1	Three distinct Holocene intervals of stalagmite deposition and non-
2	deposition revealed in NW Madagascar, and their paleoclimate
3	implications
4 5 6 7	Ny Riavo G. Voarintsoa <sup>1*†</sup> , L. Bruce Railsback <sup>1</sup> , George A. Brook <sup>2</sup> , Lixin Wang <sup>2</sup> , Gayatri Kathayat <sup>3</sup> , Hai Cheng <sup>3,4</sup> , Xianglei Li <sup>3</sup> , R. Lawrence Edwards <sup>4</sup> , Amos Fety Michel Rakotondrazafy <sup>5</sup> , Marie Olga Madison Razanatseheno <sup>5</sup>
9 10 11 12 13 14	<ul> <li><sup>1</sup> Department of Geology, University of Georgia, Athens, GA 30602-2501 U.S.A.</li> <li><sup>2</sup> Department of Geography, University of Georgia, Athens, Georgia, 30602-2502 U.S.A.</li> <li><sup>3</sup> Institute of Global Environmental Change, Xi'an Jiaotong University, Xi'an, Shaanxi 710049, P.R. China</li> <li><sup>4</sup> Department of Earth Sciences, University of Minnesota, Minneapolis, Minnesota 55455, U.S.A.</li> <li><sup>5</sup> Mention Sciences de la Terre et de l'Environnement, Domaine Sciences et Technologie, University d'Antananarivo, Madagascar</li> </ul>
16	*Correspondence to: Ny Riavo Voarintsoa ( <u>nv1@uga.edu</u> or <u>nyriavony@gmail.com</u> )
17 18 19	<sup>+</sup> Current address: Institute of Earth Sciences, The Hebrew University in Jerusalem, A. Safra Campus, 91904, Jerusalem, Israel (nyriavo.voarintsoa@mail.huji.ac.il)
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21	ABSTRACT
22	Petrographic features, mineralogy, and stable isotopes from two stalagmites, ANJB-2 and
23	MAJ-5, respectively from Anjohibe and Anjokipoty Caves, allow distinction of three intervals of the
24	Holocene in NW Madagascar. The Malagasy <i>early</i> Holocene (between c. 9.8 and 7.8 ka) and <i>late</i>
25	Holocene (after c. 1.6 ka) intervals (MEHI and MLHI, respectively) record evidence of stalagmite
26	deposition. The Malagasy <i>middle</i> Holocene interval (MMHI, between c. 7.8 ka and 1.6 ka) is
27	marked by a depositional hiatus of c. 6500 years.
28	Deposition of these stalagmites indicates that the two caves were sufficiently supplied with
29	water to allow stalagmite formation. This suggests that the MEHI and MLHI intervals may have
30	been comparatively wet in NW Madagascar. In contrast, the long-term depositional hiatus during
31	the MMHI implies it was relatively drier than the MEHI and the MLHI.
32	The alternating "wet/dry/wet" conditions during the Holocene may have been linked to
33	the long-term migrations of the Inter-Tropical Convergence Zone (ITCZ). When the ITCZ's mean
21	position is farther south NW Madagascar experiences wetter conditions, such as during the MEHL

and MLHI, and when it moves north, NW Madagascar climate becomes drier, such as during the
 MMHI. A similar wet/dry/wet succession during the Holocene has been reported in neighboring
 locations, such as southeastern Africa. Beyond these three subdivisions, the records also suggest
 wet conditions around the cold 8.2 ka event, suggesting a causal relationship. However, additional
 Southern Hemisphere high-resolution data will be needed to confirm this.

# 40 1. Introduction

41 Although much is known about Holocene climate change worldwide (Mayewski et al., 2004; 42 Wanner and Ritz, 2011; Wanner et al., 2011; 2015), high-resolution climate data for the Holocene 43 period is still regionally limited in the Southern Hemisphere (SH) (e.g., Wanner et al., 2008; Marcott 44 et al., 2013; Wanner et al., 2015), including Madagascar. This uneven distribution of data hinders 45 our understanding of the spatio-temporal characteristics of Holocene climate change, and the 46 forcings involved. For example, some forcings would have influenced the behavior of the Inter-Tropical Convergence Zone (ITCZ) as well as monsoonal responses in low- to mid-latitude regions 47 48 (e.g., Wanner et al., 2015; Talento and Barreiro, 2016). In fact, Madagascar is ideally located to 49 provide data on SH Holocene climate changes because of its location in Southwestern Indian 50 Ocean and because it is seasonally visited by the ITCZ (Fig. 1a). Furthermore, a karst belt with caves 51 extends from the north to the south of the island (Fig. 1c), crossing latitudinal climate belts, and 52 this could potentially be a source of stalagmite data. Thus, Madagascar is a natural laboratory to 53 study ITCZ dynamics over time. New records from Madagascar could fill gaps in paleoclimate 54 datasets for the SH that might help refine paleoclimate simulations, and thus provide a better 55 understanding of global circulation and land-atmosphere-ocean interactions during the 56 Holocene.

In this paper, we present records of stable isotopes, petrography, mineralogy, variability of layer-specific width (or LSW) from stalagmites from Anjohibe and Anjokipoty caves. Stalagmites are used because of their potential to store significant climatic information (e.g., Fairchild and Baker, 2012, p. 9–10), and in Anjohibe Cave, recent studies have shown the replicability of paleoclimate records from stalagmites (e.g., Burns et al., 2016).

Two stalagmites were investigated, and these allowed us to characterize Holocene climate change in NW Madagascar. First, we developed a record of climate change using the stalagmite proxy data. With a better understanding of Madagascar's paleoclimate, we then investigated
 possible drivers of tropical climate change to isolate the major factors controlling the hydrological
 cycle in NW Madagascar and surrounding regions during the Holocene.

67 **2.** Setting

# 68 2.1. Stalagmites and their setting

69 Stalagmites are secondary cave deposits that are  $CaCO_3$  precipitates from cave dripwater. 70 Calcium carbonate precipitation occurs mainly by  $CO_2$  degassing, which increases the pH of the 71 dripwater and thus increases the concentration of  $CO_3^{2^-}$ . In some cases, evaporation may also 72 contribute to increased  $Ca^{2+}$  and/or  $CO_3^{2^-}$  concentrations in dripwater.  $CO_2$  degassing occurs when 73 high-*P*CO<sub>2</sub> water from the epikarst encounters low-*P*CO<sub>2</sub> cave air. Evaporation occurs when 74 humidity inside the cave is relatively low. The fundamental equation for stalagmite deposition is:

75 
$$Ca_{(aq)}^{2+} + 2HCO_{3(aq)}^{-} \rightleftharpoons CaCO_{3(s)} + CO_{2(g)} + H_2O_{(l)}$$
 (Eq. 1)

76 Growth and non-growth of stalagmites depends on conditions that affect Eq. 1. An increase in  $Ca^{2+}$  drives the equation to the right (towards precipitation) and an increase in  $CO_2$  of the cave air 77 78 and/or  $H_2O$  drives it to the left (towards dissolution). All components of the equation are 79 influenced by the supply of water to the cave, which is generally climate-dependent. More water 80 enters the cave during warm/rainy seasons than during cold/dry seasons. Stalagmites will form 81 when cave dripwater is saturated with respect to calcite and/or aragonite. If the water passes 82 through the bedrock too quickly to dissolve significant carbonate rock, and/or enters the cave and reaches the stalagmite too quickly to degas significant CO<sub>2</sub>, it will not be saturated with respect to 83 84 CaCO<sub>3</sub>, inhibiting stalagmite formation. Stalagmite growth will slow as dripwater declines and will 85 stop entirely if flow ceases. Vegetation provides CO<sub>2</sub> to the soil via root respiration so the 86 vegetation cover above the cave and the type of vegetation can promote or limit stalagmite 87 growth. Overall, the karst hydrological system plays a crucial role in the deposition and non-88 deposition of stalagmites, and this is closely linked to changes in local and regional environment 89 and climate.

#### 2.2. Regional environmental setting

Stalagmites ANJB-2 and MAJ-5 were collected from Anjohibe and Anjokipoty caves, respectively, in the Majunga region of NW Madagascar (Fig. 1). Sediments and fossils from these caves have already provided many insights about the paleoenvironmental and archaeological history of NW Madagascar (e.g., Burney et al., 1997, 2004; Brook et al., 1999; Gommery et al., 2011; Jungers et al., 2008; Vasey et al., 2013; Burns et al., 2016; Voarintsoa et al., 2017b).

97 Anjohibe (S15° 32' 33.3"; E046° 53' 07.4") and Anjokipoty (S15° 34' 42.2"; E046° 44' 03.7") 98 are about 16.5 km apart (Fig. 1c). Their location in the zone visited by the ITCZ (e.g., Nassor and 99 Jury, 1998) makes them ideal sites to test the hypothesis that latitudinal migration of the ITCZ 100 influenced the Holocene climate of NW Madagascar (e.g., Chiang and Bitz, 2005; Broccoli et al., 101 2006; Chiang and Friedman, 2012; Schneider et al., 2014). Majunga has a tropical savanna climate 102 (Aw) according to the Köppen-Geiger climate classification, with a distinct wet summer (from 103 October to April) and dry winter (May-September). The mean annual rainfall is around 1160 mm. 104 The mean maximum temperature in November, the hottest month in the summer, is about 32°C. 105 The mean minimum temperature in July, the coldest month of the dry winter, is about 18°C (Fig. 106 1b).

107

# 108 2.3. Climate of Madagascar

109 The climate of Madagascar is unique because of its varied topography and its position in the 110 Indian Ocean (Figs. S1–S2; also see for e.g., Jury, 2003; DGM, 2008, Douglas and Zinke, 2015, p. 111 281-299; Voarintsoa et al., 2017b, p.138-139; Scroxton et al., 2017). Regionally distinct rainfall 112 gradients from east to west and from north to south are evident across the country (Jury, 2003; 113 Dewar and Richard, 2007), and these are linked to easterly trade winds in winter (May-October) 114 and northwesterly tropical storms in summer, respectively. In NW Madagascar, summer rainfall is 115 monsoonal and it is in phase with the seasonal southward migration of the ITCZ. Chiang and Bitz' 116 (2005) and Broccoli et al.'s (2006)'s have suggested that cooler/warmer intervals bring the ITCZ 117 south/north, thus regions in tropical SH are wet/dry. Generally, the ITCZ migrates towards the 118 Earth's warmer hemisphere (Frierson and Hwang, 2012; Kang et al., 2008; McGee et al., 2014; 119 Sachs et al., 2009). In fact, longer-term ITCZ migration appears to have affected climate in NW

Madagascar between c. 370 CE and 800 CE (see Fig, 8 of Voarintsoa et al., 2017b). This relationship
was inferred from changes in global climate conditions.

122 The climate of Madagascar is also influenced by changes in Indian Ocean sea surface 123 temperature (SST) (Zinke et al., 2004; see also Kunhert et al., 2014) and changes in SST the Aghulas 124 Current off southwestern Madagascar (Lutjeharms, 2006; Beal et al., 2011; Zinke et al., 2014). The 125 most immediate signal is the Indian Ocean Dipole (IOD), or Indian Ocean Zonal Mode (Li et al., 126 2003; Zinke et al., 2004), but El-Niño Southern Oscillation (ENSO) may also influence its climate 127 (e.g., Brook et al., 1999). IOD has been linked to Holocene climate variability in tropical Indian 128 Ocean (Abram et al., 2009; Tierney et al., 2013). However, its linkages to ENSO is still debated (e.g., 129 Saji et al., 1999; Li et al., 2003; Lee et al., 2008; Brown et al., 2009; Schott et al., 2009; Shinoda et 130 al., 2004; Venzke et al., 2000; Abram et al., 2008; Saji and Yagamata, 2003; Meyers et al., 2007). 131 The complex interactions between these inter-annual climatic factors make them an ideal topic 132 for further investigation using high-resolution records, and thus they will not be the focus of this 133 paper. However, their possible effects are referred to briefly in Supplementary text no. 4.

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#### 2.4. The Holocene in NW Madagascar

136 Little is hitherto known about Holocene climate change in NW Madagascar nor about the 137 major drivers of long-term climatic changes there. Most paleoclimate information from this region 138 covers the last two millennia with more focus on the anthropogenic effects on the Malagasy 139 ecosystems (e.g., Crowley and Samonds, 2013; Burns et al., 2016; Voarintsoa et al., 2017b). This is 140 because several studies show that megafaunal extinctions in Madagascar coincide with the arrival 141 of humans around 2-3 thousand (ka) BP (e.g., see Table 1 of Virah-Sawmy et al., 2010; MacPhee 142 and Burney, 1991; Burney et al., 1997; Crowley, 2010). There are even fewer long-term 143 paleoclimate records for the NW region, with only sediments from Lake Mitsinjo (3.5 ka BP; 144 Matsumoto and Burney, 1994) and stalagmites from Anjohibe Cave (40 ka BP; Burney et al. 1997) 145 providing records of more than 3 ka. Even though these records provide useful information about 146 paleoenvironmental changes in NW Madagascar, links to global climatic changes, particularly the links to changes in ITCZ, are not yet fully understood. 147

#### 148 **3.** Methods

# 149 3.1. Radiometric dating

150 A total of 22 samples were drilled from Stalagmite ANJB-2 and 9 samples for Stalagmite 151 MAJ-5 for U-series dating (Table S1 and S2). Each sample is a long (~5 to 20 mm), narrow (~1-152 2mm), and shallow (~1 mm) trench, allowing us to extract 50-250 mg of CaCO<sub>3</sub> powder. We 153 followed the chemical procedures described in Edwards et al. (1987) and Shen et al. (2002) when 154 separating uranium and thorium. U/Th measurements were performed on the multi-collector ICP-155 MS of the University of Minnesota, USA and on a similar instrument in the Stable Isotopes 156 Laboratory of Xi'an, in Jiaotong, China. Instrument details are provided in Cheng et al. (2013). Corrected <sup>230</sup>Th ages assume an initial <sup>230</sup>Th/<sup>232</sup>Th atomic ratio of  $4.4 \pm 2.2 \times 10^{-6}$ . This is the ratio 157 for "bulk earth" or crustal material at secular equilibrium with a  $^{232}$ Th/ $^{238}$ U value of 3.8. The 158 159 uncertainty in the "bulk earth" value is assumed to be ±50% (see footnotes to Tables S1 and S2). 160 The error in the final "corrected age" incorporates this uncertainty. The radiometric data are 161 reported as years BP, where BP is Before Present, and "Present" is A.D. 1950. Stalagmite 162 chronologies were constructed using the StalAge1.0 algorithm of Scholz and Hoffman (2011) and 163 Scholz et al. (2012), an algorithm using a Monte-Carlo simulation. The algorithm can identify major 164 and minor outliers and age inversions. The StalAge scripts were run on the statistics program R 165 version 3.2.2 (2015-08-14). The age models were adjusted considering hiatal surfaces identified in 166 the samples, using the approach of Railsback et al. (2013; see their Fig. 9).

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### 168 3.2. Petrography and mineralogy

169 Petrography and mineralogy of the two stalagmites were investigated 1) by examining both 170 the polished surfaces and the scanned images of the sectioned stalagmites, and by identifying any 171 diagenetic fabrics (e.g., Zhang et al., 2014) that could potentially affect stable isotope values, 2) by 172 observing eleven oversized thin sections (3x2 in) under the Leitz Laborlux 12 Pol microscope and 173 the Leica DMLP equipped with QCapture in the Sedimentary Geochemistry Lab at the University 174 of Georgia, 3) by using scanning electron microscopy (SEM) to better understand the 175 mineralogical fabrics at locations of interest (Fig. S13), and 4) by analyzing about 30–100 mg of 176 powdered spelean layers (n=15) on a Bruker D8 X-ray Diffractometer in the Department of 177 Geology, University of Georgia. For calcite and aragonite identification, we used CoK $\alpha$  radiation at 178 a 2 $\theta$  angle between 20° and 60°.

179 Layer-specific width (LSW) of clearly-defined layers was measured at selected locations on 180 the stalagmite polished surfaces (Fig. S4; Sletten et al., 2013; Railsback et al., 2014; Voarintsoa et 181 al., 2017b). LSW is the horizontal distance between two points on the flanks of the stalagmite 182 where convexity is greatest. It is the width near the top of the stalagmite when the layer being 183 examined was deposited. LSW is measured at right angles to the growth axis of the stalagmite; it 184 is the horizontal distance between two points on the layer growth surface, at which a virtual line 185 inclined at 35° to the growth axis becomes tangent to the layer growth surface as shown in Fig. 186 S4. LSW may vary along the length of the stalagmite, with smaller values suggesting drier 187 conditions and larger values wetter conditions.

188

# 189 3.3. Stable isotopes

190 Samples of 50–100  $\mu$ g were drilled along the stalagmite's growth axis for stable isotope 191 analysis. The trench size is very small (1.5 x 0.5 x 0.5 mm). Since a small mixture of calcite and aragonite could potentially change the  $\delta^{18}$ O and  $\delta^{13}$ C of the measured spelean layers (see for 192 193 example Frisia et al., 2002), drilling and sample extraction were carefully done on individually 194 discrete layers using the smallest drill-bit head (SSW-HP-1/4) to avoid potential mixing between 195 calcite and aragonite. The polished surface of the two stalagmites were examined to see if features 196 of diagenetic alteration are present (see for example fig. 2 of Zhang et al., 2014), but none was 197 found. During sampling, the mineralogy at the crest, where stable isotope samples were extracted, 198 was recorded for future mineralogical correction.

199 Aragonite oxygen and carbon isotopic corrections were performed to compensate for 200 aragonite's inherent fractionation of heavier isotopes (e.g., Romanek et al., 1992; Kim et al., 2007; 201 McMillan et al., 2005) and to remove the mineralogical bias in isotopic interpretation between 202 calcite and aragonite. The correction consists of subtracting 0.8‰ for  $\delta^{18}$ O (Kim and O'Neil, 1997; 203 Tarutani et al., 1969; Kim et al., 2007; Zhang et al., 2014) and 1.7‰ for  $\delta^{13}$ C (Rubinson and Clayton, 204 1969; Romanek et al., 1992) for the aragonite, as has been done previously (e.g., Holmgren et al., 205 2003; Sletten et al., 2013; Liang et al., 2015; Railsback et al., 2016; Voarintsoa et al., 2017a), as 206 shown in equations 2 and 3 below (where  $R_{A/C}$  is the aragonite percentage if not 100%).

207  $\delta^{18}O_{\text{corr.}}$  (‰, VPDB) =  $\delta^{18}O_{\text{uncorr.}}$  (‰, VPDB) – [R<sub>A/C</sub> x 0.8 (‰, VPDB)] (Eq. 2)

208  $\delta^{13}C_{corr.}$  (‰, VPDB) =  $\delta^{13}C_{uncorr.}$  (‰, VPDB) – [R<sub>A/C</sub> x 1.7 (‰, VPDB)] (Eq. 3)

209 Supplementary Figures S6–S8 show both the corrected and uncorrected isotopic records.

210 For the analytical methods, oxygen and carbon isotope ratios were measured using the 211 Finnigan MAT-253 mass spectrometer fitted with the Kiel IV Carbonate Device of the Xi'an Stable 212 Isotope Laboratory in China (ANJB-2; n=654) and using the Delta V Plus at 50°C fitted with the 213 GasBench-IRMS machine of the Alabama Stable Isotope Laboratory in USA (MAJ-5; n=286). 214 Analytical procedures using the MAT 253 are identical to those described in Dykoski et al. (2005), with isotopic measurement errors of less than 0.1 % for both  $\delta^{13}$ C and  $\delta^{18}$ O. Analytical methods 215 216 and procedures using the GasBench-IRMS machine are identical to those described in Skrzypek 217 and Paul (2006), Paul and Skrzypek (2007), and Lambert and Aharon (2011), with ±0.1 ‰ errors for both  $\delta^{13}$ C and  $\delta^{18}$ O. In both techniques, the results are reported relative to Vienna PeeDee 218 219 Belemnite (VPDB) and with standardization relative to NBS19. An inter-lab comparison of the 220 isotopic results was conducted, and it involved replicating every tenth sample of Stalagmite MAJ-221 5 at both labs. This exercise showed a strong correlation between the lab results (Fig. S5).

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#### **4.** Results

# 224 4.1. Radiometric data

225 Results from radiometric analyses of the two stalagmites are presented in Tables S1 and S2. Corrected <sup>230</sup>Th ages suggest that Stalagmite ANJB-2 was deposited between c. 8977±50 and 226 227 c. 161±64 years BP, and Stalagmite MAJ-5 was deposited between c. 9796±64 and c. 150±24 years 228 BP. These ages collectively indicate stalagmite deposition at the beginning (between c. 9.8 and 7.8 229 ka BP) and at the end of the Holocene (after c. 1.6 ka BP). In both stalagmites, the older ages have 230 small  $2\sigma$  errors and they generally fall in correct stratigraphic order, except sample ANJB-2-120 231 and its replicate ANJB-2-120R, which were not used because of the sample's high porosity and 232 high detritals content. In contrast, many of the younger ages have larger uncertainties. This is 233 mainly because many of the younger samples have very low uranium concentration and the

234 detrital thorium concentration is also high, similar to what Dorale et al. (2004) reported. We also understand that the value for initial  $^{230}$ Th correction, i.e. the initial  $^{230}$ Th/ $^{232}$ Th atomic ratio of 4.4 235  $\pm 2.2 \times 10^{-6}$  for a bulk earth with a <sup>232</sup>Th/<sup>238</sup>U value of 3.8, in these samples could have slightly 236 altered the <sup>230</sup>Th age of these younger samples, leading to larger uncertainties (such as discussed 237 238 in Lachniet et al., 2012). We encountered similar problems while working on other younger 239 samples from the same cave, but we compared the stable isotope profile with other published 240 records using isochron dating methods, and results did not differ significantly (see Fig. 9 of 241 Voarintsoa et al., 2017b). Since this work does not focus on decadal or centennial interpretation 242 of the Late Holocene stable isotope data, additional chronology adjustment has not been made, 243 and we used the chronology from StalAge to construct the time series. However, in Figures 5 and 244 6, age uncertainties are given below the stable isotope profiles so that comparisons with other 245 records can accommodate these uncertainties.

The key finding from our age and petrographic data for the two stalagmites is that they indicate three distinct intervals of growth and non-growth during the Holocene (Figs. 2–4, 7). The evidence for this includes: (1) CaCO<sub>3</sub> deposition between c. 9.8 and 7.8 ka BP, (2) a long depositional hiatus between c. 7.8 and 1.6 ka BP, and (3) resumption of CaCO<sub>3</sub> deposition after c. 1.6 ka BP. In the rest of the paper, we will refer to these intervals as the Malagasy Early Holocene Interval (MEHI), Malagasy Mid-Holocene Interval (MMHI), and Malagasy Late Holocene Interval (MLHI), respectively.

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# 4.2. Stable isotopes

Raw values of  $\delta^{18}$ O and  $\delta^{13}$ C for Stalagmite ANJB-2 range from -8.9 to -2.3‰ (mean = -5.0‰), and from -11.0 to +5.2‰ (mean = -4.2‰), respectively, relative to VPDB. Raw values of  $\delta^{18}$ O and  $\delta^{13}$ C for Stalagmite MAJ-5 range from -8.8 to -0.9‰ (mean = -4.9‰), and from -9.4 to +2.6‰ (mean = -4.4‰), respectively, relative to VPDB. Mean  $\delta^{18}$ O and  $\delta^{13}$ C values are distinguishable between the MEHI and the MLHI. In both stalagmites, the amplitude of  $\delta^{18}$ O fluctuations was fairly constant throughout the Holocene; whereas the  $\delta^{13}$ C records show a dramatic shift toward higher values (i.e. from -10.9‰ to +3.8‰, VPDB) at c. 1.5 ka BP.

262 The MEHI and MLHI are isotopically distinct (Fig. 4). The MEHI is characterized by statistically correlated  $\delta^{18}$ O and  $\delta^{13}$ C (r<sup>2</sup>=0.65 and 0.53), and much depleted  $\delta^{13}$ C values (c -11.0 to -4.0 ‰). 263 A prominent isotopic excursion is evident between c. 8.1 and c. 8.3 ka BP (Fig. 5), when stalagmite 264  $\delta^{18}$ O and  $\delta^{13}$ C ratios reach their lowest values of -6.8 and -10.9‰, respectively. In contrast to the 265 MEHI, the values of  $\delta^{18}$ O and  $\delta^{13}$ C during the MLHI are poorly correlated (r<sup>2</sup>=0.25 and 0.17), and 266  $\delta^{13}$ C values are more enriched (Figs. 4, 6). Since Stalagmites ANJB-2 and MAJ-5 were collected 267 268 from two different caves, 16 km apart, it is not surprising to see discrepancies between the stable 269 isotopes during similar intervals, suggesting that local karst conditions could be one of the 270 discrepancy factors. Another potential source for the discrepancy is the larger uncertainty of the 271 younger ages due to low uranium and high detrital concentrations. This U-Th aspect has been a 272 challenge for several young stalagmites (e.g., Dorale et al., 2004; Lachniet et al., 2012) including 273 samples from NW Madagascar (this study). While the utility of speleothems as a climate proxy 274 largely depends on replication of stable isotope values, this study specifically highlights the 275 replication of stalagmite deposition and non-deposition and the isotopic characteristics of each 276 depositional intervals of the Holocene.

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### 4.3. Mineralogy, petrography, and layer-specific width

In both stalagmites ANJB-2 and MAJ-5, the hiatus of deposition is characterized by a welldeveloped Type L surface (Figs. 2, 3, S15). Petrography and mineralogy are distinct before and after this hiatus (Figs. 3, 5–6). Below the hiatus, laminations are well preserved in both stalagmites. Above the hiatus, laminations are not well-preserved, although noted in some intervals.

In Stalagmite ANJB-2, LSW varies from 37 to 26.5 mm with a mean of 30 mm. It decreases to 28 mm at the hiatus (Fig. 3). The mineralogy is dominated by aragonite below the hiatus, although there are also a few thick layers of primary calcite. A thin (~2-3 mm) layer of white, very soft, and porous aragonite is identified just below the hiatus (Fig. S15). This layer is also calcite and aragonite, with calcite dominant, and the calcite layers contain macro-cavities that are mostly offaxis macroholes (Shtober-Zisu et al., 2012).

As noted in previous section (4.2), there is a prominent isotopic excursion at c. 8.2 ka BP, and this excursion is in the calcite layer in Stalagmite ANJB-2 at 195–202 mm from its top. X-ray

291 diffraction spectrum from this layer suggests that the mineralogy is 100% calcite (Figs. S14, S16-292 S17). We believe the calcite to be primary and not a diagenetic product of aragonite for three 293 reasons. First, the laminations in the thick layer of calcite were not altered (Figs. S16–S17). Second, 294 the polished surface of the stalagmite shows no evidence of fiber relicts and textural ghosts such 295 as observed in Juxtlahuaca Cave in southwestern Mexico (Lachniet et al., 2012) and in Shennong 296 Cave in southeastern China (Zhang et al., 2014). Third, petrographic comparison with known 297 examples of primary and secondary calcite observed under microscope (e.g., Railsback, 2000; 298 Perrin et al., 2014) suggests that there is no strong evidence of aragonite-to-calcite 299 transformation.

In Stalagmite MAJ-5, LSW varies from 50 to 22 mm with a mean of 35.5 mm. It decreases to 22 mm at the hiatus (Fig. 3). Below the hiatus, mineralogy is an even mixture between calcite and aragonite. Above the hiatus, mineralogy is mainly calcite, except the uppermost 2 mm where mineralogy is 75% calcite and 25% aragonite. Macro-cavities are also present throughout this upper part of Stalagmite MAJ-5.

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# 306 4.4. Summary of results

307 The various proxy climate records from Stalagmites ANJB-2 and MAJ-5 suggest three 308 distinct climate/hydrological intervals during the Holocene. The MEHI (c. 9.8 to 7.8 ka BP), with evidence of stalagmite deposition, is characterized by statistically correlated  $\delta^{18}$ O and  $\delta^{13}$ C 309  $(r^2=0.65 \text{ and } 0.53)$  and more negative  $\delta^{13}$ C values (c. -11.0 to -4.0 ‰). The MMHI (c. 7.8 to 1.6 ka 310 311 BP) is marked by a long-term hiatus in deposition, which is preceded by a well-developed Type L 312 surface in both Stalagmite ANJB-2 and MAJ-5 (Figs. 3, S15). The Type L surface is observed as an 313 upward narrowing of stalagmite width and layer thickness. It is best developed in Stalagmite MAJ-314 5 (Fig. S15). In Stalagmite ANJB-2, the hiatus at the Type L surface is additionally preceded by a c. 315 3 mm thick layer of highly porous, very soft, and fibrous white crystals of aragonite (the only 316 aragonite with such properties). This aragonite is topped by a thin and well-defined layer of detrital 317 materials (Fig. S15), further evidence of a hiatus. Finally, the MLHI (after c. 1.6 ka BP) is characterized by poorly correlated  $\delta^{18}$ O and  $\delta^{13}$ C (r<sup>2</sup>=0.25–0.17), and by a marked shift toward 318 higher  $\delta^{13}$ C values (Figs. 4, 6). 319

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### 321 5. Discussion

#### 322 5.1. Paleoclimate significance of stalagmite growth and non-growth: implications for

323 paleohydrology

324 Growth and non-growth of stalagmites depends on several factors linked to water 325 availability, which is largely determined by climate (more water during warm/rainy seasons and 326 less water during cold/dry seasons). Water is the main dissolution and transporting agent for most 327 chemicals in speleothems. Cave hydrology varies significantly over time in response to climate, 328 and this variability influences the formation or dissolution of  $CaCO_3$ . In this regard, calcium 329 carbonate does not form if there is little or no water entering the cave, or if there is too much (see 330 Sect. 2.1). Absence of groundwater recharge most typically occurs during extremely dry 331 conditions, whereas excessive water input to the cave occurs during extremely wet conditions. In 332 the latter scenario, water is undersaturated and flow rates are too fast to allow degassing. Often, 333 water availability is reflected in the extent of vegetation above and around the cave, as plants 334 require soil moisture or shallow groundwater to survive and propagate, and this contributes to 335 the stalagmites' processes of formation. The link between stalagmite growth/non-growth and 336 cave dripwater and soil CO<sub>2</sub> is broadly influenced by changes in climate.

337 Major hiatuses in stalagmite deposition could be marked by a variety of features, including 338 the presence of erosional surfaces, chalkification, dirt bands/detrital layers, offsetting of the 339 growth axis, and/or sometimes by color changes (e.g., Holmgren et al., 1995; Dutton et al., 2009; 340 Railsback et al., 2013; Railsback et al., 2015; Voarintsoa et al., 2017a). Railsback et al. (2013) were 341 able to identify significant features in stalagmites that allow distinction between non-deposition 342 during extremely wet (Type E surfaces) and non-deposition during extremely dry conditions (Type 343 L surfaces; Fig. 3). Physical properties of stalagmites that are evidence of extreme dry and wet 344 events are summarized in Table 1 of Railsback et al. (2013) and the mechanisms are explained in 345 their Figure 5.

Type E surfaces are layer-bounding surfaces between two spelean layers when the underlying layers show evidence of truncation. The truncation results from dissolution or erosion (thus the name "E") of previously-formed stalagmites layers by abundant undersaturated water. Type E surfaces are commonly capped with a layer of calcite (Railsback et al., 2013). This mineralogical trend is not surprising as calcite commonly forms under wetter conditions (e.g., Murray, 1954; Pobeguin, 1965; Siegel, 1965; Thrailkill, 1971; Cabrol and Coudray, 1982; Railsback et al. 1994; Frisia et al., 2002). Additionally, non-carbonate detrital materials are commonly abundant with varying grain size (i.e., from silt- to sand-size; Railsback et al., 2013).

354 Type L surfaces, on the other hand, are layer-bounding surfaces where the layers become 355 narrower upward and thinner towards the flanks of the stalagmite. Decreases in layer thickness 356 and stalagmites width upward are indications of lessening deposition (thus the name "L"; Railsback 357 et al., 2013). Aragonite is a very common mineralogy below a Type L surface, especially in warmer 358 settings. Layers of aragonite commonly form under drier conditions (Murray, 1954; Pobeguin, 359 1965; Siegel, 1965; Thrailkill, 1971; Cabrol and Coudray, 1982; Railsback et al., 1994; Frisia et al., 360 2002). Non-carbonate detrital materials are scarce, and if present, they tend to form a very thin 361 horizon of very fine dust material (Railsback et al., 2013). Identification of Type L surfaces is aided 362 by measuring the LSW (e.g., Sletten et al., 2013; Railsback et al., 2014; Fig. S4).

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# 5.2. Holocene climate in NW Madagascar

The age models and petrographic features of Stalagmites ANJB-2 and MAJ-5 suggest three
 distinct Holocene climate intervals (MEHI, MMHI, and MLHI; see Sect. 4.1) in NW Madagascar.
 Possible conditions during these intervals are illustrated in the sketches of Figure 4.

- 368
- 369 5.2.1. Malagasy early Holocene interval (c. 9.8 7.8 ka BP)

370 Stalagmite deposition during the early Holocene suggests that the chambers where 371 stalagmites ANJB-2 and MAJ-5 were collected were sufficiently supplied with water to allow CaCO<sub>3</sub> 372 precipitation, in accord with Eq.1. This in turn implies relatively wet conditions that could indicate 373 longer summer rainy seasons relative to modern climate, or wet years in NW Madagascar. The 374 correlative  $\delta^{13}$ C and  $\delta^{18}$ O values further suggest that vegetation consistently responded to changes 375 in moisture availability, which in turn was dependent on climate.

376 The prominent negative  $\delta^{18}$ O and  $\delta^{13}$ C excursions in Stalagmite ANJB-2 (Sect. 4.2; Figs. 5 377 and 10) are parallel to the  $\delta^{18}$ O excursion of the Greenland ice core records at c. 8.2 ka BP (e.g. Alley et al., 1997). The decrease in  $\delta^{18}$ O and  $\delta^{13}$ C values and the presence of calcite mineralogy at the same interval combine to suggest a wet 8.2 ka BP event in NW Madagascar. The 8.2 ka BP event was triggered by a release of freshwater from the melting Laurentide Ice Sheet into the North Atlantic basin, bringing cooler conditions in several NH regions (e.g., Alley et al., 1997; Barber et al., 1999), and via global teleconnections, this may have affected climate in NW Madagascar (see Sect. 5.5).

384 The MEHI terminated when conditions became much drier, as suggested by increasing  $\delta^{18}$ O and  $\delta^{13}$ C values in Stalagmite ANJB-2, by decreasing LSW in both stalagmites, and by the presence 385 386 of a major Type L surfaces in both stalagmites. The thin (c. 3 mm), porous, and white aragonite 387 layer in Stalagmite ANJB-2, a very similar deposit to that described in Niggemann et al. (2003), 388 suggests that the terminal drought was at times severe. Aragonite is a CaCO<sub>3</sub> polymorph that forms 389 preferentially under drier conditions (Murray, 1954; Pobeguin, 1965; Siegel, 1965; Thrailkill, 1971; 390 Cabrol and Coudray, 1982; Railsback et al. 1994; Frisia et al., 2002). The porous aragonite layer in 391 Stalagmite ANJB-2 is capped by a very thin layer of non-carbonate, brown detritus, which may 392 have been transported to the stalagmite as an aerosol and accumulated on the dry stalagmite 393 surface over time. Accumulation of the detritus must take place in the absence of dripwater (e.g., 394 Railsback et al., 2013). A shift to drier conditions is also supported by isotopic data from Stalagmite ANJ94-5 from Anjohibe Cave (Wang and Brook, 2013; Wang, 2016) in which relatively low  $\delta^{13}$ C 395 and  $\delta^{18}$ O values prior to 7.6 ka BP give way to episodically greater values thereafter. 396

- 397
- 398 5.2.2. Malagasy mid-Holocene interval (c. 7.8–1.6 ka BP)

The MMHI was a long (~6.5 ka) depositional hiatus in both stalagmites (Figs. 2–3), potentially suggesting dry conditions. The question is why did neither stalagmite grow during the MMHI? Here, we try to explain the factors and the climatic conditions that may have been responsible.

The documented severe dry conditions at the end of the MEHI (see Sect. 5.2.1) could have had a significant influence (1) on the cave hydrological system (e.g., Fig. 5 of Asrat et al., 2007; Bosák, 2010), such as the water conduits (primary or secondary porosity) to the chambers, and (2) on the vegetation cover above the caves, particularly above the chambers where Stalagmites

407 ANJB-2 and MAJ-5 were collected. On one hand, it is possible that the dry conditions late in the 408 MEHI could not only bring lesser water recharge to the cave, but also lowered the hydraulic head, 409 and increased the rate of evapo-transpiration in the vadose zone. This condition possibly allowed 410 more air to penetrate the aquifer, perhaps enhancing prior carbonate precipitation (PCP) in pores 411 and conduits above the caves (e.g., Fairchild and McMillan, 2007; Fairchild et al., 2000; Johnson 412 et al., 2006; Karmann et al., 2007; McDonald et al., 2007). This process must have blocked water 413 moving towards Stalagmites ANJB-2 and MAJ-5. On the other hand, the late MEHI drying trend 414 (Sect. 5.2.1) could have challenged vegetation to grow, and we assume that some areas above 415 Anjohibe and Anjokipoty caves must have been devoid of vegetation. Consequently, biomass 416 activities could have been reduced. Because vegetation contributes CO2 to the carbonic acid 417 dissolving CaCO<sub>3</sub>, its absence in certain areas above the cave could decrease the pH of the 418 percolating water, and perhaps dissolution did not occur. Under these conditions, even if water 419 reached the stalagmites, it may not have precipitated carbonate.

420 Whatever factors were responsible for the long term-depositional hiatus in Stalagmite 421 ANJB-2 and MAJ-5, we believe that the hiatus was caused by disturbances to water catchments 422 that feed the chambers at Anjohibe and Anjokipoty caves. The disturbances could be inherited 423 from the very dry conditions at the end of the MEHI, and/or due to the lack of water supply, 424 perhaps associated with an increase in epikarst ventilation, and/or by the absence of vegetation. 425 Water and vegetation are two components of the karst system that play an important role in 426 CaCO<sub>3</sub> dissolution and precipitation (see Eq. 1). Their disturbance may have limited limestone 427 dissolution in the epikarst and then carbonate precipitation in the cave zone.

428 Other evidence supports the idea of at least episodic dryness during the MMHI. A work on 429 a 2-meter long stalagmite (ANJ94-5) from Anjohibe Cave suggests episodic dryness during the 430 MMHI and a depositional hiatus around the time when Stalagmites ANJB-2 and MAJ-5 stopped 431 growing (Wang and Brook, 2013; Wang, 2016). For regional comparison, dry spells were also felt 432 in Central and Southeastern Madagascar (e.g., Gasse and Van Campo, 1998; Virah-Sawmy et al., 433 2009).

In summary, several lines of evidence suggest relatively drier climate in NW Madagascarduring the MMHI compared to the MEHI. Drier intervals generally imply drier summer seasons

with less rainfall (Fig. 8), perhaps reflecting shorter visits by the ITCZ. In this regard, even though
the region received rainfall, the necessary conditions could not have been attained to activate the
growth of Stalagmites ANJB-2 and MAJ-5, thus the hiatuses.

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### 5.2.3. Malagasy Late Holocene Interval (c. 1.6 ka BP–present)

441 Resumption of stalagmite deposition after c. 1.6 ka BP suggests a wetter climate in NW 442 Madagascar with reactivation of the previous epikarst hydrologic system. Climatic conditions must 443 have been similar to those of the early Holocene. The sudden beginning of stalagmite growth during the MLHI and the large  $\delta^{13}$ C shift from depleted to enriched values at c. 1.5 ka BP (Fig. 6), 444 445 after such long hiatuses may have been associated with changes in vegetation cover above the 446 cave linked to human activities (e.g., Burns et al., 2016; Crowley and Samonds, 2013; Crowther et al., 2016; Voarintsoa et al., 2017b). Lower  $\delta^{13}$ C values in Stalagmite MAJ-5 after 0.8 ka BP (Fig. 3), 447 448 compared to higher values in Stalagmite ANJB-2, may suggest different local karst conditions, 449 either natural, human-induced, or something else, at each site. Further investigations will be 450 necessary to better understand this.

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# 452 5.3. Holocene climate in NW Madagascar: implications for ITCZ dynamics

453 In NW Madagascar, stalagmite deposition during the MEHI and the MLHI suggests there 454 was sufficient dripwater for stalagmite growth and therefore wetter conditions. This may indicate 455 a more southerly mean position of the ITCZ. Factors that could influence the mean position of the 456 ITCZ include changes in insolation (e.g., Haug et al., 2001; Wang et al., 2005; Cruz et al., 2005; 457 Fleitmann et al., 2003, 2007; Schefuß et al., 2005; Suziki, 2011; Kutzbach and Liu, 1997; Partridge 458 et al., 1997; Verschuren et al., 2009; Voarintsoa et al., 2017a) and difference in temperature 459 between the two hemispheres (e.g., Chiang and Bitz, 2005; Broccoli et al., 2006; Chiang and 460 Friedman, 2012; Kang et al., 2008; McGee et al., 2014; Talento and Barreiro, 2016).

In contrast, the depositional hiatuses during the MMHI could suggest overall drier conditions, and thus a northward migration of the mean ITCZ. It may agree with the paleoclimate simulation of Braconnot et al. (2007), although the simulation is of shorter term than the MMHI hiatus, but additional paleoclimate records are needed to improve its spatial and temporal resolution. A northward shift in the mean position of the ITCZ is consistent with drier conditions
in the southern tropics, e.g., weaker South American Summer Monsoon (Cruz et al., 2005; Seltzer
et al., 2000; Wang et al., 2007; but see also Fig. 9 of Zhang et al., 2013), and with wetter conditions
in the northern tropics (e.g., Dykoski et al., 2005; Fleitmann et al., 2007; Gasse, 2000; Haug et al.,
2001; Weldeab et al., 2007; Zhang et al., 2013).

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### 5.4. Regional comparisons

472 Records from neighboring locations (Figs. 8–9; Table S3) show that the Holocene wet/dry/wet 473 succession reported here for NW Madagascar also affected other locations. For example, 474 hydrogen isotope compositions of the n-C31 alkane in GeoB9307-3 from a 6.51 m long marine 475 sediment core retrieved about 100 km off the Zambezi delta show a similar wet/dry/wet climate 476 during Early, Middle, and Late Holocene respectively (Schefuß et al., 2011). These changes 477 correspond to changes in temperature from ~26.5° to 27.25° to 27°C, respectively, in the 478 Mozambique Channel, as suggested by alkenone SST records from sediment cores MD79257 (Bard 479 et al., 1997; Sonzogni et al., 1998). The Zambezi catchment is specifically relevant here because it 480 is located at the southern boundary of the modern ITCZ, and so has a similar climatic setting as 481 NW Madagascar, and its sensitivity to the latitudinal migration of the ITCZ could parallel that of 482 Madagascar. Likewise, temperature reconstruction from the Mozambique Channel could be used 483 to link regional changes in paleorainfall with regional changes in temperature. A general overview 484 of the Holocene climate in the African neighboring locations to Madagascar suggests a roughly 485 consistent wetter and drier climate during the early and middle Holocene, respectively (Fig. 9, 486 Table S3, also see Gasse, 2000; Singarayer and Burrough, 2015). However, Late Holocene 487 paleoclimate reconstructions vary. A simple explanation to this Late Holocene variability is 488 unlikely, but several interacting factors, including the latitudinal migration of the ITCZ, changes in 489 ocean oscillations and sea surface temperatures, volcanic aerosols, and anthropogenic influences 490 may have played a role (e.g., Nicholson, 1996; Gasse, 2000; Tierney et al., 2008; Truc et al., 2013). 491 Assessing these factors is beyond the scope of this study.

### 493 5.5. The 8.2 ka BP event in Madagascar: linkage to ITCZ and AMOC

494 The 8.2 ka BP event, a widespread cold event in the NH (e.g., Alley et al., 1997), is apparent in 495 the stalagmite records (Figs. 5, 10). Stalagmite ANJB-2  $\delta^{18}$ O and  $\delta^{13}$ C ratios reach their lowest 496 values of -6.8 and -10.9‰, respectively during that interval, and mineralogy is primary calcite. 497 These proxies suggest wet interval in NW Madagascar.

498 The 8.2 ka event was triggered by an abrupt freshwater influx from the melting Laurentide Ice 499 Sheet into the North Atlantic (Alley et al., 1997; Barber et al., 1999; Kleiven et al., 2008; Carlson et 500 al., 2008; Renssen et al, 2010; Wiersma et al., 2011; Wanner et al., 2015). This influx of meltwater 501 altered the density and salinity of the NADW (e.g., Thornalley et al., 2009), weakening the Atlantic 502 Meridional Overturning Circulation (AMOC, e.g., Barber et al., 1999; Clark et al., 2001; Daley et al., 503 2011; Vellinga and Wood 2002; Dong and Sutton 2002, 2007; Dahl et al. 2005; Zhang and Delworth 504 2005; Daley et al., 2011; Renssen et al., 2001). Weakening of the AMOC would cause a widespread 505 cooling in the NH regions (e.g., Clark et al., 2001; Thomas et al., 2007) but warming in the SH 506 regions (Wiersma et al., 2011; Wiersma and Renssen, 2006), creating a "bipolar seesaw" effect 507 (e.g., Crowley, 1992; Broecker, 1998). The interhemispheric temperature difference between the 508 NH and SH from this effect may be the driver of the southward displacement of the mean position 509 of the ITCZ during the 8.2 ka abrupt cooling event. This may have intensified the Malagasy 510 monsoon in NW Madagascar during austral summers, similar to what happened to the South 511 American Summer Monsoon in Brazil (e.g., Cheng et al., 2009). In contrast, regions in the NH 512 monsoon regions became drier at 8.2 ka BP as the Asian Monsoon and the East Asian Monsoon 513 weakened (e.g., Wang et al., 2005; Dykoski et al., 2005; Cheng et al., 2009; Liu et al., 2013). The 514 cold NH climate conditions and the wet climate conditions in NW Madagascar at 8.2 ka BP (Fig. 515 10) could suggest causal relationships. However, further research and data will be needed to 516 confirm this possibility.

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# 518 6. Conclusions

Petrography, mineralogy, and stable isotope records from Stalagmite ANJB-2, from Anjohibe Cave, and Stalagmite MAJ-5, from Anjokipoty Cave, combine to suggest three distinct intervals of changing climate in Madagascar during the Holocene: relatively wet conditions during the MEHI, 522 relatively drier conditions, possibly due to episodic dry intervals, during the MMHI, and relatively 523 wet conditions during the MLHI. The timing of stalagmite deposition during the MEHI and the MLHI 524 in NW Madagascar could be attributed to a more southward migration and/or an expanded ITCZ, 525 increasing the duration of the summer rainy seasons, perhaps linked to a stronger Malagasy 526 monsoon. This could have been tied to the temperature gradient between the two hemispheres 527 and weakening of the AMOC. In contrast, the c. 6500-year depositional hiatus during the MMHI 528 could indicate a northward migration of the ITCZ, leading to relatively drier conditions in NW 529 Madagascar. The evidence of the 8.2 ka event in the Malagasy records may further suggest a close 530 link between paleoenvironmental changes in Madagascar and abrupt climatic events in the NH, 531 suggesting that during the MEHI Madagascar's climate was very sensitive to abrupt oceanatmosphere events in the NH. 532

533 Although the ITCZ is unquestionably one of the climatic drivers influencing climate in 534 Madagascar and surrounding locations, several climatic factors need to be investigated in more 535 detail. For example, we do not fully understand if the latitudinal migration is paired with the 536 expansion and/or contraction of the ITCZ, which would affect the strengths of the associated 537 monsoon systems. In addition, the interplay between ITCZ and other factors involving changes in 538 sea surface temperatures, particularly IOD-ENSO, needs to be investigated in details. Data-model 539 comparison (for example at the 8.2 ka event) and improved spatial and temporal resolution of 540 paleoclimate datasets could be an approach to address this challenge.

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## 542 Author Contribution

543 N.R.G.V. conceived the research and experiments. N.R.G.V, G.K, A.F.M.R, and M.O.M.R did the 544 fieldwork and collected the samples. X.L., G.K., H.C., R.L.E, and N.R.G.V contributed to the <sup>230</sup>Th 545 dating analyses. N.R.G.V provided detailed investigation of the two stalagmites, provided stable 546 isotope measurements, prepared thin sections, and conducted X-ray diffraction analyses. G.K. also 547 assisted with the isotopic measurements on Stalagmite ANJB-2. N.R.G.V. wrote the first draft of 548 the manuscript and led the writing. L.B.R. and G.A.B. provided a thorough review of the draft. 549 N.R.G.V. and L.B.R. discussed and revised the manuscript, with additional comments from L.W. 550 N.R.G.V revised the paper with input from all authors, reviewers, and editors.

# 552 Competing Interests

553 The authors declare no conflict of interest.

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1036 Figure 1: Climatological and geographic setting of Madagascar and the study area. (a) Global 1037 rainfall maps recorded by NASA's Tropical Rainfall Measuring Mission (TRMM) satellite showing 1038 the total monthly rainfall in millimeters and the overall position of the ITCZ during November, 1039 2006. Darker shades of blue indicate regions of higher rainfall (source: NASA Earth Observatory, 1040 2016). (b) Barplots of monthly precipitation, and monthly average of daily maximum, minimum, 1041 and mean temperature in NW Madagascar, based on 1971-2000 climate data. Source: 1042 http://iridl.ldeo.columbia.edu/ (accessed August 31, 2016). (c) Simplified map showing the

southwest part of the Narinda karst and the location of the study areas. Inset figure is a map of
Madagascar showing the extent of the Tertiary limestone outcrop that makes up the Narinda karst.
(d-e) Maps of Anjohibe (ANJB) and Anjokipoty (ANJK) caves (St-Ours, 1959; Middleton and
Middleton, 2002), with approximate location for sample collection (red dots). See Figs. S1–S3 for
additional information about the study locations.

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1050Figure 2: Age model of Stalagmite MAJ-5 (left) and ANJB-2 (right) using the StalAge1.0 algorithm1051of Scholz and Hoffman (2011) and Scholz et al. (2012). Scanned image of the two samples are

- 1052 shown for reference and to indicate the three distinct Holocene intervals.
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1057 Figure 3: a) Scanned image of Stalagmite ANJB-2 and the corresponding variations in layer-specific

1058 width (LSW). b) Scanned image of Stalagmite MAJ-5 and the corresponding layer-specific width

1059 (LSW). c) Sketches of typical layer-bounding surfaces (Type E and Type L) of Railsback et al. (2013).

1060 Close-up photographs of the hiatuses are shown in Fig S6.



Figure 4: **Stable isotope data**. Scatterplots of  $\delta^{13}$ C and  $\delta^{18}$ O for Stalagmite MAJ-5 (green) and ANJB-(red) during the Malagasy early Holocene interval (circle) and the Malagasy late Holocene interval (triangle). The plot shows distinctive early and late Holocene conditions (roughly highlighted in gray and light blue, respectively).

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1072 Figure 5: Variations in  $\delta^{13}$ C,  $\delta^{18}$ O, and mineralogy in Stalagmite ANJB-2 and Stalagmite MAJ-5 1073 during the Malagasy Early Holocene Interval. Supplementary Fig. S6 shows both the corrected and 1074 uncorrected isotope values. 1075



1077 Figure 6: Variations in  $\delta^{13}$ C,  $\delta^{18}$ O, and mineralogy in Stalagmite ANJB-2 and Stalagmite MAJ-5 1078 during the Malagasy Late Holocene Interval. Supplementary Fig. S7 shows both the corrected and 1079 uncorrected isotope values, and Fig. S8 compares the corrected  $\delta^{18}$ O values for both stalagmites. 1080



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1084 Figure 7: Simplified models portraying Holocene climate change in NW Madagascar and the 1085 possible climatic conditions linked to the ITCZ. a) Wetter conditions during the early Holocene with 1086 the ITCZ further south (prior to c 7.8 ka BP), favorable for stalagmite deposition. b) Periodic dry 1087 conditions during the mid-Holocene (between c. 7.8 and 1.6 ka BP) with the ITCZ further north 1088 leading to no stalagmite formation (refer to Sect. 5.2.2). c) Wetter conditions during the late 1089 Holocene (after c. 1.6 ka BP) with the ITCZ further south, favorable for stalagmite deposition. 1090 Drawings are not to scale. The bottom figures are from the same source as Fig. 1a, and they are 1091 only used here to give a perspective of the possible position of the ITCZ during the early, mid, and 1092 late Holocene. Madagascar is indicated with a red ellipse.





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1096 Figure 8: Regional comparison. Google Earth image showing the location of sites reported in Table 1097 S3 and in Figure 9. Most site records are from lake sediments, except for GeoB9307-3 (onshore 1098 off delta sediments), MD79257 (alkenone from marine sediment core), and Cold Air, Anjohibe, 1099 and Anjokipoty caves (stalagmites  $\delta^{18}$ O).



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1102 Figure 9: Regional comparison. a) Lake Challa BTI index (Verschuren et al., 2009). b) Lake 1103 Tanganyika  $C_{28} \delta D$  (Tierney et al., 2008, 2010). c) Lake Masoko low field magnetic susceptibility  $(10^{-6}.m^{3}kg^{-1})$  (Garcin et al., 2006). d) Lake Malawi C<sub>28</sub>  $\delta$ D (Konecky et al., 2011). e) Lake Chilwa OSL 1104 1105 dates of shoreline (Thomas et al., 2009). f) Wonderkrater reconstructed paleoprecipitation, 1106 PWetQ (Precipitation of the Wettest Quarter; Truc et al., 2013). g) Cold Air Cave corrected (corr.) and uncorrected (uncorr.)  $\delta^{18}$ O profiles from Stalagmite T8 (Holmgren et al., 2003). h) Tswaing 1107 Crater paleo-rainfall derived from sediment composition (Partridge et al., 1997). i) Indian Ocean 1108 1109 SST records from alkenone (Bard et al., 1997; Sonzogni et al., 1998). j-k) Zambezi  $\delta D$  n-C<sub>31</sub> alkane  $\delta^{13}$ C n-C<sub>31</sub> alkane (Schefuß et al., 2011). I) Lake Tritrivakely stacked magnetic susceptibility 1110 1111 (Williamson et al., 1998). m) NW Madagascar (Anjohibe and Anjokipoty) interval of deposition of 1112 Stalagmite ANJB-2 and Stalagmite MAJ-5 (this study). The two vertical dashed lines indicate the 1113 boundary of the Early, Middle, and Late Holocene by Walker et al. (2012) and Head and Gibbard 1114 (2015).



1116Figure 10: The 8.2 ka BP event in Madagascar. Oxygen isotope record from Greenland (GRIP and1117NGRIP) ice cores (Vinther et al., 2009) compared with Stalagmite ANJB-2  $\delta^{18}$ O and  $\delta^{13}$ C.