetries and continental configuration, likely resulting in very  
33 different ocean circulation (Thomas et al., 2003; Voigt et al., 2013; Winguth et al., 2012; Zachos et al., 2001). In  
34 particular, the opening or closing of major oceanic gateways (such as the Drake Passage or the Panama Seaway) during  
35 the Late Paleogene and Neogene have been shown to exert a strong influence on the ocean circulation and its  
36 associated heat transport (England et al., 2017; Ladant et al., 2018; Nong et al., 2000; Sijp and England, 2004;  
37 Toggweiler and Bjornsson, 2000; Yang et al., 2014). Additionally, proxy-based reconstructions and results from  
38 Eocene model simulations suggest that the Meridional Overturning Circulation (MOC) was also very distinct from  
39 our present day MOC, with no evidence for deep water formation in the North Atlantic until the early Oligocene  
40 (Ferreira et al., 2018). Instead, formation of deep water was found to happen in the North Pacific (Hutchinson et al.,  
41 2018; Winguth et al., 2012) or only in the Southern Ocean (Sijp et al., 2014). Different ocean circulation resulting  
42 from different bathymetry are expected to result in different Ocean Heat Transport (OHT). For instance, Sijp and  
43 England (2004) found a 0.5 PW decrease in OHT in the Southern Hemisphere in response to the opening of the Drake  
44 Passage, but other factors such as different radiative forcing induced by different level of CO<sub>2</sub> in the atmosphere may  
45 also contribute to the different OHT (see Huber, 2012 for a review).

46 In our present climate, the ocean is an important actor for the Earth energy balance, as it contributes about one third  
47 to the total redistribution of heat from the Equator to the Poles (e.g. Trenberth and Caron, 2001). Although  
48 modifications of both the atmosphere and the Ocean Heat Transport (OHT) tend to compensate (Trenberth and Caron,  
49 2001), subtle changes in OHT could trigger large changes in atmospheric extratropical convection, modifying the  
50 water vapor greenhouse (Rose and Ferreira, 2013), which in turn can affect surface temperature. The OHT itself results  
51 from different contributions, and several attempts have been made in the literature to disentangle the relative roles of  
52 the horizontal and overturning ocean circulations for the meridional OHT (Ganachaud and Wunsch, 2003). In the  
53 North Atlantic, where the strong Atlantic Meridional Overturning Circulation (AMOC) is fed by the formation of



54 dense water by convection at high latitudes, the AMOC contributes up to 90% of the meridional OHT at 26.5°N  
55 (Msadek et al., 2013), where the RAPID monitoring array is located (McCarthy et al., 2015). Based on hosing  
56 experiments performed with a climate model, Yang et al. (2013) found that the meridional OHT decreases rapidly in  
57 response to an artificial shutdown of the AMOC, although Drijfhout and Hazeleger (2006) suggested that on longer  
58 (decadal) timescale, the OHT might recover its initial level as the gyre contribution tends to compensate the decrease  
59 in OHT associated with the AMOC shutdown. In contrast to the North Atlantic, Volkov et al. (2010) found that, in  
60 the Southern Ocean, the OHT results roughly equally from the gyre and overturning contributions.

61 Given the importance of both horizontal and overturning ocean circulations for the meridional OHT, one can expect  
62 that the different MOC and horizontal gyre constrained by the Eocene bathymetry would result in different  
63 contributions to the meridional OHT, potentially contributing to the warm climate at 55 Ma, and in particular to the  
64 much warmer temperatures found at high-latitudes. Based on the analysis of early Eocene and modern simulations  
65 performed with the IPSL climate model, the goal of this study is to better understand what sets the ocean circulation  
66 during the Eocene and the importance of the ocean circulation for the poleward heat transport. The model and  
67 simulations used are described in Section 2. In Sections 3 and 4, we examine the MOC and the horizontal circulation,  
68 respectively, for the present-day and 55 Ma configurations, and these circulations are then linked to the OHT in Section  
69 5. The sensitivity of the ocean circulation and heat transport to the level of CO<sub>2</sub> in the atmosphere is discussed in  
70 section 6. A summary and conclusions are given in Section 7.

## 71 **2. Numerical model data**

### 72 **2.1. Model setup and simulations**

73 The simulations used in this study are performed with the IPSL-CM5A2 earth system model (Sepulchre et al., 2019).  
74 The oceanic component of IPSL-CM5A2 is NEMOv3.6 (Madec and the NEMO team, 2016), which includes the LIM2  
75 sea ice model (Fichefet and Maqueda, 1997) and the PISCES biogeochemical model (Aumont et al., 2015). The  
76 atmospheric component is the LMDz model (Hourdin et al., 2013), which is coupled to the land surface model  
77 ORCHIDEE (Krinner et al., 2005). Here, we use IPSL-CM5A2 in its standard resolution. NEMO is thus run at a  
78 nominal resolution of 2°, increased down to 0.5° at the equator, with 31 levels that vary in thickness with depth. LMDZ  
79 is run at a horizontal resolution of 3.75° longitude×1.875° latitude and 39 vertical levels. A full description and  
80 evaluation of the IPSL-CM5A2 model can be found in Sepulchre et al. (2019).

81 Two sets of simulations are performed. The first set is composed of the reference PI simulation (referred to as PI-1x  
82 and described in Sepulchre et al., 2019) and another PI simulation in which the atmospheric CO<sub>2</sub> concentration is  
83 doubled (PI-2x). The second set consists in a baseline simulation of the Early Eocene (hereafter 55 Ma-3x), using a  
84 setup following the DeepMIP protocol described by Lunt et al. (2017), and two sensitivity experiments to CO<sub>2</sub> (55  
85 Ma-1.5x) and tidal mixing (55 Ma-noM2). In the following, we first briefly describe the baseline Early Eocene  
86 simulation (as the DeepMIP guidelines from Lunt et al. (2017) give several options for implementing the Early Eocene  
87 boundary conditions), and then precise the boundary condition differences of the two other simulations.



88 2.1.1. General considerations

89 Most of the boundary conditions that were adapted for these IPSL-CM5A2 Early Eocene simulations are described in  
90 Herold et al. (2014, hereafter H14). Following Lunt et al. (2017), the solar constant and orbital parameters are kept  
91 to PI values, so are greenhouse gas concentrations with the exception of atmospheric CO<sub>2</sub>. The latter is set to 840 ppm  
92 (3x the PI value), so that the simulation is representative of the pre-PETM (following the terminology of Lunt et al.,  
93 2017).

94 2.1.2. Oceanic boundary conditions

95 The ocean component (NEMO) is commonly run on a tripolar mesh grid (Madec and the NEMO team, 2016), which  
96 avoids singularity points in the ocean domain. Because the implementation of the Eocene land-sea mask on the  
97 ORCA2 grid would have shifted the singularity points into the ocean domain, we have constructed a new PALEORCA  
98 grid, which is suitable to run paleo-simulations with IPSL-CM5A2 (see more details in Sepulchre et al., 2019). The  
99 bathymetry is obtained by masking out the H14 topography and remapping the resulting bathymetry onto the  
100 PALEORCA grid using near-neighbor interpolation. This type of interpolation indeed ensures that small but crucial  
101 features of the H14 dataset such as islands and seaways, which may strongly impact the modeled ocean circulation,  
102 remain present in the interpolated bathymetry file. Handmade corrections were then applied at some locations (e.g. in  
103 the West African region) to retain oceanic straits that are sufficiently large to allow for exchanges. The Early Eocene  
104 bathymetry is shown on Fig. 1.

105 Modern boundary conditions of NEMO also include forcings of the dissipation associated with internal wave energy  
106 from the M2 and K1 tidal components (de Lavergne et al., 2019). The parameterization follows Simmons et al. (2004),  
107 with refinements in the modern Indonesian Through Flow (ITF) region according to Koch-Larrouy et al. (2007). To  
108 create an Early Eocene tidal dissipation forcing, we directly interpolate the H14 M2 tidal field (obtained from the tidal  
109 model simulations of Green and Huber 2013) onto the NEMO grid using bilinear interpolation. In the absence of any  
110 estimation of the K1 tidal component for the Early Eocene, we ignore this contribution. In addition, the  
111 parameterization of Koch-Larrouy et al. (2007) is not used here because the ITF does not exist in the Early Eocene.

112 The geothermal heating distribution  $q$  is created from the 55 Ma global crustal age distribution of Müller et al. (2008),  
113 on which the age-heatflow relationship of the Stein and Stein (1992) model is applied:

$$114 \quad q(t) = 510 t^{-1/2} \quad t \leq 55 \text{ Ma}$$

$$115 \quad q(t) = 48 + 96 e^{-0.0278 t} \quad t > 55 \text{ Ma}$$

116 In regions of subducted seafloor where age information is not available, we prescribe the minimal heatflow value  
117 derived from known crustal age. The 1°x1° resulting field is then bilinearly interpolated on the NEMO grid. It must  
118 be noted that the Stein and Stein parameterization becomes singular for young crustal ages, which yields unrealistically  
119 large heatflow values. We thus set an upper limit of 400 mW.m<sup>-2</sup> on heatflow values, following Emile-Geay and  
120 Madec (2009).

121 Salinity is initialized as globally constant to a value of 34.7 psu following Lunt et al. (2017). The initialization of the  
122 model with the proposed DeepMIP temperature distribution (Lunt et al., 2017) led to severe instabilities of the model  
123 during the spin-up phase. The initial temperature distribution has thus been modified to follow:



124  $T(^{\circ}\text{C}) = \frac{(1000-z)}{1000} 25 \cos(\varphi) + 10 \quad z \leq 1000 \text{ m}$

125  $T(^{\circ}\text{C}) = 10 \quad z > 1000 \text{ m}$

126 with  $\varphi$  the latitude and  $z$  the depth of the ocean. This new equation gives an initial globally constant temperature of  
127  $10^{\circ}\text{C}$  below 1000 m and a zonally symmetric distribution above, reaching surface values of  $35^{\circ}\text{C}$  at the equator and  
128  $10^{\circ}\text{C}$  at the poles. This corresponds to a  $5^{\circ}\text{C}$  surface temperature reduction compared to the DeepMIP equation (Lunt  
129 et al., 2017). No sea ice is prescribed at the beginning of the simulations.

130 The IPSL-CM5A2 model includes PISCES biogeochemical model. Biogeochemical cycles and marine biology are  
131 directly forced by dynamical variables of the physical ocean and may affect the ocean physics via their influence on  
132 chlorophyll production, which modulates light penetration in the ocean. However, because this feedback does not  
133 affect the ocean state significantly (Kageyama et al., 2013) and because the Early Eocene mean ocean color is  
134 unknown, we have prescribed a constant chlorophyll value to  $0.05 \text{ g.Chl/L}$  for the computation of light penetration in  
135 the ocean. As a consequence, marine biogeochemical cycles and biology do not alter the dynamics of the ocean and  
136 as such, biogeochemical initial forcings have been kept to modern.

### 137 2.1.3. Continental boundary conditions

138 The atmospheric (LMDZ) and land surface (ORCHIDEE) models run on a low resolution grid but require input  
139 forcings at higher resolution. The topographic field is created by masking out ocean points in the H14 reference file  
140 and upscaling the  $1^{\circ}\times 1^{\circ}$  masked H14 file to the required LMDZ input topographic resolution ( $1/6^{\circ}$ ), as LMDZ includes  
141 a subgrid scale orographic drag parameterization requiring high-resolution surface orography (Lott and Miller, 1997;  
142 Lott, 1999). A similar procedure is applied to the standard deviation of orography proposed by H14.

143 Following Lunt et al. (2017), the soil properties are prescribed as globally constant to the global mean of the PI  
144 simulation. There is no lake module in this version of IPSL-CM5A2. The river routing proposed by H14 is passed to  
145 ORCHIDEE at its original resolution of  $1^{\circ}\times 1^{\circ}$ , which ensures an appropriate downscaling to the model resolution.  
146 The vegetation cover is prescribed from the BIOME4 reconstruction of H14, using a lookup table (given in Table S2)  
147 to convert the 10 megabiomes into ORCHIDEE Plant Functional Types (PFTs). Aerosol distributions are left identical  
148 to PI values.

### 149 2.1.4. Sensitivity experiments and equilibrium

150 We perform two additional Early Eocene experiments. One has the same boundary conditions as the baseline Early  
151 Eocene experiment (55 Ma-3x) but an atmospheric  $\text{CO}_2$  concentration of 420 ppm (1.5x the PI levels, 55 Ma-1.5x).  
152 The other differs from the baseline Early Eocene experiment by the absence of tidal dissipation forcing (55 Ma-3x-  
153 noM2).

154 The 55 Ma-3x simulation is initialized from rest and run for 4000 years (Fig. S1). The 55 Ma-1.5x simulation is  
155 branched from 55 Ma-3x at year 1500 and run for 4000 years. The 55 Ma-3x-noM2 is branched from 55 Ma-3x at  
156 year 3000 and run for 2000 additional years. The two PI simulations are initialized from the Levitus climatology  
157 (Boyer et al., 2005), and run for more than 2700 years (Table 1). At the end of all the simulations, the ocean has



158 reached a quasi-equilibrated state and trends in deep ocean temperatures over the final 1000 years of all simulations  
159 are smaller than  $0.05^{\circ}\text{C}/\text{century}$ .

160 We use the monthly outputs of the last 100 years of each simulation to create a climatological year for each simulation.  
161 In the following, we will mostly focus on the comparison between the baseline Early Eocene simulation (55 Ma-3x)  
162 and the PI control simulation (PI-1x). The other simulations are analyzed in Section 3 to estimate the contribution of  
163 tidal mixing to the oceanic overturning circulation, and in Section 6 to examine the sensitivity of the ocean conditions  
164 to different levels of  $\text{CO}_2$  in the atmosphere.

## 165 **2.2. Evaluation of the simulated ocean temperature**

166 The mean state and the seasonal variation of ocean temperature in the 55 Ma simulations are examined. Further, we  
167 evaluate the ability of IPSL-CM5A2 to reasonably simulate the early Eocene sea surface temperature (SST).

168 The annual mean SST in the 55 Ma-3x simulation varies from  $10\text{--}15^{\circ}\text{C}$  in the Southern Ocean, to  $37.2^{\circ}\text{C}$  near the  
169 Equator (Fig. 2A), with a global mean of  $27.5^{\circ}\text{C}$ . During summer (defined as July-August-September for the Northern  
170 Hemisphere and January-February-March for the Southern Hemisphere), the simulated SST reaches  $\sim 20^{\circ}\text{C}$  over most  
171 of Southern Ocean (south of  $60^{\circ}\text{S}$ ), and up to  $38^{\circ}\text{C}$  in parts of equatorial Indian and Atlantic oceans (Fig. 2A). In the  
172 55 Ma-1.5x simulation, both SST and global mean temperature are  $\sim 5^{\circ}\text{C}$  lower than in the 55 Ma-3x simulation (Table  
173 1).

174 These simulated SSTs are further compared with proxy-based SST estimates for the early Eocene provided by a recent  
175 data compilation performed within the DeepMIP framework (Hollis et al., 2019). The dataset includes 32 records in  
176 total, from 4 proxy types ( $\text{TEX}^{86}$ ,  $\delta^{18}\text{O}$ , Mg/Ca and Clumped Isotope data). The spatial pattern of the model SST is  
177 overall consistent with the proxy based SST, although significant differences can be seen for some specific proxy data  
178 point (Fig. 2A and S2A). More details on the model-proxy comparison can be found in Supplementary Material.

179 In order to further compare the simulations with the proxy-based reconstructions, we also calculate the root-mean-  
180 square deviations (RMSD) between the simulated SST and the reconstructions (Table 2). Although large, the RMSD  
181 values are overall of the same order of magnitude as the uncertainty of proxy-based SST estimates, suggesting a  
182 reasonable model-data consistency. More importantly, the RMSD values are smaller for the 55 Ma-3x simulation than  
183 for the 55 Ma-1.5x simulation, suggesting that the 55 Ma-3x simulation captures better the signal of proxy-based SST  
184 reconstructions. This is also consistent with proxy reconstructions suggesting that the  $\text{CO}_2$  atmospheric concentrations  
185 during the early Eocene were most likely three to four times the PI level (Foster et al., 2017).

186 The zonal mean SSTs in the 55 Ma-3x simulation range from  $30^{\circ}\text{C}$  to  $37^{\circ}\text{C}$  in the tropics and decreases toward the  
187 high latitudes (Fig. 2B). Within the  $40^{\circ}\text{S}\text{--}40^{\circ}\text{N}$  latitudinal band, the summer SSTs remain above  $30^{\circ}\text{C}$ . Around  $60\text{--}$   
188  $70^{\circ}\text{S}$  the annual mean SSTs are  $\sim 13^{\circ}\text{C}$  with a seasonal amplitude of 10 to  $15^{\circ}\text{C}$ . Those zonal mean SSTs are overall  
189  $\sim 10^{\circ}\text{C}$  warmer than in PI-1x simulation, with the largest differences of  $12^{\circ}\text{C}$  found in the Southern Ocean (Fig. 2B).  
190 The warm SSTs found in the Southern Ocean in the 55 Ma simulations also extend at depth (Fig. 2C), with a mean  
191 global temperature of  $11.3^{\circ}\text{C}$  in the 55 Ma-3x run, compared to  $3.3^{\circ}\text{C}$  in PI-1x (Table 1). The very warm temperatures  
192 found at depth are compatible with several proxy-based temperature estimates (bottom temperatures at 1000-5000m



193 between 10 and 15°C, e.g. Huber et al., 2000, their Fig. 4-1). More specifically, Mg/Ca-based temperature estimates  
194 suggest that the bottom-water (below 1000 m) temperatures were around 15°C during the early Eocene (Cramer et al.  
195 2011), and Dunkley Jones et al. (2013) found a similar value based on  $\delta^{18}\text{O}$ -Mg/Ca thermometry data for the PETM  
196 time-window.

### 197 **3. The overturning circulation**

198 Here we describe the simulated MOC in the different simulations, and investigate the links between the MOC and  
199 deep-water formation.

#### 200 **3.1. Meridional overturning circulation**

201 The global MOC is represented through the vertical streamfunction  $\psi$  computed from the zonally-integrated  
202 meridional volume transport as:

$$203 \quad \psi(z) = \int_W^E \int_{-H}^0 v(x, z) dz dx \quad (1)$$

204 where  $v$  is the meridional velocity,  $z$  is the vertical coordinate,  $H$  is the ocean depth,  $x$  is the zonal coordinate  
205 integrated from West (W) to East (E) boundary.

206 In the 55 Ma-3x simulation, a single clockwise inter-hemispheric MOC cell fills the whole deep ocean, with a  
207 maximum of 40 Sv at 1500 m depth and 60°S (Fig. 3). This strong MOC cell (referred to as SOMOC for Southern  
208 Ocean MOC) is associated with the formation of (Eocene) Antarctic Bottom Water (AABW) in the Southern Ocean,  
209 flowing northward below about 2000 m. The SOMOC is associated with an upwelling branch extending over the  
210 whole Northern Hemisphere, with almost 26 Sv crossing the equator northward (22 Sv at 20°N and only ~5 Sv at  
211 60°N). There is no deep-water formation in the Northern Hemisphere, neither in the North Atlantic nor in the North  
212 Pacific.

213 In contrast to the 55 Ma-3x run, the MOC in the PI-1x simulation is composed of the traditional upper and lower cells.  
214 The upper cell is clockwise and associated with the AMOC, with a maximum strength of 11–12 Sv reached at 800 m  
215 depth around 50–60°N. This cell is fed by the formation of North Atlantic Deep Water (NADW) around 60°N. NADW  
216 is transported southward all the way to the Southern Ocean at depth between 1000 and 3000 m, where it is brought  
217 back to the surface through wind-induced upwelling, forming the quasi-adiabatic pole-to-pole overturning circulation  
218 regime (Marshall and Speer, 2012; Wolfe and Cessi, 2014). The lower cell is anticlockwise, with a maximum strength  
219 of 15–16 Sv at 3000 m depth. This anticlockwise circulation is fed by the Antarctic Bottom Water formed by surface  
220 buoyancy loss related to ocean-sea-ice interaction (Abernathy et al., 2016), and consumed through mixing process  
221 induced by breaking of internal wave at the sea floor and geothermal heating (Nikurashin and Ferrari, 2013; de  
222 Lavergne et al., 2016).

223 The 55 Ma simulated deep-water formation in the Southern Ocean is compatible with currently available proxy-based  
224 reconstructions of the MOC. Although proxy data constraining the ocean circulation are very limited for the early





225 Eocene, these data seem to point to a common deep-water source around the Austral Ocean during that period (Abbott  
226 et al., 2016; Batenburg et al., 2018; Frank, 2002; Thomas et al., 2003, 2014). For instance, Batenburg et al. (2018)  
227 revealed a convergence of Nd-isotopes signature across the Atlantic basin from 59 Ma onward, likely resulting from  
228 an intensification of the intermediate and deep ocean circulation in the Atlantic Ocean, with the dominant deep-water  
229 masses originating from the high southern latitudes. In addition, the reconstructed deep-sea carbonate ion  
230 concentration ( $[\text{CO}_3^{2-}]$ ) during the Eocene shows a reversed inter-basin gradient compared to the present-day,  
231 suggesting a reverse ocean circulation at depth compared to the present-day circulation (Zeebe and Zachos, 2007).  
232 Therefore, these proxy data and the simulated ocean circulation are compatible, at least on the direction of the deep  
233 circulation and the source of the deep water masses.

### 234 3.2. Convection and deep-water formation

235 The abyssal circulation described in Section 3.1 is fed by deep convection processes that mainly occur in winter. We  
236 examine the simulated Mixed Layer Depth (MLD) at the end of the winter season (Fig. 4), which is an efficient  
237 indicator of convection. In the 55 Ma-3x simulation, deep convection occurs only at high-latitude in the Southern  
238 Hemisphere, with maximum MLD reaching up to 4000 m in the Weddell Sea, and large areas of MLDs deeper than  
239 2000 m around Antarctica in the Ross and Amundsen seas (Fig. 4). In contrast, MLDs remain shallow in the Northern  
240 Hemisphere, suggesting the absence of any deep convection there. In this hemisphere, the maximum MLDs are found  
241 in the North Pacific (350 m) over the poleward western boundary current between 35–50°N, and in the North Atlantic  
242 between 35–40°N (300 m), but the deepest MLD at high latitude are only 200 m over the northwest Pacific. Note that  
243 the early Eocene North Atlantic basin is limited to a narrow region west of Greenland and poleward of 52°N. In the  
244 absence of other sources of deep waters, waters sourced in the Southern Ocean around Antarctica fill the whole abyssal  
245 ocean (Fig. 3).

246 In contrast, deep-water formation occurs both in the northern North Atlantic and in the Weddell Sea in the PI-1x  
247 simulation (Fig. 4, right column). The deepest MLD (up to 1500 m) are found in the Nordic Seas between 70–75°N  
248 and just south of Iceland around 60°N. MLD larger than 1200 m are also found in the eastern Weddell Sea around  
249 65°S, indicative of the formation of AABW. This pattern is consistent with the MLD simulated by the CMIP5  
250 ensemble for modern conditions (Heuzé et al., 2015).

251 In 55 Ma-3x, the permanent absence of deep convection in the North Pacific is at odds with a few previous model  
252 studies and proxy-based reconstruction of the ocean circulation of the early Eocene that have suggested that deep  
253 water could form in the North Pacific (Lunt et al., 2010; Abbott et al., 2016). The reasons why deep water forms in  
254 the Southern Ocean in the 55 Ma-3x simulation and not in the North Pacific can be linked to the different ocean  
255 thermo-dynamical properties of these two regions. Table 3 summarizes key ocean surface variables in the convective  
256 regions of the Weddell Sea, where the deepest MLD are found, and of the North Pacific. These two regions roughly  
257 correspond to closed Sea Surface Height (SSH) contours around deep convection regions, which are associated with  
258 a minimum of SSH— the geostrophic ocean surface currents circulate cyclonically along the SSH contours, so that  
259 the interior regions remain largely isolated from surrounding waters. The Weddell Sea is characterized by a strongly





260 reduced vertical stratification compared to the North Pacific, the surface signature of which being a much larger  
261 surface density ( $+0.51 \text{ kg/m}^3$ ), which provides favorable conditions for the emergence of deep convection. According  
262 to the equation of state of seawater, the larger surface density found in the Weddell Sea convection region is due to  
263 higher salinity ( $+0.73 \text{ psu}$ ) contributing to a density increase of  $0.57 \text{ kg/m}^3$ , partly balanced by warmer temperature  
264 ( $+0.36^\circ\text{C}$ ) contributing to a  $0.06 \text{ kg/m}^3$  density decrease.

265 Two major causes may explain these large differences in salinity between the two regions: (i) the atmospheric  
266 circulation and freshwater fluxes to the ocean and (ii) the ocean circulation and positive salt-advection feedback. The  
267 surface freshwater budget for two extended regions over the Weddell Sea ( $78^\circ\text{S}$ – $61^\circ\text{S}$  and  $62^\circ\text{W}$ – $8^\circ\text{E}$ ) and the North  
268 Pacific ( $48^\circ\text{N}$ – $67^\circ\text{N}$  and  $124^\circ\text{E}$ – $143^\circ\text{W}$ ) is shown in Table 3. Averaged precipitation and runoff over the box in the  
269 North Pacific (respectively  $1.48$  and  $0.37 \text{ m/yr}$ ) exceed those in the Weddell Sea ( $1.01$  and  $0.22 \text{ m/yr}$ ) by roughly 50%,  
270 whereas the evaporation rates are almost similar ( $0.75$  vs  $0.71 \text{ m/yr}$ ). Overall, the average net surface freshwater input  
271 is  $0.53 \text{ m/yr}$  in the Weddell Sea compared to  $1.11 \text{ m/yr}$  in the North Pacific. Reduced freshwater input by precipitations  
272 and continental runoff is related to different atmospheric circulation in the Northern and Southern Hemispheres. In  
273 the Southern Ocean, winds largely follow the Antarctic orography (as shown by geopotential height at  $850 \text{ hPa}$ , Fig.  
274 S3), and induce almost no precipitation by orographic uplift over Antarctica coastal regions and no runoff to the  
275 Southern Ocean. In contrast, in the North Pacific, the westerlies are blocked by the paleo-Rocky Mountains, especially  
276 in the high-altitude region between  $50^\circ\text{N}$  and  $70^\circ\text{N}$ . The orographic uplift of moist air masses induces high  
277 precipitation (up to  $2$ – $3 \text{ m/yr}$ ) and runoff into the North Pacific (as found in several other models, Carmichael et al.,  
278 2016), leading to low sea surface salinity (below  $30 \text{ psu}$ ) along the Pacific coast of North America (not shown) hence  
279 increased surface stratification. The upper branch of the MOC and the associated poleward advection of saline  
280 subtropical waters constitutes the other contribution to the larger salinities found in the Southern Ocean relative to the  
281 North Pacific. This process was dubbed as a positive salt-advection feedback by Ferreira et al. (2018)

282 In contrast, present-day circulation, characterized by deep-water formation in the North Atlantic, is maintained by  
283 higher salinities in the North Atlantic than in the North Pacific, which are partly sustained by atmospheric fluxes and  
284 the salt-advection feedback (Ferreira et al., 2018). A recent sensitivity study of the impact of topography on modern  
285 ocean circulation reveals that the presence of the Rocky Mountains influences the global salinity pattern and the  
286 regions where deep convection occurs, through the adjustment of the freshwater transfers from the Pacific to the  
287 Atlantic Oceans (Maffre et al., 2018).

### 288 3.3. Factors contributing to the vigorous SOMOC

289 Different factors contribute to the intense SOMOC simulated by the 55 Ma-3x simulation ( $40 \text{ Sv}$ ) in comparison with  
290 typical present-day MOC, the intensity of which reaches around  $18 \text{ Sv}$  and  $20 \text{ Sv}$  for the upper and lower cells  
291 respectively (Lumpkin and Speer, 2007). Deep-water formation occurs in the three sectors of the Southern Ocean  
292 (Pacific, Atlantic and Indian) but the zonal connections between the different basins hamper a clear quantification of  
293 the contribution of each deep-water formation sector to the SOMOC. However, the contributions of the Weddell Sea  
294 and the Pacific sector of the Southern Ocean can be estimated because the narrow width of Drake and Tasman passages



295 at 55 Ma (Fig. 1) creates latitudinal continental boundaries on western and eastern sides of these regions. The Weddell  
296 Sea, which exhibits the largest MLD in 55 Ma-3x, and the Pacific sector contribute roughly equally to the SOMOC  
297 intensity ( $\sim 19$  Sv). The South Indian sector contribution is more difficult to assess directly because of the large open-  
298 ocean zonal connection with the Atlantic sector.

299 The shallow Drake Passage at 55 Ma provides a western boundary for the development of a subpolar gyre in the  
300 Weddell Sea (see Section 4 for more details). This clockwise gyre produces a favorable environment to trigger deep-  
301 water formation (known as preconditioning) through isopycnals doming in the center of the gyre, thereby bringing  
302 weakly stratified waters of the ocean interior close to the surface (Marshall and Schott, 1999). Clockwise subpolar  
303 gyres, and associated deep-water formation by winter convection, are also present in the Pacific sector of the Southern  
304 Ocean, in the Ross and Amundsen seas. Previous numerical investigations of the effects of a closed Drake Passage on  
305 ocean dynamics have revealed that the closure of the Drake Passage tends to promote the existence of subpolar gyres  
306 in the Southern Ocean and vigorous deep-water formation (Nong et al., 2000; Sijp and England, 2004; Ladant et al.,  
307 2018). Additionally, it has been recently suggested that the effects of the closure and opening of the Drake Passage  
308 and the Panama Seaway may not be independent (Yang et al., 2014; England et al., 2017; Ladant et al., 2018). For  
309 instance, Yang et al. (2014) found that closing the Drake Passage tends to suppress the AMOC and to promote the  
310 emergence of a strong SOMOC when the Panama gateway is open, whereas the AMOC may remain intense when the  
311 Panama gateway is closed. It is thus very likely that, in our 55 Ma simulations, the very shallow Drake Passage,  
312 together with the opened Panama gateway, both contribute to the strong SOMOC.

313 In addition to the influence of the different gateways, tidally-induced mixing, which represents the enhanced vertical  
314 diffusivity resulting from the breaking internal waves generated by the interaction of tidal currents with rough bottom  
315 topography (St. Laurent et al., 2002), is another factor that contributes to the strong SOMOC found in the 55 Ma-3x  
316 simulation. A twin experiment of the 55 Ma-3x simulation, in which no tidal-induced mixing is prescribed (55 Ma-  
317 3x-noM2), simulates a SOMOC with a similar structure but an intensity that is 7 Sv weaker (33 Sv, compared to 40  
318 Sv in the reference 55 Ma-3x simulation, Fig. S4). It should be noted that the additional simulation has only been run  
319 for 2000 years, so that a small part of the difference between the two runs could arise from different equilibrated states.  
320 Yet, the difference between these two simulations is consistent with the recent results of Weber and Thomas (2017),  
321 who find a 10 Sv MOC enhancement in an early Eocene simulation with the ECHAM5/MPIOM model that explicitly  
322 simulates tides. This suggests that our parameterization for tidally-induced mixing based on the  $M_2$  dissipation fields  
323 of Green and Huber (2013) reasonably represents the effect of early Eocene tides on global ocean circulation. The  
324 strengthening of the MOC induced by tidal mixing can be directly related to the driving role of diapycnal mixing on  
325 the overturning circulation. Numerous studies have demonstrated that both the magnitude and the vertical distribution  
326 of the diapycnal mixing largely affects the strength of the MOC (Bryan, 1987; Manabe and Stouffer, 1999). As already  
327 pointed out by Green and Huber (2013), the Eocene MOC may be much more sensitive to the intensity of the abyssal  
328 mixing than the present day AMOC, that is largely isolated from the ocean floor by the presence of AABW and  
329 sustained quasi-adiabatically by the wind-driven upwelling in the Antarctic Circumpolar Current (ACC; Marshall and  
330 Speer, 2012). Diapycnal mixing is the only process that can warm up the dense waters formed in the Eocene Southern



331 Ocean, and these dense waters are directly exposed to the tidally-induced mixing at the bottom, such that the abyssal  
332 dissipation becomes the main controlling factor of the SOMOC intensity. In this respect, the large tidal dissipation  
333 rate suggested by Green and Huber (2013) in the Pacific is particularly important.

334 Another often-mentioned factor affecting ocean circulation is the climate state, in response to the atmospheric CO<sub>2</sub>.  
335 Based on highly simplified models, theoretical studies have suggested that ocean ventilation tends to increase under a  
336 warmer climate state due to a higher seawater density sensitivity to temperature (e.g. de Boer et al., 2007). However,  
337 the deepwater formation is a very regional phenomenon and thus this idealized relation might be complicated by other  
338 regional scale factors. Indeed, we did not see any systematic change in SOMOC between two CO<sub>2</sub> levels (1.5x and  
339 3x) simulations (see Section 6 for further discussion).

#### 340 **4. Horizontal circulation and winds**

341 In the present day North Atlantic, vertical and horizontal circulations are intimately connected, especially in the  
342 subpolar gyre (Marshall and Schott, 1999). For instance, the warm North Atlantic Current flowing northward and the  
343 cold East Greenland Current flowing southward are found roughly at the same depth at 60°N, such that the MOC in  
344 z-coordinates does not capture the associated water mass transformation at high latitude, whereas the MOC in density  
345 coordinates does (Zhang, 2010). It is only when the cold branch deepens in the Labrador Sea and becomes the Deep  
346 Western Boundary Current off Cape Hatteras that the overturning streamfunction in z-coordinates provides a good  
347 estimate of the water mass transformation. Hence, the horizontal subpolar gyre is really part of the North Atlantic  
348 thermohaline circulation. By analogy, one would expect that similar connections exist in the Southern Ocean during  
349 the Eocene, where the near-closure of the Drake and Tasmanian Passages allows for the emergence of intense subpolar  
350 gyres, that will precondition and feed the formation of deep water sustaining a strong SOMOC. We thus examine in  
351 the following the horizontal ocean circulation during the Eocene.

#### 352 **4.1. Gyre circulations**

353 In the 55 Ma-3x run, several horizontal gyres are well developed in both hemispheres, as shown by the barotropic  
354 streamfunction (Fig. 5). In particular, the closed Drake and Tasmanian passages support the western boundary currents  
355 necessary for the buildup of intense subpolar gyres in each sub-basin of the Southern Ocean, with intensity of 40 Sv  
356 in the Weddell Sea, 35 Sv in the Indian sector, and 28 Sv in the Ross and Amundsen Seas in the Pacific sector, and  
357 winter convection and deep water formation is occurring in the center of these gyres, as it is the case in the modern  
358 Labrador Sea (Marshall and Schott, 1999). The formation of deep water in the subpolar gyres is also promoted by the  
359 advection of saline subtropical waters from the subtropical gyres southward flowing branch (visible for instance on  
360 sea surface salinity), as shown by Ferreira et al. (2018). Compared to the PI conditions, the subtropical gyres are  
361 strongly perturbed by the numerous open gateways connecting the different basins. In the Southern Hemisphere, due  
362 to the large opening between Australia and Asia, the main subtropical gyre extends over both the paleo-Indian and  
363 Pacific oceans, with a western boundary current leaning partly on the northern coast of Australia (up to 52 Sv),  
364 Madagascar (60 Sv) and Africa. This ‘super’-subtropical gyre is partly fed by a 31 Sv eastward flow south of the Cape



365 of Good Hope originating from the South Atlantic subtropical gyre. Fig. 5 also reveals the existence of a strong anti-  
366 clockwise subtropical gyre south of a clockwise subpolar gyre in the North Pacific, with maxima of ~42 and 13 Sv  
367 respectively. In contrast, the gyres in the tectonically-restricted North Atlantic basin (Fig. 1) are weak, with a  
368 maximum of 22 Sv and 2 Sv for the subtropical and subpolar cells respectively. The numerous gateways in the tropical  
369 band clearly complicate the traditional gyre pattern in each basin, and increase their global connectivity.

370 In the PI-1x simulation, the maximum of the streamfunction in the North Atlantic subtropical and subpolar gyres are  
371 respectively 37 and 19 Sv, which is comparable with the intensity of the gyres found in the Southern Hemisphere in  
372 the 55 Ma simulations. The most salient feature of the PI-1x simulation is the existence of the Antarctic Circumpolar  
373 Current (ACC), with an eastward transport of 108 Sv through the Drake Passage, which totally disrupts the subtropical-  
374 subpolar gyres circulation in the different basins of the Southern Hemisphere. In many respects, the South Atlantic  
375 and the Weddell Sea during the Eocene are thus an analogue of the present-day North Atlantic in terms of subtropical,  
376 subpolar and overturning connections.

#### 377 **4.2. Wind stress**

378 The modern ocean circulation, in particular in the surface layers, is largely driven by the surface winds (e.g. Munk,  
379 1950), and it is thus interesting to examine the difference between the PI and 55 Ma horizontal circulation in light of  
380 changes in the wind pattern.

381 Overall, the patterns of the wind stress at the ocean surface in the 55 Ma-3x and PI-1x simulations are similar (Fig.  
382 6), although the magnitude differs between the simulations, and the sign of the difference depends on the hemisphere  
383 considered. Indeed, the wind stress is about 30% weaker in 55 Ma-3x than in PI-1x in the Southern Hemisphere (by  
384  $0.05 \text{ N m}^{-2}$  at  $45^\circ\text{S}$ ) whereas it is slightly stronger in the Northern Hemisphere. In the Southern Hemisphere, the  
385 difference is particularly striking in the South Atlantic and Indian basins and is largely due to the blocking position of  
386 Australia in the early Eocene. The paleogeographic context also explains the increased symmetry in zonal wind stress  
387 fields between the Eocene hemispheres relative to the PI.

388 The more symmetrical pattern of wind stress at the ocean surface in the 55 Ma-3x simulation (compared to PI-1x) is  
389 found in the zonal wind fields from the surface to 500 hPa in the atmosphere (Fig. 6D). The zonal wind strength is  
390 largely determined by the meridional temperature gradient in the atmosphere through the thermal wind relation  
391 (Holton and Staley, 1973). Indeed, the meridional temperature gradient in the 55 Ma-3x run (compared to PI-1x) is  
392 much reduced in the Southern Hemisphere south of  $40^\circ\text{S}$  in a large part of the air column, from the surface up to at  
393 least 500 hPa, in good agreement with the weaker westerly winds found in the 55 Ma-3x run (Fig. 6D). Moreover, the  
394 positions of Australia, Africa, and South America, all much more south during the Eocene than now, result in a  
395 blocking effect on the zonal winds, reducing the wind at 500 hPa by a maximum of 18% at  $40^\circ\text{S}$ . In the Northern  
396 Hemisphere, the meridional temperature gradient in the 55 Ma-3x run is reduced at the surface only north of  $60^\circ\text{N}$ ,  
397 and similar to the PI-1x run at 500 hPa, whereas the zonal winds are slightly stronger at 55 Ma throughout the air



398 column. At the surface, the maximum zonal wind stress are 19% stronger and shifted poleward by 2°, mainly due to  
399 land-sea distribution.

400 Changes in the ocean gyres circulation between 55 Ma and PI configurations are mostly due to the large changes in  
401 the ocean basin geometry and gateways, and not to the moderate changes in the strength and patterns of the wind stress  
402 (and its curl). A full understanding of what sets the intensity of the gyres in the 55 Ma simulations (as well as in the  
403 PI simulations) would require further investigations, which are beyond the scope of the present paper.

## 404 5. Oceanic heat transport and its decomposition

405 The modern ocean circulation plays a key role in the regulation of the climate through its contribution to the  
406 redistribution of heat from the Equator to the Poles. In the tropics, the ocean transports roughly 50% of the 3 PW  
407 carried northward by the ocean–atmosphere system, but less than 10% of the total at high latitude (Trenberth and  
408 Caron, 2001). The relative contributions of the horizontal and overturning ocean circulations to the meridional heat  
409 transport also vary greatly over different latitudes and between oceanic basins (Ganachaud and Wunsch, 2003). In  
410 light of the different ocean circulations found in the Eocene and PI conditions, we further investigate in the following  
411 how efficient was the ocean at transporting heat across latitudes in the early Eocene.

### 412 5.1. Oceanic heat transport

413 The total meridional OHT at a given latitude  $y$  is defined as the sum of advective and diffusive contributions:

$$414 \quad OHT_{total} = \rho_0 C_p \int_W^E \int_{-H}^0 \left( v\theta + K_H \frac{d\theta}{dy} \right) dz dx \quad (2)$$

415 where  $\rho_0$  is the seawater density,  $C_p$  is the specific heat capacity of sea water,  $v$  is the meridional velocity,  $\theta$  is the  
416 potential temperature,  $K_H$  is the horizontal diffusivity coefficient,  $H$  is the ocean depth. Here the computation is  
417 performed from the model output at each model time step. For the 55 Ma simulation, there is a significant OHT  
418 increase in the Southern Hemisphere compared to PI simulations, but a decrease in the Northern Hemisphere (Fig. 7).  
419 As a result, the simulated OHT in the 55 Ma-3x experiment is remarkably asymmetric between hemispheres. The  
420 mean OHT difference between the 55 Ma-3x and PI-1x simulations is of the order of 0.2 PW (1 PW =  $10^{15}$  Watt),  
421 peaking at 0.5 PW around 35°S. The 55 Ma-3x OHT reaches a maximum of 1.7 PW at 15–20°S, that is ~0.3 PW  
422 larger than in the PI simulations at the same latitude, but also ~0.5 PW larger than the maximum PI value in the  
423 Northern Hemisphere. This larger OHT in the Southern Ocean contributes to maintain the Southern Hemisphere  
424 particularly warm in the Eocene, especially south of 50°S, as can be seen on the SST distribution (Fig. 2).

425 Previous studies have examined the role of OHT during the Eocene with a particular focus on the response of the OHT  
426 to the opening of Southern Ocean gateways in either true Eocene paleogeography or more idealized modern  
427 configurations. The OHT simulated by our 55 Ma-3x experiment lies within the range of values found in the literature.  
428 For instance, using the NCAR CCSM ocean model with surface heat flux boundary conditions mimicking an energy-



429 balanced atmospheric model, Nong et al. (2000) found that closing the Drake Passage in a modern configuration  
430 results in a stronger SOMOC (24 vs 12 Sv), associated with an increased poleward OHT in the Southern Hemisphere  
431 (+0.2 PW, from 1 to 1.2 PW) and a decreased OHT in the Northern Hemisphere. Similarly, using the UVic  
432 intermediate complexity Earth System Climate Model, Sijp and England (2004) found a strongly enhanced heat  
433 transport (from 1.6 to 2.4 PW) in the Southern Hemisphere in response to the closure of Drake Passage in a modern  
434 configuration. Using the fully coupled NCAR model with a closed Drake Passage but a closed Tasmanian gateway in  
435 a realistic Eocene configuration, Huber et al. (2004) found a rather weak poleward OHT in the Southern Hemisphere  
436 during the Eocene (with a maximum of 0.9 PW at ~10°S), likely because of the absence of a strong SOMOC in their  
437 simulations compared to ours. When closing the Drake passage in the GFDL model, Yang et al. (2014) found that the  
438 change in OHT was much larger when the Panama Seaway was open, with a strong increase of the OHT in the  
439 Southern Hemisphere. More recently, Baatsen et al. (2018) used the higher resolution CESM coupled model to  
440 simulate the 38 Ma climate, prescribing levels of CO<sub>2</sub> and CH<sub>4</sub> in the atmosphere 2 and 4 times the PI levels. In these  
441 two simulations, they found a maximum of ~1.5 PW OHT at 20°S in the Southern Hemisphere, associated with a 14–  
442 16 Sv SOMOC (compared to the 40 Sv and 1.7 PW at the same latitudes in our 55 Ma-3x simulation). Although these  
443 previous studies are all based on results from different models of various complexity and resolution, they all  
444 consistently suggest that the OHT is largely coupled to the structure and strength of the MOC cells. For instance, the  
445 difference in the MOC intensity simulated by coupled and uncoupled models could be primarily caused by the positive  
446 salt-advection feedback and the self-stabilizing thermal feedback (Sijp and England, 2004). It is noteworthy that the  
447 strengthening influence of tidal-induced mixing on the MOC (+7 Sv) is associated with a rather weak increase in  
448 OHT, lower than 0.03 PW on average, in agreement with Weber and Thomas (2017). Such a nonlinear relationship  
449 between OHT and MOC has also been suggested by Boccaletti (2005), who stressed that, locally, the shallow  
450 circulation can be as important as the deep overturning for determining the OHT.

## 451 5.2. Decomposition of the meridional ocean heat transport

452 In order to understand the differences in OHT among our simulations (Fig. 7A), we further split up the OHT into an  
453 advective contribution (OHT<sub>adv</sub>) and a diffusive contribution, corresponding respectively to the first and second term  
454 on the RHS of Eq. 2 (Fig. 7B and 7C). This decomposition reveals that, in both the 55 Ma and the PI runs, the advective  
455 part dominates the OHT at all latitudes, except at 40°S/N in PI where the presence of large temperature gradient (Fig.  
456 2B) results in larger diffusive heat transports. Figs. 7B and 7C also reveals that different OHT between PI and 55 Ma  
457 are mainly due to differences in the advective components, the differences in diffusive OHT being rather small.

458 The advective OHT<sub>adv</sub> can be decomposed further into an overturning (OHT<sub>MOC</sub>) and a gyre (OHT<sub>gyre</sub>) component,  
459 following for instance Bryan (1982) or Volkov et al. (2010):

$$460 \quad OHT_{adv} = \rho_0 C_p \iint \bar{v} \bar{\theta} dx dz + \rho_0 C_p \iint v' \theta' dx dz \quad (3)$$

461 where  $\bar{v}$ ,  $\bar{\theta}$  represent the zonal averages of the velocity and temperature, respectively, and  $v'$ , and  $\theta'$  the deviations  
462 from these zonal means. The first term of the RHS of Eq. 3 corresponds to the overturning component (OHT<sub>MOC</sub>) and





463 the second term corresponds to the horizontal transport associated with the large-scale gyre circulation ( $OHT_{gyre}$ ).  
464 Note that due to limitations on the availability of model outputs, the different terms of Eq. 3 presented on Figs. 7D  
465 and 7E are computed from monthly means. This explains why the sum of the two terms does not completely equal  
466  $OHT_{adv}$  shown on Fig. 7C, the latter being computed at each model time step during the simulations. The differences  
467 between the two computations can be seen on Fig. 7C.

468 The decomposition reveals that the enhanced Southern Hemisphere  $OHT_{adv}$  at 55 Ma (compared to PI) is overall due  
469 to differences in  $OHT_{MOC}$  (Fig. 7D, 7E). The contribution from the gyre circulation varies with latitude, with a  
470 compensation effect between  $OHT_{gyre}$  and  $OHT_{MOC}$  in the low-latitudes, and an enhancement in mid-to-high latitudes.  
471 Consequently, in the tropics, the strong Eocene  $OHT_{MOC}$  (up to 2 PW) is  $\sim 0.3$  PW larger than in the PI simulations,  
472 leading to an overall larger  $OHT_{adv}$  in the 55 Ma simulations. In the mid-latitudes of the Southern Hemisphere, where  
473 the OHT is overall smaller than in the tropics, the  $OHT_{MOC}$  is almost 1 PW stronger than in PI. By contrast, south of  
474  $60^\circ S$ , enhanced OHT in 55 Ma simulations (compared to PI) results from a combination of stronger  $OHT_{gyre}$  and  
475  $OHT_{MOC}$ . It is worth noticing that  $OHT_{gyre}$  at  $\sim 40^\circ S$  in the PI-1x simulation is very likely underestimated in our  
476 decomposition computed from monthly mean data, because higher-frequency processes (resulting for instance from  
477 atmospheric synoptic variability) could contribute significantly to  $OHT_{gyre}$  in regions where the mean meridional  
478 currents are weak (Volkov et al., 2010).

479 It is obvious from Figs. 3 and 7 that the vigorous SOMOC simulated in the 55 Ma experiment drives a strong net OHT  
480 toward the South Pole. This strong SOMOC is associated with a poleward transport of warm waters at shallow depths  
481 where zonal oceanic temperature gradients are larger, and a returning equatorward transport of colder water at depth  
482 where ocean temperature tends to be more homogenous (Fig. 2C). Remarkably, although the ACC is absent from the  
483 Eocene simulation because of different Drake and Tasmanian passages configurations that constitute latitudinal  
484 barriers (Munday et al., 2015), the contribution of the gyre circulation and diffusive process to poleward heat transport  
485 ( $OHT_{gyre}$ ) is smaller than in PI simulations. This small  $OHT_{gyre}$  in the 55 Ma simulation is unexpected, given previous  
486 hypotheses on the climatic effects of the ACC (e.g. Nong et al., 2000; Toggweiler and Bjornsson, 2000; Sijp and  
487 England, 2004). The ACC has indeed been suggested to be a barrier for poleward heat transport, so that the onset of  
488 the ACC could be a potential driver for the Eocene-Oligocene Antarctica cooling around 34 Ma. Yet, these studies  
489 may not have captured the full complexity of the links between the ACC and the OHT in the Southern Ocean. Indeed,  
490 the analysis of both in-situ observations (Watts et al., 2016) and the CESM1.0 model (Yang et al., 2015) have revealed  
491 that the ACC is composed of meridional excursions of the mean geostrophic horizontal shear flow, energetic eddies  
492 and large diffusive heat transport, which balances out the equatorward OHT due to Ekman transport and leads to a net  
493 poleward OHT in the Southern Ocean (Volkov et al., 2010).

## 494 **6. Sensitivity of the ocean response to a doubling of the levels of atmospheric CO<sub>2</sub>**

495 Our analysis has so far focused on the comparison between the 55 Ma-3x and the PI-1x simulations, as the former is  
496 performed with atmospheric CO<sub>2</sub> levels thought to be representative of the early Eocene (Foster et al., 2017). The





497 analysis of two additional simulations (55 Ma-1.5x and PI-2x) allows us to investigate the robustness of the ocean  
498 circulation in this range of atmospheric CO<sub>2</sub> concentration. The set of simulations based on both the 55 Ma and PI  
499 configurations also help us to quantify the sensitivity of the oceanic conditions to a doubling of the level of CO<sub>2</sub> in the  
500 atmosphere for early Eocene and modern setting.

501 The mean ocean temperatures are very sensitive to the atmospheric CO<sub>2</sub> concentration in the Eocene configuration  
502 (Table 1). Global mean ocean temperature in 55 Ma simulations increases by 4.9°C in response to CO<sub>2</sub> doubling from  
503 1.5x to 3x, which is much larger than the 1°C increase in PI simulations from 1x to 2x. SST also shows a much larger  
504 increase at 55 Ma (+4.7°C) than in PI simulations (+2.6°C), especially at high latitudes and in the regions of deep-  
505 convection of the Southern Ocean (Fig. 8), which is similar to the changes in air temperature at 2 m (+5.6°C vs +3.5°C  
506 respectively, this difference is known as the climate sensitivity). Such contrasted values of climate sensitivity between  
507 55 Ma and PI are in good agreement with the recent results of Farnsworth et al. (2019) when analyzing a series of  
508 climate models. In the absence of sea-ice at 55 Ma, the winter SST in deep-convection regions is largely influenced  
509 by the air-sea interactions and thus directly related to air temperature, that rarely decreases below 10°C (resp. 5°C) in  
510 55 Ma-3x (resp. 1.5x), such that the deep waters filling the whole ocean vary accordingly with temperature. This is  
511 not the case in the present-day configuration where deep-water formation is tied to the marginal ice zones, such that  
512 the dense water formed through the effect of the brine rejection have initial temperatures close to freezing. Intuitively,  
513 we could have expected a larger sensitivity of ocean temperature to the level of CO<sub>2</sub> in the atmosphere in the PI runs,  
514 induced by the ice-albedo feedback. Yet, the effect of this feedback appears to be limited to the high latitudes (Fig.  
515 8), and only plays a marginal role for the changes in global mean SSTs or temperatures of the deeper water masses  
516 (Table 1).

517 The ocean circulation only shows a minor response to a CO<sub>2</sub> doubling in both the 55 Ma and PI configurations,  
518 although a regional response still exists (Table 4). In the 55 Ma simulations, doubling CO<sub>2</sub> enhances the maximum  
519 abyssal SOMOC by only 0.3 Sv out of 40 Sv in total, while the intensity of the shallow MOC cell and of the barotropic  
520 streamfunction are slightly reduced. In the PI simulations, doubling the level of atmospheric CO<sub>2</sub> has the opposite  
521 effect on the MOC cell occupied by the AABW, whose maximum intensity reduces by 1.2 Sv (out of ~15 Sv), whilst  
522 the AMOC slightly increases by 0.3 Sv (out of ~11 Sv). This small increase in the steady-state AMOC is quite  
523 interesting and needs to be contrasted with the transient response of the AMOC intensity to global warming (Gent,  
524 2018; Jansen et al., 2018). Indeed, CMIP-type climate models consistently project a strong decline of the AMOC  
525 strength when forced with a range of increasing greenhouse gas emission scenarios (Schmittner et al., 2005; Cheng et  
526 al., 2013). Yet, when run for longer integrations until full equilibrium, models suggest that the AMOC tends to recover,  
527 so that the AMOC is not very sensitive to the level of CO<sub>2</sub> (Jansen et al., 2018; Thomas and Fedorov, 2019), as is the  
528 case in our PI simulations. Once equilibrium is reached, the most significant effect of a doubling in the CO<sub>2</sub>  
529 concentration in our experiments is a sharp increase in the ACC transport (+22 Sv out of 108 Sv in PI-1x), in response  
530 to stronger westerlies in the Southern Ocean. In contrast, the barotropic circulation remains almost the same in the  
531 Pacific, but varies in the North Atlantic with a 5 Sv stronger (weaker) subpolar (subtropical) gyre in the PI-2x  
532 simulation compared to PI-1x.



533 A number of paleoclimate studies have investigated the influence of CO<sub>2</sub> levels on ocean circulation in coupled  
534 models, with contrasted responses depending on the period considered and the model used. For instance, the deep  
535 overturning circulation in the HadCM3L climate model shows an overall high sensitivity to CO<sub>2</sub> concentrations in  
536 simulation of the Paleocene-Eocene period (Lunt et al., 2010), whilst the GENIE model only exhibits a small response  
537 in the Cretaceous simulations of Monteiro et al. (2012). Winguth et al. (2010) further suggest that, in a given simulation  
538 of the Paleocene-Eocene period, the different MOC cells (e.g. in the Northern and Southern hemispheres) could  
539 respond differently to a change of CO<sub>2</sub> levels. These various responses can be attributed to different factors. First, it  
540 is clear that the resolution and overall complexity of the model used for these studies may partly control the sensitivity  
541 of the MOC to CO<sub>2</sub> levels, as in models of modern climate (e.g. Bryan et al., 2006). Second, a variety of time scales  
542 are intertwined in the adjustment of the ocean circulation to external perturbations, from decades for the dynamical  
543 adjustment to millennia for the thermo-dynamical response of bottom waters through vertical advective-diffusive  
544 balance (e.g. Donnadieu et al., 2016). The transient response of the ocean can therefore differ, or even be in the  
545 opposite direction, from the final equilibrium response.

546 The OHT response to a doubling of CO<sub>2</sub> in the Eocene simulations is also rather small, with a slight decrease, in  
547 contrast to the OHT increase seen in the PI simulations (although the magnitude of the change is smaller; Fig. 9A).  
548 The OHT in the 55 Ma-3x simulation is about 0.15 PW smaller than in 55 Ma-1.5x simulation over most of latitudes.  
549 Both overturning and horizontal components contribute to this overall smaller OHT in the 55 Ma-3x simulation (Fig.  
550 9D, 9E). A weak OHT in the tropics is due to weaker MOC at the same latitudes, while a weak OHT at high-latitude  
551 can be attributed to the small amplitude of horizontal gyres. For the PI simulations, the OHT in the PI-2x simulation  
552 is ~0.1 PW larger than in PI-1x at high latitudes. This larger OHT in PI-2x is mostly due to the gyre component, which  
553 is in good agreement with the ~5 Sv stronger North Atlantic subpolar gyre and almost no change in the AMOC for  
554 instance.

555 Our results therefore support a stable, yet rather small, response of global ocean circulation and heat transport to the  
556 doubling of atmospheric CO<sub>2</sub> levels. Nevertheless, we only investigate a limited range of CO<sub>2</sub> levels (from 1.5x to 3x)  
557 and cannot exclude that the sensitivity of ocean circulation to CO<sub>2</sub> concentrations may change at more extreme CO<sub>2</sub>  
558 levels, as the response of the ocean conditions is highly non-linear (Lunt et al., 2010). Given that the levels of CO<sub>2</sub> in  
559 the Eocene atmosphere are relatively poorly constrained by proxy reconstructions, additional experiments are  
560 underway to explore higher values of CO<sub>2</sub> concentration for the early Eocene configuration.

## 561 7. Conclusions

562 The early Eocene (~55 Ma) was most likely the warmest period in the Cenozoic. During that period, paleogeographic  
563 restrictions of certain modern basins and gateways, such as the North Atlantic and the Drake and Tasmanian passages,  
564 suggest fundamental differences with the modern large-scale ocean circulation. It has been proposed that the distinct  
565 mode of ocean circulation operating during the early Eocene may have contributed to the significant polar warmth  
566 recorded by observational evidence. There is however no consensus on the modes of early Eocene ocean circulation



567 or on the relative influence of the overturning and horizontal circulation on the poleward heat transport. Here we  
568 revisit this question by analyzing the ocean circulation and its contribution to the meridional OHT using simulations  
569 of the Early Eocene performed with the IPSL-CM5A2 coupled climate model set up with the recent paleogeographic  
570 reconstructions of this Eocene time slice distributed as part of the DeepMIP project. Our main results are summarized  
571 hereafter.

572 A strong abyssal overturning circulation is found in the 55 Ma simulation, with deep water formed only in the Southern  
573 Ocean (mainly in the Weddell Sea), whereas there is no deep water formation in the Northern Hemisphere, in contrast  
574 to some previous work on the early Eocene (e.g. Winguth et al., 2012). This situation is favored by orographically-  
575 induced freshwater fluxes (precipitation and runoff) and maintained by a salt-advection feedback. Indeed, the  
576 atmospheric circulation around Antarctica induces relatively low precipitation rates in the Southern Ocean, resulting  
577 in higher salinity, and hence larger surface density than in the North Pacific, where the large precipitations and runoff  
578 induced by the orographic uplift of the westerlies above the paleo-Rockies tend to reduce the surface salinity and  
579 inhibit deep water formation.

580 The paleogeography (and paleobathymetry) and tidally-induced mixing are the main drivers of the strong SOMOC  
581 (up to 40 Sv) during the early Eocene. The (nearly-)closed Drake and Tasmanian Passages are of fundamental  
582 importance for sustaining the SOMOC, via their effect on the horizontal ocean circulation. More specifically, with the  
583 (nearly-)closed Drake and Tasmanian Passages serving as a western boundary, clockwise subpolar gyres are well-  
584 developed (~40 Sv) in the Weddell and Ross Seas, favoring the emergence of deep convection and deep-water  
585 formation through isopycnal doming and salt-advection feedback. Tidal-induced mixing also contributes to 7 Sv (out  
586 of 40 Sv) to this SOMOC, but with only a limited impact on the heat transport.

587 The vigorous SOMOC simulated for the Eocene is associated with a larger poleward heat transport (by a maximum  
588 of 0.5 PW relative to PI) in the Southern Hemisphere, that largely contributes to maintain a warm Southern Ocean and  
589 Antarctica. Perturbation experiments have been conducted in present-day coupled models to evaluate the impact of an  
590 AMOC shutdown. In their model, Vellinga and Wood (2008) found that a 10 Sv reduction in AMOC, associated with  
591 a change of its structure, leads to a 1.7°C cooling of the Northern Hemisphere, with a local stronger cooling by 5°C  
592 in the northern North Atlantic. This gives credit to the importance of the 40 Sv SOMOC for maintaining the Southern  
593 Ocean warm in the 55 Ma simulations. However, other factors than a strong SOMOC and associated OHT could  
594 contribute to the warm Southern Ocean. Indeed, Rose and Ferreira (2013) have shown that changes in OHT can induce  
595 changes in global mean temperature and meridional temperature gradient through convective adjustment of the  
596 extratropical troposphere and increased greenhouse effect. According to their results, the magnitude of those changes  
597 could be up to 1°C and 2.6°C for every 0.5 PW enhancement in OHT. Further investigations would be required to  
598 examine if this mechanism is also at play in our simulation, and contributes significantly to the Southern Ocean  
599 warmth.

600 A further decomposition of the OHT reveals that the different overturning circulations between 55 Ma and PI explain  
601 most of the increase of the OHT in the Southern Hemisphere. The contribution of gyre circulation to the OHT is only



602 secondary, and varies with latitude, with a compensating effect between the MOC and gyre circulations in low- and  
603 mid-latitudes, whereas the two contributions add up in high latitudes. More importantly, the latitudinal distribution of  
604 the gyre contribution to the OHT only marginally varies between the 55 Ma and PI simulations, despite the absence  
605 of an ACC in the Eocene experiment. Given that the 55 Ma paleo-bathymetry does not allow the existence of a strong  
606 ACC, this questions strongly the idea that the ACC could be a strong barrier for the OHT in our modern climate, and  
607 suggests that the meridional excursions of the ACC might indeed play an important role in the gyre-related OHT  
608 (Volkov et al., 2010; Watts et al., 2016; Yang et al., 2015).

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## Tables

**Table 1** Summary of the simulation setup and key diagnostics in the different simulations used in this study. All the values presented are averaged over the last 100 years of each simulation.

Simulation	Setup			Ocean surface			Full depth ocean			Atmos.
	CO <sub>2</sub> (ppmv)	Bathymetry	Duration (yr)	SST (°C)	SSS (psu)	Sigma (kg/m <sup>3</sup> )	T (°C)	S (psu)	Sigma (kg/m <sup>3</sup> )	T at 2m (°C)
55 Ma-3x	840	55Ma	4000	27.51	34.04	21.33	11.30	34.68	26.24	25.12
55 Ma-3x-noM2	840	55Ma	2000	27.47	34.16	21.44	10.85	34.68	26.34	24.98
55 Ma-1.5x	420	55Ma	4000	22.82	34.36	22.94	6.38	34.68	26.94	19.56
PI-2x	560	PI	2910	20.14	34.40	23.63	4.32	34.61	27.27	16.88
PI-1x	280	PI	2790	17.51	34.43	24.32	3.34	34.61	27.40	13.33

**Table 2** Root-mean-square-deviation (RMSD) of simulated annual mean SST and proxy-based SST estimates (in °C). RMSD metrics are defined in Supplementary material. The proxy-based SST estimates are from the DeepMIP dataset for the early Eocene (Hollis et al. 2019), and the number of data points and the uncertainty for each proxy type is also indicated. The uncertainty range is defined as the 2σ deviations for δ<sup>18</sup>O, and as the range between 5% and 95% percentile SST estimates for TEX<sup>86</sup>, Mg/Ca and Clumped isotope data.

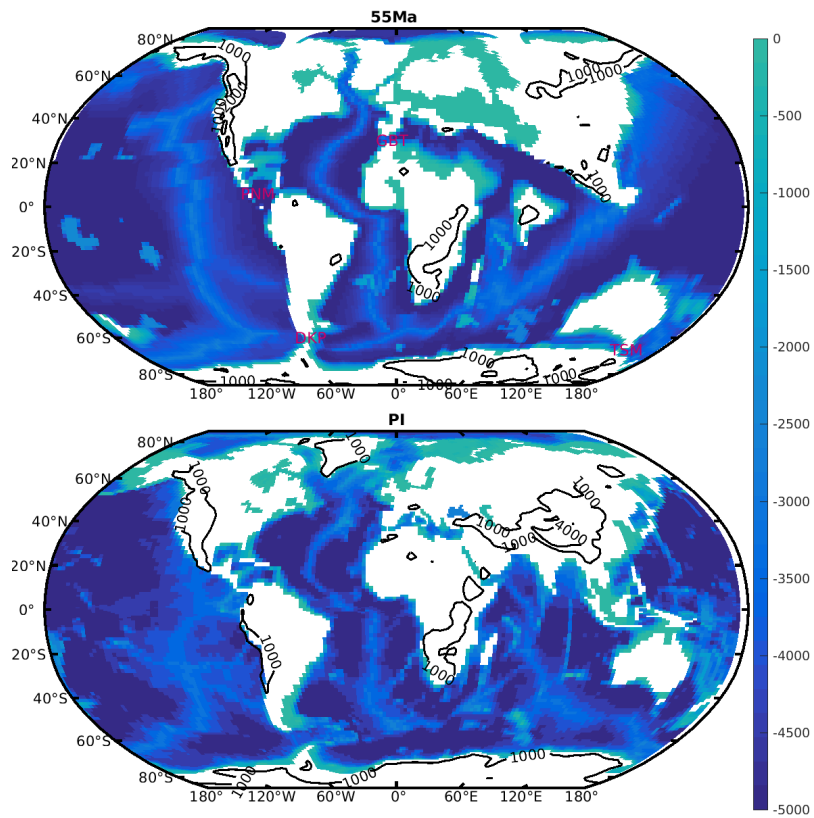
Type of proxy	TEX <sup>86</sup>	δ <sup>18</sup> O	Mg/Ca	Clum. isotope
Number of data points	10	10	7	5
Uncertainty range of proxy-data	15.1	3.4	6.7	5.1
RMSD	55Ma-3x	13.7	7.5	6.4
	55Ma-1.5x	18.2	7.5	10.6

**Table 3** Key ocean surface parameters in the North Pacific and Weddell Sea in 55Ma-3x simulation. The Weddell Sea and the North Pacific convection regions are defined by the deepest mixed layer depth at high-latitudes and the extended regions are defined as boxes over the Weddell Sea (78°S-61°S and 62°W-8°E) and the North Pacific (48°N-67°N and 124°E-143°W) roughly corresponding to closed contours of sea surface height. SST, SSS, sigma are winter average (January-February-March for the Northern Hemisphere and July-August-September for the Southern Hemisphere), MLD is given for the end of winter (March for the Northern Hemisphere and September for the Southern Hemisphere), while precipitation, runoff and evaporation are annual means.

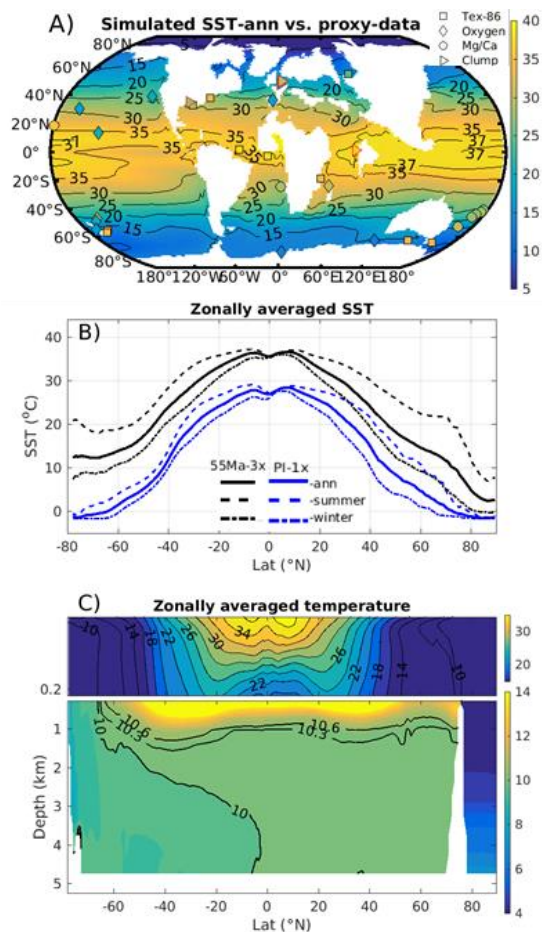
	Convection region				Extended region						
	SST (°C)	SSS (psu)	sigma (kg/m <sup>3</sup> )	MLD (m)	Area (km <sup>2</sup> )	Precipitation (P) (m/yr)	Runoff (R) (m/yr)	Evaporation (E) (m/yr)	P+R-E (m/yr)	SST (°C)	SSS (psu)
North Pacific	9.21	34.06	26.22	111	9.23e6	1.48	0.37	0.74	1.12	10.33	32.89
Weddell Sea	9.57	34.79	26.73	3129	4.70e6	1.01	0.22	0.69	0.54	10.13	34.77

**Table 4** Intensity of the gyres (SPG<sub>SA</sub>: South Atlantic subpolar gyre for 55 Ma; STG<sub>NA</sub>: North Atlantic subtropical gyre; SPG<sub>NA</sub>: North Atlantic subpolar gyre; STG<sub>NP</sub>: North Pacific subtropical gyre; SPG<sub>NP</sub>: North Pacific subpolar gyre; ACC: Antarctic circumpolar current for PI) and overturning cells (SOMOC: Southern Ocean MOC for 55Ma; AABW: Antarctic bottom water for PI; NADW: North Atlantic deep water for PI).

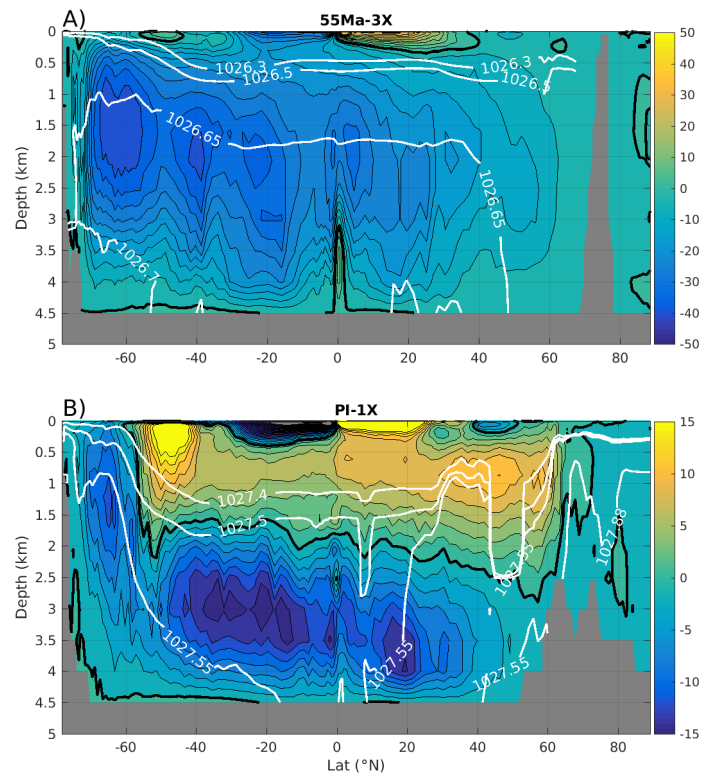
Simulation	Overturning (Sv)		Gyre intensity (Sv)						
	SOMOC/ AABW	NADW	SPG <sub>SA</sub> / ACC	SPG <sub>SI</sub>	SPG <sub>SP</sub>	STG <sub>NA</sub>	SPG <sub>NA</sub>	STG <sub>NP</sub>	SPG <sub>NP</sub>
55 Ma-3x	-40	—	40	35	28	22	-2	42	-13
55 Ma-1.5x	-40	—	38	46	35	29	0	53	-8
PI-2x	-15	11.6	130	—	—	32	-24	48	-19
PI-1x	-16	11.3	108	—	—	37	-19	48	-20



**Figure 1.** Bathymetry/topography (m) boundary conditions used in the 55 Ma (based on Herold et al. 2014) and PI simulations. The black contours indicate the 1000 and the 4000m altitude. DKP indicates the Drake Passage, TSM the Tasmanian Passage, PNM the Panama passage and GBT Gibraltar.

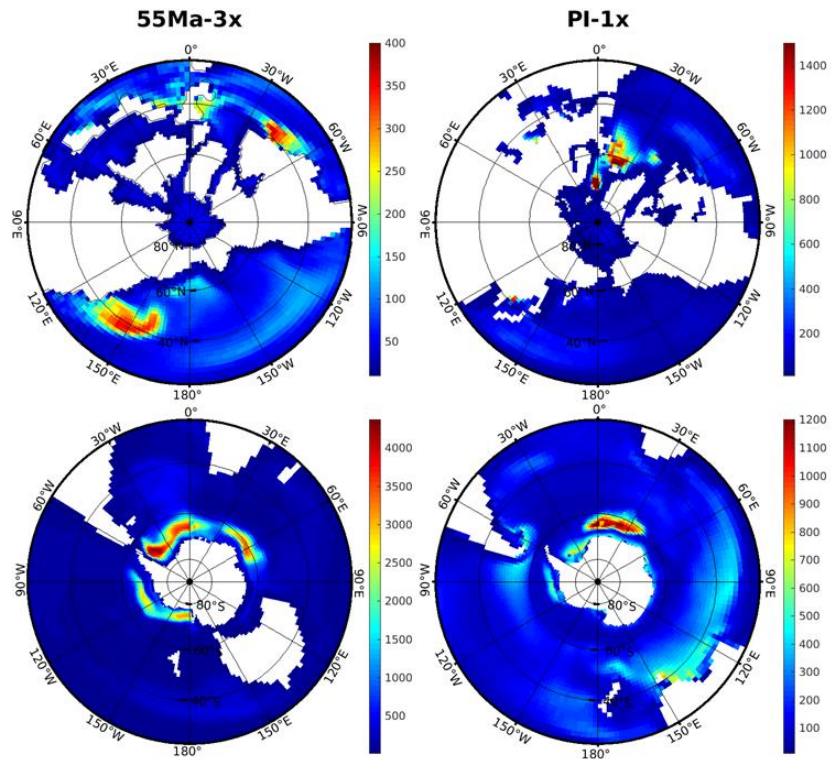


**Figure 2.** (A) Annual mean SST (in °C) in the 55 Ma-3x simulation, and point-to-point comparisons with proxy-based SST estimates from the DeepMIP dataset for the early Eocene (Hollis et al. 2019). The different symbols represent different proxies. (B) Simulated zonally-averaged annual mean SST (in °C, solid lines) in the 55 Ma-3x (black) and PI-1x (blue) simulations. The dashed and dotted lines are indicating the means for Summer and Winter, respectively. (C) Zonally-averaged ocean temperature in the 55 Ma-3x simulation, with a zoom in the upper 200 m (contour interval: 2°C and 0.3°C for the top and bottom panels, respectively).

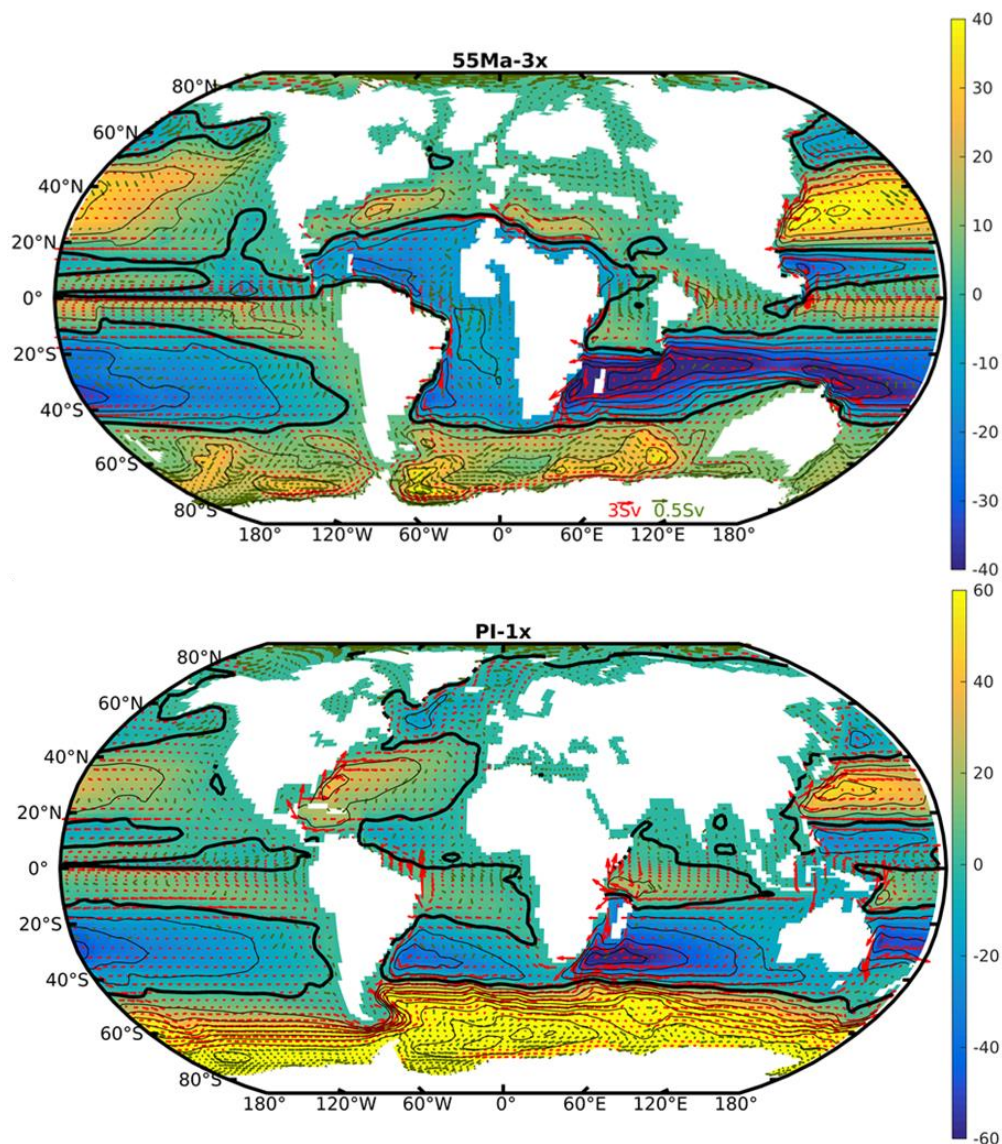


**Figure 3.** Streamfunction of the meridional overturning circulation (in Sv) in the 55 Ma-3x (A, contour interval: 4 Sv) and PI-1x (B, contour interval: 2 Sv) simulations. Note the different colorbars of the two subplots. The black thick lines indicate the zero-contour, with positive values indicating clockwise circulation, and negative values anti-clockwise circulation. White lines show selected zonally-averaged isopycnal contours (potential density in  $\text{kg/m}^3$ ).

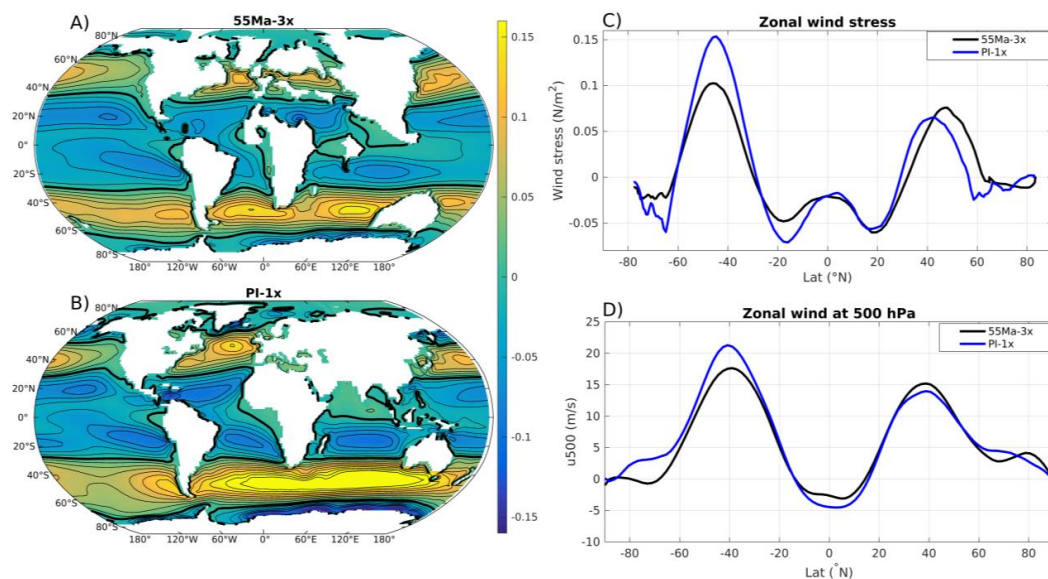




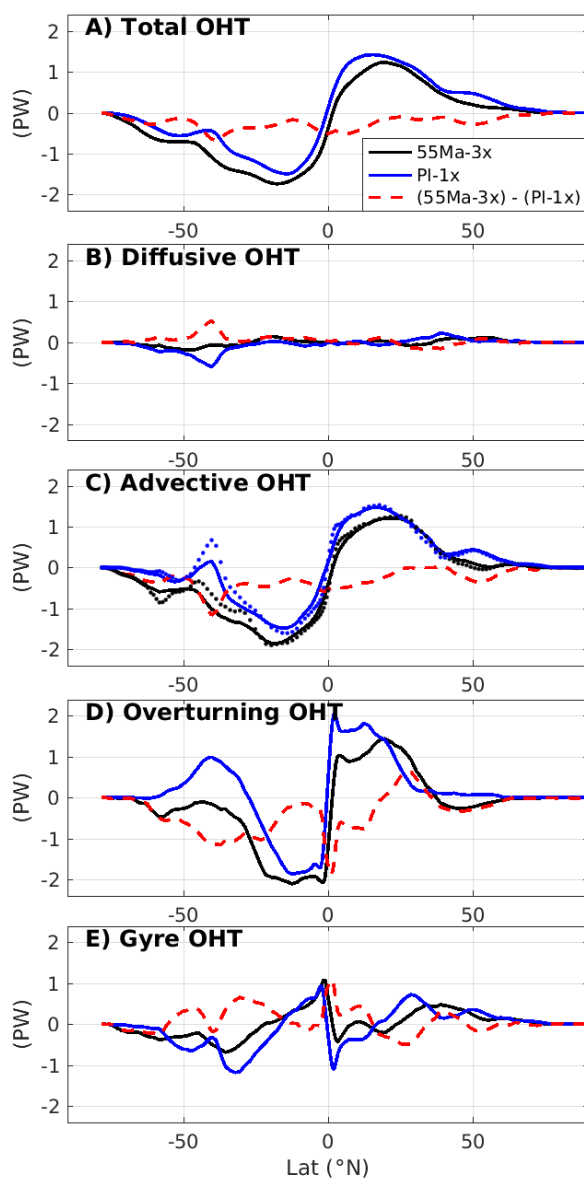
**Figure 4.** Winter (i.e. March in the Northern Hemisphere – top panels; and September in the Southern Hemisphere – bottom panels) mixed layer depth (in m) in the 55 Ma-3x (left) and PI-1x (right) simulations. Note the very different colorbars among plots. Mixed layer depth is defined by the potential density difference of  $0.3 \text{ kg/m}^3$  with reference to the surface.



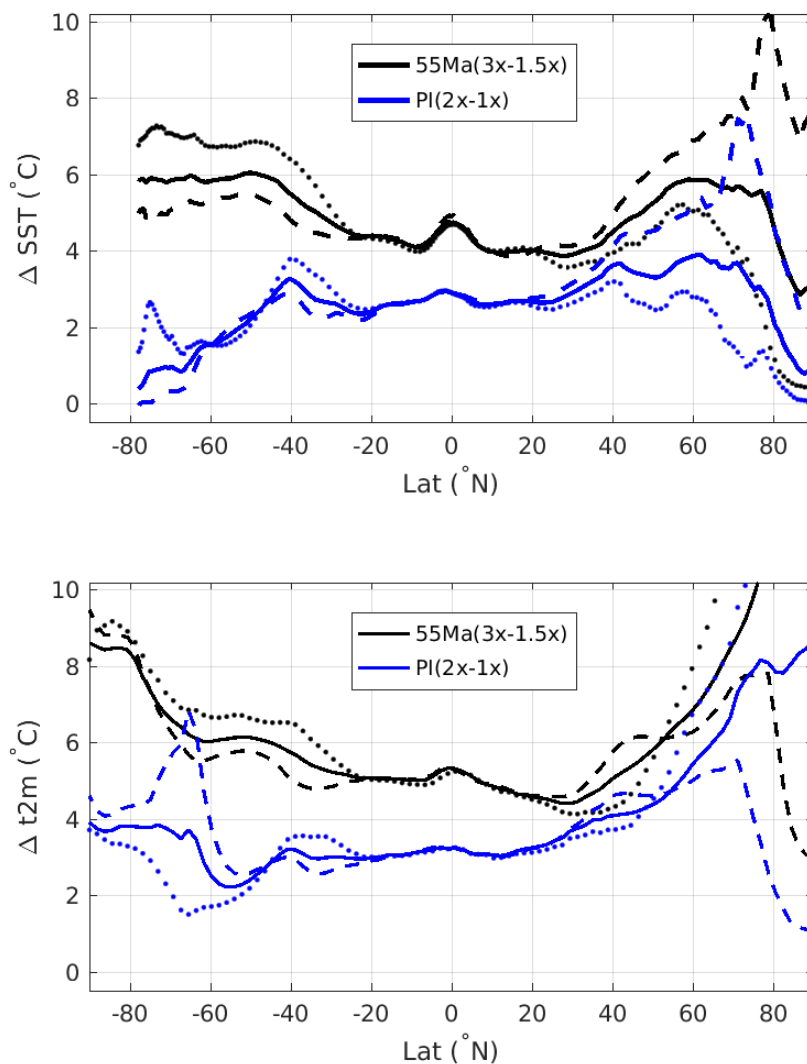
**Figure 5.** Barotropic streamfunction (in Sv) in the 55 Ma-3x (top, contour interval: 10 Sv) and PI-1x (bottom, contour interval: 15 Sv) simulations, integrated northward from Antarctica. Note the different colorbars of the two subplots. The black thick lines indicate the zero-contour. The mean transport (in Sv) integrated over the top 300m is indicated with vectors (note that only one every two point is plotted to increase the readability). Two scales are used to represent transport larger (in red) or lower (in green) than 0.5 Sv. The transports through the key gateways in the 55 Ma-3x simulation are respectively: Drake Passage 3.1 Sv, Tasmania 1.3 Sv, Panama 4.7 Sv, Gibraltar -14.3 Sv (positive transports are eastward, negative westward). The Drake Passage throughflow, corresponding to ACC, in the PI-1X simulation is ~108 Sv.



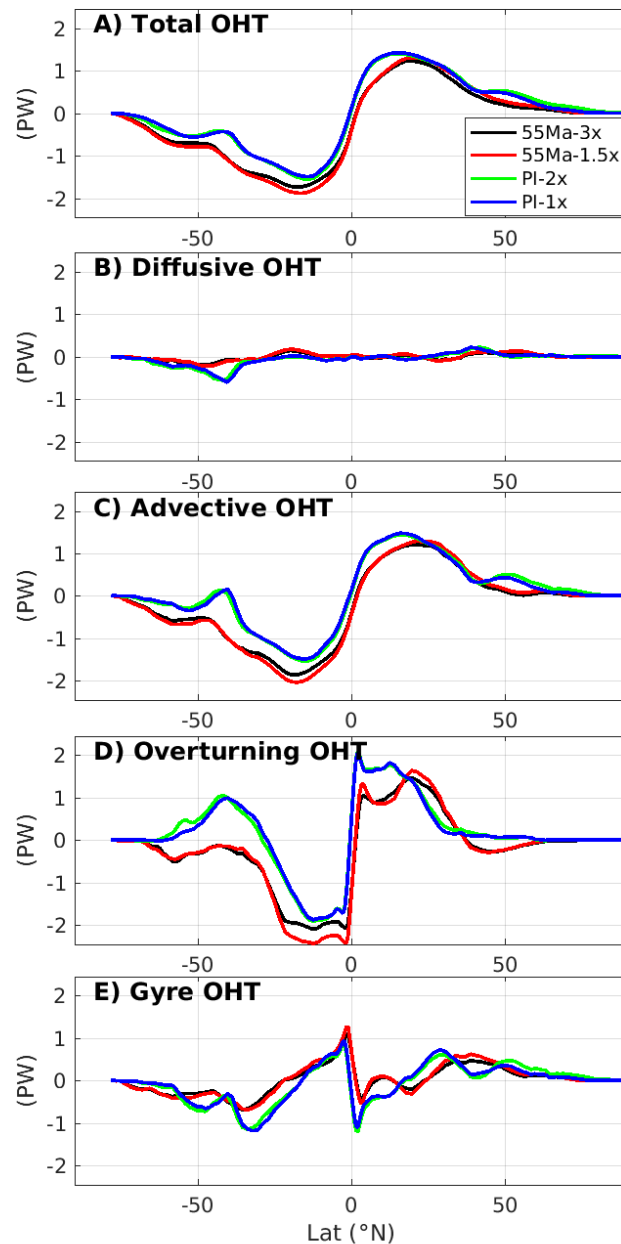
**Figure 6.** Zonal wind stress (in  $N/m^2$ ) in the 55 Ma-3x (A) and the PI-1x simulation (B). (C) Zonally averaged zonal wind stress (in  $N/m^2$ ) as a function of latitude in the 55 Ma-3x and PI-1x simulations. (D) Zonally averaged zonal wind (in  $m/s$ ) at 500 hPa in the atmosphere as a function of latitude in the 55 Ma-3x and PI-1x simulations.



**Figure 7.** Meridional oceanic heat transport (in PW) as a function of latitude (positive contribution is northward) and its decomposition according to Eqs. 2 and 3. Results from the 55 Ma-3x and the PI-1x runs are shown in black and blue, respectively, and the difference between the two is in red. On panel C, the dotted lines indicate the sum of  $OHT_{MOC}$  and  $OHT_{gyre}$  estimated from monthly means, while the solid lines are computed at the model time step.



**Figure 8.** Difference of the zonally-averaged SST (top, in °C) and air temperature at 2m (bottom, in °C) as a function of latitude between the 55 Ma-3x and the 55 Ma-1.5x runs in black, and the PI-2x and PI-1x in blue. The solid line indicates the annual mean, the dashed line the mean over July-August-September, and the dotted line the mean over January-February-March.



**Figure 9.** Meridional ocean heat transport (in PW) as a function of latitude (positive contribution is northward) and its decomposition according to Eqs. 2 and 3 in the 4 simulations.