



1 **Hydrological variations of the intermediate water masses of**  
2 **the western Mediterranean Sea during the past 20 ka inferred**  
3 **from neodymium isotopic composition in foraminifera and**  
4 **cold-water corals**

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27 **Abstract.** The neodymium isotopic composition ( $\epsilon\text{Nd}$ ) of mixed planktonic foraminifera species and  
28 scleractinian cold-water corals (CWC; *Madrepora oculata*, *Lophelia pertusa*) collected at 280-620 m water  
29 depth in the Balearic Sea, the Alboran Sea and the south Sardinian continental margin was investigated to  
30 constrain hydrological variations at intermediate depths in the western Mediterranean Sea during the last 20 ka.  
31 Planktonic (*Globigerina bulloides*) and benthic (*Cibicides pachyderma*) foraminifera were also analyzed for  
32 stable oxygen ( $\delta^{18}\text{O}$ ) and carbon ( $\delta^{13}\text{C}$ ) isotopes. The foraminiferal and coral  $\epsilon\text{Nd}$  values from the Balearic Sea  
33 and the Alboran Sea are comparable over the past ~13 ka, with mean values of  $-8.94 \pm 0.26$  ( $1\sigma$ ;  $n=24$ ) and -  
34  $8.91 \pm 0.18$  ( $1\sigma$ ;  $n=25$ ), respectively. Before 13 ka BP, the foraminiferal  $\epsilon\text{Nd}$  values are slightly lower (-  
35  $9.28 \pm 0.15$ ) and tend to reflect a higher mixing between intermediate and deep waters, characterized by more  
36 unradiogenic  $\epsilon\text{Nd}$  values. The slight  $\epsilon\text{Nd}$  increase after 13 ka BP is associated to a marked difference in the  
37 benthic foraminiferal  $\delta^{13}\text{C}$  composition of intermediate and deeper depths, which started at ~16 ka BP. This  
38 suggests an earlier stratification of the water masses and a subsequent reduced contribution of unradiogenic  $\epsilon\text{Nd}$   
39 from deep waters. The CWC from the Sardinia Channel show a much larger scattering of  $\epsilon\text{Nd}$  values, from -  
40  $8.66 \pm 0.30$  to  $-5.99 \pm 0.50$ , and a lower average ( $-7.31 \pm 0.73$ ;  $n=19$ ) compared to the CWC and foraminifera from  
41 the Alboran Sea and Balearic Sea, indicative of intermediate waters sourced from the Levantine basin. At the  
42 time of sapropel S1 deposition (10.2 to 6.4 ka), the  $\epsilon\text{Nd}$  values of the Sardinian CWC become more unradiogenic  
43 ( $-8.38 \pm 0.47$ ;  $n=3$  at ~8.7 ka BP), suggesting a significant contribution of intermediate waters originated from the  
44 western basin. Accordingly, we propose here that western Mediterranean intermediate waters replaced the  
45 Levantine Intermediate Water (LIW), which was strongly reduced during the mid-sapropel (~8.7 ka BP). This



46 observation supports a notable change of Mediterranean circulation pattern centered on sapropel S1 that needs  
47 further investigations to be confirmed.

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#### 49 **1. Introduction**

50 The Mediterranean Sea is a mid-latitude semi-enclosed basin, characterized by evaporation exceeding  
51 precipitation and river runoff, where the inflow of fresh and relatively warm surface Atlantic water is  
52 transformed into saltier and cooler (i.e. denser) intermediate and deep waters. Several studies have demonstrated  
53 that the Mediterranean thermohaline circulation was highly sensitive to both the rapid climatic changes  
54 propagated into the basin from high latitudes of the Northern Hemisphere (Cacho et al., 1999, 2000, 2002;  
55 Moreno et al., 2002, 2005; Paterne et al., 1999; Martrat et al., 2004; Sierro et al., 2005; Frigola et al., 2007,  
56 2008) and orbitally-forced modifications of the eastern Mediterranean freshwater budget mainly driven by  
57 monsoonal river runoff from the south (Rohling et al., 2002; 2004; Bahr et al., 2015). A link between the  
58 intensification of the Mediterranean Outflow Water (MOW) and the intensity of the Atlantic Meridional  
59 Overturning Circulation (AMOC) was proposed (Cacho et al., 1999, 2000, 2001; Bigg and Wadley, 2001; Sierro  
60 et al., 2005; Voelker et al., 2006) and recently supported by new geochemical data in sediments of the Gulf of  
61 Cádiz (Bahr et al., 2015). In particular, it has been suggested that the intensity of the MOW and, more generally,  
62 the variations of the thermohaline circulation of the Mediterranean Sea could play a significant role in triggering  
63 a switch from a weakened to an enhanced state of the AMOC through the injection of saline Mediterranean  
64 waters in the intermediate North Atlantic at times of weak AMOC (Rogerson et al., 2006; Voelker et al., 2006;  
65 Khélifi et al., 2009). Since the Mediterranean intermediate waters, notably the Levantine Intermediate Water  
66 (LIW), represent today up to 80 % in volume of the MOW (Kinder and Parilla, 1987) and are therefore a key  
67 driver of MOW-derived salt into the North Atlantic, it is crucial to gain a more complete understanding of the  
68 variability of the Mediterranean intermediate circulation in the past and its impact on the outflow.

69 Previous studies have mainly focused on the glacial variability of the deep-water circulation in the western  
70 Mediterranean basin (Cacho et al., 2000, 2006; Sierro et al., 2005; Frigola et al., 2007, 2008). During the Last  
71 Glacial Maximum (LGM), strong deep-water convection took place in the Gulf of Lions, producing cold, well-  
72 ventilated western Mediterranean Deep Water (WMDW) (Cacho et al., 2000, 2006; Sierro et al., 2005), while  
73 the MOW flowed at greater depth in the Gulf of Cádiz (Rogerson et al., 2005; Schönfeld and Zahn, 2000). With  
74 the onset of the Termination 1 (T1) at about 15 ka, the WMDW production declined until the transition to the  
75 Holocene due to the rising sea level, with a relatively weak mode during the Heinrich Stadial 1 (HS1) and the  
76 Younger Dryas (YD) (Sierro et al., 2005; Frigola et al., 2008), that led to the deposition of the Organic Rich  
77 Layer 1 (ORL1; 14.5-8.2 ka BP; Cacho et al., 2002).

78 Because of the disappearance during the Early Holocene of specific epibenthic foraminiferal species, such as  
79 *Cibicides* spp., which are commonly used for paleohydrological reconstructions, information about the  
80 Holocene variability of the deep-water circulation in the western Mediterranean are relatively scarce and are  
81 mainly based on grain size analysis and sediment geochemistry (Frigola et al., 2007). These authors have  
82 identified four distinct phases representing different deep-water overturning conditions in the western  
83 Mediterranean basin during the Holocene, as well as centennial- to millennial-scale abrupt events of overturning  
84 reinforcement.



85 Faunal and stable isotope records from benthic foraminifera located at intermediate depths in the eastern basin  
86 reveal uninterrupted well-ventilated LIW during the last glacial period and deglaciation (Kuhnt et al., 2008;  
87 Schmiedl et al., 2010). A grain-size record obtained from a sediment core collected within the LIW depth range  
88 (~500 m water depth) at the east Corsica margin also reveals enhanced bottom currents during HS1 and the YD  
89 (Toucanne et al., 2012). The Early Holocene is characterized by a collapse of the LIW (Kuhnt et al., 2008;  
90 Schmiedl et al., 2010; Toucanne et al., 2012) synchronous with the sapropel S1 deposition (10.2 – 6.4 cal ka BP;  
91 Mercone et al., 2000). Proxies for deep-water conditions reveal the occurrence of episodes of deep-water  
92 overturning reinforcement in the eastern Mediterranean basin at 8.2 ka BP (Rohling et al., 1997, 2015; Kuhnt et  
93 al., 2007; Abu-Zied et al., 2008, Siani et al., 2013; Tachikawa et al; 2015), responsible for the interruption of the  
94 sapropel S1 in the eastern Mediterranean basin (Mercone et al., 2001; Rohling et al., 2015).

95 It has recently been shown that the neodymium (Nd) isotopic composition, expressed as  $\epsilon\text{Nd} =$   
96  $\left( \left( \frac{{}^{143}\text{Nd}/{}^{144}\text{Nd}}{\text{sample}} / \frac{{}^{143}\text{Nd}/{}^{144}\text{Nd}}{\text{CHUR}} \right) - 1 \right) \times 10000$  (CHUR: Chondritic Uniform Reservoir [Jacobsen and  
97 Wasserburg, 1980]) of living and fossil scleractinian CWC faithfully traces intermediate and deep-water mass  
98 provenance and mixing of the ocean (e.g. van de Fliedert et al., 2010; Colin et al., 2010; López Correa et al.,  
99 2012; Monterro-Serrano et al., 2011, 2013; Copard et al., 2012). Differently from the CWC, the  $\epsilon\text{Nd}$   
100 composition of fossil planktonic foraminifera is not related to the ambient seawater at calcification depths but  
101 reflects the bottom and/or pore water  $\epsilon\text{Nd}$ , due to the presence of authigenic Fe-Mn coatings precipitated on their  
102 carbonate shell (Roberts et al., 2010; Elmore et al., 2011; Piotrowski et al., 2012; Tachikawa et al., 2014; Wu et  
103 al., 2015). Therefore, the  $\epsilon\text{Nd}$  composition of planktonic foraminiferal tests can be used as a useful tracer of  
104 deep-water circulation changes in the past, although the effect of pore water on foraminiferal  $\epsilon\text{Nd}$  values could  
105 potentially complicate the interpretation (Tachikawa et al., 2014).

106 In the Mediterranean Sea, modern seawater  $\epsilon\text{Nd}$  values display a large range from ~-11 to ~-5, and a clear  
107 vertical and longitudinal gradient, with more radiogenic values encountered in the eastern basin and typically at  
108 intermediate and deeper depths (Spivack and Wasserburg 1988; Henry et al., 1994; Tachikawa et al., 2004;  
109 Vance et al., 2004). Considering this large  $\epsilon\text{Nd}$  contrast,  $\epsilon\text{Nd}$  recorded in fossil CWC and planktonic  
110 foraminifera from the Mediterranean offers great potential to trace intermediate and deep-water mass exchange  
111 between the two basins, especially during periods devoid of key epibenthic foraminifera, such as the sapropel S1  
112 and ORL1 events.

113 Here, the  $\epsilon\text{Nd}$  of planktonic foraminifera from a sediment core collected in the Balearic Sea and CWC samples  
114 from the Alboran Sea and the Sardinia Channel was investigated to establish past changes of the  $\epsilon\text{Nd}$  values at  
115 intermediate depths and constrain hydrological variations of the LIW during the past ~20 ka. The  $\epsilon\text{Nd}$  values  
116 have been combined with stable oxygen ( $\delta^{18}\text{O}$ ) and carbon ( $\delta^{13}\text{C}$ ) isotope measurements of benthic (*Cibicidoides*  
117 *pachyderma*) and planktonic (*Globigerina bulloides*) foraminifera and sea-surface temperature estimates by  
118 modern analogue technique (MAT). Results reveal significant changes of the E-W gradient of  $\epsilon\text{Nd}$  values for the  
119 LIW of the western basin interpreted by a drastic reduction of the hydrological exchanges between the western  
120 and eastern Mediterranean Sea and the subsequent higher proportion of intermediate water produced in the Gulf  
121 of Lions during the time interval corresponding to the sapropel S1 deposition.

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## 125 2. Seawater $\epsilon\text{Nd}$ distribution in the Mediterranean Sea

126 The Atlantic Water (AW) enters the Mediterranean Sea as surface inflow through the Strait of Gibraltar with an  
127 unradiogenic  $\epsilon\text{Nd}$  signature of  $\sim -9.7$  in the strait (Tachikawa et al., 2004) and  $\sim -10.4$  in the Alboran Sea  
128 (Tachikawa et al., 2004, Spivack and Wasserburg, 1988) for depths shallower than 50 m. During its eastward  
129 flowing, AW mixes with upwelled Mediterranean Intermediate Water forming the Modified Atlantic Water  
130 (MAW) that spreads within the basin (Millot and Taupier-Letage, 2005) (Fig.1). The surface water  $\epsilon\text{Nd}$  values  
131 (shallower than 50 m) range from  $-9.8$  to  $-8.8$  in the western Mediterranean basin (Henry et al., 1994; Montagna  
132 et al., in prep) and  $-9.3$  to  $-4.2$  in the eastern basin, with seawater off the Nile delta showing the most radiogenic  
133 values (Tachikawa et al., 2004; Vance et al., 2004; Montagna et al., in prep). The surface waters in the eastern  
134 Mediterranean basin become denser due to strong mixing and evaporation caused by cold and dry air masses  
135 flowing over the Cyprus-Rhodes area in winter, and eventually sink leading to the formation of LIW  
136 (Ovchinnikov, 1984; Lascaratos et al., 1993, 1998; Malanotte-Rizzoli et al., 1999; Pinardi and Masetti, 2000).  
137 The LIW spreads throughout the entire Mediterranean basin at depths between  $\sim 150$ - $200$  m and  $\sim 600$ - $700$  m,  
138 and is characterized by more radiogenic  $\epsilon\text{Nd}$  values ranging from  $-7.9$  to  $-4.8$  (average value  $\pm 1\sigma$ :  $-6.6 \pm 1$ ) in  
139 the eastern basin and from  $-10.4$  to  $-7.58$  ( $-8.7 \pm 0.9$ ) in the western basin (Henry et al., 1994; Tachikawa et al.,  
140 2004; Vance et al., 2004; Montagna et al., in prep). The LIW acquires its  $\epsilon\text{Nd}$  signature mainly from the partial  
141 dissolution of Nile River particles (Tachikawa et al., 2004), which have an average isotopic composition of  $-3.25$   
142 (Weldeab et al., 2002), and the mixing along its path with overlying and underlying water masses with different  
143  $\epsilon\text{Nd}$  signatures. The LIW finally enters the Atlantic Ocean at intermediate depths through the Strait of Gibraltar  
144 with an average  $\epsilon\text{Nd}$  value of  $-9.2 \pm 0.2$  (Tachikawa et al., 2004; Montagna et al., in prep).

145 The WMDW is formed in the Gulf of Lions due to winter cooling and evaporation followed by mixing between  
146 the relative fresh surface water and the saline LIW and spreads into the Balearic basin and Tyrrhenian Sea  
147 between  $\sim 2000$  m and  $3000$  m (Millot, 1999; Schroeder et al., 2013) (Fig. 1). The WMDW is characterized by an  
148 average  $\epsilon\text{Nd}$  value of  $-9.4 \pm 0.9$  (Henry et al., 1994; Tachikawa et al., 2004; Montagna et al., in prep). Between  
149 the WMDW and the LIW (from  $\sim 700$  to  $2000$  m), the Tyrrhenian Deep Water (TDW) has been found (Millot et  
150 al., 2006), which is produced by the mixing between WMDW and Eastern Mediterranean Deep Water (EMDW)  
151 that cascades in the Tyrrhenian Sea after entering from the Strait of Sicily (Millot, 1999, 2009; Astraldi et al.,  
152 2001). The TDW has an average  $\epsilon\text{Nd}$  value of  $-8.1 \pm 0.5$  (Montagna et al., in prep).

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## 154 3. Material and methods

### 155 3.1. Cold-water coral and foraminifera samples

156 Forty-four CWC samples belonging to the species *Lophelia pertusa* and *Madrepora oculata* collected from the  
157 Alboran Sea and the Sardinia Channel were selected for this study (Fig. 1). Nineteen fragments were collected at  
158 various core depths from a coral-bearing sediment core (RECORD 23;  $38^{\circ}42.18'$  N;  $08^{\circ}54.75'$  E; Fig. 1)  
159 retrieved from  $414$  m water depth in the "Sardinian Cold-Water Coral Province" (Taviani et al., 2015) during the  
160 R/V Urania cruise "RECORD" in 2013. The Sardinian CWC samples were used for U-series dating and Nd  
161 isotopic composition measurements. For the southern Alboran Sea, twenty-five CWC samples were collected at  
162 water depths between  $280$  and  $442$  m in the "eastern Melilla Coral Province" (Fig. 1) during the R/V Poseidon  
163 cruise "POS-385" in 2009 (Hebbeln et al. 2009). Eleven samples were collected at the surface of two coral  
164 mounds (New Mound and Horse Mound) and three coral ridges (Brittlestar ridges I, II and III), using a box corer



165 and a remotely operated vehicle (ROV). In addition, fourteen CWC samples were collected from various core  
166 depths of three coral-bearing sediment cores (GeoB13728, 13729 and 13730) retrieved from the Brittlestar ridge  
167 I. Details on the location of surface samples and cores collected in the southern Alboran Sea and details on the  
168 radiocarbon ages obtained from these coral samples are reported in Fink et al. (2013). Like the CWC sample set  
169 from the Sardinia Channel, the dated Alboran CWC samples were also used for further Nd isotopic composition  
170 analyses in this study.

171 In addition, a sediment core (barren of any CWC fragments) was recovered in the Balearic Sea at 622 m water  
172 depth during the R/V Le Suroît cruise "PALEOCINAT II" in 1992 (SU92-33; 35°25.38' N; 0°33.86' E; Fig. 1).  
173 The core was sub-sampled continuously at 5-10 cm intervals for the upper 2.1 m for a total number of 24  
174 samples used for further multi-proxy analyses.

175

### 176 **3.2. Analytical procedures on cold-water coral samples**

#### 177 *3.2.1. U/Th dating*

178 The nineteen CWC samples collected from the sediment core RECORD 23 (Sardinia Channel) were analysed for  
179 uranium and thorium isotopes to obtain absolute dating using a Thermo Scientific™ Neptune<sup>plus</sup> MC-ICPMS  
180 installed at the Laboratoire des Sciences du Climat et de l'Environnement (LSCE, Gif-sur-Yvette, France). Prior  
181 to analysis, the samples were carefully cleaned using a small diamond blade to remove any visible contamination  
182 and sediment-filled cavities. The fragments were examined under a binocular microscope to ensure against the  
183 presence of bioeroded zones and finally crushed into a coarse-grained powder with an agate mortar and pestle.  
184 The powders (~60-100 mg) were transferred to acid cleaned Teflon beakers, ultrasonicated in MilliQ water,  
185 leached with 0.1N HCl for ~ 15 s and finally rinsed twice with MilliQ water. The physically and chemically  
186 cleaned samples were dissolved in 3-4 ml dilute HCl (~10%) and mixed with an internal triple spike with known  
187 concentrations of <sup>229</sup>Th, <sup>233</sup>U and <sup>236</sup>U, calibrated against a Harwell Uraninite solution (HU-1) assumed to be at  
188 secular equilibrium. The solutions were evaporated to dryness at 70°C, redissolved in 0.6 ml 3N HNO<sub>3</sub> and then  
189 loaded into 500 µl columns packed with Eichrom UTEVA resin to isolate uranium and thorium from the other  
190 major and trace elements of the carbonate matrix. The U and Th separation and purification followed a  
191 procedure slightly modified from Douville et al. (2010). The U and Th isotopes were determined following the  
192 protocol recently revisited at LSCE (Pons-Branchu et al., 2014). The <sup>230</sup>Th/U ages were calculated from  
193 measured atomic ratios through iterative age estimation (Ludwig and Titterton, 1994), using the <sup>230</sup>Th, <sup>234</sup>U  
194 and <sup>238</sup>U decay constants of Cheng et al. (2013) and Jaffey et al. (1971). Due to the low <sup>232</sup>Th concentration (< 1  
195 ng/g; see Table 1), no correction was applied for the non-radiogenic <sup>230</sup>Th fraction.

196

#### 197 *3.2.2 Nd isotopic composition analyses on cold-water coral fragments*

198 Sub-samples of the CWC fragments from the Sardinia Channel used for U-series dating in this study (Table 1) as  
199 well as sub-samples of the twenty-five CWC fragments originating from the Alboran Sea, which were already  
200 radiocarbon-dated by Fink et al. (2013) (Table 2), were used for further Nd isotopic composition analyses. The  
201 fragments (350 to 600 mg) were subjected to a mechanical and chemical cleaning procedure. The visible  
202 contaminations, such as Fe-Mn coatings and detrital particles, were carefully removed from the inner and  
203 outermost surfaces of the coral skeletons using a small diamond blade. The physically cleaned fragments were  
204 ultrasonicated for 10 min with 0.1 N ultra-clean HCl, followed by several MilliQ water rinses and finally



205 dissolved in 2.5 N ultraclean HNO<sub>3</sub>. Nd was separated from the carbonate matrix using Eichrom TRU and LN  
206 resins, following the analytical procedure described in detail in Copard et al. (2010).  
207 The <sup>143</sup>Nd/<sup>144</sup>Nd ratios of all purified Nd fractions were analyzed using the ThermoScientific Neptune<sup>Plus</sup> Multi-  
208 Collector Inductively Coupled Plasma Mass Spectrometer (MC-ICP-MS) hosted at LSCE. The mass-  
209 fractionation correction was made by normalizing <sup>146</sup>Nd/<sup>144</sup>Nd to 0.7219 and applying an exponential law.  
210 During each analytical session, samples were systematically bracketed with analyses of JNdi-1 and La Jolla  
211 standard solutions, which are characterised by accepted values of 0.512115±0.000006 (Tanaka et al., 2000) and  
212 0.511855±0.000007 (Lugmair et al., 1983), respectively. Standard JNdi-1 and La Jolla solutions were analysed  
213 at concentrations similar to those of the samples (5-10 ppb) and all the measurements affected by instrumental  
214 bias were corrected, when necessary, using La Jolla standard. The external reproducibility (2σ) for time resolved  
215 measurement, deduced from repeated analyses of La Jolla and JNdi-1 standards, ranged from 0.1 to 0.5 εNd  
216 units for the different analytical sessions. The analytical error for each sample analysis was taken as the external  
217 reproducibility of the La Jolla standard for each session. Concentrations of Nd blanks were negligible compared  
218 to the amount of Nd of CWC investigated in this study.

219

### 220 3.3. Analyses on sediment of core SU92-33

#### 221 3.3.1. Radiocarbon dating

222 Radiocarbon dating was measured at UMS-ARTEMIS (Pelletron 3MV) AMS (CNRS-CEA Saclay, France).  
223 Seven AMS radiocarbon (<sup>14</sup>C) dating were performed in core SU92-33 on well-preserved calcareous tests of the  
224 planktonic foraminifera *G. bulloides* in the size fraction >150 μm (Table 3). The age model for the core was  
225 derived from the calibrated planktonic ages by applying a mean reservoir effect of ~400 years (Siani et al., 2000,  
226 2001). All <sup>14</sup>C ages were converted to calendar years (cal. yr BP, BP = AD 1950) by using the INTCAL13  
227 calibration data set (Reimer et al., 2013) and the CALIB 7.0 program (Stuiver and Reimer, 1993).

228

#### 229 3.3.2. Stable isotopes

230 Stable oxygen (δ<sup>18</sup>O) and carbon (δ<sup>13</sup>C) isotope measurements were performed in core SU92-33 on well-  
231 preserved (clean and intact) samples of the planktonic foraminifera *G. bulloides* (250-315 μm fraction) and the  
232 epibenthic foraminifera *C. pachyderma* (250-315 μm fraction) using a Finnigan MAT-253 mass spectrometer at  
233 the State Key Laboratory of Marine Geology (Tongji University). Both δ<sup>18</sup>O and δ<sup>13</sup>C values are presented  
234 relative to the Pee Dee Belemnite (PDB) scale by comparison with the National Bureau of Standards (NBS) 18  
235 and 19. The mean external reproducibility was checked by replicate analyses of laboratory standards and is better  
236 than ±0.07‰ (1σ) for δ<sup>18</sup>O and ±0.04‰ for δ<sup>13</sup>C.

#### 237 3.3.3 Nd isotope measurements on planktonic foraminifera

238 Approximately 25 mg of mixed planktonic foraminifera species were picked from the >63 μm size fraction of  
239 each sample already used for stable isotope measurements (Table 4). The samples were gently crushed between  
240 glass slides under the microscope to ensure that all chambers were open, and ultrasonicated with MilliQ water.  
241 Samples were allowed to settle between ultrasonication steps before removing the supernatant. Each sample was  
242 rinsed thoroughly with MilliQ water until the solution was clear and free of clay. The cleaned samples were  
243 dissolved in 1N acetic acid and finally centrifuged to ensure that all residual particles were removed, following





244 the procedure described in Roberts et al. (2010). Nd was separated following the analytical procedure reported in  
245 Wu et al. (2015). For details on the measurement of Nd isotopes see the section above.

246

#### 247 3.3.4. Modern analogue technique (MAT)

248 The palaeo-sea surface temperatures (SST) were estimated using the modern analogue technique (MAT)  
249 (Hutson, 1980; Prell, 1985), implemented by Kallel et al. (1997) for the Mediterranean Sea. This method directly  
250 measures the difference between the faunal composition of a fossil sample with a modern database, and it  
251 identifies the best modern analogues for each fossil assemblage (Prell, 1985). Reliability of SST reconstructions  
252 is estimated using a square chord distance test (dissimilarity coefficient), which represents the mean degree of  
253 similarity between the sample and the best 10 modern analogues. When the dissimilarity coefficient is lower than  
254 0.25, the reconstruction is considered to be of good quality (Overpeck et al., 1985; Kallel et al., 1997). For core  
255 SU92-33, good dissimilarity coefficients are <0.2, with an average value of 0.13.

256

## 257 4. Results

### 258 4.1. Cold-water coral ages

259 The good state of preservation for the CWC samples from the Sardinia Channel (RECORD 23; Fig. 1) is attested  
260 by their initial  $\delta^{234}\text{U}$  values (Table 1), which is in the range of the modern seawater value ( $146.8 \pm 0.1$ ; Andersen  
261 et al., 2010). If the uncertainty of the  $\delta^{234}\text{U}$  is taken into account, all the values fulfill the so-called “strict”  $\pm 4$   
262 ‰ reliability criterion and the U/Th ages can be considered strictly reliable. The coral ages range from  
263  $0.091 \pm 0.011$  to  $10.904 \pm 0.042$  ka BP (Table 1), and reveal three distinct clusters of coral age distribution during  
264 the Holocene representing periods of sustained coral occurrence. These periods coincide with the Early Holocene  
265 encompassing a 700-years-lasting time interval from  $\sim 10.9$  to  $10.2$  ka BP, the very late Early Holocene at  $\sim 8.7$   
266 ka BP, and the Late Holocene starting at  $\sim 1.5$  ka BP (Table 1).

267 Radiocarbon ages obtained for CWC samples collected in the Alboran Sea were published by Fink et al. (2013)  
268 (Table 2). They also document three periods of sustained CWC occurrence coinciding with the Bølling–Allerød  
269 (B-A) interstadial ( $13.5$ – $12.9$  cal ka BP), the Early Holocene ( $11.2$ – $9.8$  cal ka BP) and the Mid- to Late Holocene  
270 ( $5.4$ – $0.3$  cal ka BP).

271

### 272 4.2 Chronological framework for core SU92-33

273 The stratigraphy of core SU92-33 was derived from the  $\delta^{18}\text{O}$  variations of the planktonic foraminifera  
274 *G. bulloides* (Fig. 2b). The last glacial/interglacial transition and the Holocene encompasses the upper 2.1 m of  
275 the core (Fig. 2b). The  $\delta^{18}\text{O}$  record of *G. bulloides* shows higher values ( $\sim 3.5$  ‰) during the late glacial  
276 compared to the Holocene (from  $\sim 1.5$  to  $0.8$  ‰) exhibiting a pattern similar to those observed in nearby deep-sea  
277 cores from the Western Mediterranean Sea (Sierro et al., 2005; Melki et al., 2009).

278 The age model of core SU92-33 is based on 7 AMS- $^{14}\text{C}$  age measurements for the upper 1.2 m of the core  
279 and by a linear interpolation between these ages (Table 3, Fig. 2). Below, a control point has been established for  
280 the onset of the last deglaciation that presents a coeval age in the western and central Mediterranean Sea at about  
281  $17$  cal ka BP (Sierro et al., 2005; Melki et al., 2009; Siani et al., 2001). The upper 2.1 m of core SU92-33 spans  
282 the last 19 ka, with an estimated average sedimentation rate between  $9$  to  $15$  cm ka $^{-1}$ , with the lowest values  
283 observed during the Holocene.



284

285 **4.3 SST reconstructions of core SU92-33**

286 April-May SST reconstruction was derived from MAT to define the main climatic events recorded in  
287 core SU92-33 during the last 19 ka. SSTs vary from 8.5°C to 17.5°C with high amplitude variability over the last  
288 19 ka BP (Fig. 2a). The LGM (19-18 ka BP) is characterized by SST values centered at around 12°C. Then, a  
289 progressive decrease of ~4°C between 17.8 ka to 16 ka marks the Heinrich Stadial 1 (HS1) (Fig. 2a). A warming  
290 phase (~14°C) between 14.5 ka BP and 13.8 ka BP coincides with the B-A interstadial and is followed by a  
291 cooling (~11°C) between 13.1 ka BP and 11.8 ka BP largely corresponding to the YD (Fig. 2). During the  
292 Holocene, SSTs show mainly values of ~16°C, with one exception between 7 ka BP and 6 ka BP pointing to an  
293 abrupt cooling of ~3°C (Fig. 2a). From the late glacial to the Holocene, SST variations show a similar pattern to  
294 that previously observed in the Gulf of Lions and Tyrrhenian Sea (Kallel et al., 1997; Melki et al., 2009) and  
295 globally synchronous for the main climatic transitions to the well dated South Adriatic Sea core MD90-917  
296 confirming the robustness of the SU92-33 age model (Fig. 2a).

297

298 **4.4 Benthic stable oxygen and carbon isotope records of core SU92-33**

299 The  $\delta^{18}\text{O}$  and  $\delta^{13}\text{C}$  records obtained from the benthic foraminifera *C. pachyderma* display significant variations  
300 at millennial time scales (Figs. 2c and 2d). The  $\delta^{18}\text{O}$  values decrease steadily from ~4.5 ‰ during the LGM to  
301 ~1.5 ‰ during the Holocene, without showing any significant excursion during HS1 and the YD events (Fig.  
302 2c), in agreement with results obtained for the neighbor core MD99-2343 (Sierro et al., 2005).

303 The  $\delta^{13}\text{C}$  record obtained from *C. pachyderma* shows a decreasing trend since the LGM with a low variability  
304 from ~1.6 ‰ to ~0.6 ‰ (Fig. 2d). The heaviest  $\delta^{13}\text{C}$  values are related to the LGM (~1.6 ‰) while the lightest  
305 values (~0.6 ‰) characterize the Early Holocene and in particular the period corresponding to the sapropel S1  
306 event in the eastern Mediterranean basin (Fig. 2d).

307

308 **4.5 Nd isotopic composition of planktonic foraminifera and cold-water corals**

309  $\epsilon\text{Nd}$  values of planktonic foraminifera of core SU92-33 collected from the Balearic Sea vary within a relatively  
310 narrow range between  $-9.50 \pm 0.30$  and  $-8.61 \pm 0.30$ , with an average value of  $-9.06 \pm 0.28$  (Table 2, Fig. 3b). The  
311 record shows a slight increasing trend since the LGM, with the more unradiogenic values (average  $-9.28 \pm 0.15$ ;  
312  $n=7$ ) observed in the oldest part of the record (between 18 and 13.5 ka BP), whereas Holocene values are  
313 generally more radiogenic (average  $-8.84 \pm 0.22$ ;  $n=17$ ) (Fig. 3b).

314 The  $\epsilon\text{Nd}$  record obtained for the CWC samples from the Alboran Sea displays a narrow range from  $-9.22 \pm 0.30$  to  
315  $-8.59 \pm 0.3$ , which is comparable to the  $\epsilon\text{Nd}$  record obtained on planktonic foraminifera from the Balearic Sea  
316 over the last 13.5 ka (Table 2, Fig. 3b). Most of the CWC  $\epsilon\text{Nd}$  values are similar within error and the record does  
317 not reveal any clear difference over the last ~13.5 ka.

318 Finally, the CWC samples from the Sardinia Channel display  $\epsilon\text{Nd}$  values ranging from  $-5.99 \pm 0.50$  to  $-7.75 \pm 0.10$   
319 during the Early and Late Holocene, and values as low as  $-8.66 \pm 0.30$  during the the mid-sapropel S1 deposition  
320 (S1a) (~8.7 ka BP) (Table 1, Fig. 3c).

321

322

323





324

## 5. Discussion

325 As first observations, the CWC and foraminiferal  $\epsilon\text{Nd}$  values measured for this study indicate a pronounced  
326 dispersion at intermediate depth in terms of absolute values and variability in Nd isotopes during the Holocene  
327 between the Alboran and Balearic Seas and the Sardinia Channel. In addition the foraminiferal  $\epsilon\text{Nd}$  record  
328 reveals an evolution towards more radiogenic values at intermediate water depth in the Balearic Sea over the last  
329  $\sim 19$  ka (Fig. 3).

330 A prerequisite to properly interpret such  $\epsilon\text{Nd}$  values differences and variations through time consists in  
331 characterizing the present-day  $\epsilon\text{Nd}$  of the main water-mass end-members circulating in the western  
332 Mediterranean basin. This is possible by evaluating the temporal changes in  $\epsilon\text{Nd}$  of the end-members since the  
333 LGM, and assessing the potential influences of lithogenic Nd input and regional exchange between the  
334 continental margins and seawater (“boundary exchange”; Lacan and Jeandel, 2001, 2005) on the  $\epsilon\text{Nd}$  values of  
335 intermediate water masses.

336 During its westward flow, the LIW continuously mixes with surrounding waters with different  $\epsilon\text{Nd}$  signatures  
337 lying above and below. For the western Mediterranean basin, these water masses are the MAW/Western  
338 Intermediate Water (WIW) and the TDW/WMDW, respectively. Accordingly, a well-defined and gradual  $\epsilon\text{Nd}$   
339 gradient exists at intermediate depth between the eastern and western Mediterranean basins, with LIW values  
340 becoming progressively more unradiogenic towards the Strait of Gibraltar, from  $-4.8 \pm 0.2$  at 227 m in the  
341 Levantine basin to  $-10.4 \pm 0.2$  at 200 m in the Alboran Sea (Tachikawa et al., 2004). Such an  $\epsilon\text{Nd}$  pattern implies  
342 an effective vertical mixing with more unradiogenic water masses along the E-W LIW trajectory ruling out  
343 severe isotopic modifications of the LIW due to the local exchange between the continental margins and  
344 seawater. Unfortunately, no information exists on the potential temporal variability in  $\epsilon\text{Nd}$  of the Mediterranean  
345 water-mass end-members since the LGM.

346 It has been demonstrated that eolian dust input can modify the surface and sub-surface  $\epsilon\text{Nd}$  distribution of the  
347 ocean in some areas (Arsouze et al., 2009). The last glacial period was associated with an aridification of North  
348 Africa (Sarnthein et al., 1981; Hooghiemstra et al., 1987; Moreno et al., 2002; Wienberg et al., 2010) and higher  
349 fluxes of Saharan dust to the NE tropical Atlantic (Itambi et al., 2009) and the western Mediterranean Sea  
350 characterized by unradiogenic  $\epsilon\text{Nd}$  values (between  $-10 \pm 0.4$  and  $-17 \pm 0.4$ ; Grousset et al., 1992, 1998; Grousset  
351 and Biscaye, 2005). Bout-Roumzeilles et al. (2013) documented a dominant role of eolian supply in the Siculo-  
352 Tunisian Strait during the last 20 ka, with the exception of a significant riverine contribution (from the Nile  
353 River) and a strong reduction of eolian input during the sapropel S1 event. Such variations in the eolian input to  
354 the Mediterranean Sea are not associated to a significant change in the seawater  $\epsilon\text{Nd}$  record obtained for the  
355 Balearic Sea (core SU92-33) during the sapropel S1 event (Fig. 3). Furthermore, the  $\epsilon\text{Nd}$  signature of the CWC  
356 from the Sardinia Channel (core RECORD 23) shifts to more unradiogenic values ( $-8.66 \pm 0.30$ ) during the  
357 sapropel S1 event, which is opposite to what expected if related to a strong reduction of eolian sediment input.  
358 Thus, these results suggest that changes of eolian dust input since the LGM were not responsible for the  
359 observed  $\epsilon\text{Nd}$  variability at intermediate water depths.

360 Consequently, assuming that the Nd isotopic budget of the western Mediterranean Sea has not been strongly  
361 modified since the LGM, the reconstructed variations of the E-W gradient of  $\epsilon\text{Nd}$  values in the western  
362 Mediterranean Sea for the past and notably during the sapropel S1 event (Fig. 3) are indicative of a major  
363 reorganization of intermediate water circulation.



364

365 **5.1 Hydrological changes in the Alboran and Balearic Seas since the LGM**

366 The range in  $\epsilon\text{Nd}$  for the CWC from the Alboran Sea (from  $-9.22\pm 0.30$  to  $-8.8.59\pm 0.30$ ; Table 2) is very close to  
367 the one obtained for the planktonic foraminifera from the Balearic Sea (from  $-9.50\pm 0.3$  to  $-8.61\pm 0.3$ ; Table 4,  
368 Fig. 3c), suggesting that both sites are influenced by the same intermediate water masses at least for the last 13.5  
369 ka BP. Today, LIW occupies a depth range between  $\sim 200$  and  $\sim 700$  m in the western Mediterranean basin  
370 (Millot, 1999; Sparnocchia et al., 1999). More specifically, the salinity maximum corresponding to the core of  
371 LIW is found at around 400 m in the Alboran Sea (Millot, 2009) and up to 550 m in the Balearic Sea (López-  
372 Jurado et al., 2008). The youngest CWC sample collected in the Alboran Sea with a rather "recent" age of 0.34  
373 cal ka BP (Fink et al. 2013) displays an  $\epsilon\text{Nd}$  value of  $-8.59\pm 0.30$  (Table 2) that is similar to the present-day value  
374 of the LIW at the same site ( $-8.3\pm 0.2$ ) (Dubois-Dauphin et al., submitted) and is significantly different from the  
375 WMDW  $\epsilon\text{Nd}$  signature in the Alboran Sea ( $-10.7\pm 0.2$ , 1270 m water depth; Tachikawa et al., 2004). Considering  
376 the intermediate depth range of the studied CWC and foraminifera samples, we can reasonably assume that  
377 samples from both sites, in the Balearic Sea (622 m water depth) and in the Alboran Sea (280 to 442 m water  
378 depth), record  $\epsilon\text{Nd}$  variations of the LIW. The  $\epsilon\text{Nd}$  record obtained on planktonic foraminifera generally displays  
379 more unradiogenic and homogenous values before  $\sim 13$  cal ka BP (range:  $-9.46$  to  $-9.12$ ) compared to the most  
380 recent part of the record (range:  $-9.50$  to  $-8.61$ ), with the highest value of  $-8.61\pm 0.3$  in the Early and Late  
381 Holocene.

382 The  $\delta^{18}\text{O}$  record obtained on *G. bulloides* indicates an abrupt ‰ excursion towards lighter values centered at  
383 about 16 cal ka BP (Table 4), synchronous with the HS1 (Fig. 2b), which is similar to the  $\delta^{18}\text{O}$  shift reported by  
384 Sierro et al. (2005) for a core collected at 2391 m water depth NE of the Balearic Islands (MD99-2343; Fig. 1).  
385 As the Heinrich events over the last glacial period are characterized by colder and fresher surface water in the  
386 Alboran Sea (Cacho et al., 1999; Pérez-Folgado et al., 2003; Martrat et al., 2004) and dry climate on land over  
387 the western Mediterranean Sea (Allen et al., 1999; Combourieu-Nebout et al., 2002; Sanchez Goni et al., 2002;  
388 Bartov et al., 2003), lighter  $\delta^{18}\text{O}$  values of planktonic *G. bulloides* are thought to be the result of the inflow of  
389 freshwater derived from the melting of icebergs in the Atlantic Ocean into the Mediterranean Sea (Sierro et al.,  
390 2005; Rogerson et al., 2008).

391 During this time interval, the  $\delta^{13}\text{C}$  record of *C. pachyderma* from the Balearic Sea (core SU92-33) displays a  
392 decreasing  $\delta^{13}\text{C}$  trend after  $\sim 16$  cal ka BP (from 1.4 ‰ to 0.9 ‰; Table 4; Fig. 4a). Moreover, the  $\delta^{13}\text{C}$  record  
393 obtained on benthic foraminifera *C. pachyderma* from the deep Balearic Sea (core MD99-2343) reveals similar  
394  $\delta^{13}\text{C}$  values before  $\sim 16$  cal ka BP suggesting well-mixed and ventilated water masses during the LGM and the  
395 onset of the deglaciation (Sierro et al., 2005).

396 The slightly lower foraminiferal  $\epsilon\text{Nd}$  values before  $\sim 13$  cal ka BP could reflect a stronger influence of water  
397 masses deriving from the Gulf of Lions as WMDW ( $\epsilon\text{Nd}$ :  $-9.4\pm 0.9$ ; Henry et al., 1994; Tachikawa et al., 2004;  
398 Montagna et al., in prep). This is in agreement with  $\epsilon\text{Nd}$  results obtained by Jiménez-Espejo et al. (2015) from  
399 planktonic foraminifera collected from deep-water sites (1989 m and 2382 m) in the Alboran Sea (Fig. 4c).  
400 Jiménez-Espejo et al. (2015) documented lower  $\epsilon\text{Nd}$  values (ranging from  $-10.14\pm 0.27$  to  $-9.58\pm 0.22$ ) during the  
401 LGM, suggesting an intense deep-water formation. This is also associated to an enhanced activity of the deeper  
402 branch of the MOW in the Gulf of Cádiz (Rogerson et al., 2005; Voelker et al., 2006) linked to the active  
403 production of the WMDW in the Gulf of Lions during the LGM (Jiménez-Espejo et al., 2015).



404 The end of the HS1 (14.7 cal ka BP) is concurrent with the onset of the B-A warm interval characterized by  
405 increased SST identified for various sites in the Mediterranean Sea (Cacho et al., 1999; Martrat et al., 2004;  
406 Essallami et al., 2007), in agreement with the SST record obtained for the Balearic Sea (SU92-33; Fig. 3a). The  
407 B-A interval is associated to the so-called melt-water pulse 1A (e.g. Weaver et al., 2003) occurring at around  
408 14.5 cal ka BP. This led to a rapid sea-level rise of about 20 m in less than 500 years and large freshwater  
409 discharges in the Atlantic Ocean due to the melting of continental ice sheets (Deschamps et al., 2012), resulting  
410 in an enhanced Atlantic inflow across the Strait of Gibraltar. Synchronously, cosmogenic dating of Alpine  
411 glacier retreat throughout the western Mediterranean hinterland suggests maximum retreat rates (Ivy-Ochs et al.,  
412 2007; Kelly et al., 2006). Overall, these events are responsible for freshening Mediterranean waters and reduced  
413 surface water density, and hence, weakened ventilation of intermediate (Toucanne et al., 2012) and deep-water  
414 masses (Cacho et al., 2000; Sierro et al., 2005). Similarly, lower benthic  $\delta^{13}\text{C}$  values obtained for the Balearic  
415 Sea (Fig. 4a) point to less ventilated intermediate water relative to the late glacial. In addition, a decoupling in  
416 the benthic  $\delta^{13}\text{C}$  values is observed between deep (MD99-2343) and intermediate (core SU92-33) waters after  
417  $\sim 16$  cal ka BP (Sierro et al. 2005), suggesting an enhanced stratification of the waters masses (Fig. 4a). At this  
418 time, the shallowest  $\epsilon\text{Nd}$  record from the deep Alboran Sea (core 300G) shifted towards more radiogenic values,  
419 while the deepest one (core 304G) remained close to the LGM values (Jimenez-Espejo et al., 2015) (Fig. 4c).  
420 Furthermore, results from the UP10 fraction (particles  $> 10 \mu\text{m}$ ) of the MD99-2343 sediment core (Fig. 4d),  
421 indicate a declining bottom-current velocity at 15 ka (Frigola et al., 2008). Rogerson et al. (2008) have  
422 hypothesized that during deglacial periods the sinking depth of dense waters produced in the Gulf of Lions was  
423 shallower resulting in new intermediate water (WIW) rather than new deep-water (WMDW) as observed today  
424 during mild winters (Millot, 1999; Schott et al., 1996). Therefore, intermediate depths of the Balearic Sea could  
425 have been isolated from the deep-water with the onset of the T1 (at  $\sim 15$  ka BP). The reduced convection in the  
426 deep western Mediterranean Sea together with the shoaling of the nutricline (Rogerson et al., 2008) led to the  
427 deposition of the ORL 1 (14.5 to 8.2 ka B.P; Cacho et al., 2002) and dysoxic conditions below 2000 m in  
428 agreement with the absence of epibenthic foraminifera such as *C. pachyderma* after 11 cal ka BP in MD99-2343  
429 (Sierro et al., 2005) (Fig. 4a).

430 After 13.5 ka BP, planktonic foraminifera  $\epsilon\text{Nd}$  values from the Balearic Sea (core SU92-33) become more  
431 radiogenic and are in the range of CWC  $\epsilon\text{Nd}$  values from the Alboran Sea (Fig. 4b). These values may reveal a  
432 stronger influence of the LIW in the Balearic Sea during the Younger Dryas, as also supported by the sortable  
433 silt record from the Tyrrhenian Sea (Toucanne et al., 2012) (Fig. 4e). Deeper depths of the Alboran Sea also  
434 record a stronger influence of the LIW with an  $\epsilon\text{Nd}$  value of  $-9.1 \pm 0.4$  (Jimenez-Espejo et al., 2015). In addition,  
435 a concomitant activation of the upper MOW branch, as reconstructed from higher values of Zr/Al ratio in  
436 sediments of the Gulf of Cádiz, can be related to the enhanced LIW flow in the western Mediterranean Sea (Fig.  
437 4f) (Bahr et al., 2015).

438 The time of sapropel S1 deposition (10.2 – 6.4 ka) is characterized by a weakening or a shutdown of  
439 intermediate- and deep-water formation in the eastern Mediterranean basin (Rossignol-Strick et al., 1982; Cramp  
440 and O'Sullivan, 1999; Emeis et al., 2000; Rohling et al., 2015). At this time, planktonic foraminifera  $\epsilon\text{Nd}$  values  
441 from intermediate water depths in the Balearic Sea (core SU92-33) remain high (between  $-9.15 \pm 0.3$  and -  
442  $8.61 \pm 0.3$ ) (Fig. 4b). On the other hand, the deeper Alboran Sea provides a value of  $-9.8 \pm 0.3$  pointing to a



443 stronger contribution of WMDW (Jimenez-Espejo et al., 2015), coeval with the recovery of deep-water activity  
444 from core MD99-2343 (Frigola et al., 2008).

445

#### 446 *5.2 Hydrological changes in the Sardinia Channel during the Holocene*

447 The present-day hydrographic structure of the Sardinia Channel is characterized by four water masses, with the  
448 surface, intermediate and deep-water masses being represented by MAW, LIW and TDW/WMDW, respectively  
449 (Astraldi et al., 2002a; Millot and Taupier-Lepage, 2005). In addition, the WIW, flowing between the MAW and  
450 the LIW, has also been observed along the Channel (Sammari et al., 1999). The core of the LIW is located at  
451 400-450 m water depth in the Tyrrhenian Sea (Hopkins, 1988; Astraldi et al., 2002b), which is the depth range of  
452 CWC samples from the Sardinia Channel (RECORD 23; 414 m) (Taviani et al., 2015). The youngest CWC  
453 sample dated at ~0.1 ka BP has an  $\epsilon\text{Nd}$  value of  $-7.70\pm 0.10$  (Table 1, Fig. 5), which is similar within error to the  
454 value obtained from a seawater sample collected at 451 m close to the coral sampling location ( $-8.0\pm 0.4$ ;  
455 Montagna et al., in prep).

456 The CWC dating from the Sardinia Channel shows three distinct periods of sustained coral occurrence in this  
457 area during the Holocene, with each displaying a large variability in  $\epsilon\text{Nd}$  values. CWC from the Early Holocene  
458 (10.9-10.2 ka BP) and the Late Holocene (<1.5 ka BP) exhibit similar ranges of  $\epsilon\text{Nd}$  values (ranging from -  
459  $5.99\pm 0.50$  to  $-7.75\pm 0.20$ ; Table 1, Fig 5c). Such variations are within the present-day  $\epsilon\text{Nd}$  range being  
460 characteristic for intermediate waters in the eastern Mediterranean Sea ( $-6.6\pm 1.0$ ; Tachikawa et al., 2004; Vance  
461 et al., 2004). However, the CWC  $\epsilon\text{Nd}$  values are more radiogenic than those observed at mid-depth in the  
462 present-day western basin (ranging from  $-10.4\pm 0.2$  to  $-7.58\pm 0.47$ ; Henry et al., 1994; Tachikawa et al., 2004;  
463 Montagna et al., in prep), suggesting a stronger LIW component in the Sardinia Channel during the Early and  
464 Late Holocene. The Sardinian CWC  $\epsilon\text{Nd}$  variability also reflects the sensitivity of the LIW to changes in the  
465 eastern basin such as rapid variability of the Nile River flood discharge (Revel et al., 2014; 2015; Weldeab et al.,  
466 2014) or a modification through time in the proportion between the LIW and the Cretan Intermediate Water  
467 (CIW). Today, the intermediate water outflowing from the Strait of Sicily is composed by ~66 to 75 % of LIW  
468 and 33 to 25 % of CIW (Manca et al., 2006; Millot, 2014). As the CIW is formed in the Aegean Sea, this  
469 intermediate water mass is generally more radiogenic than LIW (Tachikawa et al., 2004; Montagna et al., in  
470 prep). Following this hypothesis, a modification of the mixing proportion between the CIW and the LIW may  
471 potentially explain values as radiogenic as about -6 in the Sardinia Channel during the Early and Late Holocene  
472 (Fig. 5c). However, a stronger LIW and/or a CIW contribution cannot be responsible for  $\epsilon\text{Nd}$  values as low as -  
473  $8.66\pm 0.30$  observed during the sapropel S1 event at 8.7 ka BP (Table 1, Fig. 5c). Considering that such  
474 unradiogenic value is not observed at intermediate depth in the modern eastern Mediterranean basin, the most  
475 plausible hypothesis suggested here is that the CWC were bathed in intermediate waters which were more  
476 marked by the western basin.

477

#### 478 *5.3 Hydrological implications for the intermediate water masses of the western Mediterranean Sea*

479 The  $\epsilon\text{Nd}$  records of the Balearic Sea, Alboran Sea and Sardinia Channel document a temporal variability of the  
480 east-west gradient in the western Mediterranean basin during the Holocene. The magnitude of the gradient  
481 ranges from ~1.5 to ~3  $\epsilon$  units during the Early and Late Holocene and it is strongly reduced at 8.7 ka BP,



482 coinciding with the sapropel S1 event affecting the eastern Mediterranean basin (Fig. 5). Such variations could  
483 be the result of a modification of the Nd isotopic composition of intermediate water masses due to intensity  
484 changes of the the LIW through time and a higher contribution of the western-sourced intermediate water  
485 towards the Sardinia Channel coinciding with the sapropel S1 event.

486 The LIW acquires its radiogenic  $\epsilon\text{Nd}$  in the Mediterranean Levantine basin mainly from Nd exchange between  
487 seawater and lithogenic particles originating mainly from Nile River (Tachikawa et al., 2004). A higher sediment  
488 supply from the Nile River starting at ~15 ka BP was documented by a shift to more radiogenic  $\epsilon\text{Nd}$  values of  
489 the terrigenous fraction obtained from a sediment core having been influenced by the Nile River discharge  
490 (Revel et al., 2015) (Fig. 5e). However, others studies pointed to a gradual enhanced Nile River runoff as soon as  
491 14.8 ka and a peak of Nile discharge from 9.7 to 8.4 ka recorded by large increase in sedimentation rate from 9.7  
492 to 8.4 ka (>120 cm/ka) (Revel et al., 2015; Weldeab et al., 2014; Castaneda et al., 2016). The increase in Nile  
493 River discharge has been related to the African Humid Period (14.8–5.5 ka BP; Shanahan et al., 2015), which in  
494 turn was linked to the precessional increase in Northern Hemisphere insolation during low eccentricity  
495 (deMenocal et al., 2000; Barker et al., 2004; Garcin et al., 2009). An increasing amount of radiogenic sediments  
496 dominated by the Blue/Atbara Nile River contribution (Revel et al., 2014) could have modified the  $\epsilon\text{Nd}$  of  
497 surface water towards more radiogenic values (Revel et al., in prep). This signature was likely transferred to  
498 intermediate depth as a consequence of the LIW formation in the Rhodes Gyre, and it might have been  
499 propagated westwards towards the Sardinia Channel.

500 The Nile River runoff was also strongly enhanced during the sapropel S1 event (Revel et al., 2010; Weldeab et  
501 al., 2014; Revel et al., 2014). Scrivner et al. (2004) have reported very high foraminifera  $\epsilon\text{Nd}$  values (-3 to -3.5)  
502 corresponding to the sapropel S1 event in the eastern Levantine Basin (ODP site 967; 34°04.27'N, 32°43.53'E;  
503 2553 m water depth), pointing to a maximum Nile discharge at this time. Hence, considering the more  
504 unradiogenic value of the CWC samples from the Sardinia Channel during the sapropel S1a event, it is very  
505 unlikely that eastern-sourced water flowed at intermediate depth towards the Sardinia Channel. A possible  
506 explanation could be the replacement of the radiogenic LIW that was no longer produced in the eastern basin  
507 (Rohling, 1994) by less radiogenic western intermediate water (possibly WIW). Such a scenario could even  
508 support previous hypotheses that invoke a potential circulation reversal in the eastern Mediterranean from anti-  
509 estuarine to estuarine during sapropel formation (Huang and Stanley, 1972; Calvert, 1983; Sarmiento et al.,  
510 1988; Buckley and Johnson, 1988; Thunell and Williams, 1989).

511

## 512 6. Conclusions

513 The foraminiferal  $\epsilon\text{Nd}$  record from the intermediate Balearic Sea reveals a relatively narrow range of  $\epsilon\text{Nd}$  values  
514 varying between -9.50 and -8.61 since the LGM (~20 ka). Between 18 and 13.5 cal ka BP, the more  
515 unradiogenic  $\epsilon\text{Nd}$  values support a vigorous deep overturning in the Gulf of Lions while  $\delta^{18}\text{O}$  and  $\delta^{13}\text{C}$  values  
516 indicate a stratification of the water masses after 16 cal ka BP. The stratification together with a decrease of the  
517 deep-water intensity led to more radiogenic values after ~13 cal ka BP. The  $\epsilon\text{Nd}$  record from planktonic  
518 foraminifera, supplemented by CWC from the intermediate depths of the Alboran Sea, show only minor changes  
519 in  $\epsilon\text{Nd}$  values from 13.5 cal ka BP to 0.34 cal ka BP, suggesting that the westernmost part of the western  
520 Mediterranean basin is not very sensitive to hydrological variations of the LIW.



521 On the contrary, CWC located at the depth of the LIW in the Sardinia Channel indicate high amplitude variations  
522 of the  $\epsilon\text{Nd}$  values (between  $-7.75\pm 0.10$  and  $-5.99\pm 0.50$ ) during the Holocene, which could highlight either the  
523 role of the Nile River in changing the  $\epsilon\text{Nd}$  of the LIW in the eastern Mediterranean basin or a different  
524 LIW/CIW mixing of the water outflowing from the Strait of Sicily. Coinciding with the sapropel S1 event at  
525  $\sim 8.7$  ka BP, CWC display a shift toward lower values ( $-8.66\pm 0.30$ ), similar to those obtained at intermediate  
526 depths in the westernmost part of the western basin. This suggests that western-sourced intermediate water likely  
527 filled mid-depth of the southern Sardinia, replacing LIW that was no longer produced (or heavily reduced) in the  
528 eastern basin. These results could potentially support a reversal of the Mediterranean circulation, although this  
529 assumption needs further investigation to be confirmed.

530

### 531 Acknowledgements

532 The research leading to this study has received funding from the French National Research Agency  
533 “Investissement d’Avenir” (n°ANR-10-LABX-0018), the HAMOC project ANR-13-BS06-0003, the  
534 MISTRALS/PALEOMEX/COFIMED and ENVIMED/Boron Isotope and Trace Elements project. This work  
535 contributes to the RITMARE project. We thank Hiske Fink for selecting and kindly providing the cold-water  
536 corals samples from the Alboran Sea. We further thank François Thil and Louise Bordier for their support with  
537 Nd isotopic composition analyses. Paolo Montagna is grateful for financial support from the Short Term  
538 Mobility Program (CNR). Thanks are also extended to the captains, crews, chief scientists, and scientific parties  
539 of research cruises RECORD (R/V Urania), POS-385 (R/V Poseidon) and PALEOCINAT II (R/V Le Suroît).

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### 938 Table captions

939

940 **Table 1.** U-series ages and  $\epsilon\text{Nd}$  values obtained for cold-water coral samples collected from sediment core RECORD 23  
 941 (Sardinia Channel).

942

943 **Table 2.**  $\epsilon\text{Nd}$  values obtained for cold-water corals from the southern Alboran Sea. The AMS  $^{14}\text{C}$  ages published by Fink et  
 944 al. (2013) are also reported as Median probability age (ka BP).

945

946 **Table 3.** AMS  $^{14}\text{C}$  ages of samples of the planktonic foraminifer *G. bulloides* from ‘off-mound’ sediment core SU92-33. The  
 947 AMS  $^{14}\text{C}$  ages were corrected for  $^{13}\text{C}$  and a mean reservoir age of 400 yrs, and were converted into calendar years using the  
 948 INTCAL13 calibration data set (Reimer et al., 2013) and the CALIB 7.0 program (Struiver et al., 2005).

949

950 **Table 4.** Multiproxy data obtained for the upper 2.1 m of sediment core SU92-33 (Balearic Sea). Stable oxygen and carbon  
 951 isotopes were measured on benthic (*C. pachyderma*) and planktonic (*G. bulloides*) foraminifera;  $\epsilon\text{Nd}$  values were obtained on  
 952 mixed planktonic foraminifera samples. The age results from a combination of 7 AMS- $^{14}\text{C}$  age measurements for the upper  
 953 1.2 m of the core and by a linear interpolation between these ages as well as the  $\delta^{18}\text{O}$  variations of the planktonic  
 954 foraminifera *G. bulloides*.

955

### 956 Figure captions

957

958 **Figure 1.** Map of the western Mediterranean Sea showing the locations of samples investigated in this study. Yellow dot  
 959 indicates the sampling location of the sediment core from the Balearic Sea (SU92-33); yellow stars indicate the locations of  
 960 the CWC-bearing cores from the Sardinia Channel (RECORD 23) and the southern Alboran Sea (for further details on the  
 961 CWC from the Alboran Sea refer also to Fink et al., 2013). The cores discussed in this paper (Gulf of Cádiz: IODP site  
 962 U1387, Balearic Sea: MD09-2343, northern Tyrrhenian Sea: MD01-2472, Adriatic Sea: MD90-917) are indicated by black  
 963 dots, and seawater stations are marked by open squares. Arrows represent the main oceanographic currents. The black line  
 964 shows the general trajectory of the Modified Atlantic Water (MAW) flowing at the surface from the Atlantic Ocean toward  
 965 the western and eastern Mediterranean. The orange line represents the Levantine Intermediate Water (LIW) originating from  
 966 the eastern basin. The black dashed line shows the trajectory of the Western Mediterranean Deep Water (WMDW) flowing  
 967 from the Gulf of Lions toward the Strait of Gibraltar.

968

969 **Figure 2.** (a) Sea Surface Temperature (SST) records of cores SU92-33 (red line) and MD90-917 (green line), (b)  $\delta^{18}\text{O}$   
 970 record obtained on planktonic foraminifer *G. bulloides* for core SU92-33, (c)  $\delta^{18}\text{O}$  record obtained on benthic foraminifer *C.*  
 971 *pachyderma* for core SU92-33, (d)  $\delta^{13}\text{C}$  record obtained on benthic foraminifer *C. pachyderma* for core SU92-33. LGM: Last



972 Glacial Maximum; HS1: Heinrich Stadial 1; B-A: Bølling-Allerød; YD: Younger Dryas. Black triangles indicate AMS  $^{14}\text{C}$   
973 age control points.

974

975 **Figure 3.** (a) Sea Surface Temperature (SST) record of core SU92-33 (red line), (b)  $\epsilon\text{Nd}$  records obtained on mixed  
976 planktonic foraminifers from core SU92-33 (open circles) and from cold-water coral fragments collected in the Alboran Sea  
977 (red squares), (c)  $\epsilon\text{Nd}$  values of cold-water corals from core RECORD 23 (Sardinia Channel).

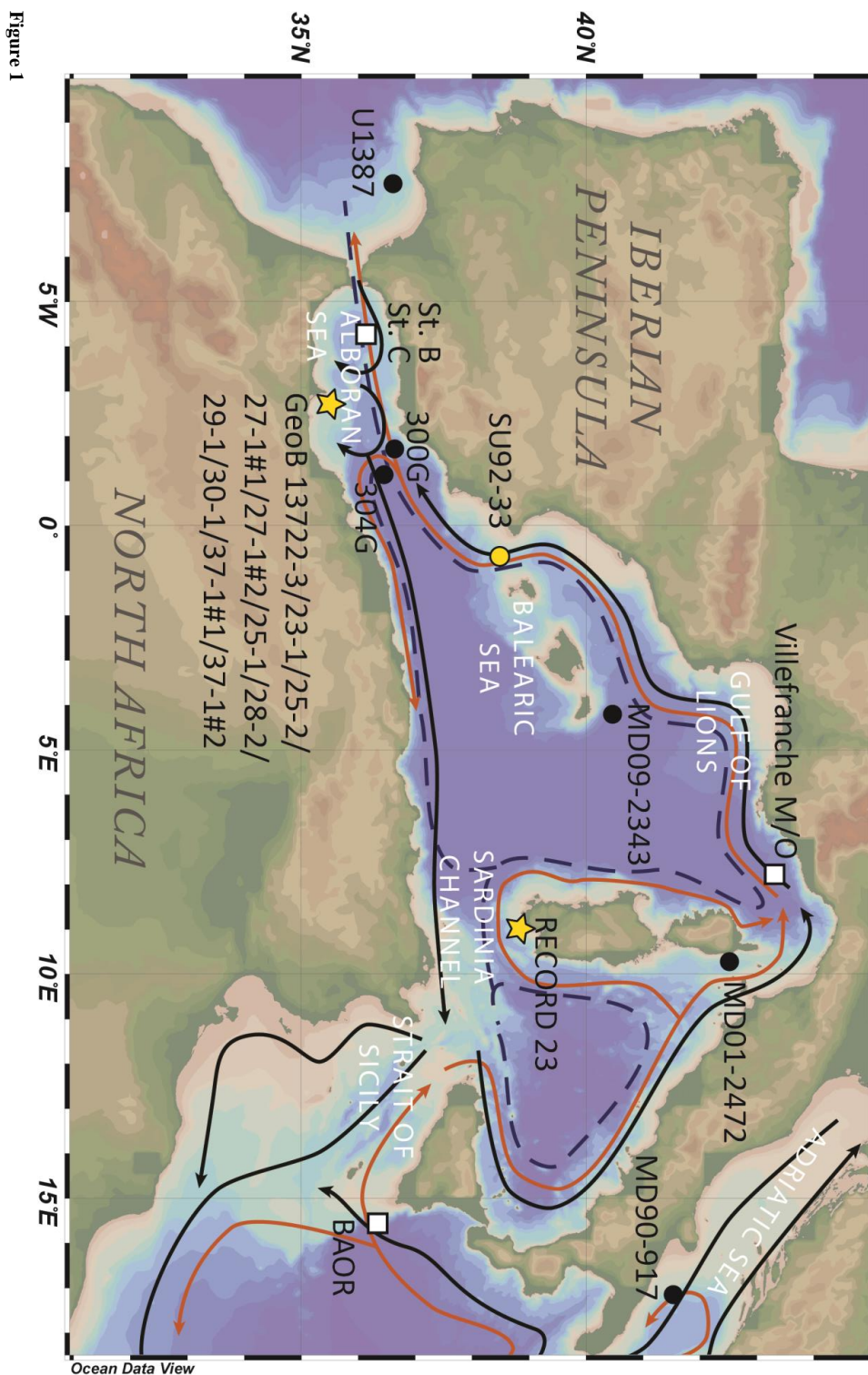
978

979 **Figure 4.** (a)  $\delta^{13}\text{C}$  records obtained on benthic foraminifer *C. pachyderma* for cores SU92-33 (red line) and MD99-2343  
980 (blue line; Sierro et al., 2005). (b)  $\epsilon\text{Nd}$  records obtained on mixed planktonic foraminifers from core SU92-33 (open circles)  
981 and from cold-water coral fragments collected in the Alboran Sea (red squares). Modern  $\epsilon\text{Nd}$  values for LIW (orange dashed  
982 line) and WMDW (blue dashed line) are also reported for comparison. (c)  $\epsilon\text{Nd}$  values obtained for planktonic foraminifera  
983 with Fe-Mn coatings at sites 300G (36°21.532' N, 1°47.507' W; 1860 m; open dots) and 304G (36°19.873' N, 1°31.631' W;  
984 2382 m; black dots) in Alboran Sea (Jimenez-Espejo et al., 2015). (d) UP10 fraction (>10  $\mu\text{m}$ ) from core MD99-2343  
985 (Frigola et al., 2008). (e) Sortable silt mean grain-size of core MD01-2472 (Toucanne et al., 2012). (f) Ln Zr/Al ratio at IODP  
986 site U1387 (36°48.3' N 7°43.1' W; 559 m) (Bahr et al., 2015).

987

988 **Figure 5.** (a)  $\delta^{18}\text{O}$  record obtained on planktonic foraminifer *G. bulloides* for core SU92-33, (b)  $\delta^{13}\text{C}$  records obtained on  
989 benthic foraminifer *C. pachyderma* for core SU92-33, (c)  $\epsilon\text{Nd}$  values of cold-water corals from core RECORD 23 (Sardinia  
990 Channel), (d)  $\epsilon\text{Nd}$  values records obtained on mixed planktonic foraminifera from core SU92-33 (open circles) and from  
991 cold-water coral fragments collected in the Alboran Sea (red squares), (e)  $\epsilon\text{Nd}$  values obtained on terrigenous fraction of  
992 MS27PT located close the Nile River mouth in the eastern Mediterranean basin (Revel et al., 2015).

993





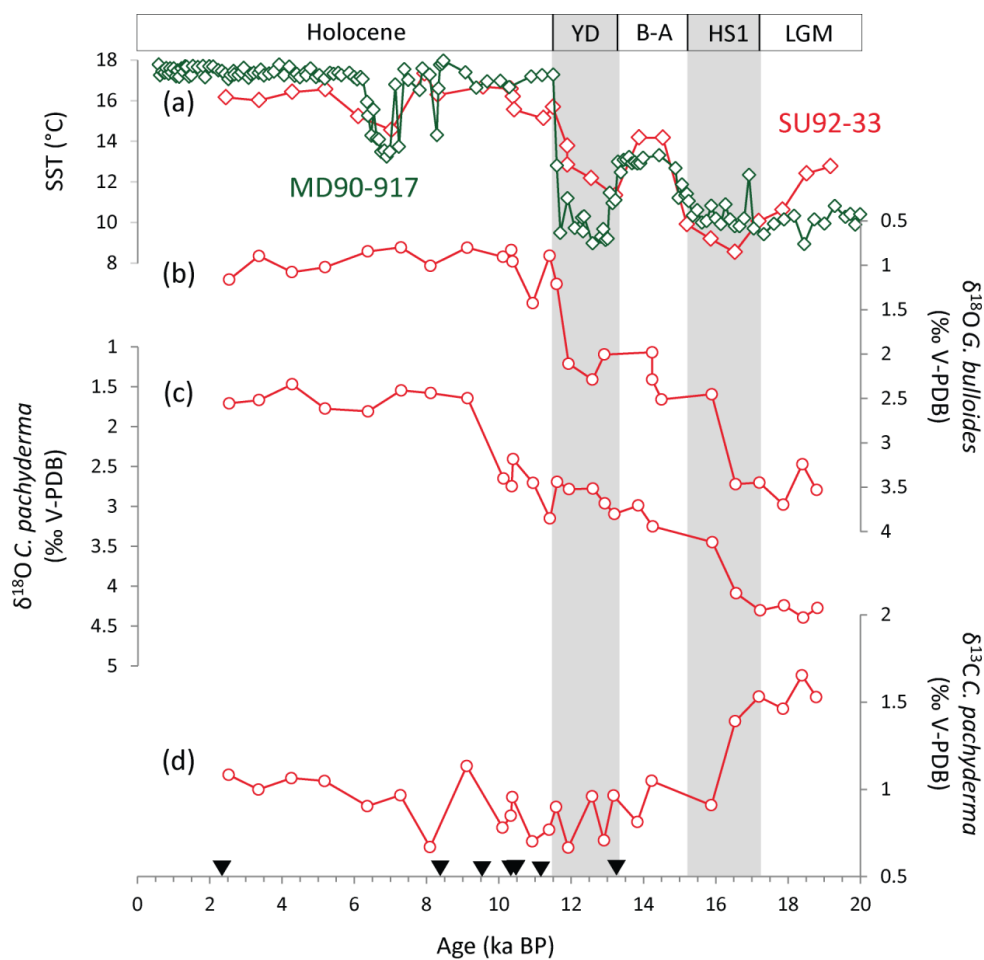


Figure 2



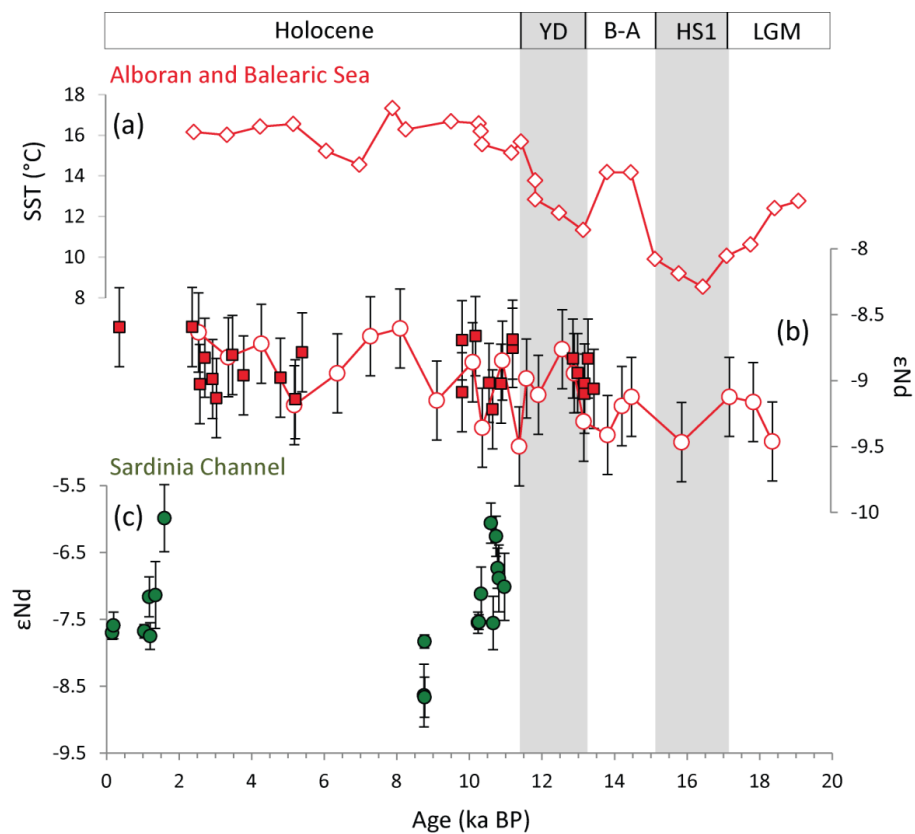


Figure 3

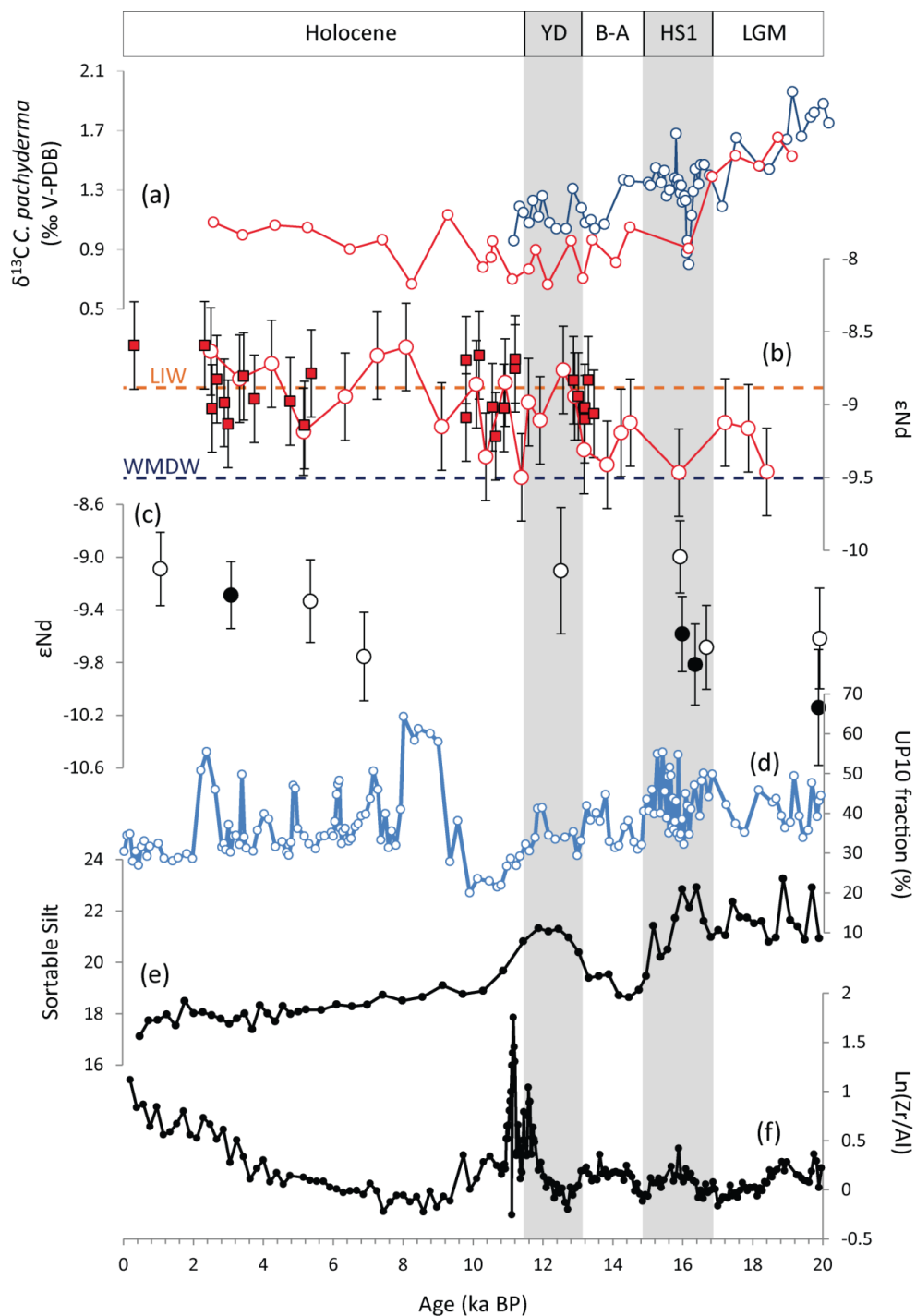


Figure 4

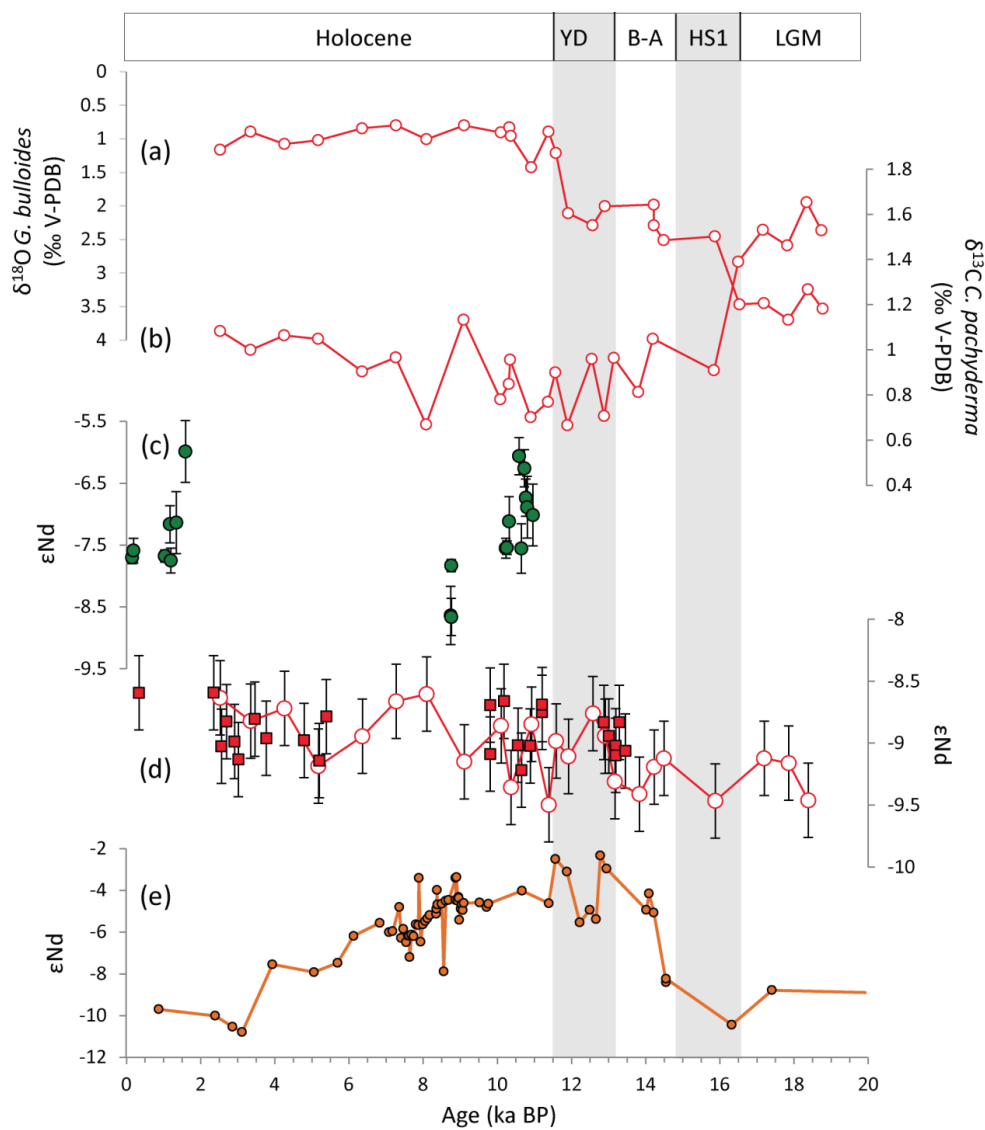


Figure 5



Sample ID	Depth in core (cm)	Corals species	$^{238}\text{U}$ ( $\mu\text{g/g}$ )	$^{232}\text{Th}$ ( $\text{ng/g}$ )	$\delta^{234}\text{U}_n$ (‰)	$^{230}\text{Th}/^{238}\text{U}$	$^{230}\text{Th}/^{232}\text{Th}$	Age (ka BP)	$\delta^{234}\text{U}_{(0)}$ (‰)	$^{143}\text{Nd}/^{144}\text{Nd}$	$\epsilon\text{Nd}$
RECORD_23_V	0-3.5	<i>Madræpora oculata</i>	3.31 ±0.005	0.68 ±0.014	151.85 ±1.7	0.00163 ±0.00011	25 ±1.7	0.091 ±0.011	151.92 ±1.7	0.512243 ±0.000005	-7.70 ±0.10
RECORD_23_V	3-7	<i>Madræpora oculata</i>	3.23 ±0.002	0.52 ±0.001	147.11 ±0.6	0.00199 ±0.00006	38 ±1.1	0.127 ±0.006	147.19 ±0.6	0.512249 ±0.000010	-7.59 ±0.20
RECORD_23_V	7-10	<i>Madræpora oculata</i>	3.99 ±0.007	0.25 ±0.002	147.52 ±1.7	0.01227 ±0.00022	640 ±11.6	1.110 ±0.023	148.01 ±1.7	0.512244 ±0.000015	-7.68 ±0.30
RECORD_23_V	8-10	<i>Madræpora oculata</i>	3.79 ±0.005	0.41 ±0.001	147.77 ±0.7	0.01253 ±0.00007	350 ±2.0	1.135 ±0.008	148.27 ±0.7	0.512271 ±0.000010	-7.16 ±0.20
RECORD_23_IV	6-9	<i>Madræpora oculata</i>	4.06 ±0.006	0.35 ±0.001	148.47 ±1.2	0.01366 ±0.00011	480 ±3.8	1.243 ±0.012	149.02 ±1.2	0.512241 ±0.000010	-7.75 ±0.20
RECORD_23_IV	27-30	<i>Madræpora oculata</i>	4.06 ±0.003	1.09 ±0.001	146.91 ±1.3	0.01405 ±0.00013	159 ±1.4	1.283 ±0.014	147.47 ±1.3	0.512272 ±0.000026	-7.14 ±0.50
RECORD_23_IV	37-40	<i>Madræpora oculata</i>	3.52 ±0.005	0.08 ±0.000	148.25 ±1.1	0.01663 ±0.00012	2308 ±16.4	1.529 ±0.013	148.92 ±1.1	0.512331 ±0.000026	-5.99 ±0.50
RECORD_23_III	55-57	<i>Madræpora oculata</i>	3.63 ±0.002	0.27 ±0.000	145.30 ±0.7	0.08832 ±0.00020	3530 ±8.1	8.685 ±0.027	148.93 ±0.8	0.512195 ±0.000026	-8.64 ±0.50
RECORD_23_III	58-61	<i>Madræpora oculata</i>	4.24 ±0.004	0.36 ±0.001	146.71 ±1.2	0.08859 ±0.00037	3336 ±14.0	8.702 ±0.048	150.39 ±1.2	0.512237 ±0.000010	-7.83 ±0.20
RECORD_23_III	63-66	<i>Lophelia pertusa</i>	4.15 ±0.005	0.42 ±0.002	147.19 ±0.8	0.08863 ±0.00054	2783 ±17.1	8.703 ±0.063	150.89 ±0.9	0.512194 ±0.000015	-8.66 ±0.30
RECORD_23_I	0-2	<i>Lophelia pertusa</i>	3.35 ±0.002	0.37 ±0.000	147.02 ±0.7	0.10283 ±0.00018	2788 ±4.8	10.173 ±0.025	151.34 ±0.7	0.512251 ±0.000010	-7.55 ±0.20
RECORD_23_II	62-65	<i>Lophelia pertusa</i>	3.27 ±0.003	0.39 ±0.002	144.75 ±1.2	0.10289 ±0.00061	1046 ±6.2	10.260 ±0.079	149.69 ±1.6	0.512273 ±0.000021	-7.12 ±0.40
RECORD_23_II	50-52	<i>Lophelia pertusa</i>	2.92 ±0.003	0.92 ±0.003	145.39 ±1.6	0.10351 ±0.00061	1046 ±6.2	10.201 ±0.075	149.01 ±1.2	0.512251 ±0.000010	-7.54 ±0.20
RECORD_23_I	12-14	<i>Lophelia pertusa</i>	3.07 ±0.002	0.49 ±0.000	145.22 ±0.7	0.10609 ±0.00023	1971 ±4.3	10.531 ±0.031	149.64 ±0.7	0.512327 ±0.000015	-6.06 ±0.30
RECORD_23_I	5-7	<i>Lophelia pertusa</i>	3.50 ±0.002	0.42 ±0.000	146.35 ±0.9	0.10677 ±0.00016	2654 ±4.0	10.591 ±0.025	150.82 ±0.9	0.512251 ±0.000021	-7.55 ±0.40
RECORD_23_II	94-98	<i>Lophelia pertusa</i>	3.14 ±0.003	0.62 ±0.002	146.42 ±1.0	0.10755 ±0.00047	1737 ±7.6	10.672 ±0.059	150.94 ±1.0	0.512317 ±0.000015	-6.26 ±0.30
RECORD_23_I	15-17	<i>Lophelia pertusa</i>	3.40 ±0.003	0.46 ±0.000	146.01 ±0.9	0.10790 ±0.00021	2409 ±4.6	10.713 ±0.031	150.53 ±0.9	0.512293 ±0.000015	-6.73 ±0.30
RECORD_23_II	96-100	<i>Lophelia pertusa</i>	3.61 ±0.004	0.35 ±0.001	145.50 ±0.8	0.10821 ±0.00044	3579 ±14.7	10.750 ±0.055	150.02 ±0.8	0.512285 ±0.000026	-6.89 ±0.50
RECORD_23_II	93-95	<i>Lophelia pertusa</i>	3.19 ±0.003	0.24 ±0.000	143.33 ±0.8	0.10947 ±0.00032	4381 ±12.7	10.904 ±0.042	147.85 ±0.9	0.512279 ±0.000026	-7.01 ±0.50

**Table 1**



Sample ID	Core depth (cm)	Species	Water Depth (m)	Median probability age (ka BP)	$^{143}\text{Nd}/^{144}\text{Nd}$	$\epsilon\text{Nd}$
GeoB 13727-1#1	Surface	<i>Lophelia pertusa</i>	363	0.339	0.512198 ±0.000015	-8.59 ±0.30
GeoB 13727-1#2	Surface	<i>Madrepora oculata</i>	353	2.351	0.512198 ±0.000015	-8.59 ±0.30
GeoB 13730-1	6	<i>Lophelia pertusa</i>	338	2.563	0.512175 ±0.000015	-9.03 ±0.30
GeoB 13728-1	Bulk (0-15)	<i>Lophelia pertusa</i>	343	2.698	0.512185 ±0.000015	-8.83 ±0.30
GeoB 13728-2	2	<i>Lophelia pertusa</i>	343	2.913	0.512177 ±0.000015	-8.99 ±0.30
GeoB 13722-3	Bulk (0-15)	<i>Madrepora oculata</i>	280	3.018	0.512170 ±0.000015	-9.13 ±0.30
GeoB 13722-3	Bulk (15-30)	<i>Madrepora oculata</i>	280	3.463	0.512186 ±0.000015	-8.81 ±0.30
GeoB 13735-1	Bulk (0-15)	<i>Madrepora oculata</i>	280	3.770	0.512179 ±0.000015	-8.96 ±0.30
GeoB 13723-1	Bulk (0-8)	<i>Madrepora oculata</i>	291	4.790	0.512178 ±0.000015	-8.98 ±0.30
GeoB 13725-2	Surface	<i>Madrepora oculata</i>	355	5.201	0.512169 ±0.000015	-9.14 ±0.30
GeoB 13723-1	Bulk (8-20)	<i>Madrepora oculata</i>	291	5.390	0.512187 ±0.000015	-8.79 ±0.30
GeoB 13729-1	2.5	<i>Lophelia pertusa</i>	442	9.810	0.512172 ±0.000015	-9.09 ±0.30
GeoB 13729-1	2.5	<i>Lophelia pertusa</i>	442	9.810	0.512193 ±0.000015	-8.69 ±0.30
GeoB 13729-1	49	<i>Lophelia pertusa</i>	442	10.181	0.512194 ±0.000015	-8.66 ±0.30
GeoB 13730-1	102	<i>Lophelia pertusa</i>	338	10.556	0.512176 ±0.000015	-9.02 ±0.30
GeoB 13730-1	194	<i>Lophelia pertusa</i>	338	10.652	0.512165 ±0.000015	-9.22 ±0.30
GeoB 13729-1	315	<i>Lophelia pertusa</i>	442	10.889	0.512176 ±0.000015	-9.02 ±0.30
GeoB 13729-1	375	<i>Lophelia pertusa</i>	442	11.206	0.512189 ±0.000015	-8.75 ±0.30
GeoB 13730-1	298	<i>Lophelia pertusa</i>	338	11.208	0.512193 ±0.000015	-8.69 ±0.30
GeoB 13728-2	191	<i>Lophelia pertusa</i>	343	12.874	0.512185 ±0.000015	-8.83 ±0.30
GeoB 13737-1#2	Surface	<i>Lophelia pertusa</i>	297	13.005	0.512180 ±0.000015	-8.94 ±0.30
GeoB 13728-2	295	<i>Lophelia pertusa</i>	364	13.194	0.512176 ±0.000015	-9.02 ±0.30
GeoB 13728-2	295	<i>Lophelia pertusa</i>	364	13.194	0.512171 ±0.000015	-9.10 ±0.30
GeoB 13730-1	427	<i>Lophelia pertusa</i>	338	13.291	0.512185 ±0.000015	-8.83 ±0.30
GeoB 13737-1#1	Surface	<i>Lophelia pertusa</i>	299	13.452	0.512174 ±0.000015	-9.06 ±0.30

Table 2



Core	Depth in core (cm)	<sup>14</sup> C-age (years)	±1σ (years)	Median probability age (ka BP)
SU92-33	0	2770	70	2437
SU92-33	64	7870	90	8280
SU92-33	70	8670	80	9528
SU92-33	74	9510	100	10295
SU92-33	84	9610	90	10389
SU92-33	90	10180	100	11192
SU92-33	120	11710	110	13172

**Table 3**



Depth in core (cm)	Age (ka BP)	$\delta^{13}\text{C}$		$\delta^{18}\text{O}$		$^{143}\text{Nd}/^{144}\text{Nd}$	$\epsilon\text{Nd}$
		<i>C. pachyderma</i> (‰ VPDB)	<i>C. pachyderma</i> (‰ VPDB)	<i>G. bulloides</i> (‰ VPDB)	<i>G. bulloides</i> (‰ VPDB)		
1	2.53	1.08	1.71	-0.6	1.16	0.512195 ±0.000015	-8.64 ±0.30
10	3.35	1.00	1.67	-0.82	0.90	0.512186 ±0.000015	-8.82 ±0.30
19.5	4.26	1.06	1.47	-0.55	1.08	0.512191 ±0.000015	-8.72 ±0.30
29.5	5.18	1.05	1.78	-0.55	1.02	0.512167 ±0.000015	-9.19 ±0.30
42.5	6.36	0.90	1.81	-0.91	0.84	0.512179 ±0.000015	-8.95 ±0.30
52.5	7.28	0.97	1.55	-0.80	0.80	0.512194 ±0.000015	-8.66 ±0.30
61.5	8.10	0.67	1.58	-0.95	1.01	0.512197 ±0.000015	-8.61 ±0.30
67.5	9.11	1.13	1.65	-1.07	0.80	0.512169 ±0.000015	-9.15 ±0.30
72.5	10.10	0.78	2.65	-1.27	0.91	0.512184 ±0.000015	-8.86 ±0.30
77.5	10.33	0.85	2.75	-1.10	0.83	-	-
81.5	10.37	0.96	2.41	-1.21	0.96	0.512158 ±0.000015	-9.36 ±0.30
87.5	10.92	0.70	2.71	-0.11	1.43	0.512184 ±0.000015	-8.85 ±0.30
92.5	11.39	0.77	3.15	-1.00	0.89	0.512151 ±0.000015	-9.50 ±0.30
95.5	11.59	0.90	2.69	-1.14	1.21	0.512178 ±0.000015	-8.98 ±0.30
100.5	11.92	0.67	2.78	-0.44	2.11	0.512171 ±0.000015	-9.11 ±0.30
110.5	12.58	0.96	2.78	-0.86	2.29	0.512189 ±0.000015	-8.76 ±0.30
115.5	12.91	0.71	2.96	-0.54	2.01	0.512180 ±0.000015	-8.94 ±0.30
119.5	13.17	0.96	3.09	-	-	0.512161 ±0.000015	-9.31 ±0.30
129.5	13.83	0.81	2.99	-	-	0.512156 ±0.000015	-9.41 ±0.30
135.5	14.23	1.05	3.25	-1.16	1.98	0.512167 ±0.000015	-9.19 ±0.30
135.5	14.23	-	-	-0.94	2.29	-	-
139.5	14.49	-	-	-0.96	2.51	0.512170 ±0.000015	-9.12 ±0.30
159.5	15.88	0.91	3.45	-0.81	2.45	0.512153 ±0.000015	-9.47 ±0.30
169.5	16.54	1.39	4.09	-0.76	3.47	-	-
179.5	17.20	1.53	4.30	-0.98	3.45	0.512170 ±0.000015	-9.12 ±0.30
190	17.86	1.46	4.24	-1.10	3.70	0.512168 ±0.000015	-9.16 ±0.30
198	18.39	1.65	4.39	-1.24	3.24	0.512153 ±0.000015	-9.46 ±0.30
206	18.78	1.53	4.28	-0.90	3.53	-	-

Table 4