

1 **Links between MIS 11 millennial to sub-millennial climate variability and**
2 **long term trends as revealed by new high resolution EPICA Dome C**
3 **deuterium data – A comparison with the Holocene.**

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22 **Abstract**

23

24 We expand here the description of the Antarctic temperature variability during the long
25 interglacial period occurring ~400 thousand years before present (Marine Isotopic Stage, MIS
26 11). Our study is based on new detailed deuterium measurements conducted on the EPICA
27 Dome C ice core, Antarctica, with a ~50 year temporal resolution. Despite an ice diffusion
28 length reaching ~8cm at MIS 11 depth, the data allow to highlight a variability at multi-
29 centennial scale for MIS 11, as it has already been observed for the Holocene period (MIS 1).
30 Differences between MIS 1 and MIS 11 are analysed regarding the links between multi-
31 millennial trends and sub-millennial variability. The EPICA Dome C deuterium record shows
32 an increased variability and the onset of millennial to sub-millennial periodicities at the
33 beginning of the final cooling phase of MIS 11. Our findings are robust with respect to
34 sensitivity tests on the somewhat uncertain MIS 11 duration.

35 1. Introduction

36

37 Past interglacials, free from human impact on climate, are nowadays well documented by
38 long available climatic records such as marine sediment (Lisiecki and Raymo, 2005) or ice
39 cores (Jouzel et al., 2007; Loulergue et al., 2008; Spahni et al., 2005; Lüthi et al., 2008;
40 Siegenthaler et al., 2005) and offer the possibility to study the natural climate variability
41 during warm periods (Tzedakis et al., 2009). Improving the knowledge of their dynamics is
42 expected to provide a better understanding of the past and future evolution of our present
43 warm climatic period: the Holocene, whose natural course is disturbed by anthropogenic
44 forcings (IPCC, 2007). In this context, the challenge relies on finding the most appropriate
45 past interglacial for a comparison with the Holocene period. Occurring in an orbital
46 configuration close to the recent one (low eccentricity) around 400 thousand years before
47 present day (kyr BP, hereafter noted ka), Marine Isotopic Stage (MIS) 11 was proposed to be
48 a good candidate, according to a high correlation between the mean 65°N June insolations of
49 MIS 1 (or Holocene) and 11 (Loutre and Berger, 2000, 2003). Although the Antarctic
50 temperature derived from the EPICA Dome C (EDC) isotopic data (Jouzel et al., 2007)
51 exhibits values up to +2°C higher than the mean value for present day at MIS 11 maximum
52 (dated around 406 ka using the EDC3 chronology, Parrenin et al., 2007), CO₂ (Siegenthaler et
53 al., 2005) and CH₄ (Spahni et al., 2005) concentrations reach similar levels, around 280 ppm
54 and 710 ppb respectively, during MIS 11 and preindustrial period. Moreover recent sea-level
55 records (Bowen, 2010; Rohling et al., 2010) converge towards an estimation of sea-level at
56 ~400 ka comparable with the present one, as it was previously suggested by McManus et al.
57 (2003) and Waelbroeck et al. (2002) and modelled by Bintanja et al. (2005).

58 The climatic contexts of MIS 1 and 11 thus present interesting similarities. Controversies
59 have however emerged regarding the orbital alignment of the two interglacials with

60 implications for the prediction of the MIS 1 duration. The debate summarized in Tzedakis
61 (2010) arises from the choice of aligning either precession (Loutre and Berger, 2000, 2003;
62 Ruddimann, 2005, 2007) or obliquity (EPICA-community-members, 2004; Masson-Delmotte
63 et al., 2006) for the synchronisation of the two interglacials. A recent marine study (Dickson
64 et al., 2008) tends to support the alignment of obliquity; nevertheless, the double-peak
65 precession configuration of MIS 11 still contrasts with the orbital context of MIS 1, which is
66 marked by a single precession maximum. A careful comparison with earlier past interglacials
67 now appoints MIS 19 as the warm climatic period with the closest orbital configuration to the
68 Holocene one (Pol et al., 2010; Rohling et al., 2010; Tzedakis, 2010). The study of MIS 19,
69 remains, however, difficult due to age scale uncertainties as well as the lack of high resolution
70 records (Pol et al., 2010). Instead, MIS 11 offers the ability to document at high resolution
71 natural climate variability along the longest interglacial recorded since one million years ago
72 and the establishment of the 100 kyr glacial-interglacial cycles (Bintanja et al., 2005; Jouzel et
73 al., 2007; Lisiecki and Raymo, 2005).

74 While earlier comparisons of MIS 11 and Holocene focused on the analysis of their trends or
75 amplitudes (EPICA-community-members, 2004; Masson-Delmotte et al., 2006), we propose
76 here to analyse the Antarctic high frequency climate variability within these two periods. Our
77 study relies on new high resolution measurements of water stable isotope ratios
78 (deuterium/hydrogen ratio expressed as δD) conducted on the EDC ice core, which have
79 improved the temporal resolution for MIS 11 (Sect. 2). This new resolution, close to the
80 Holocene one (~ 20 years), allows us to better document and compare MIS 1 and 11 Antarctic
81 temperature fluctuations at sub-millennial scale into two different ways: (i) by studying the
82 δD record variance changes in relationship with long term trends, which was not possible with
83 the previous MIS 11 δD bag data because of the lack of a sufficient number of points; (ii) by
84 performing spectral analyses of our δD signals. Section 3 is dedicated to the description of the

85 methods before presenting the results (Sect. 4) and the variability analyses (Sect. 5) in detail.
86 Spectral analyses are highly dependent on the age-scale and on the estimation of MIS 11
87 duration. As differences have been reported between the EDC3 chronology (Parrenin et al.,
88 2007) used in Jouzel et al. (2007) and other age-scales (Kawamura et al., 2010; Lisiecki and
89 Raymo, 2005), we perform sensitivity tests for different MIS 11 durations, compatible with
90 orbital constrains derived from available and new EDC air data records, and discuss the
91 robustness of our variability analyses. At last, we investigate the possible mechanisms at the
92 origin of the observed δD variability during MIS 1 and 11 periods (Sect. 6).

93

94 **2. Material**

95

96 The EDC site in East Antarctica ($75^{\circ}06'$ S, $123^{\circ}2'$ E) has provided ~ 3260 m of ice core.
97 Derived from the measurements of 5800 samples coming from the continuous cut every 55
98 cm of the core (“bag samples”), a first long δD record unveiling ~ 800 ka of local temperature
99 variations has been provided (Jouzel et al., 2007). In central Antarctica, the climatic
100 information imprinted in surface snow stable isotope composition is affected by post
101 deposition processes such as firn diffusion and wind scouring. Detailed signal to noise studies
102 conducted at Vostok have shown an effective preservation of the climate signal at a temporal
103 resolution of ~ 20 years (Ekaykin et al., 2002). In EDC, the 55 cm sampling allows to
104 document Holocene climate variability at this temporal step of ~ 20 years (Masson-Delmotte
105 et al., 2004). But, due to ice thinning, it describes past interglacials at a lower temporal
106 resolution. A second cut of the EDC core providing 11 cm long samples, called “fine
107 samples” (Pol et al., 2010), increases the depth resolution by a factor of 5, thus improving the
108 temporal resolution for stable isotope records over past interglacials.

109 Referring to the threshold of -403‰ as an arbitrary definition of Antarctic warm periods
110 (related to the lowest 300 year average δD value observed over the past ~ 12 ka, EPICA-
111 community-members, 2004), the MIS 11 warm Antarctic phase is found in the depth interval
112 from ~ 2699 to 2779 m. This interval is dated between ~ 395 and 426.7 ka, according to the
113 official time-scale for the EDC core (EDC3 by Parrenin et al., 2007), with an uncertainty of
114 ~ 6 kyr on absolute ages and of $\pm 20\%$ on MIS11 duration (estimated at ~ 32 kyr). Here, we
115 have extended the study up to ~ 2694 m, thus covering a time interval from ~ 392.5 to 426.7
116 ka, in order to also depict the glacial inception. The MIS 11 available temporal resolution
117 derived from the δD bag data (Jouzel et al., 2007) has been contained between 170 and 300
118 years in the studied interval (the evolution of bag resolution with respect to depth is displayed
119 in Pol et al., 2010). With the new high resolution δD measurements conducted on 770 fine
120 samples, the stable isotope variability is now documented at a resolution ranged between ~ 35
121 and 60 years

122

123 **3. Methods**

124 3.1. Deuterium measurements

125

126 The method for deuterium analysis is the same as for the original bag samples measurements.
127 Water is reduced on uranium to form H_2 gas as described in Vaughn et al. (1998) for
128 measuring fine samples, including $\sim 30\%$ replicate measurements. Data are given with an
129 analytical accuracy of $\pm 0.5\text{‰}$ at 1σ . The coherency between bag and fine samples during the
130 MIS 11 period can be checked with the calculation of an average signal on 5 fine cut data.
131 The signal derived from the individual bag samples is statically less accurate (0.5‰ at 1σ)
132 than the average signal ($\pm 0.23\text{‰}$ at 1σ), as it benefits from an experimental noise reduced by

133 a factor of $\sqrt{5}$; this can be explained by the use of 5 measurements for its establishment
134 instead of one for the bag samples profile, over the same 55 cm depth interval.

135

136 3.2. Correction for isotopic diffusion

137

138 Water stable isotopes undergo firn and ice isotopic diffusion. After snow deposition, such
139 processes gradually smooth isotope profiles by removing the highest frequency climatic
140 information first during the firnification (Johnsen, 1975, or more recently Neumann and
141 Waddington, 2004) and then in the solid ice (Ramseier, 1967). In the upper part of the firn,
142 direct exchanges between snow water molecules and vapour, involving sublimation-
143 condensation processes, erase high frequency isotopic variations. In solid ice, smoothing
144 results from the temperature-dependent molecular diffusivity of water stable isotopes causing
145 self diffusion inside ice crystals.

146 Diffusion models can be applied to a given ice core to evaluate the smoothing of isotope
147 profiles using the diffusion length σ_l (characteristic length in cm of an ice layer affected by
148 the smoothing at a given depth), and to reconstruct the original amplitude of climatic
149 variations (back diffused signals). Here we use the Johnsen et al. (2000) method, with the
150 implementation of the parameters suitable for the EDC core (see Pol et al., 2010), on our new
151 δD data. A spectral analysis with respect to depth (cycles/m) of the high resolution signal is
152 performed and the associated red noise of the power spectrum is translated into a diffusion

153 length. This relies on the following equation $A = A_0 \cdot \exp\left(-\frac{1}{2} \sigma_l^2 \cdot k^2\right)$, which links the
154 amplitude of a given harmonic cycle A recorded in the data and altered by the diffusion
155 within the ice, to the initial amplitude A_0 (with σ the diffusion length and k the wave number
156 associated to the harmonic cycle). The empirical diffusion length at the MIS 11 depth,

157 estimated here to be ~8 cm according to our high resolution data, allows to reconstruct the
158 original amplitude of climatic variations recorded in the isotopic signal.

159 We cannot expect to preserve climatic information below ~20 years in central Antarctica
160 considering post depositional processes (Ekaykin et al., 2002), even though the bag samples
161 are supposed to describe the Holocene period at a better temporal resolution (from 8 years at
162 the top of the core to 18 years at ~12 ka depth). The isotopic diffusion occurring in the upper
163 part of the core only affects the part of the signal that is supposed to highlight periodicities
164 lower than 20 years. Therefore, the isotopic diffusion is considered insignificant for the
165 Holocene record studied here.

166

167 3.3. Variance analysis

168

169 To characterize the new information about MIS 11 climatic variability revealed by the high
170 resolution EDC δD data, we first resample our signals (original and back-diffused ones) on a
171 regular time-step. In the MIS 11 time interval dedicated to the variability study (see Sect. 5),
172 the lowest available temporal resolution is of ~50 years. The 50 year resampling is thus
173 imposed to avoid extrapolation. For Holocene δD , we have chosen to keep the 20 year time-
174 step in order to highlight all information than can be accessed at this optimal resolution. The
175 long term trends, calculated using a Singular Spectrum Analysis (SSA method) and
176 representing multi-millennial scale climatic variations, are then subtracted from the signals to
177 focus on the millennial to sub-millennial scale variability (<5000 years). The variance of the
178 signals is then described using a running standard deviation calculated on the detrended
179 signals, over 3 kyr from the past to the next 1.5 kyr at a given time point. This time period is
180 an arbitrary choice for describing the high frequency variability on a millennial scale; it is
181 constrained by the duration of MIS 1 and the temporal resolution available for MIS 11. Given

182 the normal distribution of the deuterium variability, the significance of variance changes can
183 be assessed using a Fischer F-test. Significance thresholds differ for the MIS 1 (149 degrees
184 of freedom over 3 kyr intervals) and MIS 11 (59 degrees of freedom). At the 95% confidence
185 level, ratios of standard deviations are significant when they are larger than 15% (MIS 1) and
186 22% (MIS 11). The main changes in variance described in Sect. 5.1 can therefore be
187 considered as significant. Other tests have been performed using shorter or longer reference
188 lengths without changing the principal features. In order to compare the variability with the
189 trends, we have calculated coherently with the running standard deviation a running average
190 of the trends over each 3 kyr interval.

191

192 3.4. Spectral analysis

193

194 To examine the frequency distribution of the isotopic record, we also perform spectral
195 analyses of our resampled and detrended original δD signals. The same analysis on MIS 11
196 back-diffused signal will not be discussed as it does not provide any supplementary
197 information on the power spectra. Here we use the wavelet analysis method particularly well
198 adapted for describing non-stationarities or changes in frequency and magnitude (Torrence
199 and Compo, 1998), common characteristics of climatic records. This method (mathematical
200 formalization described in Mudelsee, 2010) is used to decompose, for different exploratory
201 scales, a signal in a sum of small wave functions of finite length that are highly localized in
202 time, unlike the classical Fourier transform which explores the complete length of the signal
203 and separates it into infinite-length sine-wave functions. Resulting in a loss of time
204 information, the Fourier analysis thus fails to detect the time variable statistical properties of
205 stochastic processes.

206 To avoid edge effects and spectral leakage produced by the finite length of the time series,
207 these last ones were zero-padded to twice the data length. But, this leads to underestimate the
208 lowest frequencies near the edges of the spectrum. It is thus necessary to assess the areas
209 (known as cones of influence) which delineate the parts of the spectrum where estimated
210 energy bands are likely to be less powerful than they actually are. For all local wavelet
211 spectra, Monte Carlo simulations are then used to assess the statistical significance of peaks.
212 The background noise for each signal is firstly separated and estimated using singular
213 spectrum analysis. Secondly an autoregressive simulation is performed for each noise time
214 series to determine the AR(1) stochastic process against which the initial time series has to be
215 tested. The estimated power spectra are here tested against a background red noise
216 (AR(1)=0.7); the confidence levels are taken above 99%, consistently with the recommended
217 level of $1-1/(1-n)$ (Thomson, 1990), where n is the number of points in the time interval of
218 interest (580 and 570 for MIS 1 and 11 respectively considering the resampled signals).

219

220 **4. Results**

221

222 Figure 1 (panel a) displays the results from our new high resolution δD measurements (grey)
223 for MIS 11 depths (from 2694 to 2799m), confronted to the initial low resolution δD signal
224 (black), published in EPICA-community-members (2004) and in Jouzel et al. (2007). It is
225 important to notice here that the new high resolution data confirm the patterns originally
226 exhibited by the black curve. An abrupt event - also exhibited by the CO_2 (Spahni et al., 2005)
227 and CH_4 (Loulergue et al., 2008) records - is observable at ~ 2775 m. It is followed by a slow
228 warming and then a small decrease between depths of ~ 2770 and 2750 m, the maximum of
229 δD being reached at ~ 2735 m before the beginning of the cooling phase. The comparison over
230 the same 55 cm depth interval between the “calculated” signal (Fig. 1b, grey) obtained from

231 the average of 5 high resolution data (see Sect. 3.1) and the previous low resolution profile
232 (Fig. 1b, black) shows a good agreement over the full period, within their respective accuracy
233 (refer to the Sect. 3.1), and confirms the robustness of the measurements. Differences of 1‰
234 on average and up to 2‰ can be nevertheless observed during the warming phase, over depth
235 intervals ranging between 2740 and 2746 m or 2760 and 2766 m. In parallel, the new data
236 have been used for the calculation of the EDC δD -excess during MIS 11. The quality of the
237 new high resolution δD measurements is confirmed by stronger correlation with $\delta^{18}O$ bag data
238 (B. Stenni, pers. comm) over the MIS 11 period ($R^2=0.98$ using the average of the detailed δD
239 data versus 0.97 using the initial bag δD data), as well as by the smaller dispersion with
240 respect to this linear regression (~ 0.8 versus 1.1‰). This comforts us in the reliability of our
241 new deuterium measurements.

242 Thereafter, we focus on the added information brought by the high resolution δD data, the
243 signal being examined with respect to time. The reference time-scale for the EDC core is the
244 EDC3 chronology established by Parrenin et al. (2007). Figure 1c displays the high resolution
245 δD signal on the corresponding time interval from ~ 392 to ~ 427 ka. The signal corrected for
246 isotopic diffusion is shown in grey and exhibits increased variability up to 2 ‰.

247

248 **5. Variability analysis: MIS 1 and 11 comparison**

249

250 Our new δD data for MIS 11 now enable a detailed comparison of climate variability below
251 millennial scale during MIS 1 and 11, referring to comparable temporal resolutions of ~ 50
252 years and ~ 20 years respectively. Holocene is scrutinized from the present day to 11.7ka.
253 (beginning of the plateau just after the Antarctic Cold Reversal, Jouzel et al., 1995, 2001);
254 MIS 11 from 397 to 421 ka (Fig. 1c, grey area), to avoid the difficulty of correctly capturing
255 both the first abrupt event around ~ 425 ka and the abrupt cooling after 397 ka using the SSA

256 method. Panels a of Fig. 2 display the resampled δD signals (initial ones in black and MIS 11
257 back diffused one in grey) and the calculated trends (red) over the two interglacials. The
258 signals are centred on their respective mean value, $\sim -396\text{‰}$ for Holocene and $\sim -391\text{‰}$ for
259 MIS 11, with a δD difference of 5‰ corresponding to a $\sim 0.8^\circ\text{C}$ temperature gradient
260 according to the modern spatial slope of 6‰ per $^\circ\text{C}$ (Masson-Delmotte et al., 2008) in East
261 Antarctica. While this modern spatial slope is of current use for interpreting glacial-
262 interglacial changes (Jouzel et al., 2003, 2007), the magnitude of the variability for present-
263 day or warmer conditions may be underestimated (by typically 30%), as suggested by isotope
264 modelling studies for present day interannual variability (Schmidt et al., 2007) or for
265 projections towards a warmer CO_2 world (Sime et al., 2008).

266 The comparison of the long term trends first allows to characterize the corresponding multi-
267 millennial climate variability (>5000 years) over the two focused periods and to depict two
268 different evolutions. The Holocene signal exhibits two successive plateaus, one between 10
269 and 11.7 ka characterized by a δD anomaly of $\sim +4\text{‰}$ and the second one from the present
270 day to 5.5 ka, at Holocene mean level. In contrast, MIS 11 presents a slow δD increase
271 between 413 and 421 ka followed by a rapid warming that reaches a δD optimum at ~ 407 ka
272 of $\sim 7\text{‰}$ above MIS 11 mean value level, before finally entering its cooling phase. The
273 relationship between these long term trends and the high frequency climate variability of our
274 two interglacials is then documented by a variance and a spectral analysis, following the
275 methods described in Sects. 3.3 and 3.4.

276

278

279 By subtracting the red signal from the black one (Fig. 2 panels a), one gains access to the
280 millennial to sub-millennial scale variability as represented on panels b of Fig. 2. The
281 remaining signal gives information on the amplitude of variations characteristic of both
282 periods. Even after the correction for isotopic diffusion (see Sect. 3.2), the MIS 11 δD signal
283 characterized by a maximal amplitude of variations of $\sim 7\text{‰}$ does not reach the level of
284 variability exhibited during Holocene (up to $\sim 10\text{‰}$). This difference partly arises from the
285 Holocene temporal resolution more than twice higher than for MIS 11. When resampling the
286 Holocene signal every 50 years as done for MIS 11, the amplitude of MIS 1 variation is
287 reduced to 8‰ (not shown).

288 In order to go beyond the problem of variability levels, we compare the evolution of sub-
289 millennial climate variability (Fig. 2 panels c) by calculating a running standard deviation
290 over 3 ka of the panels b signals (see Sect. 3.3). The lower variability for MIS 11 is again
291 clearly depicted with values oscillating around 2‰ (up to 2.5‰ after back diffusion
292 correction in grey) against 3.5‰ for MIS 1. For the description of the variability evolution,
293 the noticeable points of standard deviation slope changes are labelled by letters ordered from
294 the past to the present. Holocene variability first shows a progressive increase of $\sim 0.6\text{‰}$ from
295 point 1.A to 1.B. Then the variability decreases (until the point 1.C) before reaching a quite
296 stable level (albeit with a weak increase during the last 5 ka). The MIS 11 pattern of
297 variability is characterized by a non-stable region followed by a quickly increasing variability
298 (by 1‰ from 11.A to 11.B) with a maximal running standard deviation value of $\sim 2.5\text{‰}$ hold
299 until the point 1.C. The variability then decreases by 0.5‰ (from 11.C to 11.D) before
300 progressively increasing again at the end of MIS 11. Except for the overall level of variability,

301 the pattern remains unchanged when taking into account isotopic diffusion (grey curve).
302 Therefore only the original signal (in black) is discussed in the rest of the study.
303 These changes of millennial to sub-millennial-scale variability can be linked to the long term
304 trend by plotting (panels d Fig. 2) the running standard deviation with respect to a running
305 average over 3 kyr of each interglacial δD signal trend. This approach highlights the
306 progressive increase of Holocene variability (from 1.A to 1.B) occurring during the cooling
307 phase between the two plateaus. In contrast, its decrease until 1.C is linked to the slow Mid-
308 Holocene warming. For MIS 11, the noticeable increase of variability between 11.A and 11.B
309 begins just before the maximum of the δD signal. The highest value of the standard deviation
310 is maintained stable during the beginning of the cooling phase. After an abrupt decrease (from
311 11.C to 11.D), the variability keeps increasing during the final cooling phase at the end of
312 MIS 11. Despite a symmetrical aspect of the δD trend on each side of the MIS 11 optimum
313 (comparable increasing and decreasing trends), the sub-millennial variability exhibits a clear
314 shift between the warming and the cooling phases. MIS 11 presents thus a higher level of
315 variability during all the cooling phases, even after its abrupt decrease (from 11.C to 11.D).
316 This feature is comparable to Holocene increasing variability observed during the short
317 cooling between its two plateaus. This highlights a difference in terms of climate dynamics
318 between cooling and warming phases.

319

320 5.2. Spectral Analysis

321

322 We performed a spectral analysis of the detrended signals (displayed on Fig. 2 panels b,
323 black) for each interglacial focus period, using a wavelet analysis (see Sect. 3.4). The
324 difference of temporal resolution between MIS 1 and 11 implies a different available range of
325 frequencies for our two interglacial periods (25 to 10 kyr⁻¹ for MIS 1 and 11 respectively

326 corresponding to 40 and 100 year cycles). For the present comparison we focus on multi-
327 centennial variability which is accessible with the 50 year resolution of MIS 11 data. Due to
328 the diffusion (characterized by a ~ 8 cm diffusion length at MIS 11 depth, see Sect. 3.2) all the
329 periodicities under ~ 130 years are lost in the spectral signal. Figure 3 displays the time
330 continuous spectra of the two interglacials.

331 We can first observe a millennial to multi-centennial scale variability for both interglacials.
332 Holocene is marked by significant periodicities from 90 to ~ 300 years punctually present over
333 the full period. In contrast, the multi-centennial scale variability is not present persistently
334 over the full MIS 11 spectrum but appears from ~ 406 ka (based on EDC3 chronology), as
335 revealed by the highlighted significant periodicities ranged between ~ 180 to ~ 500 years
336 (highlighted I Fig. 3 with a dashed contour). The establishment of this multi-centennial
337 variability is then followed by the occurrence of a ~ 1400 year significant periodicity. MIS 11
338 thus presents a transition in its variability pattern coinciding with the beginning of the long
339 term cooling phase (Fig. 2a). The previous observation of high amplitudes of variations
340 between 11.B and C points (see Sect. 5.1) can thus be attributed to the onset of multi-
341 centennial variability from ~ 406 ka. Then, the variability increasing again after the 11.D point
342 seems to be more strongly expressed at millennial scale. Altogether, these results highlight the
343 establishment of a new mode of climatic variability during the final cooling phase of MIS 11,
344 as firstly noted in the variance analysis. By comparison, a small transition in the MIS 1
345 millennial scale variability can be detected at ~ 5.5 ka with a ~ 950 year periodicity changing
346 into a ~ 800 year cycle. This corresponds to the beginning of the second plateau of Holocene
347 (see Fig. 2 panel a) which is also marked by amplitudes of variations progressively decreasing
348 (segment 1.B to 1.C) before reaching a stable level of variability (Sect. 5.1).

349

350 5.3. Sensitivity to uncertainties on MIS 11 duration

351

352 The previous comparison between MIS 1 and 11 climate variability features is constrained by
353 their durations, here based on the EDC3 chronology (Parrenin et al., 2007). Referring to the -
354 403‰ level for the definition of an interglacial in the EDC core (see Sect. 2, EPICA-
355 community-members, 2004) the EDC3 age-scale estimates that MIS 1 has lasted ~12 kyr so
356 far, consistently with the new EDC chronology established back to 50 ka (Lemieux-Dudon et
357 al., 2010), derived from a new inverse method for ice core dating (Lemieux-Dudon et al.,
358 2009). The estimate for MIS 11 duration is of ~32 kyr with a given uncertainty of $\pm 20\%$ (\pm
359 6.4 kyr). With an added difficulty of comparing different type of records that use different
360 references for delineating the interglacial periods, the Lisiecki and Raymo (2005) chronology
361 for marine sediment cores establishes the beginning of MIS 1 at ~11ka and evaluates the
362 benthic MIS 11 duration of 20 kyr between 398 and 418 ka (with an uncertainty of 4 kyr on
363 absolute ages), suggesting a shorter MIS 11 than depicted in EDC3 for Antarctic temperature.
364 In parallel, Kawamura et al. (2010) have recently extended an orbital $\delta\text{O}_2/\text{N}_2$ chronology for
365 the Antarctic Dome Fuji core, first established for the past 360 kyr (Kawamura et al., 2007),
366 back to 470 ka. They obtain a MIS 11 duration shorter by ~9 kyr in comparison with the
367 EDC3 chronology. Such differences between MIS 11 length estimates lead us to perform
368 sensitivity tests on MIS 11 duration and evaluate the impacts on MIS 11 climate variability
369 using the same variance and spectral analyses as before (see Sect. 5).

370 Considering O_2/N_2 ratios measured in trapped air of ice cores as direct tools (free from the
371 usual Δ age between gas and ice) for the establishment of orbital tuning chronology (Bender,
372 2002), we have used EDC O_2/N_2 data (Landais et al., 2010) to produce two others age scales
373 for MIS 11 (Fig. 1d and e), in a different way than a simple linear compression of MIS 11
374 duration. By synchronising arbitrary the mean 75°S insolation curve to the EDC O_2/N_2 record
375 (as done in Landais et al., 2010), the first age-scale obtained shortens our MIS 11 signal by ~4

376 kyr, now dated between 395.2 and 425.7 ka (Fig. 1d, hereafter Test1). Such a reduced
377 duration was also tested in Rohling et al. (2010) and remains within the uncertainty range of
378 the EDC3 time-scale. A second age-scale test can be then produced by fitting the EDC3 with
379 the orbital Dome Fuji chronology, reducing in larger proportions the MIS11 duration by ~8.5
380 kyr, between 400.5 and 426 ka (Fig. 1e, hereafter Test2).

381 Focusing on the same part of the MIS 11 signal as in Sect. 5 (Fi. 1d and e, grey areas), the
382 variability is then analysed on the corresponding reduced time-intervals ranging from 399 to
383 421.3 ka and from 403.7 to 420 ka for Test 1 and 2 (Fig. 2). The impacts of Test 1 on MIS 11
384 variability analysis are globally negligible compared to the previous results obtained with the
385 EDC3 age-scale. The amplitude of variations (panel b) is hardly affected; the running
386 standard deviation (panel c) and its evolution with respect to trend (panel d) exhibit an overall
387 similar pattern as the one produced using the EDC3 age-sale. By contrast, the significant
388 shortening of the MIS 11 duration by Test 2 drops the level of variability by 1.5‰ (panel b).
389 The running standard deviation is then affected and presents values of 0.3‰ lower on average
390 than using the original signal (panel c). Its evolution with respect to the trend is impacted with
391 less pronounced changes of variability levels (between 11.A and 11.B, 11.C and 11.D, panel
392 d). But, both for Test 1 and Test 2, the panels d illustrate that the main features described in
393 Sect. 5.1 remain unchanged, showing that the age-scale uncertainties do not affect our main
394 conclusions regarding changes in variance.

395 In parallel, the spectral analysis is obviously affected by dating uncertainties, with again
396 larger impacts when applying Test 2 compared to the use of Test 1 (Fig. 3). Hence, the MIS
397 11 spectrum obtained by the application of Test 1 first shows a loss of the ~200-year
398 periodicity previously highlighted with the EDC3 chronology at ~406ka; the ~1400-year one
399 presents a value slightly shifted to ~1330 years and becomes also more pronounced at the end
400 of the studied period. Second, Test 2 implies fewer significant periodicities at the multi-

401 centennial scale and make the ~1400-year periodicity value turned into ~950 years.
402 Altogether, our sensitivity tests still exhibit the same pattern of variability and confirm the
403 robustness of a changing climate dynamics at the onset of the final MIS 11 cooling phase.

404

405 **6. Discussion**

406

407 We now discuss the links between the variability features highlighted in the δD signals of
408 MIS 1 and 11, natural climate forcings and internal climate variability. While long term
409 changes have classically been attributed to the climate system response to orbital forcing (as
410 firstly hypothesised by Milankovitch, 1941), the drivers of millennial to sub-millennial
411 variability involve external forcings such as solar and volcanic activities, as well as internal
412 climate dynamics implying the oceanic and atmospheric components (as further discussed in
413 the following sub-sections). In particular one can question the influence of local processes
414 such as precipitation intermittency, moisture origin, evaporation conditions in relationship
415 with atmospheric circulation and austral ocean surface conditions on Antarctic δD records.

416 As the Holocene benefits from a substantial documentation, we first discuss the results of
417 MIS 1 spectral analysis in the context of available literature. Assuming that the patterns of
418 forcings and internal modes of variability described over the last 12 kyrs were also at play
419 during MIS 11, we can then suggest that the same mechanisms are involved during MIS 11.
420 Due to the uncertainties on MIS 11 duration which impact the significant periodicities
421 highlighted in our high resolution δD data, analogy between MIS 1 and 11 climate forcings
422 remains, however, difficult to establish. The discussion is thus limited to the comparison of
423 the general evolutions of MIS 11 δD signal and other climate records from different proxies
424 available for the MIS 11 period. .

425

426 6.1. Spectral Holocene EDC characteristics

427

428 By examining the solar activity during Holocene (as detailed in Steinhilber et al., 2009), we
429 first note that our EDC δD Holocene variance (Fig. 2, panel c) cannot simply be explained by
430 changes in solar forcing and deserves the exploration of other ones. Spectral analyses of the
431 EDC δD signal during Holocene have already been performed (Yiou et al., 1997; Masson et
432 al., 2000), but without clearly examining the possible relationships with climate actors. Here
433 the wavelet method presents the advantage to mark the onset of the significant periodicities of
434 the Holocene EDC δD signal. Then they can be compared to the results of many previous
435 studies that have discussed millennial to multi-centennial Holocene variability and its
436 signature in different climate and solar activity records.

437 The implication of solar forcing in the millennial scale variability of Holocene has been first
438 supported by Bond et al. (1997, 2001). Referring to cosmogenic nucleides (^{14}C and ^{10}Be)
439 measurements, they indeed claimed that the Holocene 1500-year (± 500) periodicity found in
440 North Atlantic drift ice records can be attributed to solar forcing. The same periodicity has
441 also been detected in other proxies of the North Atlantic region (Bianchi and McCave, 1999;
442 Campbell et al., 1998; Mayewski et al., 1997 as a non-exhaustive list), as in the South
443 Hemisphere (Lamy et al., 2001), but without underlining a persistent link with solar activity.
444 Using a wavelet analysis method on several records from the North Hemisphere, Debret et al.
445 (2007) actually highlight a decoupling of the apparent Holocene 1500-year climatic cycle into
446 three superimposed significant periodicities of 1000, 1500 and 2500 years. Whereas the
447 comparison of different marine sediment cores (Bianchi and McCave, 1999; Chapman and
448 Shackleton, 2000; Giraudeau et al., 2000) preferentially attributes the ~ 1500 -year periodicity
449 to oceanic dynamics, the link between 1000 and 2500-year climatic cycles and solar activity

450 is confirmed when confronting the spectral analyses of Bond et al. (2001) records and of the
451 Vonmoos et al. (2006) ^{10}Be data.

452 Exhibiting cyclicities close to ~ 1000 years, our EDC Holocene δD record (Fig. 3) could also
453 corroborate the solar forcing at millennial scale in the Antarctic region. But, the spectral
454 analysis actually highlights a decoupling of the ~ 1000 -year cycle into a ~ 950 -year periodicity
455 during early Holocene (similar to the Northern records of Bond et al., 2001 and Vonmoos et
456 al., 2006 according to the Debret et al., 2007 analysis and close to the 900-year cycle of Lamy
457 et al., 2001 in the South), to a ~ 800 -year one in the late Holocene (also recorded in Chapman
458 and Shackleton, 2000). This transition phase in the frequency domain, recorded at ~ 5.5 ka in
459 our EDC δD signal and phasing with the establishment of a progressive stable level of
460 variability (see Sect. 5.2), was documented in Debret et al. (2009) and Wirtz et al. (2010), in a
461 synthesis of records covering the two Hemispheres. Known as Mid-Holocene transition, it is
462 first suggested by Debret et al. (2009) to underline a change in the dominant mechanisms of
463 variability, from an external origin (essentially from solar activity) in the early Holocene, to
464 internal processes in the late Holocene. This hypothesis is in line with other studies (Wanner
465 et al., 2008; Wirtz et al. 2010), which noticed a lower magnitude of solar variability in the
466 early Holocene compared to the mid to late Holocene. At the same time, Wirtz et al. (2010)
467 observe the emergence of a more pronounced variability at centennial scale (between 200 and
468 850 years), after this Mid-Holocene transition. Our signal does exhibit a 290-year periodicity
469 (Figure 3) in the early Holocene. Thus, the establishment of a new mode of variability during
470 the mid Holocene also has to be explored at multi-centennial scale. This leads us to now
471 discuss the possible mechanisms (internal and/or external) at play in the second part of
472 Holocene that could explain the identified multi-centennial variability.

473 The multi-centennial variability recorded in the Holocene EDC δD signal is a common feature
474 of both North and South Hemisphere records. In particular our periodicity of ~ 290 years is

475 close to the ~250-year one found in a East Antarctica marine core (Crosta et al., 2007) at ~2.5
476 ka and to the ~240-year cycle of the Rouse et al. (2006) data from a North Icelandic marine
477 sediment core, identified at ~3 ka. Climate models (Park and Latif, 2008; Schulz et al., 2007)
478 show that the multi-centennial variability is a persistent feature of Atlantic Ocean circulation.
479 Park and Latif (2008) have demonstrated the implication of both hemispheres in high
480 frequency variability through large changes in the Atlantic sea ice extent, with a rapid
481 response of the North Hemisphere at decadal scale and a slower one of the South Hemisphere
482 at multi-centennial scale.

483 Invoking a sun-ocean-climate linkage, Hu et al. (2003) underline the possible forcing of the
484 multi-centennial changes in sea ice extent by the centennial solar forcing (Karlén and
485 Kuylenstierna, 1996). The same forcing is further proposed by Varma et al. (2010) to drive
486 the southern annular mode. Altogether, these studies suggest that the multi-centennial scale
487 variability found in the EDC δD record could be closely associated with changes in austral sea
488 ice extent and atmospheric circulation, in response to multi-centennial variations in solar
489 activity. Changes in volcanic forcing may also be at play (Castellano et al., 2005), as recent
490 modelling studies suggest a possible centennial response time (Stenchikov et al., 2009;
491 Schneider et al., 2009) but have not so far been explored due to the lack of quantitative
492 reconstructions beyond the last millennium (Gao et al., 2008).

493 Nevertheless, centennial variability may not necessarily be driven by external forcings and
494 may also result from modes of internal climate variability. Modelling experiments focusing
495 on North Atlantic Deep Water (NADW) formation (Jongma et al., 2007; Renssen et al., 2007)
496 have indeed highlighted that internal periodic processes such as freshwater releases could
497 provide a sensible mechanism to explain Holocene multi-centennial scale variability.
498 Focusing on the thermohaline structure of the Southern Ocean, Pierce et al. (1995) also linked

499 modelled centennial-scale oscillations with changes both in the local precipitation affecting
500 the Antarctic Circumpolar Current and in the NADW.

501 Thus, while bipolar see-saw patterns are well known to link Antarctica and Greenland stable
502 isotope records during abrupt glacial (Blunier et al., 1998) or early interglacial events
503 (Masson-Delmotte et al., 2010), even at sub-millennial scale (Capron et al., 2010), these
504 modelling experiments reinforce the hypothesis of a similar interhemispheric linkage at play
505 at multi-centennial variability during interglacial periods. Such internal mechanisms could be
506 involved in the observed variability changes at Mid-Holocene transition, as previously
507 suggested by Debret et al. (2009).

508

509 6.2. MIS 11 EDC variability

510

511 Here the spectral analysis of MIS 11 cannot be discussed in the same way as for MIS 1, due to
512 large uncertainties on MIS 11 duration and the absence of information about external forcings
513 (solar and volcanic activities) for this period, The MIS 11 still benefits from sufficient
514 documentation to allow comparisons between our new EDC δD profile and “other” proxy
515 signals.

516 In addition to the various climatic information provided by the EDC core (e.g. Jouzel et al.,
517 2007; Siegenthaler et al., 2005; Spahni et al., 2005), MIS 11 has been documented in other
518 long marine or continental records through different proxies (Lisiecki and Raymo, 2005;
519 McManus et al., 2003; Tzedakis et al., 2006). They consistently underline the general
520 comparable background climate conditions between MIS 1 and 11 (e.g. sea level, greenhouse
521 gas concentrations, local temperatures, vegetation history) and the exceptional length of MIS
522 11. One study (de Vernal and Hillaire-Marcel, 2008) emphasizes an exceptional development
523 of boreal ecosystems on the Greenland coasts, suggesting particularly reduced Greenland ice

524 sheet extent during this interglacial. Due to the lack of a sufficient temporal resolution for
525 performing reliable spectral analysis, comparisons with these records remain however,
526 restricted to the analyses of trends or intensities.

527 Still, similarities between the EDC CO₂ (Siegenthaler et al., 2005) and the 500-year resolution
528 δ¹³C record of a marine core from the Cape Basin (Dickson et al., 2008) at the end of MIS 11
529 have revealed an interesting oceanic circulation - atmospheric CO₂ concentration linkage. The
530 parallel between the observed δ¹³C gradient and CO₂ drawdown at the end of MIS 11 supports
531 the hypothesis of a close link between deep austral ocean ventilation and changes in
532 atmospheric greenhouse gas concentrations (Toggweiler, 1999; Hodell et al., 2003). The onset
533 of an increasing variability in our δD record (at ~406 ka) does not coincide with any marked
534 change in the CO₂ concentration. Its phasing with a methane concentration starting to
535 decrease (Loulergue et al., 2008) and the increase of the EDC sea salt sodium flux (Wolff et
536 al., 2010) is however robust within age-scale uncertainties. It suggests that the observed
537 increase in EDC δD variability at the end of MIS 11 occurs in parallel with: first, an East
538 Antarctica cooling trend; second, an extent of austral sea ice cover associated to a reduced
539 methane production in tropical and boreal wetlands.

540 Further discussions about the links between the EDC climate variability (derived from our δD
541 data) and ocean circulation variability requires both higher resolution marine records and
542 improved chronologies and synchronization methods. But, as a first step in the MIS 11
543 variability analysis, our data enable to suggest that the increased Antarctic variance and the
544 onset of millennial to sub-millennial variability are intimately linked with the global transition
545 between interglacial and glacial states.

546 7. Conclusion

547

548 Our δD measurements conducted on high resolution EDC samples have first confirmed the
549 patterns of East Antarctica temperatures over the full MIS 11 period, as previously described
550 by the original δD bag record. Then, our study has aimed to demonstrate the added value of
551 analysing EDC high resolution δD data for first, improving the documentation of past
552 interglacial climate variability, going beyond trend and intensity considerations, and second,
553 permitting a comparison with Holocene.

554 Our results highlight a specific variability pattern during MIS 11, with two distinguishable
555 evolutions on each side of the late MIS 11 maximum (~ 406 ka according to the EDC3
556 chronology). Indeed, the MIS 11 signal is characterized by a variability enhanced from the
557 beginning of the cooling phase, which contrasts with the lower variability exhibited during the
558 preceding warming phase. Moreover, a spectral analysis allows us to relate these MIS 11
559 variability features with the onset at ~ 406 ka of new climatic dynamic modes marked by the
560 emergence of periodicities at millennial to multi-centennial scales.

561 The Holocene signal exhibits a similar pattern with increasing variability occurring just after
562 the early Holocene plateau and persistent during the following decrease in temperatures.

563 Unlike for MIS11, this change in variance is not evidenced in the spectral analysis. The Mid-
564 Holocene transition, dated at 5.5 ka and documented in many previous studies, is still
565 imprinted in the obtained spectrum and characterized by a shift in the significant periodicities.

566 The links between Holocene variability changes on each side of this transition and external
567 forcings or internal climate system responses can be explored, thanks to the limited
568 uncertainties (~ 100 - 200 years) on Holocene EDC dating and the substantial available
569 documentation. Such discussion is however impossible for the MIS 11 signal, because of the
570 lack of records with sufficient resolution, the lack of documentation of natural forcing

571 variability, and because of the large age-scale uncertainties attached to the MIS 11 duration.
572 While our results about MIS 11 variability patterns are robust with respect to these
573 uncertainties, the length of MIS 11 impacts in a larger proportion the values of periodicities
574 revealed by the spectral analysis. It thus prevents to clearly attribute the increasing variability
575 at the glacial inception to one or other climatic component.

576 Consequently, we stress the need to: first scrutinize the MIS 11 variability with other records
577 e.g. from tropical, temperate and polar regions at sufficient temporal resolution for improving
578 the global documentation of changes in variability along the full period; second, reduce
579 uncertainties on the length of this interglacial by the building of an accurate reference time-
580 scale for the EDC core. It will help in the future to precisely specify the MIS 11 variability
581 spectrum, but also the one of other past interglacials. New detailed isotopic measurements
582 from the EDC core are indeed now available for a variety of interglacials and will allow to
583 further explore the relationships between mean state and Antarctic climate variability under
584 contrasted orbital contexts.

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586

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930 **Figure Captions**

931

932 **Figure 1:** Summary of the available EDC δD (‰) data for MIS 11 first plotted in function of
933 depth (panels a and b, in m) and then in function of time (panels c to e, in ka). (a) δD from
934 MIS 11 bag samples (black) and the new high resolution δD signal from the fine samples
935 (grey). (b) MIS 11 δD bag samples (black) in comparison with the mean signal (grey)
936 obtained from the average of 5 fine cuts. (c) High resolution δD data (black) plotted with
937 respect to the official EDC chronology (EDC3, Parrenin et al., 2007) and corrected from the
938 isotopic diffusion (grey). Panels (d) and (e) display the same signals using two other age-
939 scales (section 5.3) (Test 1 and Test 2 respectively) for sensitivity tests on MIS 11 duration.

940

941 **Figure 2:** Variability analysis (in ‰) of MIS 1 (top), MIS 11 using EDC3 chronology
942 (middle top), Test 1 age-scale (middle bottom) and Test 2 age-scale (bottom). (a) Signals
943 (black) centred on the δD mean value of the focused period: 0-12 ka for MIS 1; 397-421 ka
944 for MIS 11 – EDC3; 399-421.3 ka for MIS 11 – Test1; 403.7-420 ka for MIS 11 – Test 2. The
945 general trends are plotted in red; signals corrected from isotopic diffusion in grey for MIS 11.
946 (b) Signals minus their respective trends (red, panels a). (c) Calculated running standard
947 deviation of panel b signals over 3 kyr, from the past 1.5 kyr to the next 1.5 kyr at a given
948 time point (black: original signals; grey: correction for isotopic diffusion). Remarkable
949 changes of slope are labelled from A to C or D from past to present (the 1 or 11 numbers refer
950 to the studied interglacial). (d) Running standard deviation (panels c, ‰) plotted in function
951 of the respective general trends (panels a, red, ‰). Signals are smoothed using a binomial
952 algorithm for an easier readability. The labelled points (panels c) are reported and arrows
953 indicate the way of reading from past to present.

954

955 **Figure 3:** Spectral analysis of the detrended signals (displayed on Fig. 2, panels b) for MIS 1
956 (left) and MIS 11 (right), using EDC3 chronology (top), Test 1 age-scale (middle) and Test 2
957 (bottom). The spectral power is displayed in function of time (ka) in term of frequency (1/kyr,
958 left axis) or period (kyr, right axis). Black lines correspond to the cone of influence; dot lines
959 indicate the significant periodicities (application of the statistical test of Torrence and Compo:
960 <http://paos.colorado.edu/research/wavelets/>).