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Using ice-flow models to evaluate potential sites of million year-old ice in Antarctica

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Abstract

Finding suitable potential sites for an undisturbed record of million-year old ice in Antarctica requires a slow-moving ice sheet (preferably an ice divide) and basal conditions that are not disturbed by large topographic variations. Furthermore, ice should be thick and cold basal conditions should prevail, since basal melting would destroy the bottom layers. However, thick ice (needed to resolve the signal at sufficient high resolution) increases basal temperatures, which is a conflicting condition in view of finding a suitable drill site. In addition, slow moving areas in the center of ice sheets are also low-accumulation areas, and low accumulation reduces potential cooling of the ice through vertical advection. While boundary conditions such as ice thickness and accumulation rates are relatively well constraint, the major uncertainty in determining basal conditions resides in the geothermal heat flow (GHF) underneath the ice sheet. We explore uncertainties in existing GHF datasets and their effect on basal temperatures of the Antarctic ice sheet and propose an updated method based on Pattyn (2010) to improve existing GHF datasets in agreement with known basal temperatures and their gradients to reduce this uncertainty. Both complementary methods lead to a better comprehension of basal temperature sensitivity and a characterization of potential ice coring sites within these uncertainties.

1 Introduction

One of the major future challenges in the ice coring community is the search for a continuous and undisturbed ice core record dating back to 1.5 million years BP (Jouzel and Masson-Delmotte, 2010). The reason for such quest is that the oldest part of the EPICA Dome C ice core has revealed low values of CO₂ from 650 000 to 800 000 yr ago (Lüthi et al., 2008), which questions the strong Antarctic temperature-carbon cycle coupling on long time scales. Marine records show evidence of a reorganisation of the pattern of climate variability around 1 Myr ago, shifting from the “obliquity” dominated

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signal, characterized by 40 000 yr weak glacial-interglacial cycles to the “eccentricity”-dominated signal with longer glacial-interglacial cycles (Lisiecki and Raymo, 2005). The origin of this major climate reorganization (the so-called Mid-Pleistocene Transition, MPT) remains unknown and may be intrinsic to a series of feedback mechanisms between climate, cryosphere and carbon cycle (Jouzel and Masson-Delmotte, 2010). Alternatively, a recent study has demonstrated that climate oscillations over the past four million years can be explained by a single mechanism, i.e. the synchronization of nonlinear internal climate oscillations and the 413 000 yr eccentricity cycle (Rial et al., 2013). According to model calculations in conjunction with spectral analysis, Rial et al. (2013) find that the climate system first synchronized to this 413 000 yr eccentricity cycle about 1.2 million years ago, roughly coinciding with this MPT.

Deep ice core drillings have been carried out in the past in Antarctica, reaching back in time over several hundred thousands of years. Amongst the longest records are Vostok (Petit et al., 1999), EPICA Dome Concordia (EPICA community members, 2004), Dome Fuji (Watanabe et al., 2003), and EPICA Dronning Maud Land (EPICA community members, 2006). The longest record is EPICA Dome Concordia, going back for more than 800 000 yr. Those records have in common that they are all recovered in the center of the ice sheet, and given the fact that the Antarctic ice sheet has been relatively constant in size and over the last 13 million years (DeConto and Pollard, 2003), they are consequently undisturbed by dramatic changes in ice flow, contrary to the longest records from the Greenland ice sheet (Johnson, 2001; NEEM community members, 2013).

In theory, and in absence of basal melting, these deep Antarctic records could reach several million years back in time with layers getting infinitesimally thin near the bottom. In reality, however, all deep records lack the bottom sequence as they are all found to be at pressure melting point and lower layers are melted away or heavily disturbed due to complex basal processes. Furthermore, resolving deep records not only requires that the bottom sequence is unaltered, but that the ice is sufficiently thick so that the gas

signal can still be retrieved and analyzed with **sufficient** high resolution in the bottom layers.

In this paper we use several thermodynamic models to infer suitable areas for retrieving long ice-core records. We first investigate the most influential parameters having an effect on ice-core record length. Secondly, we apply this simple concept to evaluate uncertainties in GHF and use this uncertainty to guide the search for a suitable drilling place. Thirdly, we carry out a sensitivity analysis with a three-dimensional thermodynamical model (Pattyn, 2010) to determine the sensitivity of basal conditions to uncertainties in GHF, guided by a priori knowledge of basal conditions through the geographical distribution of subglacial lakes. **Results are discussed in the last section.**

2 Why obvious drill sites are unsuitable

Obvious places to look for oldest ice are the deepest parts of the ice sheet, where ice is thick, and accumulation rates are low. However, a thick ice cover insulates very well and keeps the geothermal heat from escaping to the surface. Furthermore, we know that at least 379 subglacial lakes exist under the Antarctic ice sheet, which implies that large portions of the bedrock should be at pressure melting point (Smith et al., 2009; Pattyn, 2010; Wright and Siegert, 2012). Most subglacial lakes occur in the so-called Lakes District (stretching between Subglacial Lake Vostok and Wilkes Land in East Antarctica), characterized by a thick ice cover and also low geothermal heat flow (Shapiro and Ritzwoller, 2004; Pollard et al., 2005). Therefore, GHF is not the main culprit in causing subglacial melt.

The interplay between GHF and accumulation rates is very subtle, as high GHF increases basal temperatures, while high accumulation rates cool down the ice mass. To illustrate this we calculate the minimum GHF needed to reach pressure melting point at the bottom of any ice mass as a function of environmental parameters. This can easily be determined analytically (Hindmarsh, 1999; Siegert, 2000). Using the simplified

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high accumulation rates, one needs a significantly higher GHF to reach melting point at the base for a given ice thickness.

Even the low GHF values in Fig. 1 in conjunction with low accumulation rates are quite common over the central part of the East Antarctic ice sheet, which hampers the retrieval of a long time sequence under thick ice conditions. Despite the simplicity of the model, it can be applied to central parts of the Antarctic ice sheet (absence of horizontal advection) to explore suitable drill sites as a function of known (or estimated) geothermal heat fluxes.

3 Uncertainties in Antarctic GHF and the oldest ice

3.1 Data sets and model setup

The above simple model is applied to central areas of the Antarctic ice sheet that are characterized by slow ice motion, hence the vicinity of ice divides. To do so, ice thickness is taken from the recent BEDMAP2 compilation (Fretwell et al., 2013) and resampled on a 5 km grid. Surface mass balance is obtained from van de Berg et al. (2006) and van den Broeke et al. (2006), based on the output of a regional atmospheric climate model for the period 1980 to 2004 and calibrated using observed mass balance rates. Surface temperatures are due to van den Broeke (2008), based on a combined regional climate model, calibrated with observed 10 m ice temperatures. Using these datasets enables us to calculate the minimum required GHF to reach pressure melting point at the bed. Since the model is along the vertical dimension, it is only valid for divide areas. The simulations are therefore only carried out for horizontal velocities smaller than 2 myr^{-1} . Ice sheet velocities in divide areas are determined based on balance velocities, stating that the mass of ice flowing out of any area within the horizontal domain (x, y) exactly equals the sum of the inflow and the ice accumulated over the area (Budd and Warner, 1996; Fricker et al., 2000; Le Brocq et al., 2006),

$$\nabla_H \mathbf{q}_s = \dot{a}, \quad (4)$$

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with

$$\mathbf{q}_s = -H \int_b^s \mathbf{v}_H(z') dz', \quad (5)$$

and where b and s are the bottom and the surface of the ice sheet (m a.s.l.), respectively. Integrating Eq. (4) over the whole surface of the ice sheet, starting at the ice divides, one obtains the vertically averaged horizontal balance velocities $\bar{\mathbf{v}}_H = (\bar{v}_x, \bar{v}_y)$. Details of this procedure are given in Pattyn (2010).

Using the above datasets, the minimum geothermal heat flow G_{\min} from Eq. (1) needed to reach pressure melting point at the bed is calculated for the areas of the Antarctic ice sheet where horizontal flow velocities are $< 2 \text{ myr}^{-1}$ and where ice thickness $H > 2000 \text{ m}$. This ice thickness is considered to be the lower limit for possibly recovery of a million-year old climate signal (H. Fischer, personal communication, 2013). The calculated values of G_{\min} are subsequently compared to known values of GHF.

Several datasets of derived GHF underneath the Antarctic ice sheet exist. The first one (G_1) uses a global seismic model of the crust and the upper mantle to guide the extrapolation of existing heat-flow measurements to regions where such measurements are rare or absent (Shapiro and Ritzwoller, 2004). The second GHF database (G_2) stems from satellite magnetic measurements (Fox-Maule et al., 2005). Values of GHF are in the same range as Shapiro and Ritzwoller (2004), but the spatial variability is contrasting. Heat flow measurements according to Fox-Maule et al. (2005) are also considerably higher than those by Shapiro and Ritzwoller (2004). A third dataset is due to Puruker (2013). This geothermal heat flux data set (G_3) is based on low resolution observations collected by the CHAMP satellite between 2000 and 2010, and produced from the MF-6 model following the technique described in Fox-Maule et al. (2005). While the technique is similar, GHF values are considerably lower than the latter, and even lower than those derived from the seismic model (Shapiro and Ritzwoller, 2004).

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3.2 Results

In view of the large uncertainty in GHF estimates, we combined all three datasets into two databases, i.e. a mean GHF, \bar{G} , and a standard deviation, σ_G . The latter is calculated based on the inter-database variability and the standard deviation given for the Shapiro and Ritzwoller (2004) dataset in the following way:

$$\sigma_G = \sigma [G_1 - \sigma(G_1), G_1 + \sigma(G_1), G_2, G_3]. \quad (6)$$

Both are depicted in Fig. 2. High values of σ_G indicate a large dispersion between the three datasets. These are essentially found in West Antarctica and along the Transantarctic Mountains. The lowest values are restricted to the central parts of the East-Antarctic continent.

The calculated values of G_{\min} are directly compared to the map of mean GHF. For $G_{\min} < \bar{G}$, the observed GHF is too elevated to prevent the bottom ice to reach pressure melting and most likely (within error bounds) the ice is temperate. For $G_{\min} > \bar{G}$ the minimum GHF needed to reach pressure melting at the base is higher than the value reported. Of course, this information needs to be further evaluated against the dispersion between the GHF datasets, represented by σ_G . The result is shown in Fig. 3, where the rectangular area points to the potentially most suitable conditions in terms of basal temperature, i.e. the largest difference between actual GHF and minimum GHF in combination with the lowest variability between the three GHF datasets. The furthest to the right in Fig. 3, the colder the bed because a significant higher GHF than observed is needed to make the bed temperate; the lower the value of σ_G , the more likely there is a small spread (hence reduced uncertainty) in GHF, so that the observed value is likely. On top of this, the color scale shows the ice thickness for each of the points. The thickest ice is obviously corresponding to zones that are temperate (negative values of ΔG), while for large positive ΔG and small σ_G , ice is also the thinnest.

These restrictions (superposed on the ice flow speed limit and minimum ice thickness) lead to a few areas in the central part of the Antarctic ice sheet that are

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considered suitable for cold-bed conditions. The largest zone is situated near Dome Argus on top of the Gamburtsev subglacial Mountains (Fig. 4). However, subglacial mountain ranges are likely characterized by an accidented bed topography (Bell et al., 2011), which may also hamper the interpretation of the paleo-climatic signal (Groote et al., 1993). Other potential areas are situated around Dome Fuji as well as on Ridge B, between Subglacial Lake Vostok and Dome Argus. The Dome Concordia area seems less prone to cold basal conditions, due to the large uncertainty in GHF and the thick ice, which makes temperate conditions more acceptable.

4 Thermomechanical ice-flow modelling

The simple thermodynamic model used in the previous section neglects horizontal advection, which – even in the interior of the Antarctic ice sheet – will play a significant role in changing the thermal properties of the ice/bed interface. In the next section we present a more advanced thermomechanical ice-sheet model to calculate basal temperatures for a set of given boundary conditions and applied to the whole Antarctic ice sheet. Moreover, we try to reduce uncertainties in GHF by incorporating actual information on bed properties, such as the geographical distribution of subglacial lakes.

4.1 Model description

The thermodynamical model used for this purpose is the same as described in detail in Pattyn (2010). The major differences are related to the way the horizontal flow field is calculated. Moreover, we use a new series of datasets on ice thickness (see Sect. 3.1) and geothermal heat flow.

The thermodynamic equation for the temperature distribution in an ice mass is given by

$$\rho c_p \frac{\partial T}{\partial t} = \nabla(k \nabla T) - \rho c_p \mathbf{v} \cdot \nabla T - 2\dot{\epsilon}\sigma, \quad (7)$$

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where T is the ice temperature (K) and $\mathbf{v} = (v_x, v_y, v_z)$ is the three-dimensional ice velocity vector (myr^{-1}). The last term on the right-hand side represents internal heating rate per unit volume (Pattyn, 2003), where $\dot{\epsilon}$ and σ are effective strain rate and effective shear stress, respectively. Horizontal diffusion is neglected, and the temperature field is considered to be in steady state ($\frac{\partial T}{\partial t} = 0$).

Boundary conditions for Eq. (7) are the surface temperature T_s and a basal temperature gradient, based on the geothermal heat flux:

$$\left. \frac{\partial T}{\partial z} \right|_b = -\frac{1}{k} (G + \tau_b \mathbf{v}_b), \quad (8)$$

where G is the geothermal heat flux entering the base of the ice sheet and the second term on the right-hand side of (8) heat produced due to basal sliding. τ_b is the basal shear stress, and can be defined as $\tau_b = -\rho g H \nabla_H s$, where $\nabla_H s$ is the surface slope. Whenever pressure melting point is reached, the temperature in the ice is kept at this value $T_{\text{mp}} = T_0 - \gamma(s - z)$.

4.2 Velocity field

Ice sheet velocities are obtained from a combination of observed velocities from satellite radar interferometry and modelled velocities. Satellite-derived velocities are available for almost the entire continent (Rignot et al., 2011), but are only relevant in the coastal areas and for fast ice flow. Generally spoken, the error associated with the slow flowing areas is substantially higher than 100 %. Furthermore, in vicinity of the South Pole, interferometric velocities are lacking due to the sun-synchronous orbit of satellites. To fill in the gaps and to guarantee a continuous flow field for our simulations, a heuristic method was implemented that uses interferometrically-derived velocities for flow speeds above 100 myr^{-1} and modelled velocities for flow speeds below 15 myr^{-1} . Modelled velocities are derived from balance velocities, described in Sect. 3. Between 15 and 100 myr^{-1} , both modelled and interferometric velocities are combined as a

fraction of flow speed, in order to keep the transition between both datasets as smooth as possible and to guarantee a correct flow direction.

The three-dimensional horizontal velocities are then determined from the shallow-ice approximation (Hutter, 1983), by

$$\mathbf{v}_H(x, y, \zeta) = \left(\frac{n+2}{n+1} \bar{\mathbf{v}}_H - \mathbf{v}_b \right) (1 - \zeta^{n+1}) + \mathbf{v}_b, \quad (9)$$

where basal sliding \mathbf{v}_b is represented by a Weertman sliding law (Pattyn, 2010). The vertical velocity field is derived from mass conservation combined with the incompressibility condition for ice. Given an ice sheet in steady state, a simple analytical expression can be obtained, based on the horizontal balance velocities (Hindmarsh, 1999; Hindmarsh et al., 2009). Expressed in local coordinates, and in the absence of subglacial melting, this leads to

$$v_\zeta(x, y, \zeta) = - \left[\frac{\zeta^{n+2} - 1 + (n+2)(1 - \zeta)}{n+1} \right] \dot{a} + \mathbf{v}_H \nabla b + (1 - \zeta) \mathbf{v}_H \nabla H. \quad (10)$$

The numerical solution of the model is detailed in Pattyn (2010). For all experiments, $n = 3$ was used, which corresponds to the isothermal case. However, in the thermomechanically-coupled case, the exponent is larger (Ritz, 1987), which results in a different shape of the vertical velocity profile, hence advection being more concentrated to the surface, leading to warmer basal conditions compared to the isothermal case. However, this effect is most pronounced in areas where horizontal velocity gradients $\partial v_x / \partial x$, $\partial v_y / \partial y$ are more important. Since we concentrate on the central areas of the ice sheet, this bias (underestimation of basal temperatures) will have a limited effect.

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4.3 Input data calibration

Major input datasets are already described in Sect. 3. In this section, we will focus on the improvements made to the initial GHF datasets in order to reduce uncertainty on GHF.

Direct measurements of GHF are very rare, and are usually obtained from temperature measurements in boreholes of deep ice core drillings. Basal temperature gradients in observed temperature profiles of deep boreholes, compared with values from the three GHF datasets, show rather large discrepancies. Therefore, the three GHF datasets were corrected using observed basal temperature gradients, surface temperature and accumulation rates, in such a way that modeled temperature profiles match as close as possible the observed ones (Pattyn, 2010).

This type of correction is made for sites where temperature profiles are available, i.e. Byrd (Gow et al., 1968), Taylor Dome (G. Clow and E. Waddington, personal communication, 2008), Siple Dome (MacGregor et al., 2007), Law Dome (Dahl-Jensen et al., 1999; van Ommen et al., 1999), Vostok (Salamat et al., 1994; Parrenin et al., 2004), South Pole (Price et al., 2002), Dome Fuji (Fujii et al., 2002; Hondoh et al., 2002), EPICA Dome C (Parrenin et al., 2007, C. Ritz, personal communication, 2008), and EPICA DML (Ruth et al., 2007). The applied method consists of determining the difference between observed (o) and corresponding database values and to adapt a Gaussian function for a sufficient large influence area. For a variable in the database X (either surface accumulation, surface temperature or geothermal heat flux), its corrected value X_c based on an observation X_o is obtained by

$$X_c(x, y) = X + [X_o - X] \exp \left[-\frac{x^2 + y^2}{\sigma^2} \right], \quad (11)$$

where (x, y) is the horizontal distance from this observed position (0,0). The influence area is dictated by σ and calculations were performed for $\sigma = 0, 20, 50, 100$, and 200 km (a discussion on the choice of these values is given below). As such, by

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tuning GHF (constraining the vertical temperature gradient) and surface mass balance (constraining vertical advection), the difference between modeled and observed temperature profiles is less than 2 K. The remaining difference is still due to horizontal advection, which is a model output, as well as past changes in surface temperature that were not taken into account in the model.

4.4 Subglacial lake correction

Numerous subglacial lakes have been identified from radio-echo sounding. An initial inventory brought their number on 145 (Siegert et al., 2005), and more than 230 have been added since (Bell et al., 2006, 2007; Carter et al., 2007; Popov and Masolov, 2007; Fricker et al., 2007; Fricker and Scambos, 2009; Smith et al., 2009), leading to a total number of 379 lakes of varying size (Wright and Siegert, 2012). Subglacial lakes are usually identified from radio-echo sounding (RES) and characterized by a strong basal reflector and a constant echo strength (corroborating a smooth surface) or identified through surface elevation changes using satellite altimetry, corroborating sudden subglacial water discharge (Fricker et al., 2007; Pattyn, 2011).

Subglacial lakes are used to constrain the GHF datasets, considering them to be at pressure melting point. As such, we calculate the minimum GHF needed to reach pressure melting point using Eq. (1) for any position of a subglacial lake. The value for G_{\min} thus obtained is a minimum value, which means that if at that spot the database contains a higher value, the latter is retained. Spatial corrections are subsequently applied using the Gaussian function defined in Eq. (11) for different influence areas as defined above.

5 Ensemble model results

The temperature field in the ice sheet was calculated for 15 different sets of boundary conditions, i.e. the three datasets of GHF (Shapiro and Ritzwoller, 2004; Fox-Maule

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et al., 2005; Puruker, 2013), and each of the datasets corrected for subglacial lakes and existing temperature profiles for influence area size $\sigma = 0$ (no correction), 20, 50, 100, and 200 km, respectively. The result is given in Fig. 5, representing the mean basal temperature of the 15 experiments, corrected for the dependence on pressure melting, and the RMSE (Root Mean Square Error) corresponding to the different experiments.

Low values of RMSE correspond to zones where correction is effective and the difference between the experiments is low, or areas that despite the variability in GHF are always at pressure melting point. This is the case for the central part of the West-Antarctic ice sheet, as well as extensive zones in the Lakes District, where the dense network of subglacial lakes keeps the bed at melting point.

Based on the ensemble experiments, the relation between accumulation (vertical advection), ice thickness and basal temperature is less straightforward than with the simple model. The focus of the full model is to reduce uncertainties on GHF using proxy data, and therefore the RMSE guides us towards more suitable sites. Figure 6 summarizes the most suitable drilling areas based on the full model for flow speeds lower than 2 myr^{-1} , ice thickness $H > 2000 \text{ m}$, and a basal temperature lower than -5°C . The color scale denotes the RMSE based on the ensemble experiments. We deliberately excluded basal temperatures higher than -5°C , a value considered to be sufficiently away from the melting point in view of our model approximations. Suitable areas characterized by low values of RMSE (hence smaller spread in basal temperatures according to the ensemble experiments) are found near existing ice core sites where a temperature gradient is at hand, i.e. Dome Concordia, Dome Fuji and Vostok. Since all three sites are at or close to **pressure melting** point at the base, suitable cold-based sites are not exactly situated at those spots, but in their vicinity where ice is slightly thinner.

Similar to the simple model, suitable sites (sufficiently low basal temperature) are found in the Gamburtsev Mountain region as well as along Ridge B. However, the ensemble analysis results in a larger spread of basal temperature range due to either the lack of basal temperature gradient constraints and/or the absence of subglacial

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Subglacial topography is a key factor in determining suitable sites for oldest ice. Given the strong relationship between basal temperatures and ice thickness, as depicted by Fig. 1, it is quite likely to find suitable cold-based spots in the vicinity of deep ice core sites that have the bottom ice at or near pressure melting point. Areas that should be avoided are those in which a large number of subglacial lakes are found, such as the Lakes District, where even low values of GHF are sufficient to keep the ice at pressure melting.

Another factor that may influence basal conditions is due to the glacial-interglacial history of the ice sheet and the time-scales needed for the ice sheet to thermally adapt to different climates. Moreover, the temperature calculations made in this study are based on present-day observed parameters of surface temperature, ice thickness and accumulation rate. To test this effect, we calculated the minimum geothermal heat flow G_{\min} needed to keep the base at pressure melting point for environmental conditions that are the mean for a longer time span. We reduced the surface temperature T_s by 6 K, reduced the surface accumulation rate \dot{a} to 60 % of its current value, and reduced ice thickness H with 100 m, which is valid for the divide areas. The results are surprisingly coherent with the previously-calculated values, and are therefore not shown separately. The main reason is that for this spread of values the reduced accumulation rate (which reduces vertical advection, hence warms the bottom ice layers) is largely counteracted by the decrease in surface temperature. However, one needs to keep in mind that both calculations (present-day and mean glacial-interglacial) relate to steady-state conditions, which in reality is not the case.

Nevertheless, one should be careful in using the above model results as a sole guide in the process of detecting suitable cold-based areas for retrieving a long ice-core record, due to a number of factors that were not taken into account:

1. Areas characterized by subglacial mountains or other bedrock variability may well be suitable from a thermal point of view, the topographic variability may well hamper the deciphering of the climate signal due to complex processes, such as ice overturning (NEEM community members, 2013) or refreezing (Bell et al., 2011).

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2. The upper limit on the flow velocity of 2 myr^{-1} may also be too high for reconstructing the climate signal without having to rely to heavily on ice flow models for correcting for upstream advection. In theory, ice could have traveled over several hundreds of kilometers before reaching the ice core site, and this without taking into account any shifts in ice divides over glacial-interglacial periods, which would also influence the flow direction over time.

3. The spatial variability of GHF may in reality be much higher than the one represented in the three GHF datasets.

4. Areas where bedrock data is unavailable (or where interpolation is based on sparse data) may be wrongly classified in the above analysis, and some suitable areas thus overseen.

While this paper gives an overview of continental-scale basal conditions of the Antarctic ice sheet, the processed datasets from both the simple (\bar{G} , σ_G) and the full model (\bar{T} , RMSE_T) will be made available online together with simple MatLab scripts to allow for a more detailed search/zoom for potential sites, based on the figures presented here.

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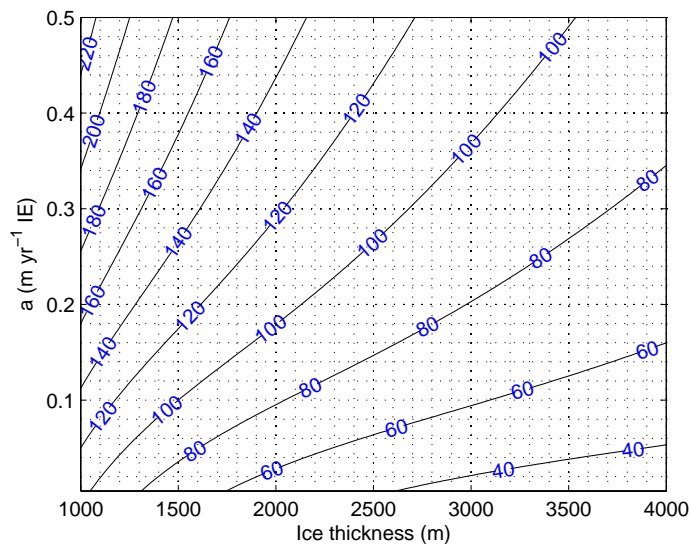


Fig. 1. Minimum GHF (mWm^{-2}) needed to keep the bed at pressure melting point as a function of surface accumulation rate (ice equivalent, IE) and ice thickness and in the absence of horizontal advection. Results are shown for a mean surface temperature of $T_s = -50^\circ\text{C}$.

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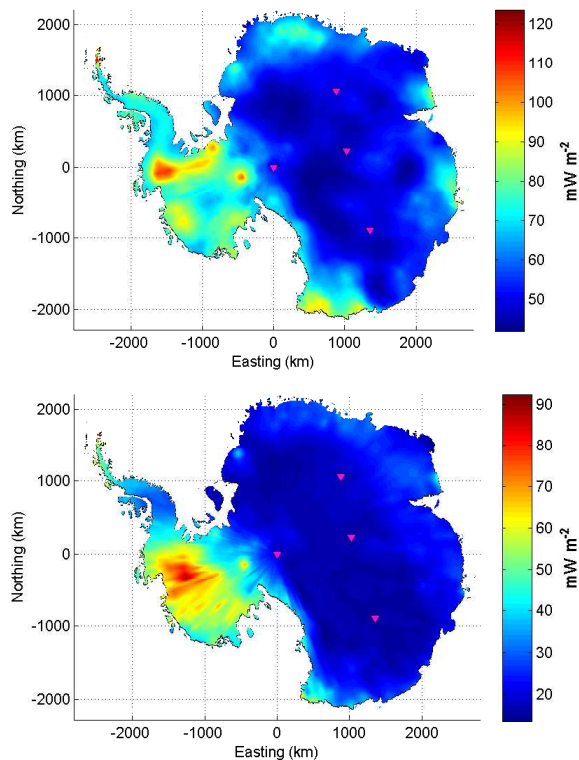



Fig. 2. Top: mean GHF \bar{G} (mW m^{-2}) based on GHF estimates by Puruker (2013), Fox-Maule et al. (2005) and Shapiro and Ritzwoller (2004). Bottom: standard deviation σ_G on the GHF datasets. The magenta triangles are the major drill **site** (from top to bottom): Dome Fuji, Dome Argus, South Pole and Dome Concordia.

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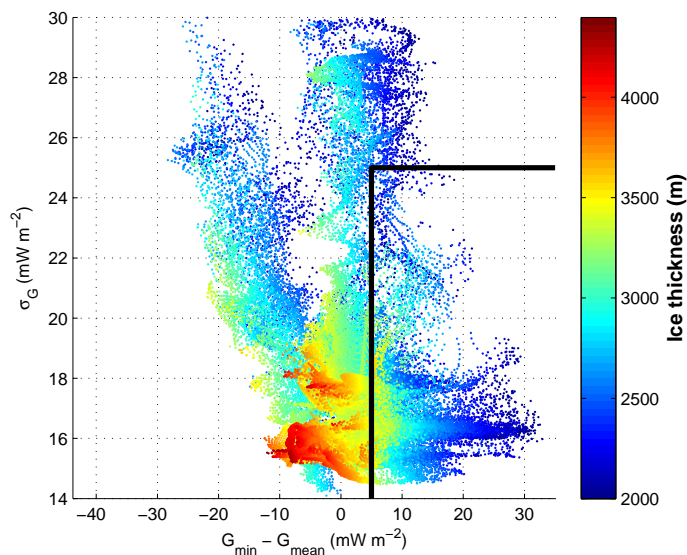


Fig. 3. Scatterplot of $\Delta G = G_{\min} - \bar{G}$ versus σ_G for all points with ice thickness $H > 2000$ m and horizontal flow speed $< 2 \text{ myr}^{-1}$. The colorscale depicts ice thickness for each of the grid points. Negative values of ΔG means that pressure melting point is reached, hence basal melt occurs. Positive values means that the minimum required heat flow to reach pressure melting point is higher than the mean of the three GHF datasets. Points lying within the rectangle are likely to be cold based, taken into account the variability of GHF.

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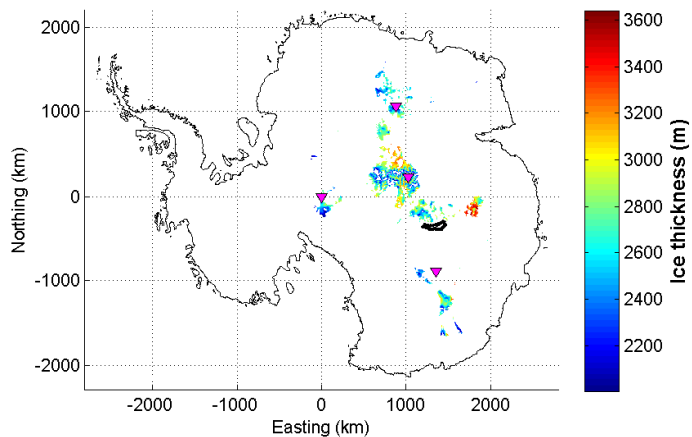



Fig. 4. Potential locations of cold basal conditions in areas with ice thickness $H > 2000$ m (colorbar) and horizontal flow speeds smaller than 2 myr^{-1} , for $\Delta G < 10 \text{ mWm}^{-2}$ and $\sigma_G < 10 \text{ mWm}^{-2}$, and as calculated with the simple model.

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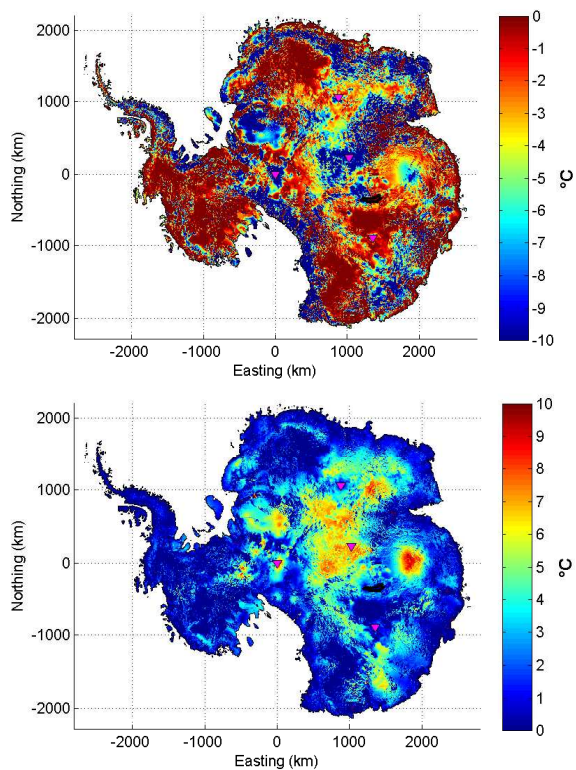


Fig. 5. Top: mean temperature according to the ensemble of 15 experiments (see text for more details), corrected for the dependence on pressure. The lower limit has been cut of at -10°C . Bottom: Root Mean Square Error (RMSE, $^{\circ}\text{C}$) according to the same ensemble.

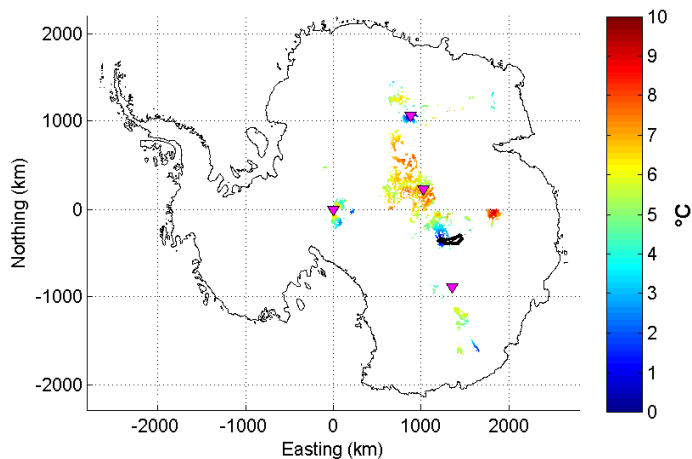


Fig. 6. Potential locations of cold basal conditions in areas with ice thickness $H > 2000$ m, horizontal flow speeds are smaller than 2 myr^{-1} and basal temperatures as calculated with the full model are lower than -5°C . The colorbar denotes the RMSE ($^\circ\text{C}$) based on the ensemble calculations.

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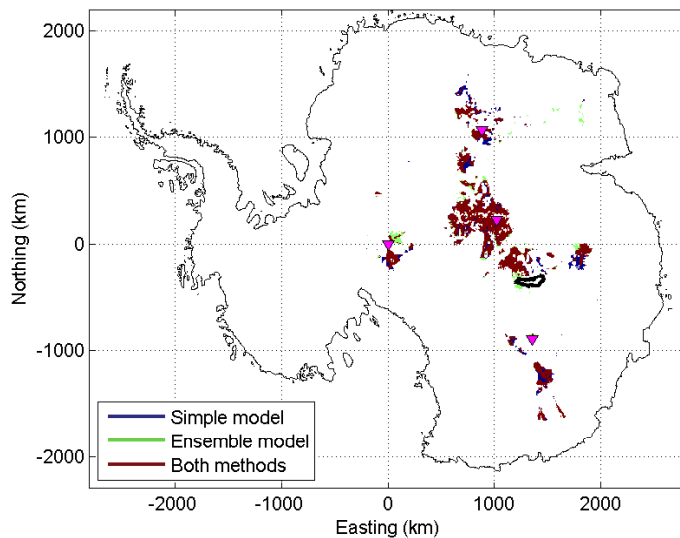


Fig. 7. Potential locations of cold basal conditions in areas with ice thickness $H > 2000$ m, horizontal flow speeds are smaller than 2 m yr^{-1} according to the simple model (depicted in Fig. 4) and the ensemble model (depicted in Fig. 6).

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