

# Atmospheric methane control mechanisms during the early Holocene

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**Abstract.** Understanding processes controlling atmospheric methane (CH<sub>4</sub>) mixing ratio is crucial to predict and mitigate the future climate change. In spite of recent studies using various approaches for the last ~1000 to 2000 years, control mechanisms of CH<sub>4</sub> still remain unclear, partly because the late Holocene CH<sub>4</sub> budget is comprised of natural and anthropogenic emissions. In contrast, the early Holocene was a period when human influence should have been substantially smaller, so that it allows us to elucidate the natural controls under interglacial conditions. Here we present new high resolution CH<sub>4</sub> records from Siple Dome, Antarctica, covering from 11.6 to 7.7 thousands of years before 1950 AD (ka). We observe several local CH<sub>4</sub> minima on a roughly 1000-year spacing, which correspond to cool periods in Greenland. We hypothesize that the cooling in Greenland forced the Intertropical Convergence Zone (ITCZ) to migrate southward, reducing rainfall in northern tropical wetlands although there is no obvious change was observed in low latitude hydrology corresponding to abrupt CH<sub>4</sub> reduction at ~10.3 ka. Inter-polar difference (IPD) of CH<sub>4</sub> shows a gradual increase from the onset of the Holocene to ~9.9 ka, which implies growth of boreal source strength following the climate waring in the northern extratropics during that period. Finally, we find that amplitude of centennial- to millennial scale CH<sub>4</sub> variability of the early Holocene is larger on average than that of the earlier part of the late Holocene (3.5 – 1.2 ka).

## 20 **1 Introduction**

Methane (CH<sub>4</sub>) is a potent greenhouse gas whose atmospheric mixing ratio has increased more than 2.5 times since the Industrial Revolution (Dlugokencky et al., 2009). Although lower in abundance compared to carbon dioxide (CO<sub>2</sub>), CH<sub>4</sub> has ~28 times higher global warming potential (GWP) on centennial timescales, and even higher GWP on shorter time scales due to the shorter lifetime in the atmosphere (Stocker et al., 2013). Hence the knowledge of control mechanisms of CH<sub>4</sub> is important to predict and mitigate the future climatic and environmental changes.

Naturally, CH<sub>4</sub> is mainly produced by microbial decomposition by methanogens in anaerobic environments, such as waterlogged soil, wetlands, or sediments of lakes and rivers. Even though CH<sub>4</sub> can be oxidized and emitted as CO<sub>2</sub>, a considerable amount of CH<sub>4</sub> is still released into the atmosphere through vascular plants, diffusion and ebullition processes (e.g., Joabsson and Christensen, 2001). Geological CH<sub>4</sub> released from mud volcanoes and gas seepages through faults is the

second most important natural source (e.g., Etiope et al., 2008 and references therein). Additionally, a portion of CH<sub>4</sub> is produced by termites and wild animals via microbial digestive process (e.g., Sanderson, 1996), and by pyrogenic sources such as wildfire and biomass burning (e.g., Andreae and Merlet, 2001; Ferretti et al., 2005; Hao and Ward, 1993). The oceanic CH<sub>4</sub> flux is considered as too small to create a significant change in global budget compared to the other sources (Rhee et al., 2009).

5 The major sink of atmospheric CH<sub>4</sub> is photochemical reactions (oxidation) with the hydroxyl radical (OH), which is mainly controlled by atmospheric temperature, humidity, and mixing ratio of non-methane volatile organic compound (NMVOC) (e.g., Levine et al., 2011 and references therein). The air temperature affects air humidity, limiting the production of OH. Both NMVOCs and CH<sub>4</sub> are competing for OH to be oxidized, that is, increase in NMVOC emission reduces the available OH, so it increases the lifetime of CH<sub>4</sub> in the atmosphere (Valdes et al., 2005). Further, since the OH is produced by photo-dissociation

10 reaction, the CH<sub>4</sub> sink strength is affected by light availability and tropospheric ozone (e.g., Levy, 1971). Polar winters may affect the CH<sub>4</sub> sink strength by reduced OH production rate, but the seasonal-scale cycles are not resolvable in ice core records due to gas dispersion in firn layers. However, recent model studies suggested the dominant role of source changes rather than sink in controlling atmospheric CH<sub>4</sub> during the past climate changes (Weber et al., 2010; Levine et al., 2011). Therefore, the sink changes are not considered here.

15 Since the direct CH<sub>4</sub> monitoring of modern air samples only covers the late 20<sup>th</sup> and early 21<sup>st</sup> centuries (Dlugokencky et al., 1994, 2011), investigation further back in time requires the unique archive of polar ice that preserves the ancient atmospheric air. Paleoatmospheric CH<sub>4</sub> levels have been reconstructed for the last 800 ka from Antarctic- and Greenland ice cores. Given the relatively long lifetime in troposphere ( $11.2 \pm 1.3$  years at present, e.g., Prather et al., 2012) compared to atmospheric mixing time, ice core CH<sub>4</sub> records represent well-mixed global signatures. From the 800 ka records, it was revealed that the

20 past CH<sub>4</sub> change generally followed the glacial-interglacial cycles, being low during glacial periods and high in interglacials, as well as the shorter orbital cycles of obliquity and precession (e.g., Spahni et al., 2005; Loulergue et al., 2008). Those earlier studies have suggested that the changes in climate and hydrology on tropical wetlands induced by the orbital changes controlled the CH<sub>4</sub> emissions. The resemblance between water stable isotopes from Greenland ice cores, a proxy for Greenland climate change, and global CH<sub>4</sub> mixing ratio has been largely reported. This implies that local temperature change around Greenland

25 could affect the major CH<sub>4</sub> sources in low latitudes (e.g., Brook et al., 1996; Chappellaz et al., 1993; Huber et al., 2006). The similarity is also held in short time scale climate events. Previous works reported that rapid CH<sub>4</sub> increases were coincident with abrupt climate changes in Dansgaard-Oeschger (DO) events during the glacial period (e.g., Brook et al., 1996; EPICA Community Members, 2006; Grachev et al., 2007, 2009). The coincidence between abrupt CH<sub>4</sub> and Greenland climate change was also found in 8.2k cooling event, Preboreal Oscillation (PBO), Younger Drays (YD), and Bølling-Allerød (BA) periods

30 (Brook et al., 2000; Kobashi et al., 2007, 2008).

Intensive precipitation changes on the low latitude summer monsoon area by insolation changes (e.g., Asian monsoon) have been suggested as an important control during the glacial period (e.g., Chappellaz et al., 1990). From time series analysis of past CH<sub>4</sub> records, Guo et al. (2012) found that the tropical monsoon circulations are primary control of relatively shorter (millennial) time scale variability, while long-term (multi-millennial to orbital scale) variations are dominated by solar

insolation changes. It has been found that tropical monsoon activities were closely related to orbital-scale CH<sub>4</sub> change (e.g., Brook et al., 1996; Chappellaz et al., 1990), especially reported are Asian monsoon (e.g., Loulergue et al., 2008) and South American monsoon (e.g., Cruz et al., 2005). However, no direct correlation between CH<sub>4</sub> and tropical monsoon signals has been reported for the early Holocene, although demonstrated were the positive relationships between Greenland climate and tropical monsoons (e.g., Chiang et al., 2008), and between Greenland climate and CH<sub>4</sub> (e.g., Spahni et al., 2005; Wang et al., 2005; Mitchell et al., 2011) have been discussed. Given that waterlogged wetlands are the largest natural CH<sub>4</sub> source, this complex relationship may imply that tropical monsoons are not the sole, primary controls; wetlands in northern high latitude and southern hemisphere might act as a secondary role.

Relationship between the latitudinal shift of ITCZ and CH<sub>4</sub> emission varies with time scales. Landais et al. (2010) and Guo et al. (2012) suggested that ITCZ migration is not a dominant control of glacial-interglacial CH<sub>4</sub> cycle because long-term CH<sub>4</sub> trend does not follow well the precessional insolation change in the northern hemisphere. Modelling studies found the southward shift of ITCZ coincides with reduced CH<sub>4</sub> in Last Glacial Maximum (LGM) and Heinrich Stadial (HS) events, but changes in wetland area and surface hydrology were small (Weber et al., 2010; Hopcroft et al., 2011). They instead suggested that changes in temperature and/or plant productivity affected CH<sub>4</sub> production during those events. Rather, ITCZ migration appears to be related with millennial- or sub-millennial scale CH<sub>4</sub> change. Brook et al. (2000) found that submillennial-scale CH<sub>4</sub> minima during the last deglacial period correspond with reduced precipitation recorded in Cariaco Basin sediment data, which indicates southward displacement of ITCZ (Hughen et al., 1996). It is supported by spectral analysis of CH<sub>4</sub> during the past 800 ka record that found that ITCZ change becomes an important driver of millennial scale CH<sub>4</sub> change (Tzedakis et al., 2009; Guo et al., 2012).

The ice core scientists started to apply high-resolution CH<sub>4</sub> mixing ratio and stable isotope data to discern governing mechanisms of Holocene CH<sub>4</sub> variation, but currently the high-resolution records cover a part of the Holocene. High-resolution CH<sub>4</sub> records from Law Dome and WAIS Divide ice cores (Antarctica) show characteristic variability in multi-decadal to centennial time scale, apart from long-term gradual increasing trend (MacFarling-Meure et al., 2006; Mitchell et al., 2011). The high-resolution records were compared with various temperature- and precipitation proxies, but the previous works found no strong correlations that explain the observed decadal- to centennial scale variabilities. Limitation was that the late Holocene CH<sub>4</sub> budget may have been comprised of both natural and anthropogenic terms, making it difficult to distinguish between them. Mitchell et al. (2011) pointed out that some of abrupt CH<sub>4</sub> decreases could have resulted by historical events, such as Mongolian invasion, Plagues, or Spanish invasion. Later, Mitchell et al. (2013) made simultaneous measurement of Antarctic (WAIS Divide) and Greenland (GISP2) ices to derive IPD, and extended their high-resolution records back to ~4 ka. They used eight-box atmospheric methane model (EBAMM) and anthropogenic- and natural emission scenarios to investigate CH<sub>4</sub> control factors. Their results showed that the late Holocene CH<sub>4</sub> evolution can be explained by a combination of natural- and anthropogenic emissions. In the other hand, stable isotope ratios of CH<sub>4</sub> help us to distinguish the types of sources – biogenic, pyrogenic, and geologic. Sowers (2010) reconstructed CH<sub>4</sub> mixing ratio and stable isotope ratio of CH<sub>4</sub> ( $\delta^{13}\text{C-CH}_4$  and  $\delta\text{D-CH}_4$ ) throughout the entire Holocene and suggested several possible control factors, such as boreal wetlands and thermokarst

lakes, changing C<sub>3</sub>/C<sub>4</sub> plant ratio of CH<sub>4</sub>-emitting ecosystems, and changing composition of methanogenic communities. However, temporal resolution of the data (~138 years in average during 7.0 – 11.3 ka) was not sufficient to capture sub-millennial scale variability. Former studies have shown the reduction of pyrogenic emission and increased agricultural emission during the last millennia (Ferretti et al., 2005; Mischler et al., 2009). In later work using  $\delta^{13}\text{C}$ -CH<sub>4</sub> records from NEEM ice core, Sapart et al. (2012) found that the centennial-scale variations during the last two millennia were caused by changes in pyrogenic- and biogenic emissions. Ruddiman et al. (2011) and Sapart et al. (2012) estimated CH<sub>4</sub> emission change due to anthropogenic land use changes, which shows a good agreement with long-term CH<sub>4</sub> increasing trend. However, there is no high-resolution reconstruction of past population and land use area, and consequently large uncertainties of CH<sub>4</sub> emission from land use change impede identification of any shorter scale changes.

10 The early Holocene is suitable period to study natural CH<sub>4</sub> controls under Holocene interglacial climate condition. Since there was only negligible human population and relevant CH<sub>4</sub>-emitting anthropogenic activities (e.g., Goldewijk et al., 2010; Kaplan et al., 2011) during this time, the early Holocene CH<sub>4</sub> changes must have occurred mostly due to the natural causes. Understanding natural controls could contribute to better constrain the human-induced CH<sub>4</sub> changes. However, high-resolution study for the early Holocene has not been carried out extensively so far, except for the prominent cooling event at 8200 years BP (Spahni et al., 2003; Kobashi et al., 2007; Ahn et al., 2014). Earlier studies mainly focused on long-term change, attributing the major control to low latitude hydrology based on regional climate records that show wetter climate in tropics during the early Holocene (Blunier et al., 1995; Brook et al., 2000; Chappellaz et al., 1993, 1997). Therefore, in this study we present a new high-resolution CH<sub>4</sub> record during the early Holocene to investigate natural control mechanisms under interglacial condition. It should be noted that environmental boundary conditions of the early Holocene were not identical to those of the late Holocene. The global sea level was rising throughout the early Holocene and there were still remnant ice sheets in the north America.

## 2 Materials and Methods

In this study we used ice samples from a Siple Dome deep ice core (SDMA) drilled from 1997 to 1999 on the Siple Coast, West Antarctica (81.65°S, 148.81°W; 621m elevation) (Taylor et al., 2004). The SDMA samples were collected and cut at National Ice Core Laboratory (NICL, Denver, Colorado, USA) from January to February of 2013. Since brittle zone of SDMA ice starts below 400 m depth (Gow and Meese, 2007) that makes some part of ices fractured and/or cracked internally. Hence the samples were carefully collected from unbroken ices during the sample preparation at NICL. The samples were packed in isothermal foam boxes with numerous eutectic gels, and shipped to South Korea via expedited air freight. Temperature loggers were enclosed within the isothermal boxes to record the temperature change inside during the shipping, and it showed the temperatures were maintained below -25°C. The boxes were picked up directly just after custom clearance at the airport and then the ice samples were stored in a walk-in freezer at Seoul National University (SNU, Seoul, South Korea) that was

maintained below  $-20^{\circ}\text{C}$ . We measured 295 individual ice samples from 156 depth intervals from 518.87 to 718.83 m, covering from 8.36 to 20.25 ka after synchronizing to the Greenland Ice Core Chronology 2005 (GICC05, Rasmussen et al., 2006). All samples were duplicated, so that our final  $\text{CH}_4$  data were presented by averaging the results of duplicate analysis from the same depth and the analytical uncertainty of each data point is estimated by standard error of the mean of duplicate pairs. The air occluded in ice was extracted by a melting and refreeze process under vacuum. Ice samples were prepared in a walk-in freezer in the morning of each experiment day, and trimmed the outermost  $>2$  mm to eliminate potential contamination by ambient air during the storage. Then the samples were moved to the laboratory and placed in glass sample containers. The sample containers (sample flask hereafter) were custom-made glass flasks welded to stainless steel flange, and attached to the vacuum line with a copper gasket. The sample flasks were partially submerged into a chilled ethanol bath during ice insert and attaching to the line for preventing temperature increase by laboratory air. After that all flasks were evacuated at least 40 minutes, the ice samples were melted by submerging the sample flasks in a warm water bath. Melting process was usually completed within 30 minutes. The sample flasks were then submerged in the cold ethanol bath chilled to around  $-82^{\circ}\text{C}$  for more than an hour to refreeze. During the refreezing, we carried out GC pre-running (20 injections) and daily calibration that normally took 90 minutes. The ethanol temperature rose up to  $-55^{\circ}\text{C}$  just after submerging the flasks, and it was chilled to below  $-65^{\circ}\text{C}$  before expansion of the air in the flasks. The extracted air in the head space was expanded into a gas chromatograph (GC) equipped with a flame ionization detector (FID) to measure  $\text{CH}_4$  mixing ratio. The GC linearity was tested by a series of inter-tank calibration using four working standard air cylinders (395.5, 721.3, 895.0, and 1384.9 ppb  $\text{CH}_4$  in NOAA04 scale, Dlugokencky et al., 2005). Daily GC calibration curve was determined by measurements of a working standard having the closest  $\text{CH}_4$  mixing ratio of expected value from the samples; in this study we used 721.3 ppb  $\text{CH}_4$  standard for samples of the early Holocene. To account for system condition change throughout experiments (i.e., influence by water vapor), we calibrated with a standard air six times before and after sample measurements. The detailed configuration of the vacuum line and GC is described in another paper (Yang et al., in preparation).

To estimate our daily blank offset, we used four bubble-free blank ices every day, instead of interpolating between days of blank analysis (Mitchell et al., 2011). The blank ice was made by chilling the degassed ultrapure water (resistivity  $> 18.2$   $\text{M}\Omega\cdot\text{cm}$  at  $25^{\circ}\text{C}$ ) slowly from bottom in a closed stainless steel chamber. The daily blank offset is calculated from the mean of the four blank results ranging from 5 to 15 ppb, and it reflects any daily offsets by contaminants, leaks, and any different GC conditions. The exact cause of this blank offset is currently not clear, but the four blank results agree well each other, yielding the intra-day standard error of the mean of  $2.0 \pm 1.0$  ppb. This ‘intra-day’ blank offset is much smaller than the ‘inter-day’ offset, which implies that conditions of each sample flask are rather constant within a day, and vary systematically day by day. Since every single data point is obtained by analysis of at least in duplicate, the intra-day blank offset for one depth is reduced by a factor of  $\sqrt{2}$ . The robustness of our final results was proven by re-analysis of eight duplicates at adjacent depths ( $< 10$  cm) 8 to 80 days after the first analysis. The difference of the mean of duplicates between the time intervals was 1.9 ppb on average (pooled standard deviation of 1.4 ppb). The good reproducibility of our results demonstrates that our blank correction method is reliable.

Mass dependent (gravitational) fractionation within firn layer (Craig et al., 1988; Schwander, 1989) was corrected by using the nitrogen isotope ratio ( $\delta^{15}\text{N}$ ) of atmospheric  $\text{N}_2$  occluded in bubbles. Siple Dome  $\delta^{15}\text{N}$  records show a mean enrichment of  $0.23 \pm 0.01$  ‰ during the early Holocene (Severinghaus et al., 2009) and result in a slight decrease of  $\text{CH}_4$  by  $1.97 \pm 0.15$  ppb, which we added to all our measurements to correct the gravitational fractionation.

5 Different solubility of each gas species cause preferential dissolution of a gas having higher solubility than others, and consequently it makes the mixing ratio of extracted air different from that trapped originally within the ice (solubility effect hear after). As solubility of  $\text{CH}_4$  is higher than the other major component of air – nitrogen ( $\text{N}_2$ ), oxygen ( $\text{O}_2$ ), Argon (Ar), the solubility effect lowers the  $\text{CH}_4$  mole fraction of the extracted air and needs to be corrected properly. At the early stage of method development, we derive theoretically the solubility effect by using Henry's Law in a closed- and chemically  
10 equilibrated condition. After applying this theoretical solubility correction, we observed that SDMA  $\text{CH}_4$  data measured at SNU are lower than SDMA  $\text{CH}_4$  records from OSU by  $\sim 3$  ppb in average. Hence we compared the theoretical solubility correction with that obtained empirically from the second gas extraction (following the method described in Mitchell et al., 2011) and found that a correction factor of 1.0058 from the theoretically- to empirically-driven solubility effect. Further details on correction method will be discussed in our manuscript in preparation (Yang et al., in preparation).

15 Our new Siple Dome  $\text{CH}_4$  data has the currently third highest temporal resolution of Antarctic  $\text{CH}_4$  records covering the early Holocene after the WAIS Divide continuous ( $\sim 2$  years, Rhodes et al., 2015) and discrete ( $\sim 20$  years, WAIS Divide members, 2015) records. To check reliability of the record we compared our data set with previous SDMA measurement at Oregon State University (OSU) for 8.4 to 9.1 ka period when the two records are overlapped. The OSU  $\text{CH}_4$  record was measured with a temporal resolution of 8 years with precision of  $\pm 2.8$  ppb (Mitchell et al., 2011; Ahn et al., 2014). Resulting  
20 mean offset between two data set is  $\sim 0.6$  ppb, which lies within analytical uncertainty range of both institutes. Therefore, we merge the two records to make the SDMA  $\text{CH}_4$  composite data. Fig. 1 presents the SDMA data points during the early Holocene period (119 depths, 518.87 – 623.38 m).

### 3. Result and Discussion

#### 3.1 Millennial scale variability

25 To extract millennial-scale variability, we carried out spectral analysis using REDFIT program (Schulze and Mudelsee, 2002) and moderate (over 90% significance level) powers were found at  $\sim 1340$ , 401, 309, and 96-year period. Given the  $\sim 42$  years of gas age distribution of SDMA (Ahn et al., 2014), it would not reliable to study centennial scale variability. Thus we produced annual data by interpolation and then calculated 250-year running means to smooth high frequency components having shorter period than 309-year. Then the smoothed time series was filtered with a high-pass window (cut off period of  
30 1800 years) to study millennial scale variability throughout the early Holocene. For comparison, the same processing scheme

was applied to WAIS Divide time series and we observed that Siple Dome and WAIS Divide CH<sub>4</sub> anomalies share similar millennial scale variability, confirming reliability of both our data and observed millennial scale changes (Fig. 2).

The CH<sub>4</sub> anomalies demonstrate millennial scale minima at ~8.2, 9.3, 10.2 and 11.0 ka, which occurred in nearly 1000-year spacing. Each minimum is accompanied by depletion of water stable isotope ratio ( $\delta^{18}\text{O}_{\text{ice}}$ ) from NGRIP ice core, which implies climate cooling in Greenland. A close relationship between CH<sub>4</sub> and Greenland  $\delta^{18}\text{O}_{\text{ice}}$  has been previously reported in glacial-interglacial cycles and Dansgaard-Oeschger (DO) events during the last glacial period (e.g., Brook et al., 1996, 2000; Blunier and Brook, 2001; Chappellaz et al., 1993, 2013; EPICA Community Members, 2006). However, it has not been confirmed at interglacial climate conditions during the Holocene. Mitchell et al. (2011) found no significant correlation with Greenland climate in multi-decadal scale during the late pre-industrial Holocene (LPIH), possibly because LPIH CH<sub>4</sub> budget is also affected substantially by anthropogenic emissions (e.g., Ferretti et al., 2005; Mischler et al., 2009; Mitchell et al., 2013; Sapart et al., 2012). In contrast, we observe a significant positive correlation ( $r = 0.57$ ,  $p = 0.06$ ) between the millennial-scale change of Siple Dome CH<sub>4</sub> and NGRIP  $\delta^{18}\text{O}_{\text{ice}}$  during the early Holocene. The correlation coefficient between the smoothed- and filtered time series of SDMA CH<sub>4</sub> (before synchronization to GICC05) and NGRIP  $\delta^{18}\text{O}_{\text{ice}}$  was calculated for the 7.8-11.5 ka by interpolating to the original ages of SDMA CH<sub>4</sub> composite, with a reduced degree of freedom.

The previous SDMA gas chronology (Brook et al., 2005; Severinghaus et al., 2009) was synchronized to GICC05 age scale by setting age tie-points with stable water isotope ( $\delta^{18}\text{O}$ ) record from the North Greenland Ice Core Project (NGRIP) ice cores during the abrupt climate change events of the Preboreal Oscillation (PBO) and the 8.2 ka event, given that both events have been proved to be synchronous with CH<sub>4</sub> change (Kobashi et al., 2007, 2008). The synchronization between the tie points was done by linear interpolation of age differences between the synchronized- and the previous ones, which range from -114 to 28 years. After synchronizing to GICC05 scale, the correlation coefficient between SDMA CH<sub>4</sub> composite and NGRIP  $\delta^{18}\text{O}_{\text{ice}}$  increases to  $r = 0.74$  ( $p < 0.01$ ) (Fig. S1).

According to atmospheric modelling studies, abrupt cooling in the North Atlantic regions can alter atmospheric circulation and to cause southward migration of the mean latitudinal position of the Intertropical Convergence Zone (ITCZ) (e.g., Chiang and Bitz, 2005; Broccoli et al., 2006; Cvijanovic and Chiang, 2012). The climatic teleconnection between northern North Atlantic and low latitude regions is shown by climate proxies. Sediment reflectance record from Cariaco Basin shows increased rainfall and humidity – which is due to southward displacement of ITCZ – corresponding to the 8.2, 9.3, and 10.9 ka abrupt cooling event, as revealed in previous studies for the different time periods (Peterson et al., 2000; Haug et al., 2001; Fleitmann et al., 2007; Deplazes et al., 2013). The southward displacement of ITCZ leads further weakening of Asian and Indian summer monsoons and probably reduces CH<sub>4</sub> emission from northern tropical wetlands. The <sup>18</sup>O enrichments of speleothem in Dongge Cave (China), Qunf Cave (Oman), and Hoti Cave (Oman, not shown, Neff et al., 2001) occurred at similar timing with abrupt cooling in Greenland at 8.2, 9.3, and 10.9 ka, which indicate the reduction of monsoonal rainfall at northern tropical wetlands. The speleothem records from Chinese and Oman caves seem to lag by ~100-200 years after the CH<sub>4</sub> change at ~9.3 ka, but this lies within chronological uncertainties of ~200-400 years at around ~9.0 ka (Dykoski et al., 2005; Fleitmann et al., 2007).

Moreover, sediment Ba/Ca ratio from Gulf of Guinea demonstrates concurrent decrease of west African monsoon (Weldeab et al., 2007). The record indicates that precipitation over the major wetland area was reduced and in turn it would lower the wetland CH<sub>4</sub> emissions in NH. In the meanwhile, an inverse relationship is observed from the Eastern Brazilian speleothem data (Lapa Grande Cave, Strikis et al., 2011) that demonstrate the increase of precipitation at the time of abrupt CH<sub>4</sub> decrease  
5 occurred as a result of southward migration of ITCZ. Rhodes et al. (2015) pointed out that strong southward migration of ITCZ could induce an abrupt CH<sub>4</sub> increase from southern hemisphere during the HS 1, 2, 4, and 5 events. Sperlich et al. (2015) also found that a sharp CH<sub>4</sub> peak at Greenland Interstadial 21.2 (~85 ka) was occurred by emission from Asian and South American wetlands. However, considering the orbital parameters that indicate maximum summer insolation in NH while minimum in SH during the early Holocene, it can be inferred that contribution of SH wetland emission was relatively weak and  
10 overcompensated by reduction of NH emission.

The possibility that the observed CH<sub>4</sub> minima were caused by reduction of northern extra-tropical sources is not supported by previous modelling study. Zürcher et al. (2013) found that abrupt cooling in Greenland and northern high latitudes by large freshwater input causes boreal peatland CH<sub>4</sub> emission to decrease substantially, which explains ~23% of abrupt CH<sub>4</sub> decrease (~80 ppb) during the 8.2 ka event. If we assume linear scaling of the model response, it implies that boreal peatland source  
15 change only accounts for ~23% of total CH<sub>4</sub> change during the rest of CH<sub>4</sub> decrease events. Given the meltwater pulses during the early Holocene before the 8.2 ka event are more than 10 times weaker (Teller and Leverington, 2004) than that corresponding to the 8.2 ka event, we consider the boreal emission change is not the major cause of CH<sub>4</sub> local minima.

Previously, Björck et al. (2001) found the climate cooling in the northern Atlantic and Santa Barbara Basin occurred with solar-forcing change at ~10.3 ka. However, the proxy data in Figure 2 show no clear indication of southward migration of  
20 ITCZ position and changes in Asian, Indian, African, and South American summer monsoon intensity associated to ~10.2 ka cooling (Fig. 2b-f). Moreover, there was no distinct change in  $\Delta\epsilon_{\text{LAND}}$  at that time, a proxy of global terrestrial respiratory fractionation of atmospheric O<sub>2</sub>, which is affected by low latitude surface hydrology (Severinghaus et al., 2009). This paleoproxy record suggests that changes in precipitation and surface hydrology in the northern tropics may have not changed significantly during around the 10.2 ka. Instead, there are two small decreases at ~9.9 and ~10.6 ka as shown in Dongge cave  
25 deposit record (Fig. 2d). These episodes are not likely associated with the CH<sub>4</sub> minimum at 10.2 ka because the timing differences between the CH<sub>4</sub> minimum and reductions of Asian summer monsoon intensity are beyond the chronological uncertainty. The age uncertainty of Dongge Cave deposits is  $\pm 77$  years (2 sigma error; Dykoski et al., 2005), and the estimated error of SDMA CH<sub>4</sub> gas age in this study is less than ~150 years (see Supplement and Fig. S3). The climate teleconnection between North Atlantic and tropical hydrology at 10.2 ka might not have been sufficiently strong enough to change the low  
30 latitude climate. Weak cooling around the North Atlantic region can be a candidate, given that NGRIP  $\delta^{18}\text{O}_{\text{ice}}$  records demonstrate smaller amplitude negative anomaly during ~10.2 ka event than those of 8.2 and 9.3 ka. The amplitude of  $\delta^{18}\text{O}_{\text{ice}}$  changes at 10.2 ka of the high-pass filtered GRIP and GISP2 records does show smaller variability than those at 8.2 and 9.3 ka cooling events, but larger than the variability at 10.9 ka (Fig. S2). Hence, the “weak cooling” speculation is not fully supported by the other Greenland ice core records. Although there appears to have been no strong change in low latitude



hydrology at 10.2 ka, the amplitude of CH<sub>4</sub> decrease at 10.2 ka is similar order to the other millennial events. Given that no clear reduction of the Asian, Indian, and African monsoon intensity is observed, it may imply the CH<sub>4</sub> decrease at 10.2 ka was controlled by other processes than the monsoon circulation change. If the climate proxy from Dongge cave reflects rather regional climate changes, monsoonal rainfalls and surface hydrology of other regions could be responsible for CH<sub>4</sub> decrease.

5 There should be an ultimate cause of the CH<sub>4</sub> and climate change in the early Holocene. Previous works have suggested an important role of solar forcing during the Holocene (e.g., Björck et al., 2001; Bond et al., 1997, 2001). Bond et al. (1997) reported four large ice-rafted debris (IRD) drifts occurred at ~8.1, 9.4, 10.3 and 11.1 ka caused by surface cooling of North Atlantic Ocean. They found that the ocean surface cooling and the IRD events are closely related to cooling over the Greenland. Figure 2 shows that each IRD event (maxima in hematite stained grain) occurred concurrently with minima of NGRIP  $\delta^{18}\text{O}_{\text{ice}}$  record within age uncertainty. Then the Greenland cooling leads southward shift of ITCZ and in turn it changes wetland CH<sub>4</sub> emission in low latitudes. Bond et al. (2001) found that IRD maxima during the Holocene coincide with solar activity minima. The authors suggested that solar forcing could affect the climate change around the North Atlantic Ocean (and Greenland), through amplification by changes in sea ice and/or deep water formation. A close interplay between solar activity and monsoon intensity has been observed in previous studies using the Chinese and Oman speleothem records during the Holocene (Neff et al., 2001; Wang et al., 2005; Gupta et al., 2005), even on multi-decadal time scales (Agnihotri et al., 2002). However, the forcing mechanism of solar activity on the North Atlantic and global climate is not well understood. Jiang et al. (2015) found positive correlations between North Atlantic SST and solar forcing inferred from plaeoproxies (<sup>14</sup>C and <sup>10</sup>Be) for the last 4000 years, while the correlation disappears during the mid- and early Holocene. They hypothesized that climate sensitivity to solar forcing is high for cooler climate. As evidenced above, the early Holocene CH<sub>4</sub> minima were likely triggered by anomalous low solar activity, but future study is needed to make it more conclusive.

20 Meanwhile, shifting to El Niño-like SST condition was suggested as another mechanism that changes tropical rainfall pattern (Marchitto et al., 2010). According to modern atmospheric observation, El Niño condition leads drying conditions in low latitude wetlands in Africa, Asia, and America (e.g., Dai and Wigley, 2000; Lyon and Barnston, 2005; Hodson et al., 2011), which reduces tropical CH<sub>4</sub> emissions. Thus, we could speculate that both the ITCZ migration and El Niño-like SST change affected the tropical surface hydrology and CH<sub>4</sub> emission. According to Holocene ENSO activity reconstructions by Moy et al. (2002), no ENSO event was recorded during the early Holocene until around 7 ka, except weak ENSO events during 10.4-10.1 ka, where abrupt CH<sub>4</sub> decrease is observed without significant changes in ITCZ and NH monsoon intensities. Mitchell et al. (2011) observed a significant positive correlation between CH<sub>4</sub> and Pacific Decadal Oscillation (PDO) variability during the late Holocene. It has been reported that PDO modulates the wet/dry impact of ENSO depending on phase relationship between ENSO and PDO (e.g., Wang et al., 2014 and references therein). The Holocene PDO reconstruction from sediment grain size analysis by Kirby et al. (2010) shows PDO-related drying intervals in North America during 9.5-9.1, 8.9-8.6, and 8.3-7.8 ka, which overlap the CH<sub>4</sub> minima at 8.2 and 9.3 ka present in this study.

### 3.2. Inter-polar difference of CH<sub>4</sub> during the early Holocene

We calculated inter-polar difference (IPD) of CH<sub>4</sub> to trace the latitudinal source distribution change during the early Holocene. The currently available high-resolution CH<sub>4</sub> records covering the early Holocene are SDMA discrete (this study), WAIS Divide discrete (WAIS Divide project members, 2015), WAIS Divide continuous (Rhodes et al., 2015), NEEM discrete (Chappellaz et al., 2013) and NEEM continuous data (Chappellaz et al., 2013). Among the Antarctic records, we consider WAIS continuous records are more reliable than WAIS discrete ones because WAIS discrete data covering the early Holocene were measured in a different institute (Penn State University, PSU), showing a pooled standard deviation of ~7.3 ppb (1 $\sigma$ ). In addition, there is an unexplained offset between WAIS CH<sub>4</sub> measured in OSU and PSU lab by ~9.9 ppb (Rhodes et al., 2015), which is larger than that observed between SNU and OSU SDMA data sets. Regarding the Greenland side, we use NEEM discrete records because there are discrepancies between continuous- and discrete data in some intervals, but also because the NEEM continuous record is not exactly “continuous”. Hence, here we regard the NEEM discrete, Siple Dome discrete, and WAIS Divide continuous data as more reliable ones than the others to reconstruct IPD during the early Holocene.

Precise synchronization is crucial for direct comparison between data sets which have high frequency variations. For synchronizing between Antarctic (Siple Dome and WAIS Divide continuous) and NEEM records, the NEEM CH<sub>4</sub> record (~11-year resolution on average) is chosen as reference. Synchronization was done by two steps: First, we made initial synchronization between the Antarctic and NEEM data by setting match points at the midpoint of abrupt CH<sub>4</sub> change, and then we linearly interpolated the age offset of each match point for the rest of data points. Then we applied a Monte Carlo simulation to find a maximum correlation. Both data sets were resampled every 30 years, and each point was randomly disturbed (assuming a normal distribution with 1 sigma of 30 years). By doing so 1000 different time series were created, and one set having a maximum correlation with NEEM data was chosen. Criteria for “best fit” is correlation coefficient of 0.8 with NEEM original age scale, so that a maximum correlation less than 0.8 was discarded. This procedure was repeated to make 20 sets of maximum correlation time series, and the mean ages of 20 replicate simulations were set to synchronized age scale. Temporal uncertainty (synchronizing error) was determined for each point as 1 standard deviation of 20 replicates and CH<sub>4</sub> uncertainty includes analytical error of the both records (4.3 ppb for NEEM discrete /1.4 ppb for SDMA / 1.5 ppb for WAIS continuous, 1 sigma). Since IPD calculation is very sensitive to high frequency variability of CH<sub>4</sub> records from both poles, and it is difficult to reconstruct reliable IPD in short time scales, all IPD records in this study were filtered by a 1000-year low-pass window to discuss multi-centennial to millennial scale change. As discrete measurements are regarded as more accurate than continuous ones in absolute sense, WAIS continuous data were calibrated against to Siple Dome data instead of WAIS discrete record, because Siple Dome records were more rigorously tested for its reproducibility, and also compared with OSU measurements that shows small offset during the early Holocene interval.

The IPDs from those three data sets are plotted in grey (NEEM discrete – Siple Dome, IPD-1 hereafter) and green (NEEM discrete – WAIS continuous, IPD-2 hereafter) in Figure 3. Resulting IPD-1 and IPD-2 show long-term increase from 11.5 to 9.9 ka, which indicates that boreal source contribution enhanced. However, IPD-1 shows a sharp increase during PBO followed

by decrease until ~10.7 ka, and in the latter case both IPDs differ beyond 95% envelope (from 10.4 to 10.8 ka). This is because Siple Dome discrete data are higher than WAIS continuous data during this interval. Similarly, the small peak of IPD-1 at ~11.1 ka that is not seen in IPD-2 may be caused by offset between Siple- and WAIS data during 11.0 to 11.3 ka. By using our new IPDs and the reliable highly resolved CH<sub>4</sub> records (NEEM discrete – SDMA discrete / WAIS continuous), we ran a simple 3-box CH<sub>4</sub> source distribution model to quantify how much the boreal and tropical source strengths were changed. Here we used the same box model employed in Chappellaz et al. (1997) and Brook et al. (2000). Briefly, the model contains 3 boxes; northern extra-tropical latitude (30-90°N, N-box), tropical (30°S-30°N, T-box), and southern extra-tropical latitude boxes (30-90°S, S-box). CH<sub>4</sub> mixing ratios in 3 boxes (in Tg box<sup>-1</sup>) were determined from CH<sub>4</sub> mixing ratio of Antarctica and Greenland. The mean CH<sub>4</sub> mole fraction of N-box (30-90N) is not identical to that of Greenland ice core record, given the latitudinal CH<sub>4</sub> distribution (e.g., Fung et al., 1991). To derive the N-box CH<sub>4</sub>, we followed the assumption of Chappellaz et al. (1997), where the authors assumed that difference between Greenland and the mean N-box CH<sub>4</sub> is 7% of IPD. Hence here the N-box CH<sub>4</sub> is calculated by subtracting 7% of IPD from the Greenland mixing ratio. T-box mixing ratio is inferred by assuming that the S-box emission is constant of 15 Tg yr<sup>-1</sup> (Fung et al., 1991). Emission from each box (Tg yr<sup>-1</sup>) is then estimated by using the mixing ratios of the boxes, lifetime of CH<sub>4</sub> in each box, and transport times among the boxes. The modelled emission changes in NH extratropical- and tropical boxes from IPD-1 are plotted in Figure 4 (see Fig. S3 for results from IPD-2). The model reveals decreasing tropical sources (accounting for the largest portion in CH<sub>4</sub> budget), while enhancing NH extratropical emissions. The tropical emission was elevated by ~8 Tg yr<sup>-1</sup> from the onset of the Holocene to its maximum at 10.6 ka, followed by ~15 g yr<sup>-1</sup> reduction to ~111 Tg yr<sup>-1</sup> at 9.5 ka. Tropical emission decrease is also observed in IPD-2 from 134 to 115 Tg yr<sup>-1</sup> during the 11.5-10.0 ka, but this change is not significant in 95% confidence range (Fig. S3 and Table S2). The long-term decreasing trend follows the NH summer insolation change. This covariation may reflect the insolation-driven changes in multi-millennial timescale (e.g., Loulergue et al., 2008; Guo et al., 2012). CH<sub>4</sub> flux from NH extratropical box increased from ~57 Tg yr<sup>-1</sup> (11.5 ka) to ~70 Tg yr<sup>-1</sup> (9.5 ka), showing a local minimum of ~63 Tg yr<sup>-1</sup> at 10.7 ka. Also plotted in Figure 4 is boreal source fraction, defined as ratio of N-box emission to total source emissions. It shows a significant increase from ~30% at 11.5 ka to ~35% at 9.5 ka. The box model results from IPD-2 demonstrate increase of NH extratropical emission from 60 to 71 Tg yr<sup>-1</sup>, and hence increase of the boreal source fraction from 29 to 35% during the 11.5 to 10.0 ka interval.

This conclusion is supported by proxy-based temperature reconstructions that indicate a gradual warming in northern extratropical region (30N-90N) until ~9.6 ka, while tropical temperature remains stable (Marcott et al., 2013). The climate warming in northern high latitudes caused ice sheet retreat (e.g., Dyke, 2004) and enhanced CH<sub>4</sub> emission from boreal permafrost by forming new wetlands in mid- to high latitudes (e.g., Gorham et al., 2007; Yu et al., 2013) and accelerating microbial decomposition of organic materials (e.g., Christensen et al., 2004; Schuur et al., 2015). Thermokarst lakes created by thawing ice wedges and ground ices in Alaskan- and Siberian permafrost has been suggested as a source of CH<sub>4</sub> (e.g., Walter et al., 2006, 2007; Brosius et al., 2012). The modelled enhancement of NH extratropical emission of ~13 Tg yr<sup>-1</sup> is greater than the CH<sub>4</sub> release of 8.2 Tg yr<sup>-1</sup> from thermokarst lake thawing, which is estimated based on present-day observations (Walter et al., 2014). Since most thermokarst lakes are located in NH high latitude regions (e.g., Walter et al., 2006; 2014), it

may indicate that other sources, such as northern peatlands or mid-latitude wetlands, should contribute to increasing NH extratropical emission. Our results are consistent with previous findings of CH<sub>4</sub> stable isotope analysis. Fischer et al. (2008) found that increase of boreal source contribution is required to explain the more depleted  $\delta^{13}\text{C-CH}_4$  during Preboreal Holocene than the Younger Dryas interval. Sowers (2010) extended the CH<sub>4</sub> isotopic ratio into the entire Holocene that displayed a gradual decrease of  $\delta^{13}\text{C-CH}_4$  by  $\sim 2$  ‰ from 10.5 to 4 kyr BP, which was attributed to progressive expansion of NH high latitude sources.

### 3.3. Comparison with late Holocene variability

We compared amplitude of CH<sub>4</sub> variability between the early- and the late Holocene in multi centennial to millennial time scales. Figure 5 shows amplitude spectrum and root mean square (RMS) amplitude for the early Holocene and the late Holocene, respectively. The amplitude of the early Holocene CH<sub>4</sub> change is  $\sim 10$  ppb and does not change greater except for PBO and the 8.2 ka event, while the late Holocene spectrum shows smaller amplitude than early Holocene for shorter-term change and larger for longer-term fluctuation. Late Pre-Industrial Holocene (LPIH) CH<sub>4</sub> amplitude is elevated to early Holocene level from  $\sim 0$  C.E. ( $\sim 2.0$  ka), and increases up to higher from  $\sim 1450$  C.E. ( $\sim 0.5$  ka).

The reason of low amplitude variability during 3.5 to 1.2 ka, or why the early Holocene CH<sub>4</sub> variability is larger than this period, is probably related to different orbital configuration in both time periods. Previous studies found covariation between CH<sub>4</sub> amplitude and NH summer insolation change, reflecting that mean temperature of the warmest seasons is an important factor of CH<sub>4</sub> emission, during the interstadial conditions (Flückiger et al., 2004; Baumgartner et al., 2014). Combined with elevated summer insolation in Northern Hemisphere (NH) and with climate warming in NH extratropics, the amplified variability of the early Holocene may suggest that CH<sub>4</sub> control by NH wetlands was likely stronger than the late Holocene period. Meanwhile, lower summer insolation during the late Holocene might induce diminished CH<sub>4</sub> amplitude. This evidence indicates the natural forcing in centennial- to millennial time scales is reduced in the late Holocene, given that the atmospheric CH<sub>4</sub> budget during 3.5-1.2 ka (604.9 ppb) is similar to that during 9.0-7.6 ka (628.6 ppb), and that anthropogenic emission is greater in later Holocene than the early Holocene. Abrupt increase of CH<sub>4</sub> amplitude since  $\sim 800$  C.E. (1.2 ka) is likely driven by increasing anthropogenic contribution, which is consistent with anthropogenic emission scenario based on past population and agricultural activity (Mitchell et al., 2013). Also superimposed are short-term cooling events during Little Ice Age, making CH<sub>4</sub> variability greater.

### 4. Conclusion and summary

We reconstructed a new high resolution CH<sub>4</sub> record during the early Holocene from Siple Dome ice core, Antarctica, to study millennial-scale CH<sub>4</sub> variability and its natural controls under Holocene interglacial condition. Since the new SDMA data agree well with previous measurements at OSU, we made SDMA CH<sub>4</sub> composite data covering  $\sim 7.7$  to 11.6 ka. We found

four millennial scale CH<sub>4</sub> minima accompanied with Greenland cooling, changes in ITCZ position and reduced Asian and Indian monsoon intensities. The observed evidences suggest that low latitude hydro climate changes were closely related to millennial scale CH<sub>4</sub> minima. Further, this study presented the millennial scale change of IPD, which was calculated from high resolution discrete dataset of NEEM and SDMA. Here we reported for the first time the IPD increase from the onset of the  
5 Holocene to ~9.9 ka following the temperature rise in NH extra-tropical regions. The three-box model demonstrates that elevated emission from NH extratropics and reduction of tropical sources, resulting the increased contribution of the NH extra-tropical sources. Finally, we observed that RMS amplitude of earlier part of the late Holocene is smaller than that of the early Holocene, which may be attributed to different orbital parameters.

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### **Data availability**

The early Holocene Siple Dome CH<sub>4</sub> data will be available on NOAA Paleoclimatology database and PANGAEA data  
20 repository.

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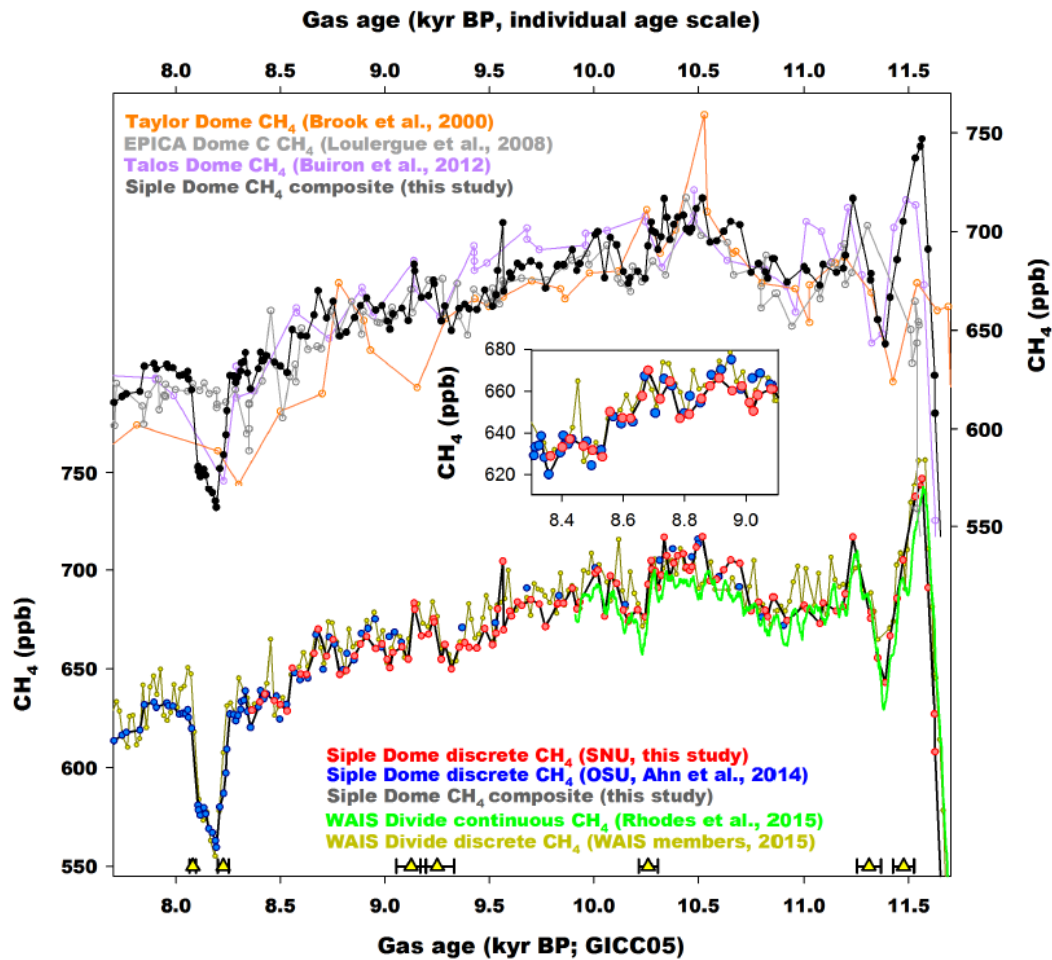


Figure 1. Atmospheric CH<sub>4</sub> concentration reconstructions during the early Holocene. Top: new high-resolution Siple Dome composite (black, this study and Ahn et al., 2014) compared with previous records from Taylor Dome (orange, Brook et al., 2000), EPICA Dome C (grey, Louergue et al., 2008), and Talos Dome (purple, Buiron et al., 2012). Bottom: Siple Dome CH<sub>4</sub> records measured in OSU (blue, Ahn et al., 2014) and SNU (red, this study). Siple Dome composite (black line) is plotted with WAIS Divide discrete (dark yellow, WAIS Divide project members, 2015) and continuous measurement records (green, Rhodes et al., 2015). Inset: Enlarged plot showing overlapped interval between OSU and SNU Siple Dome data.

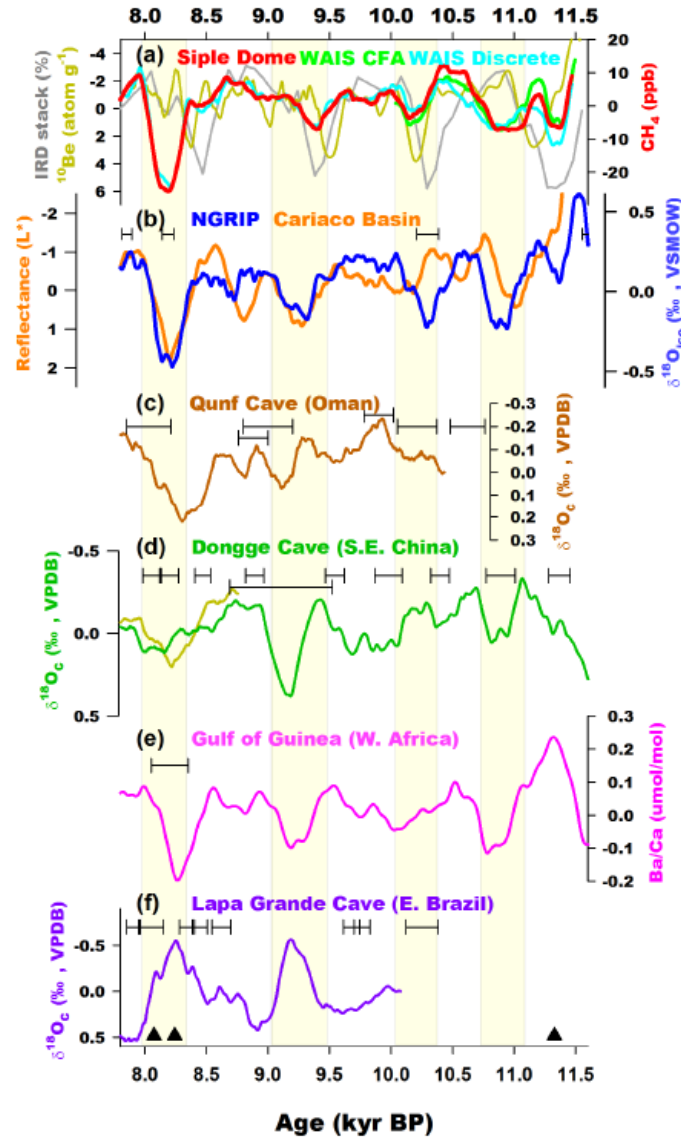


Figure 2. Millennial scale climate variability. All proxies present here were smoothed by 250-year running average and detrended by high-pass filter with 1/1800-year window. (a) Siple Dome CH<sub>4</sub> (red, this study), Greenland <sup>10</sup>Be (dark yellow, Finkel and Nishizumii, 1997), North Atlantic IRD stack (grey, Bond et al., 2001). Also shown are WAIS Divide CH<sub>4</sub> data by discrete (cyan, denoted “WAIS Discrete”, WAIS Divide project members, 2015) and continuous (yellow green, denoted “WAIS CFA”, Rhodes et al., 2015) technique. (b) NGRIP stable water isotope ratio (blue, Rasmussen et al., 2006) and Cariaco Basin reflectance (orange, Deplazes et al., 2013). (c) Qunf Cave speleothem oxygen isotope (Fleitmann et al., 2007). (d) Dongge Cave speleothem oxygen isotope (green, Dykoski et al., 2005; dark yellow, Wang et al., 2005). (e) Gulf of Guinea planktonic Ba/Ca ratio (Weldeab et al., 2007). (f) Lapa Grande Cave speleothem oxygen isotope (purple, Strikis et al., 2011). Black solid triangles are age tie-points used to adjust Siple Dome and WAIS Divide CH<sub>4</sub> data to GICC05 scale.

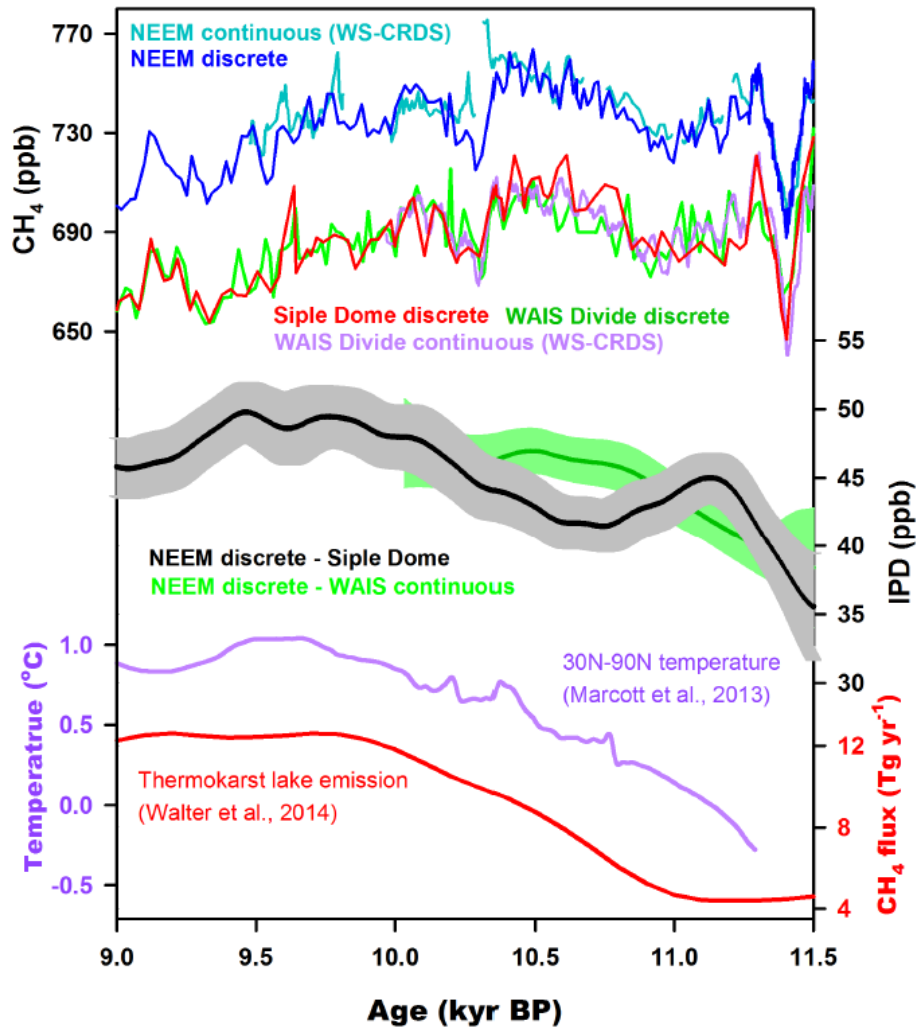


Figure 3. CH<sub>4</sub> inter-polar difference (IPD) and high latitude CH<sub>4</sub> sources. Top: High resolution CH<sub>4</sub> discrete measurements from NEEM discrete (blue, Chappellaz et al., 2013), NEEM continuous (turquoise, Chappellaz et al., 2013), WAIS Divide discrete (yellow green, WAIS Divide project members, 2015), WAIS Divide continuous (purple, Rhodes et al., 2015), and Siple Dome (red, this study) ice core records. Middle: 1000-year low-pass filtered IPD-1 (black) and IPD-2 (green) with 95 % significance interval (shaded). Bottom: Proxy-based temperature reconstruction for 30°N-90°N and 30°S-30°N latitude (blue, Marcott et al., 2013). CH<sub>4</sub> flux estimate from Siberian- and Alaskan thermokarst lakes (red, Walter-Anthony et al., 2014).

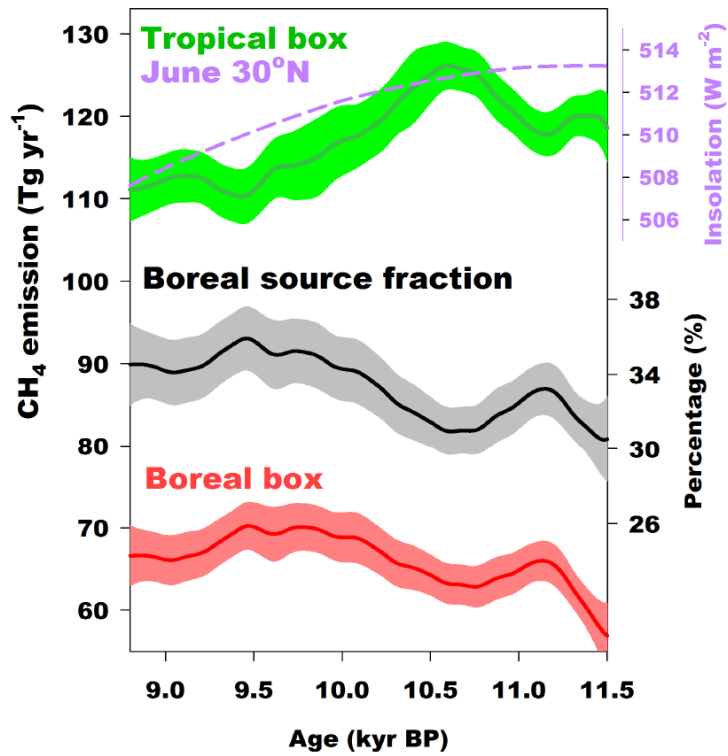


Figure4. 3-box source distribution model results of tropical (green) and boreal (red) boxes. Black line shows the [boreal to total source fraction](#) (see text). [Purple dashed line plotted with tropical emission is summer insolation in 30°N \(Berger and Loutre, 1991\).](#)



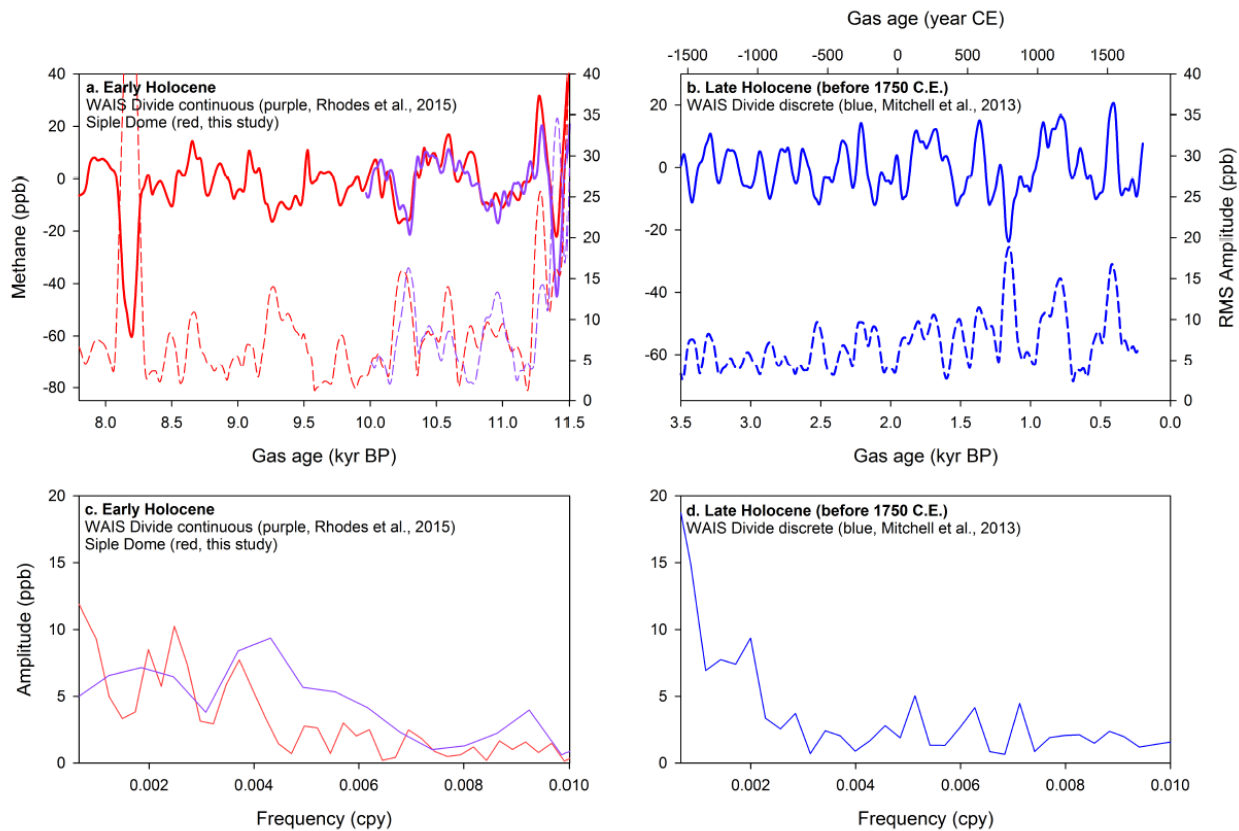


Figure 5. Upper: Detrended (75 to 1800-year band-pass filtered) CH<sub>4</sub> for the early (a) and late (b) Holocene from Siple Dome (red, this study), WAIS divide continuous (purple, Rhodes et al., 2015), and WAIS divide discrete (blue, Mitchell et al., 2013) data. Dashed lines are root mean square (RMS) amplitude running averaged by 75-year window. Lower: Amplitude spectrum of Early (c) and Late (d) Holocene CH<sub>4</sub> records. Note that CH<sub>4</sub> data before 1750 C.E. are used for the preindustrial late Holocene.

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**Table 1. 3-box source distribution model results of tropical (green, T) and boreal (red, N) boxes and boreal source fraction (N/(T+N+S)) compared with previous results. Errors denote 95% confidence interval.**

<u>Ref.</u>	<u>N box</u>	<u>T box</u>	<u>Boreal source fraction</u> <u>N/(N+T+S)</u>
<u>(ka)</u>	<u>(Tg yr<sup>-1</sup>)</u>		<u>(%)</u>
<u>Brook et al., 2000</u> <u>(9.5-11.5 ka)</u>	<u>64 ± 5</u>	<u>123 ± 8</u>	<u>32 ± 3</u>
<u>Chappellaz et al., 1997</u> <u>(9.5-11.5 ka)</u>	<u>66 ± 8</u>	<u>120 ± 9</u>	<u>33 ± 3</u>
<u>This study</u> <u>(9.5-11.5 ka)</u>	<u>66 ± 4</u>	<u>120 ± 4</u>	<u>33 ± 2</u>
<u>This study</u> <u>(11.5 ka)</u>	<u>47 ± 9</u>	<u>149 ± 10</u>	<u>22 ± 5</u>
<u>This study</u> <u>(9.9 ka)</u>	<u>65 ± 8</u>	<u>119 ± 9</u>	<u>33 ± 5</u>

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