1 Three distinct Holocene intervals of stalagmite deposition and non-

2 deposition revealed in NW Madagascar, and their paleoclimate

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17 ABSTRACT

Petrographic features, mineralogy, and stable isotopes from two stalagmites collected from Anjohibe and Anjokipoty caves allow distinction of three intervals of the Holocene in NW Madagascar. The Malagasy early Holocene (between c. 9.8 and 7.8 ka) and late Holocene (after c. 1.6 ka) intervals (MEHI and MLHI, respectively) record evidence of stalagmite deposition. The Malagasy middle Holocene interval (MMHI, between c. 7.8 ka and 1.6 ka), however, is marked by a depositional hiatus lasting for c. 6500 years.

Deposition of Stalagmites ANJB-2 and MAJ-5 from Anjohibe and Anjokipoty caves, respectively, during the MEHI and the MLHI suggests that these caves were sufficiently supplied with water to allow stalagmite formation. These MEHI and MLHI intervals may have been comparatively wet. In contrast, the long-term depositional hiatus likely suggests that the MMHI was relatively drier than the MEHI and the MLHI. This dry condition could have influenced the amount of water supplied to the cave, and thus prevented formation of the stalagmites.

The alternating "wet/dry/wet" during each of these Holocene intervals could be generally linked to the long-term migration of the Inter-Tropical Convergence Zone (ITCZ). When the ITCZ's mean position is farther south, NW Madagascar experiences wetter conditions, such as during the MEHI and MLHI, and when it moves north, NW Madagascar climate becomes drier, such as during the MMHI. A similar wet/dry/wet succession during the Holocene has been reported inneighboring locations, such as southeastern Africa.

Stable isotope records also suggest that although the MEHI and MLHI were wetter, the stronger correlation between δ^{18} O and δ^{13} C suggest that the early Holocene vegetation closely responded to changes in climate. In contrast, the weaker correlation between δ^{18} O and δ^{13} C and the positive shift in δ^{13} C suggest that the late Holocene vegetation was controlled by something other than climate, and the plausible explanation for such changes is the practice of swidden agriculture, as reported in previous literature.

42 Beyond these three subdivisions, the evidence of the 8.2 ka event in the stalagmite 43 records also suggests that climate in Madagascar was sensitive to abrupt climate changes, such 44 as the abrupt influx of the Laurentide Ice Sheet meltwater to the North Atlantic. The freshwater 45 influx into the N. Atlantic, known to have weakened the Atlantic Meridional Overturning 46 Circulation (AMOC), also led to an enhanced temperature gradient between the two hemispheres, i.e. cold NH and warm SH, shifting the mean position of the ITCZ further south. 47 48 This brought wet conditions in the SH monsoon regions, such as NW Madagascar, and dry 49 conditions in the NH monsoon regions, including the Asian Monsoon and the East Asian Summer 50 Monsoon.

51 **1.** Introduction

52 Although much is known about Holocene climate change worldwide (Mayewski et al., 53 2004; Wanner and Ritz, 2011; Wanner et al., 2011; 2015), high-resolution climate data for the 54 Holocene period is still regionally limited in the Southern Hemisphere (SH) (e.g., Wanner et al., 55 2008; Marcott et al., 2013; Wanner et al., 2015). This uneven distribution of data hinders our understanding of the spatio-temporal characteristics of Holocene climate change, and the 56 57 forcings involved. For example, some of these forcings would have an influence on Inter-Tropical 58 Convergence Zone (ITCZ) behavior and monsoonal response in low- to mid-latitude regions (e.g., 59 Wanner et al., 2015; Talento and Barreiro, 2016). The island of Madagascar, in the southwest 60 Indian Ocean (Fig. 1a), is seasonally visited by the ITCZ with a karst region crossing latitudinal 61 belts (Fig. 1c). Thus, it is a natural laboratory to study changes in the ITCZ over time. New records 62 from Madagascar could fill gaps in paleoclimate reconstruction in the SH that might help refine paleoclimate simulations, which in turn could provide better understanding of the global
circulation and the land–atmosphere–ocean interaction during the Holocene.

In this paper, we present multiproxy records (stable isotopes, petrography, mineralogy,
variability of layer-specific width, or LSW) from stalagmites from Anjohibe and Anjokipoty caves.
Stalagmites are used because of their potential to store significant climatic information (e.g.,
Fairchild and Baker, 2012, p. 9–10), and in Anjohibe Cave, recent studies have shown the
replicability of paleoclimate records from stalagmites (e.g., Burns et al., 2016).

Two stalagmites were investigated, and these allowed us to characterize Holocene climate change in NW Madagascar. First, we developed a record of climate change from the multiproxy data. With a better understanding of Madagascar's paleoclimate, we then investigated possible climate drivers of tropical climate change to draw a more comprehensive conclusion on the major factors controlling the hydrological cycle in NW Madagascar and surrounding regions during the Holocene.

76 **2.** Setting

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77 *2.1.*Stalagmites and their setting

Stalagmites are secondary cave deposits that are $CaCO_3$ precipitates from cave dripwater. Calcium carbonate precipitation occurs mainly by CO_2 degassing, which increases the pH of the dripwater and thus increases the concentration of $CO_3^{2^-}$. In some cases, evaporation may also contribute to increased Ca^{2+} and/or $CO_3^{2^-}$ concentration in dripwater. CO_2 degassing occurs when high-*P*CO₂ water from the epikarst encounters low-*P*CO₂ cave air. Evaporation occurs when humidity inside the cave is relatively low. The fundamental equation for stalagmite deposition is shown in Eq. 1.

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$$Ca_{(aq)}^{2+} + 2HCO_{3(aq)}^{-} \rightleftharpoons CaCO_{3(s)} + CO_{2(g)} + H_2O_{(l)}$$
 (Eq. 1)

Growth and non-growth of stalagmites depends on conditions that affect the reaction of Eq. 1 above. An increase in Ca^{2+} drives the equation to the right (towards precipitation) and an increase in CO_2 of the cave air and/or H_2O drives it to the left (towards dissolution). All components of the equation are influenced by the supply of water to the cave, which is generally climate-dependent. More water enters the cave during warm/rainy seasons than during cold/dry 91 seasons. Stalagmites will form when cave dripwater is saturated with respect to calcite and/or 92 aragonite. If the water passes through the bedrock too quickly to dissolve significant carbonate 93 rock, and/or enters the cave and reaches the stalagmite too quickly to degas significant CO_2 , it 94 will not be saturated with respect to CaCO₃, inhibiting stalagmite formation. Stalagmite growth will slow as dripwater declines and will stop entirely if flow ceases. Vegetation provides CO₂ to 95 96 the soil via root respiration so the vegetation cover above the cave and the type of vegetation 97 can promote or limit stalagmite growth. Overall, the karst hydrological system plays a crucial role 98 in the deposition and non-deposition of stalagmites, and this is closely linked to changes in local 99 and regional environment and climate.

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101 *2.2.*Regional environmental setting

Stalagmites ANJB-2 and MAJ-5 were collected from Anjohibe and Anjokipoty caves, respectively, in the Majunga region of NW Madagascar (Fig. 1). Sediments and fossils from these caves have already provided many insights about the paleoenvironmental and archaeological history of NW Madagascar (e.g., Burney et al., 1997, 2004; Brook et al., 1999; Gommery et al., 2011; Jungers et al., 2008; Vasey et al., 2013; Burns et al., 2016; Voarintsoa et al., 2017b).

107 Anjohibe (S15° 32' 33.3"; E046° 53' 07.4") and Anjokipoty (S15° 34' 42.2"; E046° 44' 108 03.7") are about 16.5 km apart (Fig. 1c). Their location in the zone visited by the ITCZ (e.g., 109 Nassor and Jury, 1998) makes them ideal sites to test the hypothesis that latitudinal migration of 110 the ITCZ influenced the Holocene climate of NW Madagascar (e.g., Chiang and Bitz, 2005; 111 Broccoli et al., 2006; Chiang and Friedman, 2012; Schneider et al., 2014). The ITCZ brings north 112 or northwesterly monsoon winds to Madagascar during austral summers, in a pattern that the 113 Service Météorologique of Madagascar calls the "Malagasy monsoon". Majunga has a tropical 114 savanna climate (Aw) according to the Köppen-Geiger climate classification, with a distinct wet 115 summer (from October to April) and dry winter (May-September). The mean annual rainfall is 116 around 1160 mm. The mean maximum temperature in November, the hottest month in the 117 summer, is about 32°C. The mean minimum temperature in July, the coldest month of the dry 118 winter, is about 18°C (Fig. 1b).

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120 2.3.Climate of Madagascar

121 The climate of Madagascar is unique because of its varied topography and its position in the 122 Indian Ocean. Some scientists refer Madagascar as a "laboratory" for paleoecological study (e.g., 123 Burney, 1997) because it is not only susceptible to several climatic forcing mechanisms but also 124 an island with recent anthropogenic interaction, living imprints in the geological records (e.g., 125 Burney et al., 2003, 2004; Matsumoto and Burney, 1994; Crowley and Samonds, 2013; Burns et 126 al., 2016; Voarintsoa et al., 2017b). Its climate has been reviewed in several recent works (e.g., 127 Jury, 2003; DGM, 2008, Douglas and Zinke, 2015, p. 281-299; Voarintsoa et al., 2017b, p.138-128 139; Scroxton et al., 2017). Regionally distinct rainfall gradients from east to west and from 129 north to south are evident across the country (Jury, 2003; Dewar and Richard, 2007), and these 130 are linked to easterly trade-winds in winter (May-October) and northwesterly tropical storms in 131 summer, respectively. The Malagasy monsoon is modulated by the seasonal north-south 132 migration of the ITCZ, which is the main driver of austral summer rainfall in Madagascar. The 133 ITCZ's mean position has shifted northward or southward depending on the global climate 134 conditions, but most generally it migrates towards the Earth's warmer hemisphere (Frierson and 135 Hwang, 2012; Kang et al., 2008; McGee et al., 2014; Sachs et al., 2009). A relationship between 136 this long-term migration of the ITCZ and climate in Madagascar was reported in NW Madagascar 137 between c. 370 CE and 800 CE (see Fig, 8 of Voarintsoa et al., 2017b).

Beyond ITCZ, climate of Madagascar is also influenced by changes in Indian Ocean sea surface 138 139 temperatures (SST) (Zinke et al., 2004; see also Kunhert et al., 2014) and changes in SST of the 140 adjacent current off southwestern Madagascar, the Aghulas Current (Lutjeharms, 2006; Beal et 141 al., 2011; Zinke et al., 2014). The most immediate signal is the Indian Ocean Dipole (IOD), or 142 Indian Ocean Zonal Mode (Li et al., 2003). IOD-like patterns have been proposed as possible 143 contributors to Holocene climate variability in tropical Indian Ocean (Abram et al., 2009; Tierney 144 et al., 2013). IOD is as a coupled atmosphere-ocean mode in the tropical Indian Ocean (e.g., Saji 145 et al., 1999; Webster et al., 1999; Brown et al., 2009; Yagamata et al., 2004; Behera et al., 2013). 146 It is characterized by a reversal of the climatological SST gradient and winds across the Indian 147 Ocean basin (Saji et al., 1999; Webster et al., 1999; Abram et al., 2007; Brown et al., 2009). A 148 positive IOD event starts with anomalous SST cooling along the Sumatra-Java coast in the eastern

149 Indian Ocean (Abram et al., 2007, 2008), along with positive SST anomaly in the western part of 150 the basin (e.g., Saji et al., 1999; Abram et al., 2007). Such positive IOD events are observed to 151 result in increased precipitation, sometimes causing devastating floods, over East Africa (Black et 152 al., 2003; Saji et al., 1999; Webster et al., 1999; Saji and Yagamata, 2003; Weller and Cai, 2014). 153 Such events have also enhanced precipitation over the northern part of India, the Bay of Bengal, 154 Indochina, and southern part of China in 1994 (e.g., Behera et al., 1999; Guan and Yamagata, 155 2003; Saji and Yagamata, 2003). In the eastern Indian Ocean, a positive IOD is found to intensify 156 El-Niño related drought, often as severe droughts, over Indonesia (Webster et al., 1999; Weller 157 and Cai, 2014). It is however, important to note that the relationship between IOD and El-Nino 158 Southern Oscillation (ENSO) is still debated. While some researchers found no relationships (e.g., 159 Saji et al., 1999; Li et al., 2003; Lee et al., 2008), others found some relationships (e.g., Brown et 160 al., 2009; Schott et al., 2009; Shinoda et al., 2004; Venzke et al., 2000; Abram et al., 2008; Saji 161 and Yagamata, 2003; Meyers et al., 2007).

Apart from the coral study of Zinke et al. (2004) and the stalagmite study of Scroxton et al. (2017), very little is known about the effect of the IOD on Madagascar. One objective of this stalagmite study is to better understand how such mechanisms influenced climate in Madagascar during the Holocene.

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167 *2.4.*The Holocene in NW Madagascar

168 Little is hitherto known about Holocene climate change in NW Madagascar nor about the 169 major drivers of long-term climatic changes there. Most paleoclimate information from this 170 region covers the last two millennia with more focus on the anthropogenic effects on the 171 Malagasy ecosystems (e.g., Crowley and Samonds, 2013; Burns et al., 2016; Voarintsoa et al., 172 2017b). This is because several studies show that megafaunal extinctions in Madagascar 173 coincide with the arrival of humans around 2-3 ka BP (e.g., see Table 1 of Virah-Sawmy et al., 174 2010; MacPhee and Burney, 1991; Burney et al., 1997; Crowley, 2010). There are even fewer 175 long-term paleoclimate records for the NW region, with only sediments from Lake Mitsinjo 176 (3,500 yr. BP; Matsumoto and Burney, 1994) and stalagmites from Anjohibe Cave (40,000 yr. BP; 177 Burney et al. 1997) providing records of more than 3 kyr. Even though these records provided

useful information about the paleoenvironmental changes in NW Madagascar, their linkages toglobal climatic change, such as the linkages to the ITCZ, are not yet fully understood.

180 **3.** Methods

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3.1. Radiometric dating

182 A total of 22 samples were drilled from Stalagmite ANJB-2 and 9 samples for Stalagmite 183 MAJ-5 for U-series dating (Table S1 and S2). Each sample is a long (~5 to 20 mm), narrow (~1-184 2mm), and shallow (~1 mm) trench, allowing us to extract 50–250 mg of CaCO₃ powder. We 185 followed the chemical procedures described in Edwards et al. (1987) and Shen et al. (2002) when 186 separating uranium and thorium. U/Th measurements were performed on the multi-collector 187 ICP-MS of the University of Minnesota, USA and on a similar instrument in the Stable Isotopes 188 Laboratory of Xi'an, in Jiaotong, China. Instrument details are provided in Cheng et al. (2013). Corrected ²³⁰Th ages assume an initial ²³⁰Th/²³²Th atomic ratio of $4.4 \pm 2.2 \times 10^{-6}$. This is the ratio 189 for "bulk earth" or crustal material at secular equilibrium with a 232 Th/ 238 U value of 3.8. The 190 191 uncertainty in the "bulk earth" value is assumed to be ±50% (see footnotes to Table S1 and S2). 192 The error in the final "corrected age" incorporates this uncertainty. The radiometric data are 193 reported as year BP, where BP is Before Present, and "Present" is A.D. 1950. Stalagmite 194 chronologies were constructed using the StalAge1.0 algorithm of Scholz and Hoffman (2011) and 195 Scholz et al. (2012), an algorithm using a Monte-Carlo simulation designed to construct 196 speleothem age models. The algorithm can identify major and minor outliers and age inversions. 197 The StalAge scripts were run on the statistics program R version 3.2.2 (2015-08-14). The age 198 models were adjusted considering hiatal surfaces identified in the samples, using the approach 199 of Railsback et al. (2013; see their Fig. 9).

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3.2. Petrography and mineralogy

Petrography and mineralogy of the two stalagmites were investigated 1) by examining both the polished surfaces and the scanned images of the sectioned stalagmites, and by identifying any diagenetic fabrics (e.g., Zhang et al., 2014) that could potentially affect stable isotope values, 2) by observing eleven oversized thin sections (3x2 in) under the Leitz Laborlux 12 Pol microscope and the Leica DMLP equipped with QCapture in the Sedimentary 207 Geochemistry Lab at the University of Georgia, 3) by using scanning electron microscopy (SEM) 208 to better understand the mineralogical fabrics at locations of interest (Fig. S13), and 4) by 209 analyzing about 30–100 mg of powdered spelean layers (n=15) on a Bruker D8 X-ray 210 Diffractometer in the Department of Geology, University of Georgia. For calcite and aragonite 211 identification, we used CoK α radiation at a 2 θ angle between 20° and 60°.

212 Layer-specific width (LSW) of clearly-defined layers was measured at selected locations 213 on the stalagmite polished surfaces (Fig. S4; Sletten et al., 2013; Railsback et al., 2014; 214 Voarintsoa et al., 2017b). LSW is the horizontal distance between two points on the flanks of the 215 stalagmite where convexity is greatest. It is the width near the top of the stalagmite when the 216 layer being examined was deposited. LSW is measured at right angles to the growth axis of the 217 stalagmite; it is the horizontal distance between points on the layer growth surface becomes 218 tangent to a line inclined at 35° to the growth axis (Fig. S4). LSW may vary along the length of the 219 stalagmite, with smaller values suggesting drier conditions and larger values wetter conditions.

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221 3.3. Stable isotopes

222 Stable isotope samples of $50-100 \ \mu g$ were manually drilled along the stalagmite's growth 223 layers at the crest. The trench size is very small (1.5 x 0.5 x 0.5 mm). Since a small mixture of calcite and aragonite could potentially change the δ^{18} O and δ^{13} C of the measured spelean layers 224 225 (see for example Frisia et al., 2002), drilling and sample extraction was carefully done on 226 individually discrete layers using the smallest drill-bit head (SSW-HP-1/4) to avoid potential 227 mixing between calcite and aragonite. The polished surface of the two stalagmites were 228 examined to see if features of diagenetic alteration are present (see for example fig. 2 of Zhang 229 et al., 2014), but none was found. During sampling, the mineralogy at the crest, where stable 230 isotope samples were extracted, was recorded for future mineralogical correction.

Aragonite oxygen and carbon isotopic corrections were performed to compensate for aragonite's inherent fractionation of heavier isotopes (e.g., Romanek et al., 1992; Kim et al., 2007; McMillan et al., 2005) and to remove the mineralogical bias in isotopic interpretation between calcite and aragonite. The correction consists of subtracting 0.8‰ for δ^{18} O (Kim and O'Neil, 1997; Tarutani et al., 1969; Kim et al., 2007; Zhang et al., 2014) and 1.7 ‰ for δ^{13} C (Rubinson and Clayton, 1969; Romanek et al., 1992) for the aragonite as has been done
previously (e.g., Holmgren et al., 2003; Sletten et al., 2013; Liang et al., 2015; Railsback et al.,
2016; Voarintsoa et al., 2017a) as shown in equations 2 and 3 below (where R_{A/C} is the aragonite
percentage if not 100%).

Supplementary Figures S6–S8 show both the corrected and uncorrected isotopic records.

240 $\delta^{18}O_{\text{corr.}}$ (‰, VPDB) = $\delta^{18}O_{\text{uncorr.}}$ (‰, VPDB) – [R_{A/C} x 0.8 (‰, VPDB)] (Eq. 2)

241 $\delta^{13}C_{corr.}$ (‰, VPDB) = $\delta^{13}C_{uncorr.}$ (‰, VPDB) – [R_{A/C} x 1.7 (‰, VPDB)] (Eq. 3)

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243 For the analytical methods, oxygen and carbon isotope ratios were measured using the 244 Finnigan MAT-253 mass spectrometer fitted with the Kiel IV Carbonate Device of the Xi'an Stable 245 Isotope Laboratory in China (ANJB-2; n=654) and using the Delta V Plus at 50°C fitted with the 246 GasBench-IRMS machine of the Alabama Stable Isotope Laboratory in USA (MAJ-5; n=286). 247 Analytical procedures using the MAT 253 are identical to those described in Dykoski et al. (2005), with isotopic measurement errors of less than 0.1 ‰ for both δ^{13} C and δ^{18} O. Analytical methods 248 249 and procedures using the GasBench-IRMS machine are identical to those described in Skrzypek 250 and Paul (2006), Paul and Skrzypek (2007), and Lambert and Aharon (2011), with ±0.1 ‰ errors for both δ^{13} C and δ^{18} O. In both techniques, the results are reported relative to Vienna PeeDee 251 252 Belemnite (VPDB) and with standardization relative to NBS19. An inter-lab comparison of the 253 isotopic results was conducted, and it involved replicating every tenth sample of Stalagmite MAJ-254 5 at both labs. This exercise showed a strong correlation between the lab results (Fig. S5).

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4. Results

257 *4.1*.Radiometric data

Results from radiometric analyses of the two stalagmites are presented in Tables S1 and S2. Corrected ²³⁰Th ages suggest that Stalagmite ANJB-2 was deposited between c. 8977±50 and c. 161±64 yr. BP, and Stalagmite MAJ-5 was deposited between c. 9796±64 and c. 150±24 yr. BP. These ages collectively indicate stalagmite deposition at the beginning (between 9.8 and 7.8 ka BP) and at the end of the Holocene (after c. 1.6 ka BP). In both stalagmites, the older ages have small 2 σ errors and they generally fall in correct stratigraphic order, except sample ANJB-2-120 and its replicate ANJB-2-120R, which were not used because of the sample's high porosity and

265 high detritals content. In contrast, many of the younger ages have larger uncertainties. This is 266 mainly because many of the younger samples have very low uranium concentration and the 267 detrital thorium concentration is also high, similar to what Dorale et al. (2004) reported. We also understand that the value for initial 230 Th correction, i.e. the initial 230 Th/ 232 Th atomic ratio of 4.4 268 $\pm 2.2 \times 10^{-6}$ for a bulk earth with a ²³²Th/²³⁸U value of 3.8, in these samples could have slightly 269 altered the ²³⁰Th age of these younger samples, leading to larger uncertainties (such as discussed 270 271 in Lachniet et al., 2012). We encountered similar problems while working on other younger 272 samples from the same cave, but we compared the stable isotope profile with other published 273 records using isochron corrections, and results did not differ significantly (see Fig. 9 of 274 Voarintsoa et al., 2017b). Since this work does not focus on decadal or centennial interpretation 275 of the Late Holocene stable isotope data, additional chronology adjustment has not been made, 276 and we used the chronology from StalAge to construct the time series. However, in Figures 5 and 277 6, age uncertainties are given below the stable isotope profiles so that comparisons with other 278 records can accommodate these uncertainties.

- The key finding from our age and petrographic data for the two stalagmites is that they suggest that there were three distinct intervals of growth and non-growth during the Holocene (Figs. 2–4, 7). The information suggesting this includes: (1) CaCO₃ deposition between c. 9.8 and 7.8 ka B.P., (2) a long depositional hiatus between c. 7.8 and 1.6 ka B.P., and (3) resumption of CaCO₃ deposition after c. 1.6 ka B.P. In the rest of the paper, we will refer to these intervals as the Malagasy Early Holocene Interval (MEHI), Malagasy Mid-Holocene Interval (MMHI), and Malagasy Late Holocene Interval (MLHI), respectively.
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287 *4.2.*Stable isotopes

Raw values of δ^{18} O and δ^{13} C for Stalagmite ANJB-2 range from -8.9 to -2.3‰ (mean = -289 5.0‰), and from -11.0 to +5.2‰ (mean = -4.2‰), respectively, relative to VPDB. Raw values of 290 δ^{18} O and δ^{13} C for Stalagmite MAJ-5 range from -8.8 to -0.9‰ (mean = -4.9‰), and from -9.4 to 291 +2.6‰ (mean = -4.4‰), respectively, relative to VPDB. Mean δ^{18} O and δ^{13} C values are 292 distinguishable between the MEHI and the MLHI. In both stalagmites, the amplitude of δ^{18} O fluctuations was fairly constant throughout the Holocene; whereas the δ^{13} C profile shows a dramatic shift toward higher values (i.e. from -10.9‰ to +3.8‰, VPDB) at c. 1.5 ka BP.

The MEHI and MLHI are isotopically distinct (Fig. 4). The MEHI is characterized by statistically correlated δ^{18} O and δ^{13} C (r²=0.65 and 0.53), and much depleted δ^{13} C values (c -11.0 to -4.0 ‰). The 8.2 ka event, a widespread cold event in the NH (e.g., Alley et al., 1997), is also apparent in the stalagmite records. Stalagmite δ^{18} O and δ^{13} C ratios reach their lowest values of -6.8 and -10.9‰, respectively during that interval (Figs. 5, 12). In contrast to the MEHI, the values of δ^{18} O and δ^{13} C during the MLHI are poorly correlated (r²=0.25 and 0.17), and δ^{13} C values are more enriched (Figs. 4, 6).

Since Stalagmites ANJB-2 and MAJ-5 were collected from two different caves 16 km apart, 302 303 discrepancies between the stable isotopes at the same age are expected, suggesting that local 304 conditions could be one of the discrepancy factors. Another potential source for the discrepancy 305 is the larger uncertainty of the younger ages due to low uranium and high detrital 306 concentrations. This U-Th aspect has been a challenge for several young stalagmites (e.g., Dorale 307 et al., 2004; Lachniet et al., 2012) including samples from NW Madagascar (this study). While the 308 utility of speleothems as a climate proxy largely depends on replication of stable isotope values, 309 it is important to note that perfect stable isotope replication can only occur between stalagmites 310 collected from the same cave chamber (e.g., Dong et al., 2010; Burns et al., 2016).

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2 *4.3.* Mineralogy, petrography, and layer-specific width

In both stalagmites, the hiatus of deposition is characterized by a well-developed Type L
surface (Figs. 2, 3, S15). Petrography and mineralogy are distinct before and after this hiatus (Fig.
Below the hiatus, laminations are well preserved in both stalagmites. Above the hiatus,
laminations are not well-preserved, although noted in some intervals.

In Stalagmite ANJB-2, the layer-specific width varies from 37 to 26.5 mm with a mean of 30 mm. It decreases to 28 mm at the hiatus (Fig. 3). Below the hiatus, mineralogy is dominated by aragonite, although a few thick layers of calcite are also identified. A thin (~2-3 mm) but remarkable layer of white, very soft, and porous aragonite is identified just below the hiatus (Fig. 319 S15). This layer is covered by a very thin layer of dirty carbonate. Above the hiatus, mineralogy is also composed of calcite and aragonite, with calcite dominant, and the calcite layers contain
 macro-cavities that are mostly off-axis macroholes (Shtober-Zisu et al., 2012).

In Stalagmite MAJ-5, LSW varies from 50 to 22 mm with a mean of 35.5 mm. It decreases to 22 mm at the hiatus (Fig. 3). Below the hiatus, mineralogy is a mixture of calcite and aragonite. Above the hiatus, mineralogy is mainly calcite and macro-cavities are also present throughout that upper part of the stalagmite.

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329 *4.4.*Summary of results

330 The various records from Stalagmites ANJB-2 and MAJ-5 suggest three distinct 331 climate/hydrological intervals of the Holocene. The MEHI (c. 9.8 to 7.8 ka BP), with evidence of stalagmite deposition, is characterized by statistically correlated δ^{18} O and δ^{13} C (r²=0.65 and 0.53) 332 and more negative δ^{13} C values (c. -11.0 to -4.0 ‰). The MMHI (c. 7.8 to 1.6 ka BP) is marked by 333 334 a long-term hiatus in deposition, which is preceded by a well developed Type L surface in both 335 Stalagmite ANJB-2 and MAJ-5 (Figs. 3, S15). The Type L surface is observed as an upward 336 narrowing of the stalagmite's width and layer thickness. It is particularly well developed in 337 Stalagmite MAJ-5 (Fig. S15). In Stalagmite ANJB-2, the hiatus at the Type L surface is preceded by 338 a c. 3 mm thick layer of highly porous, very soft, and fibrous white crystals of aragonite (the only 339 aragonite with such properties). This aragonite is topped by a thin and well-defined layer of 340 detrital materials (Fig. S15), further supporting the presence of a hiatus. Finally, the MLHI (after c. 1.6 ka BP) is characterized by poorly correlated δ^{18} O and δ^{13} C (r²=0.25–0.17). This interval is 341 additionally marked by a shift in δ^{13} C toward higher values (Figs. 4, 6). 342

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344 5. Discussion

345 *5.1*.Paleoclimate significance of stalagmite growth and non-growth: implications for

346 paleohydrology

347 Growth and non-growth of stalagmites depends on several factors linked to water 348 availability, which is largely determined by climate (more water during warm/rainy seasons and 349 less water during cold/dry seasons). Water is the main dissolution and transporting agent for 350 most chemicals in speleothems. Cave hydrology varies significantly over time in response to 351 climate, and this variability influences the formation or dissolution of CaCO₃. In this regard, 352 calcium carbonate does not form if there is little or no water entering the cave, or if there is too 353 much (see Sect. 2.1). Absence of groundwater recharge most typically occurs during extremely 354 dry conditions, whereas excessive water input to the cave occurs during extremely wet 355 conditions. In the latter scenario, water is undersaturated and flow rates are too fast to allow 356 degassing. Often, water availability is reflected in the extent of vegetation above and around the 357 cave, as plants require soil moisture or shallow groundwater to survive and propagate, and this 358 contributes to the stalagmites' processes of formation. The linkage of stalagmites' growth and 359 non-growth to cave dripwater and soil CO_2 is broadly influenced by changes in climate.

360 Major hiatuses in stalagmite deposition could be marked by a variety of features, 361 including the presence of erosional surfaces, chalkification, dirt bands/detrital layers, offsetting 362 of the growth axis, and/or sometimes by color changes (e.g., Holmgren et al., 1995; Dutton et al., 363 2009; Railsback et al., 2013; Railsback et al., 2015; Voarintsoa et al., 2017a). Railsback et al. 364 (2013) were specifically able to identify significant features in stalagmites that allow distinction between non-deposition during extremely wet (Type E surfaces) and non-deposition during 365 366 extremely dry conditions (Type L surfaces; Fig. 3). Physical properties of stalagmites that are 367 evidence of extreme dry and wet events are summarized in Table 1 of Railsback et al. (2013) and 368 the mechanism is explained in their Figure 5.

369 Type E surfaces are layer-bounding surfaces between two spelean layers when the 370 underlying layers show evidence of truncation. The truncation results from dissolution or erosion 371 (thus the name "E") of previously-formed layers of stalagmites by abundant undersaturated 372 water. Type E surfaces are commonly capped with a layer of calcite (Railsback et al., 2013). This 373 mineralogical trend is not surprising as calcite commonly forms under wetter conditions (e.g., 374 Murray, 1954; Pobeguin, 1965; Siegel, 1965; Thrailkill, 1971; Cabrol and Coudray, 1982; Railsback 375 et al. 1994; Frisia et al., 2002). Additionally, non-carbonate detrital materials are commonly 376 abundant with varying grain size (i.e., from silt- to sand-size; Railsback et al., 2013).

Type L surfaces, on the other hand, are layer-bounding surfaces where the layers became narrower upward and thinner towards the flanks of the stalagmite. Decreases in layer thickness and stalagmites width of the stalagmites upward are indications of lessening deposition (thus the name "L"; Railsback et al., 2013). Aragonite is a very common mineralogy below a Type L surface,
especially in warmer settings. Layers of aragonite commonly form under drier conditions
(Murray, 1954; Pobeguin, 1965; Siegel, 1965; Thrailkill, 1971; Cabrol and Coudray, 1982;
Railsback et al., 1994; Frisia et al., 2002). Non-carbonate detrital materials are scarce, and if
present, they tend to form a very thin horizon of very fine dust material (Railsback et al., 2013).
Identification of Type L surfaces is aided by measuring the LSW (e.g., Sletten et al., 2013;
Railsback et al., 2014), an approach that is also performed in this study (Fig. S4).

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*5.2.*Holocene climate in NW Madagascar

389 Although the specific boundaries between the Early, Mid, and Late Holocene have been 390 proposed for global application (Walker et al., 2012; Head and Gibbard, 2015), their use is still 391 spatially limited (e.g., Wanner et al., 2015). The age models and petrographic features of 392 Stalagmites ANJB-2 and MAJ-5 suggest three distinct but different Holocene climate intervals 393 (MEHI, MMHI, and MLHI; see Sect. 4.1) in NW Madagascar. These intervals are illustrated in the 394 sketches of Figure 4. In this paper, these Malagasy intervals are intended not to argue against 395 the previously proposed intervals of the Holocene (Walker et al., 2012; Head and Gibbard, 2015). 396 Instead, they are presented to aid discussion of the available records. For comparison, the 397 intervals are shown in Fig. 7d.

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5.2.1. Malagasy early Holocene interval (c. 9.8 – 7.8 ka BP)

Stalagmite deposition during the early Holocene suggests that the chambers, where stalagmites ANJB-2 and MAJ-5 were collected, were sufficiently supplied with water to allow CaCO₃ precipitation, in accord with Eq.1. This in turn implies relatively wet conditions that could indicate longer summer rainy seasons relative to modern climate, or wet years in NW Madagascar (see Supplementary Text 4 and Fig. 8). The correlative δ^{13} C and δ^{18} O values further suggest that vegetation consistently responded to changes in moisture availability, which in turn was dependent on climate.

407 One striking aspect of the Stalagmite ANJB-2 δ^{18} O and δ^{13} C records is that they parallel 408 the δ^{18} O of the Greenland ice core records at c 8.2 ka BP (Figs. 5 and 12). An X-ray diffraction

409 spectrum for this period, at 195-202 mm from the top of the stalagmite, suggests that the 410 mineralogy at 8.2 ka BP is 100% calcite (Figs. S14, S16–S17). This calcite is not a diagenetic 411 product of aragonite for three reasons. First, the laminations in the thick layer of calcite were not 412 altered (Figs. S16–S17). Second, the polished surface of the stalagmite shows no evidence of 413 fiber relicts and textural ghosts such as observed in Juxtlahuaca Cave in southwestern Mexico (Lachniet et al., 2012) and in Shennong Cave in southeastern China (Zhang et al., 2014). Third, 414 415 petrographic comparison with known examples of primary and secondary calcite observation under microscope (e.g., Railsback, 2000; Perrin et al., 2014) suggests that there is no strong 416 evidence of aragonite-to-calcite transformation. The decrease in δ^{18} O and δ^{13} C values and the 417 418 presence of calcite mineralogy at the same interval combine to suggest a wet 8.2 ka BP event in 419 NW Madagascar. The 8.2 ka BP event is a prominent cold event in the North Atlantic records and 420 many NH terrestrial records. It may have been triggered by a release of freshwater from the 421 melting Laurentide Ice Sheet into the North Atlantic basin (e.g., Alley et al., 1997; Barber et al., 422 1999). Freshwater influx to the Atlantic could have altered the Atlantic Meridional Overturning 423 Circulation (AMOC, e.g., Clark et al., 2001), and could eventually have influenced the climate of 424 Madagascar (Sect. 5.5). Our records reveal a strong link between paleoenvironmental changes in 425 Madagascar and abrupt climatic events in the NH records, suggesting causal relationships.

The MEHI terminated when conditions became much drier, as suggested by increasing δ^{18} O 426 and δ^{13} C values in Stalagmite ANJB-2, by decreasing LSW of both stalagmites, and by major Type 427 428 L surfaces in both stalagmites. The thin (c. 3 mm), porous, and white aragonite layer in 429 Stalagmite ANJB-2, a very similar deposit to that described in Niggemann et al. (2003), suggests 430 that the terminal drought was at times severe. Aragonite is a CaCO₃ polymorph that forms 431 preferentially under drier conditions (Murray, 1954; Pobeguin, 1965; Siegel, 1965; Thrailkill, 432 1971; Cabrol and Coudray, 1982; Railsback et al. 1994; Frisia et al., 2002). The porous aragonite 433 layer in Stalagmite ANJB-2 is capped by a very thin layer of non-carbonate, brown detritus, which 434 may have been transported to the stalagmite as an aerosol and accumulated on the dry 435 stalagmite surface over time. Accumulation of the detritus must take place in the absence of 436 dripwater (e.g., Railsback et al., 2013). A shift to drier conditions is also supported by isotopic 437 data from Stalagmite ANJ94-5 from Anjohibe Cave (Wang and Brook, 2013; Wang, 2016) in

438 which relatively low δ^{13} C and δ^{18} O values prior to 7600 BP give way to episodically greater values 439 thereafter.

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5.2.2. Malagasy mid-Holocene interval (c. 7.8–1.6 ka BP)

The only data we have for the MMHI is the long term (~6.5 ka) depositional hiatus in both stalagmites (Figs. 2–3), that potentially indicate dry conditions. The question is why did neither stalagmite grow during the MMHI? Here, we try to explain the factors and the climatic conditions that may have been responsible for it.

446 The documented severe dry conditions at the end of the MEHI (see Sect. 5.2.1) could 447 have had a significant influence (1) on the cave hydrological system (e.g., Fig. 5 of Asrat et al., 448 2007; Bosak, 2010), such as the water conduits (primary or secondary porosity) to the chambers, 449 and (2) on the vegetation cover above the caves, particularly above the chambers where 450 Stalagmites ANJB-2 and MAJ-5 were collected. On one hand, it is possible that the dry conditions 451 late in the MEHI could not only bring lesser water recharge to the cave, but also lowered the 452 hydraulic head, and increased the rate of evapo-transpiration in the vadose zone. This condition 453 possibly allowed more air to penetrate the aquifer, perhaps enhancing prior carbonate 454 precipitation (PCP) in pores and conduits above the caves (e.g., Fairchild and McMillan, 2007; 455 Fairchild et al., 2000; Johnson et al., 2006; Karmann et al., 2007; McDonald et al., 2007). This 456 process must have blocked water moving towards Stalagmites ANJB-2 and MAJ-5. On the other 457 hand, the late MEHI drying trend (Sect. 5.2.1) could have challenged vegetation to grow, and we 458 assume that some areas above Anjohibe and Anjokipoty caves must have been devoid of 459 vegetation. Consequently, biomass activities could have been reduced. Because vegetation 460 contributes CO_2 to the carbonic acid dissolving $CaCO_3$, its absence in certain areas above the 461 cave could decrease the pH of the percolating water, and perhaps dissolution did not occur. 462 Under these conditions, even if water reached the stalagmites, it may not have precipitated 463 carbonate.

464 Whatever factors were responsible for the long term-depositional hiatus in Stalagmite 465 ANJB-2 and MAJ-5, we believe that the hiatus was caused by disturbances to water catchments 466 that feed the chambers at Anjohibe and Anjokipoty caves. The disturbances could be inherited 467 from the very dry conditions at the end of the MEHI, and/or due to the lack of water supply,
468 perhaps associated with an increase in epikarst ventilation, and/or by the absence of vegetation.
469 Water and vegetation are two components of the karst system that play an important role in
470 CaCO₃ dissolution and precipitation (see Eq. 1). Their disturbance may have limited limestone
471 dissolution in the epikarst and then carbonate precipitation in the cave zone.

Other evidence supports the idea of at least episodic dryness during the MMHI. A work
on a 2-meter long stalagmite (ANJ94-5) from Anjohibe Cave suggests episodic dryness during the
MMHI and a depositional hiatus around the time when Stalagmites ANJB-2 and MAJ-5 stopped
growing (Wang and Brook, 2013; Wang, 2016). For regional comparison, dry spells were also felt
in Central and Southeastern Madagascar (e.g., Gasse and Van Campo, 1998; Virah-Sawmy et al.,
2009).

In summary, several lines of evidence suggest relatively drier climate in NW Madagascar
during the MMHI compared to the MEHI. Drier intervals generally imply drier summer seasons
with less rainfall (Fig. 8), perhaps reflecting shorter visits by the ITCZ. In this regard, even though
the region received rainfall, the necessary conditions could not have been attained to activate
the growth of Stalagmites ANJB-2 and MAJ-5, thus the hiatuses.

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484 *5.2.3.* Malagasy Late Holocene Interval (c. 1.6 ka–present)

485 Resumption of stalagmite deposition after c. 1.6 ka BP suggests a wetter climate in NW 486 Madagascar with reactivation of the previous epikarst hydrologic system. Conditions must have 487 been similar to those of the early Holocene. Wet conditions between c. 850 and 1100 AD in 488 Stalagmite ANJB-2 and Stalagmite MAJ-5, specifically coincide with glacial advances at northern 489 high latitudes (Holzhauser et al., 2005) and a cooler interval of the Medieval Climate Anomaly, as 490 suggested by a negative temperature Anomaly in the NH (e.g., Büntgen et al., 2011; Mann et al., 491 1998; Mann and Bradley, 1999, see also Fig. S18). The sudden beginning of stalagmite growth during the MLHI and the large δ^{13} C shift from depleted to enriched values at c. 1.5 ka BP (Fig. 6), 492 493 after such long hiatuses may have been associated with changes in vegetation cover above the 494 cave linked to recent human activities (e.g., Burns et al., 2016; Crowley and Samonds, 2013; Crowther et al., 2016; Voarintsoa et al., 2017b). Lower δ^{13} C values in Stalagmite MAJ-5 after 0.8 495

ka BP (Fig. 3), compared to higher values in Stalagmite ANJB-2, suggests different conditions in or
above the two caves. More human disturbance at one site could account for the different
trends, or alternatively changes in cave micro-climate, or in the hydrologic catchments of the
two stalagmites.

Although the stalagmite data indicate overall wetter conditions during the last c. 1.6 kyr, there were occasional dry periods, as suggested by several positive peaks in the stalagmite δ^{18} O records. Drier intervals during the Late Holocene are observed in the Anjohibe data between c. AD 755 and 795 (i.e., 1195–1155 yr. BP; Voarintsoa et al., 2017b). Similar conditions have been recorded in other paleoenvironmental studies, in which a peak drought c. 1300–950 cal BP was reported (Burney, 1987a, b; Burney, 1993; Matsumoto and Burney, 1994; Virah-Sawmy et al., 2009).

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508 *5.3*.Holocene climate in NW Madagascar: implications for ITCZ dynamics

509 Figures 7 and 8 depict possible conditions in NW Madagascar during the MEHI, the 510 MMHI, and the MLHI. Figure 9 summarizes the possible forcings mechanisms linked to the 511 latitudinal migration of the ITCZ.

In NW Madagascar, stalagmite deposition during the MEHI and the MLHI could suggest 512 513 there was sufficient dripwater for stalagmite growth and therefore wetter conditions. This could 514 have been linked to a more southerly mean position of the ITCZ. Factors that could influence the 515 mean position of the ITCZ include changes in insolation (e.g., Haug et al., 2001; Wang et al., 516 2005; Cruz et al., 2005; Fleitmann et al., 2003, 2007; Schefuß et al., 2005; Suziki, 2011; Kutzbach 517 and Liu, 1997; Partridge et al., 1997; Verschuren et al., 2009; Voarintsoa et al., 2017a) and 518 difference in temperature between the two hemispheres (e.g., Chiang and Bitz, 2005; Broccoli et 519 al., 2006; Chiang and Friedman, 2012; Kang et al., 2008; McGee et al., 2014; Talento and 520 Barreiro, 2016).

In contrast, the depositional hiatuses during the MMHI could suggest drier conditions, and thus a northward migration of the mean ITCZ. It seems to agree with the paleoclimate simulation of Braconnot et al. (2007) of the 6 ka event, suggesting that the NH insolation increased (Braconnot et al., 2000; see also Chiang, 2009). This northward shift in the mean position of the ITCZ is consistent with drier conditions, i.e. weaker South American Summer
Monsoon (e.g., Cruz et al., 2005; Seltzer et al., 2000; Wang et al., 2007; but see also Fig. 9 of
Zhang et al., 2013) but wetter conditions in the northern tropics (e.g., Dykoski et al., 2005;
Fleitmann et al., 2007; Gasse, 2000; Haug et al., 2001; Weldeab et al., 2007; Zhang et al., 2013).

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530 5.4. Regional comparisons

531 Despite differences in Holocene paleoclimate reconstructions for southern Africa, 532 comparison of the NW Madagascar records with records from neighboring locations (Figs. 10-533 11; Table S3) shows that the Holocene wet/dry/wet succession reported in this study has also 534 been identified at other locations. For example, hydrogen isotope compositions of the n-C31 535 alkane in GeoB9307-3 from a 6.51 m long marine sediment core retrieved about 100 km off the 536 Zambezi delta suggest a similar wet/dry/wet climate during Early, Middle, and Late Holocene 537 respectively (Schefuß et al., 2011). Those changes correspond to changes in temperature from 538 ~26.5° to 27.25° to 27°C, respectively, in the Mozambique Channel, as suggested by alkenone 539 SST records from sediment cores MD79257 (Bard et al., 1997; Sonzogni et al., 1998). The 540 Zambezi catchment is specifically relevant here because it is located at the southern boundary of 541 the modern ITCZ, and so has similar climatic setting as NW Madagascar, and its sensitivity to the 542 latitudinal migration of the ITCZ could parallel that of Madagascar. Likewise, temperature 543 reconstruction from the Mozambique Channel could be used to link regional changes in 544 paleorainfall with regional changes in temperature. A general overview of the Holocene climate 545 in the African neighboring locations to Madagascar suggests a roughly consistent wetter and 546 drier climate during the early and middle Holocene, respectively (Fig. 11, Table S3, also see 547 Gasse, 2000; Singarayer and Burrough, 2015). However, Late Holocene paleoclimate 548 reconstructions vary. A single answer to this variability is unlikely, but several overlapping 549 factors, including the latitudinal migration of the ITCZ, changes in ocean oscillations and sea 550 surface temperatures, volcanic aerosols, and anthropogenic influences could have played a 551 major role in such variability (e.g., Nicholson, 1996; Gasse, 2000; Tierney et al., 2008; Truc et al., 552 2013). Assessing these factors is beyond the scope of this study.

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554 5.5. The 8.2 ka event in Madagascar: linkage to ITCZ and AMOC

555 The 8.2 ka event was a significant short-lived cooling of the N Atlantic and NH during the 556 Early Holocene (Alley et al., 1997). It is apparent in the ANJB-2 and MAJ-5 stalagmite records as a 557 wet interval (Sect. 5.2.1; Figs. 5, 12). The 8.2 ka event is a known interval of abrupt freshwater 558 influx from the melting Laurentide Ice Sheet into the North Atlantic (Alley et al., 1997; Barber et 559 al., 1999; Kleiven et al., 2008; Carlson et al., 2008; Renssen et al, 2010; Wiersma et al., 2011; 560 Wanner et al., 2015). It is equivalent to the sharp peak of Bond cycle number 5 (Bond et al. 1997, 561 2001). This influx of meltwater altered the density and salinity of the NADW. Thornalley et al. 562 (2009) report that there was a decrease in NADW salinity to approximately 34 p.s.u. during the 563 Early Holocene.

564 Understanding the AMOC's influence on Madagascar's hydroclimate could help us better 565 understand global atmospheric and oceanic circulation, particularly in the SH. An increase in the 566 flow of freshwater to the North Atlantic decreases the formation of North Atlantic Deep Water, 567 reducing the meridional heat transport (Barber et al., 1999; Clark et al., 2001; Daley et al., 2011; 568 Vellinga and Wood 2002; Dong and Sutton 2002, 2007; Dahl et al. 2005; Zhang and Delworth 569 2005; Daley et al., 2011; Renssen et al., 2001). Weakening of the AMOC would ultimately cause a 570 widespread cooling in the NH regions (e.g., Clark et al., 2001; Thomas et al., 2007) but warming 571 in the SH regions (Wiersma et al., 2011; Wiersma and Renssen, 2006). This "cold NH-warm SH" climate response is similar to the "bipolar seesaw" effect, well-known during the last glacial (e.g., 572 573 Crowley, 1992; Broecker, 1998). The interhemispheric temperature difference between the NH 574 and SH from such effect could be the driver of the southward displacement of the mean position 575 of the ITCZ during the 8.2 ka abrupt cooling event. This in turn could have led to an intensified 576 Malagasy monsoon in NW Madagascar during austral summers, a phenomenon identical to the 577 South American Summer Monsoon identified in Brazil (e.g., Cheng et al., 2009). In contrast, 578 regions in the NH monsoon regions became dry at 8.2 ka BP as the Asian Monsoon and the East 579 Asian Monsoon became weaker (e.g., Wang et al., 2005; Dykoski et al., 2005; Cheng et al., 2009; 580 Liu et al., 2013).

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5.6. Beyond the ITCZ: IOD and ENSO influence on Madagascar's climate

583 Although the ITCZ is the main driver of rainfall availability in Madagascar, recent studies have 584 also suggested the importance of SST changes in the surrounding ocean and teleconnection with 585 other climatic phenomena. Scroxton et al. (2017) linked rainfall changes in eastern Indian Ocean 586 with expansion and contraction of the ITCZ along with positive IOD. Zinke et al. (2004) revealed 587 strong Indian Ocean subtropical dipole events that were in phase with ENSO indices between AD 1880 and 1920, and between 1930 and 1940, and after 1970 in austral summers. Brook et al. 588 589 (1999, p. 700) suggested linkages between rainfall and ENSO in NW Madagascar since AD 1550, a 590 relationship that is less clear and complicated. This complication could be associated with an 591 unclear or yet a limited understanding of the relationship between IOD and ENSO, which is not 592 yet fully understood (e.g., Saji et al., 1999; Li et al., 2003; Lee et al., 2008 versus Brown et al., 593 2009; Schott et al., 2009; Shinoda et al., 2004; Venzke et al., 2000; Abram et al., 2008; Saji and 594 Yagamata, 2003; Meyers et al., 2007).

595 Our understanding of the oceanic and atmospheric circulation is challenged because IOD and 596 ENSO share similar features in the associated SST and precipitation anomalies (e.g., Saji et al., 597 1999; Webster et al., 1999; Krishnamurty and Kirtman, 2003; Meyers et al., 2007). In addition, 598 the driving mechanisms of ENSO and IOD during the Holocene are not fully understood, even 599 though linkages with insolation were reported (e.g., Otto-Bliesner et al., 2003; Liu et al., 2000; 600 Timmermann et al., 2007; Zheng et al., 2008; Tudhope et al., 2001; Moy et al., 2002; Koutavas et 601 al., 2006; Conroy et al., 2008; Kuhnert et al., 2014; Liu et al., 2003; Abram et al., 2007). The IOD 602 signals in the tropical Indian Ocean may additionally be overridden by the global mean 603 temperature (e.g., Vecchi and Soden, 2007; Zheng et al., 2013), or the signals could be strongly 604 influenced by monsoonal changes in the surrounding landmasses (e.g., Abram et al., 2007; Qiu et 605 al., 2012).

Despite the complicated relationships, it is possible that climate of NW Madagascar has been influenced by ITCZ, IOD, and ENSO, but this is still poorly understood during the Holocene. We are aware that the temporal and spatial resolution of available records make this investigation challenging, and we understand that the range of uncertainty of radiometric ages of several paleoclimate data could be another barrier to fully evaluate such relationship (see for example Fig. 7 of Kuhnert et al., 2014).

612 6. Conclusions

613 Petrography, mineralogy, and stable isotope records from Stalagmite ANJB-2, from Anjohibe 614 Cave, and Stalagmite MAJ-5, from Anjokipoty Cave, combine to suggest three distinct intervals of 615 changing climate in Madagascar during the Holocene: relatively wet conditions during the MEHI, 616 relatively drier conditions, possibly due to episodic dryness, during the MMHI, and relatively wet 617 conditions during the MLHI. The timing of stalagmite deposition during the MEHI and the MLHI 618 in NW Madagascar could be attributed to a more southward migration and/or an expanded ITCZ, 619 increasing the duration of the summer rainy seasons, perhaps linked to a stronger Malagasy 620 monsoon. This could have been tied to insolation, the temperature gradient between the two 621 hemispheres, and weakening of the AMOC. In contrast, the c. 6500 year depositional hiatus 622 during the MMHI could indicate a northward migration of the ITCZ, leading to relatively drier conditions in NW Madagascar. The evidence of the 8.2 ka event in the Malagasy records further 623 624 suggests a strong link between paleoenvironmental changes in Madagascar and abrupt climatic 625 events in the NH, suggesting that during the MEHI Madagascar's climate was very sensitive to 626 abrupt ocean-atmosphere events in the NH.

627 Although the ITCZ is one of the climatic drivers influencing climate in Madagascar and its 628 surrounding locations, several climatic factors need to be investigated in more detail. For 629 example, we do not fully understand if the latitudinal migration is paired with the expansion 630 and/or expansion of the ITCZ, responsible to changes in several monsoon systems. In addition, 631 the interplay between ITCZ and other factors involving changes in sea surface temperatures, 632 particularly IOD-ENSO, needs to be investigated in details. Data-model comparison seems to be 633 an approach to better understand such relationship. The lack of spatial and temporal resolution 634 of paleoclimate records is still a challenge to fully understand the climate system during the 635 Holocene.

636

637 Author Contribution

N.R.G.V. conceived the research and experiments. N.R.G.V, G.K, A.F.M.R, and M.O.M.R did
 the fieldwork and collected the samples. X.L., G.K., H.C., R.L.E, and N.R.G.V contributed to the
 ²³⁰Th dating analyses. N.R.G.V provided detailed investigation of the two stalagmites, provided

- 641 stable isotope measurements, prepared thin sections, and conducted X-ray diffraction analyses.
- 642 G.K. also assisted with the isotopic measurements on Stalagmite ANJB-2. N.R.G.V. wrote the first
- 643 draft of the manuscript and led the writing. L.B.R. and G.A.B. provided a thorough review of the

644 draft. N.R.G.V. and L.B.R. discussed and revised the manuscript, with additional comments from

- 645 L.W. N.R.G.V revised the paper with input from all authors, reviewers, and editors.
- 646

647 Competing Interests

648 The authors declare no conflict of interest.

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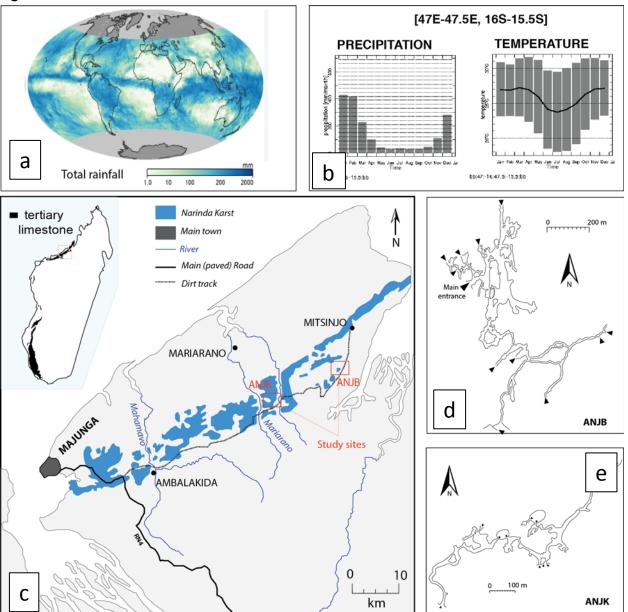
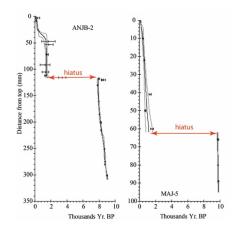


Figure 1: Climatological and geographic setting of Madagascar and the study area. (a) Global rainfall maps recorded by NASA's Tropical Rainfall Measuring Mission (TRMM) satellite showing the total monthly rainfall in millimeters and the overall position of the ITCZ during November, 2006. Darker blue shades indicate regions of higher rainfall (source: NASA Earth Observatory, 2016). (b) Barplots of the monthly climatology of precipitation, and the monthly average of daily maximum, minimum, and mean temperature in NW Madagascar. The base period used for the climatology is 1971-2000. Source: http://iridl.ldeo.columbia.edu/ (accessed August 31, 2016). (c)

Simplified map showing the southwest part of the Narinda karst and the location of the study areas. Inset figure is a map of Madagascar showing the extent of the Tertiary limestone cover that makes up the Narinda karst. (d-e) Maps of Anjohibe (ANJB) and Anjokipoty (ANJK) caves (St-Ours, 1959; Middleton and Middleton, 2002). See Figs. S1–S3 for additional information about the study locations.

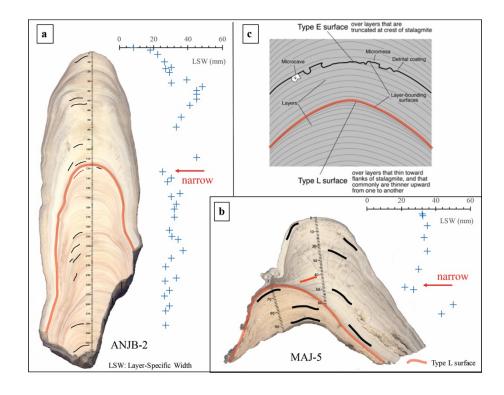
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1268 Figure 2: Age model of Stalagmite ANJB-2 and MAJ-5 using the StalAge1.0 algorithm of Scholz and

1269 Hoffman (2011) and Scholz et al. (2012).



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Figure 3: a) Scanned image of Stalagmite ANJB-2 and the corresponding variations in layerspecific width (LSW). b) Scanned image of Stalagmite MAJ-5 and the corresponding layer-specific
width (LSW). c) Sketches of typical layer-bounding surfaces (Type E and Type L) of Railsback et al.

- 1277 (2013). Close-up of photographs of the hiatuses are shown in Fig S6.
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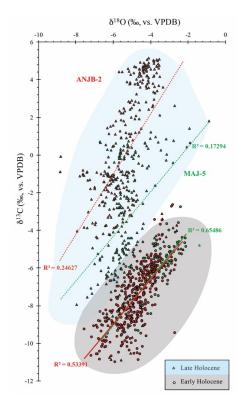
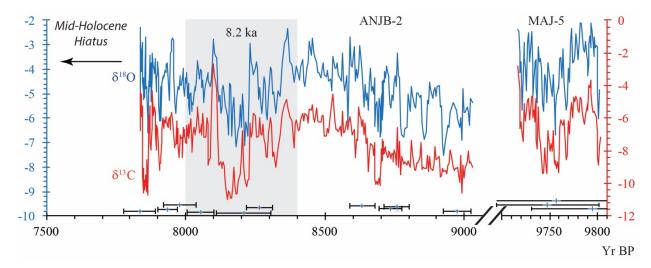




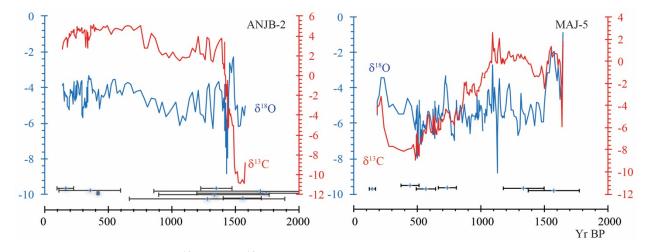
Figure 4: Stable isotope data. Scatterplots of δ^{13} C and δ^{18} O for Stalagmite MAJ-5 (green) and ANJB-2 (red) during the Malagasy early Holocene interval (circle) and the Malagasy late Holocene interval (triangle). The plot shows distinctive early and late Holocene conditions (roughly highlighted in gray and light blue shade, respectively).

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1289 Figure 5: Variations in δ^{13} C and δ^{18} O in Stalagmite ANJB-2 and Stalagmite MAJ-5 during the 1290 Malagasy Early Holocene Interval. Supplementary Fig. S6 shows both the corrected and 1291 uncorrected values.

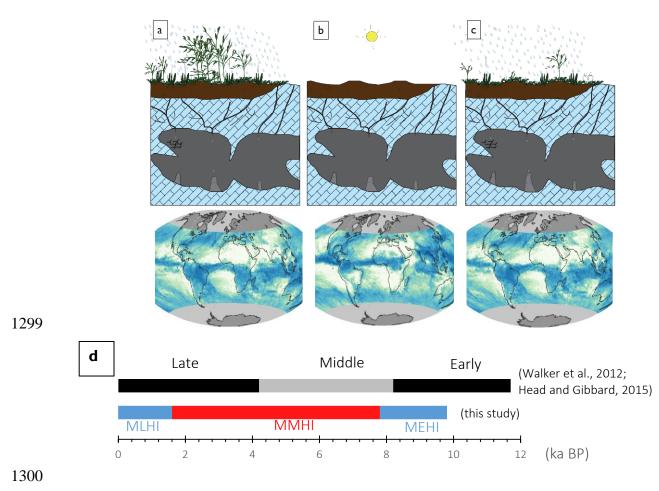




1294 Figure 6: Variations in δ^{13} C and δ^{18} O in Stalagmite ANJB-2 and Stalagmite MAJ-5 during the 1295 Malagasy Late Holocene Interval. Supplementary Fig. S7 shows both the corrected and 1296 uncorrected values, and Fig. S8 compares the corrected δ^{18} O for both stalagmites.

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1301 Figure 7: Simplified models portraying the Holocene climate change in NW Madagascar and the 1302 possible climatic conditions linked to the ITCZ. a) Wetter conditions during the early Holocene 1303 with ITCZ south (prior to c 7.8 ka), favorable for stalagmite deposition. b) Periodic dry conditions 1304 during the mid-Holocene (between c. 7.8 and 1.6 ka) with ITCZ north with no stalagmite 1305 formation (refer to Sect. 5.2.2). c) Wetter conditions during the late Holocene (after c. 1.6 ka) 1306 with ITCZ south, favorable for stalagmite deposition. For details about paleo-vegetation 1307 reconstruction. Drawings are not to scale. The bottom figures are from the same source as Fig. 1308 1a, and they are only used here to give a perspective of the possible position of the ITCZ during 1309 the early, mid, and late Holocene. d) Comparison of the three Malagasy Holocene interval with 1310 the Walker et al. (2012) and Head and Gibbard (2015) subdivision (see text for details, Sect. 5.2). 1311

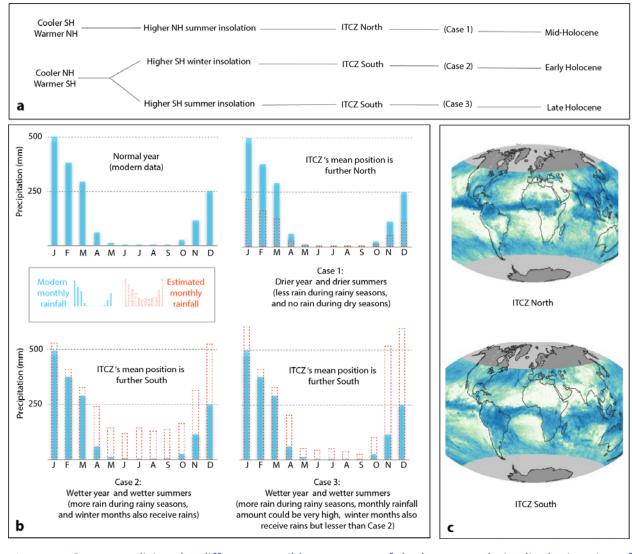




Figure 8: Conceptualizing the different possible outcomes of the long-term latitudinal migration of the ITCZ. a) Highlighting the three possible scenarios of the Holocene. b) Barplots of monthly rainfall in NW Madagascar, using the modern data as a reference to estimating the region's paleoclimate during drier and wetter conditions. c) Global rainfall maps from NASA (same source as Fig.1). These maps are modern, but they are only shown here to give a better perspective of the position of Madagascar when the ITCZ is relatively north or south. See supplementary text for details.

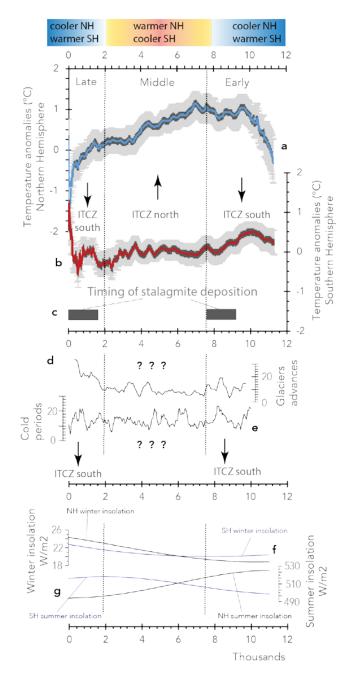


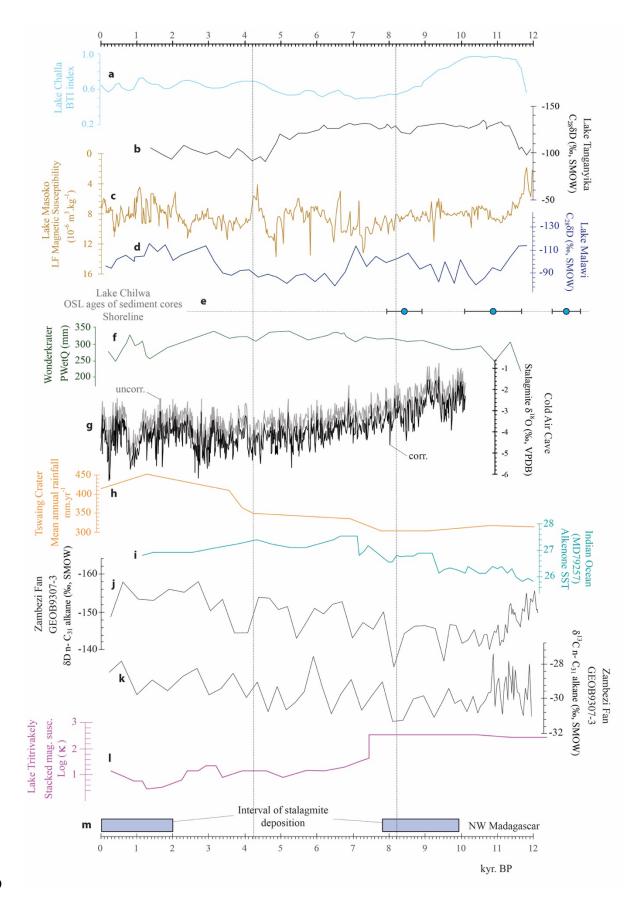
Figure 9: Possible Holocene climate forcings that influenced climate of NW Madagascar. a)
Average Holocene temperatures in the NH 90°–30°N (blue). b) Average Holocene temperatures
in the SH 90°–30°S (red). These temperature data are referenced to the 1961–1990 mean
temperature (Marcott et al., 2013), with 1σ uncertainty (gray). c) Timing of deposition of
Stalagmite ANJB-2 and MAJ-5. d) Curves representing the sum of glaciers advances from a set of
global Holocene time series compiled from natural paleoclimate archives (Wanner et al., 2011).
e) curves representing the sum of cold periods from a set of global Holocene time series

- 1329 compiled from natural paleoclimate archives (Wanner et al., 2011). f) Winter insolation curves
- 1330 (Berger and Loutre, 1991). G) Summer insolation curves (Berger and Loutre, 1991)
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- 1336 for GeoB9307-3 (onshore off delta sediments), MD79257 (alkenone from marine sediment core),
- 1337 and Cold Air, Anjohibe, and Anjokipoty caves (stalagmites δ^{18} O).
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1340 Figure 11: Regional comparison. a) Lake Challa BTI index (Verschuren et al., 2009). b) Lake 1341 Tanganyika $C_{28} \delta D$ (Tierney et al., 2008, 2010). c) Lake Masoko low field magnetic susceptibility $(10^{-6} \text{.m}^3 \text{kg}^{-1})$ (Garcin et al., 2006). d) Lake Malawi C₂₈ δ D (Konecky et al., 2011). e) Lake Chilwa 1342 1343 OSL dates of shoreline (Thomas et al., 2009). f) Wonderkrater reconstructed paleoprecipitation, 1344 PWetQ (Precipitation of the Wettest Quarter; Truc et al., 2013). g) Cold Air Cave corrected (corr.) and uncorrected (uncorr.) δ^{18} O profiles from Stalagmite T8 (Holmgren et al., 2003). h) 1345 Tswaing Crater paleo-rainfall derived from sediment composition (Partridge et al., 1997). i) 1346 1347 Indian Ocean SST records from alkenone (Bard et al., 1997; Sonzogni et al., 1998). j-k) Zambezi δD n-C₃₁ alkane $\delta^{13}C$ n-C₃₁ alkane (Schefuß et al., 2011). I) Lake Tritrivakely stacked magnetic 1348 1349 susceptibility (Williamson et al., 1998). m) NW Madagascar (Anjohibe and Anjokipoty) interval of 1350 deposition of Stalagmite ANJB-2 and Stalagmite MAJ-5 (this study). The two vertical dashed lines 1351 indicate the boundary of the Early, Middle, and Late Holocene by Walker et al. (2012) and Head 1352 and Gibbard (2015).

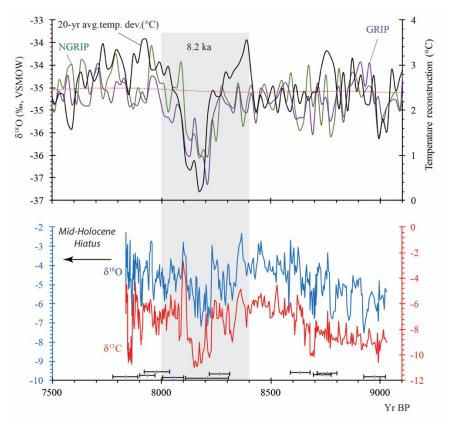


Figure 12: The 8.2 ka event in Madagascar. Oxygen isotope record from Greenland (GRIP and NGRIP) ice cores (Vinther et al., 2009) compared with Stalagmite ANJB-2 δ^{18} O and δ^{13} C.