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2 submitted in December 2015

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5 **MARKED-UP MANUSCRIPT**

6 blue: changes according to Reviewer #1

7 red: changes according to Reviewer Revel

8 green: other changes

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11 **A distal—~~145 ka~~ 140 ka sediment record of Nile** 12 **discharge and East African monsoon variability**

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15 **W. Ehrmann¹, G. Schmiedl², M. Seidel¹, S. Krüger¹, and H. Schulz³**

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17

18 ¹ Universität Leipzig, Institut für Geophysik und Geologie, Talstraße 35,
19 D-04103 Leipzig, Germany

20

21 ² Universität Hamburg, Centrum für Erdsystemforschung und Nachhaltigkeit,
22 Bundesstraße 55, D-20146 Hamburg, Germany

23

24 ³ Universität Tübingen, Fachbereich Geowissenschaften, Hölderlinstraße 12,
25 D-72074 Tübingen, Germany

26

27

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29 Correspondence to: W. Ehrmann (ehrmann@rz.uni-leipzig.de)

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31

1 **Abstract**

2

3 Clay mineral assemblages in a sediment core from the distal Nile discharge plume off
4 Israel have been used to reconstruct the late Quaternary Nile sediment discharge
5 into the Eastern Mediterranean Sea (EMS). The record spans the last ca. 140 ka 145
6 ka. Smectite abundances indicate the influence of the Blue Nile and Atbara that have
7 their headwaters in the volcanic rocks of the Ethiopian highlands. Kaolinite
8 abundances indicate the influence of wadis, which contribute periodically to the
9 suspension load of the Nile.

10

11 Due to the geographical position, the climate and the sedimentary framework of the
12 EMS is controlled by two climate systems. The long-term climate regime was
13 governed by the African monsoon that caused major humid periods with enhanced
14 sediment discharge at 132 to <126 ka (AHP5), 116 to 99 ka (AHP4), and 89 to 77 ka
15 (AHP3). ~~132 to <122 ka (AHP 5), 113 to 104 ka (AHP 4), and 86 to 74 ka (AHP 3).~~
16 They lasted much longer than the formation of the related sapropel layers S5 (>2 ka),
17 S4 (3.5 ka) and S3 (5 ka). During the last glacial period (MIS 4–2) the long-term
18 changes of the monsoonal system were superimposed by millennial-scale changes
19 of an intensified mid-latitude glacial system. This climate regime caused short but
20 pronounced drought periods in the Nile catchment, which are linked to Heinrich
21 Events and alternate with more humid interstadials.

22

23 The clay mineral record further implies that feedback mechanisms between
24 vegetation cover and sediment discharge of the Nile are detectable but of minor
25 importance for the sedimentary record in the southeastern Mediterranean Sea during
26 the investigated African Humid Periods.

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29 **1 Introduction**

30

31 The Nile in northern Africa is the longest river of the world and the dominant
32 sediment source for the Eastern Mediterranean Sea (Milliman and Syvitski, 1992). Its
33 drainage basin is about 3×10^6 km² (Garzanti et al., 2015) and extends from the
34 equator to ca. 30°N (Fig. 1). The main tributaries of the Nile are the perennial White
35 Nile originating from Lake Victoria in tropical East Africa and the highly seasonal Blue

1 Nile and Atbara originating from the subtropical Ethiopian highlands. ~~Downstream of~~
2 ~~the confluence, the Nile flows for more than 2000 km through the hyperarid Sahara.~~
3 ~~On the way from the catchment to the delta, the Nile flows through different climate~~
4 ~~zones, from a cool and humid climate in the Ethiopian highlands to a hot and arid~~
5 ~~climate in Egypt.~~

6
7 Because of this special geographical situation, the runoff of the Nile is sensitive to
8 climatic changes, and its discharge sediments are major recorders of the geological
9 and climatic conditions in the catchment areas. The Nile sediments mainly reflect the
10 intensity of rainfall in the headwaters, which has direct control on weathering, erosion
11 and transport of sediments (Krom et al., 2002; Revel et al., 2010; Box et al., 2011;
12 Garzanti et al., 2015). The present-day summer floods of the Nile are linked to a
13 northward movement of the Intertropical Convergence Zone (ITCZ) and the African
14 Rain Belt (ARB; Fig. 1) causing especially intense precipitation in July and August
15 especially in the Ethiopian highlands.

16
17 Prior to the emplacement of the Assuan High Dam in 1964, the Nile carried a
18 suspension load of about $120\text{--}160 \times 10^6 \text{ t yr}^{-1}$ to the Eastern Mediterranean Sea
19 (EMS) (Holeman, 1968; Milliman and Syvitski, 1992; Stanley and Wingerath, 1996a).
20 More than 95% of the material is derived from the Ethiopian highlands via Blue Nile
21 and Atbara. ~~In contrast, the sediment discharge whereas the load~~ of the White Nile
22 can be neglected ~~because it's load stored in Lake Victoria, Lake Albert and in the~~
23 ~~Sudd basin of South Sudan~~ (Adamson et al., 1980; Foucault and Stanley, 1989;
24 Williams et al., 2006; Williams et al., 2015; Garzanti et al., 2015). About 32.5% of the
25 Nile suspension consists of clay (Quellenec and Kruc, 1976). Most of the silt- and
26 sand-sized sediment accumulates in the Nile delta, on the Nile cone and along the
27 Mediterranean shelf, whereas the clay-sized sediment fraction is transported in
28 suspension by the surface currents of the Mediterranean Sea to the east and north
29 (Fig. 1; Venkatarathnam and Ryan, 1971; Foucault and Mélières, 2000; Hamann et
30 al., 2009).

31
32 It is well known that past changes in the amount of rainfall in the Ethiopian highlands
33 were caused by precession-driven shifts in the position of the ITCZ that influenced
34 the intensity and the spatial extent of the monsoon (e.g. Rossignol-Strick, 1983;
35 Emeis et al., 2003). During African Humid Periods (AHPs) the spreading of the

1 vegetation cover in North Africa resulted in a “green” Sahara (Renssen et al., 2006;
2 Kröpelin et al., 2008) and a reduced influx of Saharan dust into the Mediterranean
3 Sea (deMenocal et al., 2000; Ehrmann et al., 2013). The increased runoff of the Nile
4 River and of other North African river/wadi systems Nile River runoff brought huge
5 amounts of freshwater and nutrients to the Mediterranean Sea, leading to enhanced
6 productivity in the surface water, stagnation in the deep marine basin and formation
7 of sapropel layers (Rossignol-Strick et al., 1982; Rohling, 1994; Emeis et al., 2003;
8 Rohling et al., 2015). Although numerous studies concentrated on the Holocene
9 AHP, very little is known about timing and intensity of former AHPs. Marine proxy
10 data across sapropel S5 suggest that during the Eemian period (MIS 5e) the African
11 monsoon was particularly strong and the African rain belt was shifted even further
12 north than during the early Holocene. This resulted in flooding into the EMS through
13 wadi systems along the wider North African margin (Rohling et al., 2002a; Osborne
14 et al., 2008).

15

16 Different opinions exist on whether Nile sediment discharge is primarily controlled by
17 precipitation and runoff, and thus closely linked to changes in the monsoon intensity
18 (Wehausen and Brumsack, 2000; Emeis et al., 2000; Revel et al., 2010), or whether
19 feedback mechanisms play an important role, with vegetation protecting soils and
20 weathering products in the source area from being eroded (Adamson et al., 1980;
21 Krom et al., 2002; Box et al., 2011).

22

23 A comprehensive effort to reconstruct Nile discharge during the last 100 ka was
24 undertaken using the Fe content in sediments recovered from the proximal Nile delta
25 (Revel et al., 2010; Caley et al., 2011). Most other geochemical investigations
26 concentrated on shorter time spans and used different parameters such as K/Al,
27 Mg/Al, Ti/Al, Ba/Ca to characterise the Nile discharge (Wehausen and Brumsack,
28 2000; Box et al., 2011; Weldeab et al., 2014). Several studies include the isotopic
29 fingerprint of Nile sediments to reconstruct the Nile history (Krom et al., 1999a;
30 Weldeab et al., 2002; Box et al., 2011; Blanchet et al., 2014).

31

32 We choose a different approach of reconstructing the Nile River sediment supply
33 throughout the last some 140 ka 145-ka by mainly investigating the content of typical
34 Nile-derived clay minerals in sediment core GeoTü SL110, which was recovered from
35 the distal Nile discharge plume off Israel (Fig. 1). Some 80% of the terrigenous

1 sediment fraction of the surface sediments near SL110 is derived from the Nile (Krom
2 et al., 1999b), and is characterised by a high smectite content (e.g. Venkatarathnam
3 and Ryan, 1971; Foucault and Mélières, 2000). By analysing a distal core from the
4 plume, we document the Nile discharge activity as a whole rather than the activity of
5 only a single channel in the delta.

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8 **2 Material**

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10 The investigated sediment core GeoTü SL110 was recovered from the SE Levantine
11 Sea (Fig. 1, Table 1) during cruise M51/3 of the German research vessel “*Meteor*” in
12 2001 (Hemleben et al., 2003). It was retrieved at the Israeli continental slope off
13 Haifa from a water depth of 1437 m. The sediments of the 652 cm long core mainly
14 consist of muds with traces of pteropods and foraminifers. The core top contains
15 basically undisturbed surface sediments as indicated by the presence of an oxidised
16 layer. The muds show frequent diffuse colour changes, mainly between grey and
17 olive grey. Darker sapropel layers (greenish black, dark greenish grey, dark grey)
18 occur at 35–61 cm (S1, with a possible interruption at 44–45 cm), 500–514 cm (S3),
19 558–570 cm (S4) and 598–622 cm (S5, including pre-sapropelic layer). **The core**
20 **does not show any indications for slumping, debris flows or turbidites but indicates**
21 **rather constant sedimentary conditions in the distal discharge plume of the Nile. No**
22 **tephra layers were detected.**

23

24

25 **3 Methods**

26

27 The lightness (L^*) of core SL110 was determined in 1 cm steps with a Minolta colour
28 spectrophotometer immediately after opening of the core.

29

30 The content of total organic carbon (TOC) was measured on 174 ground bulk
31 sediment samples with an Eltra METALYST-CS-1000-S after removal of carbonate
32 with HCl.

33

34 The core was sampled in 1 cm intervals for investigations of the grain size
35 composition of the terrigenous sediment fraction and of the clay mineral

1 assemblages. The samples were oxidized and disaggregated by means of 5%
2 hydrogen peroxide. Carbonate was dissolved by 10% acetic acid. We isolated the
3 terrigenous sand fraction by sieving the samples through a 63 μm mesh. The silt
4 fraction (2–63 μm) was separated from the clay fraction (<2 μm) in settling tubes.

5
6 The analyses of the clay mineral composition followed standard methods (e.g.
7 Ehrmann et al., 2007). We added MoS_2 as an internal standard to the clay
8 suspension. Texturally oriented clay mounts were solvated with ethylene–glycol
9 vapour at a temperature of 60°C and then X-rayed with a Rigaku MiniFlex system
10 ($\text{CoK}\alpha$ radiation; 30 kV; 15 mA). We analysed the samples in the range 3–40° 2 θ
11 with a step size of 0.02° 2 θ and a measuring time of 2 s step⁻¹. In addition, we
12 analysed the range 27.5–30.6° 2 θ with a step size of 0.01° 2 θ and a measuring time
13 of 4 s step⁻¹ in order to better resolve the (002) peak of kaolinite and the (004) peak
14 of chlorite. We evaluated the diffractograms by using the MacDiff software
15 (Petschick, 2001). After adjusting the diffractograms to the MoS_2 peak at 6.15 Å, we
16 deconvoluted the peak doublets smectite/chlorite (17/14 Å), palygorskite/illite
17 (10.5/10 Å) and kaolinite/chlorite (3.58/3.54 Å). We based the semi-quantitative
18 estimations of the clay mineral abundances on the integrated peak areas of the
19 individual peaks and weighting factors (Biscaye, 1964; 1965; [Weaver and Beck,](#)
20 [1977](#)). ~~For palygorskite, we used the same factor as for chlorite and kaolinite.~~

21

22

23 **4 Age Model**

24

25 For establishing a stable oxygen isotope stratigraphy we analysed tests from the
26 250–355 μm size fraction of the surface-dwelling foraminifer species *Globigerinoides*
27 *ruber* (white) of about 130 samples using a Kiel IV online carbonate preparation line
28 connected to a MAT 253 mass spectrometer. The values are reported in ‰ relative
29 to VPDB using the delta notation. The reproducibility was better than $\pm 0.06\text{‰}$ (1 σ).
30 Age points were derived by a graphic correlation of our $\delta^{18}\text{O}$ record with [the *G. ruber*](#)
31 [\(white\) \$\delta^{18}\text{O}\$ data of sediment core LC21 from the southern Aegean Sea \(Grant et al.,](#)
32 [2012\), a standard stack \(Lisiecki and Raymo, 2005\)](#) aided by the software
33 AnalySeries 2.0 (Paillard et al., 1996) (Table 2, Fig. 2). [The LC21 chronology was](#)
34 [constructed by using \$^{14}\text{C}\$ dating and by tuning its \$\delta^{18}\text{O}\$ record to the U/Th-dated](#)
35 [Soreq Cave speleothem \$\delta^{18}\text{O}\$ record \(Grant et al., 2012\).](#)

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Further age control comes from four ^{14}C -accelerator mass spectrometry (AMS) datings performed by Beta Analytic Radiocarbon Dating Laboratory on well-preserved shells of planktonic foraminifera (*G. ruber*, *G. bulloides*, *G. sacculifer*, *O. universa*) that represent the age of the surface waters. We applied an eastern Levantine Sea Delta-R of 3 ± 66 years (Marine Reservoir Database). The radiocarbon ages were converted to calendar years using the Marine13 data base (Reimer et al., 2013; Table 2).

We identified a hiatus at 598 cm, at the top of the S5 interval, based on several arguments. The lower part of the dark interval between 622 and 604 cm is characterised by only about 0.8% TOC (Fig. 3a), by the occurrence of a relatively diverse benthic foraminiferal assemblage containing *G. orbicularis*, *G. translucens* and *C. laevigata*, and by relatively high $\delta^{18}\text{O}$ values (Fig. 3a). Despite the dark colour, this indicates a pre-sapropelic interval at the transition between Marine Isotope Stages (MIS) 6 and 5 rather than a full sapropel (Schmiedl et al., 2003). The TOC concentrations start to rise to ca. 1.4% at 603 cm. We therefore assume the base of sapropel S5 at 603 cm, with an age of 128.0 ka (Grant et al., 2012). ~~124.0 ka (Bar-Matthews et al., 2000; Rohling et al., 2002b; Morigi, 2009; Osborne et al., 2010).~~ The TOC values drop abruptly to 0.5% at 596 cm. Thus, the TOC concentrations remain below the typical values of 2–3% in S1, S3, S4 (Fig. 3a) and in S5 of nearby core KL83 (Weldeab et al., 2003). Also the analysis of the benthic foraminifera showed an extremely thin anoxic phase in SL110, compared to a 27 cm thick S5 in KL83 (Schmiedl et al., 2003). Thus, most of S5 is missing in SL110 due to a hiatus that was probably caused by a slump as indicated by the sharp and curved lithological boundary at 598 cm. By comparing our $\delta^{18}\text{O}$ record with standard curves and the record of MD84-641 (Table 1, Kallel et al., 2000) and considering ~~an age of 128–121 ka for S5 (Grant et al., 2012) we argue for an extrapolated age of the hiatus from 126–119 ka. the published age of 124 to 119 ka for S5 (Bar-Matthews et al., 2000; Rohling et al., 2002b; Morigi, 2009; Osborne et al., 2010) we argue for an extrapolated age of the hiatus from 122 ka to 115 ka.~~ In nearby core GeoTü KL83 (Table 1), Weldeab et al. (2002, 2003) identified a hiatus spanning the period 110–70 ka, with the base of the hiatus at 13 cm above S5.

1 For consistency reasons, we also adapted the age model of core GeoTü SL143
2 (Ehrmann et al., 2013; Fig. 4) by using basal ages of 86.0 ka and 108.0 ka for
3 sapropels S3 and S4, respectively, according to the LC21 record (Grant et al., 2012)

6 **5 Results**

8 The raw data of our investigations on sediment core GeoTü SL110 ~~are stored in the~~
9 ~~PANGAEA data base at the Alfred Wegener Institute for Polar and Marine Research,~~
10 ~~Bremerhaven, Germany (<http://doi.pangaea.de/10.1594/PANGAEA.848291>) and~~ are
11 presented in Fig. 3, which also shows the positions of the sapropel layers S1, S3, S4
12 and S5.

14 The $\delta^{18}\text{O}$ data of *G. ruber* go back to MIS 6 (Fig. 3a ~~Figs. 2, 3a~~). The Last Glacial
15 Maximum has values of ca. 3‰, whereas the minimum values during interglacials
16 MIS 1 and MIS 5 are around -1.5‰.

18 The lightness data (L^* , Fig. 3a) reflect subtle colour changes within the core. The
19 darkest intervals correlate with the sapropels (L^* ca. 37). However, also large parts of
20 the other muds are relatively dark (L^* ca. 40–45). The lightest interval occurs ~~at 625–~~
21 ~~640 cm, corresponding to~~ during the MIS 6 interval 132–135 ka.

23 TOC contents generally fluctuate between 0.4 and 0.7%, but increase to 2.5–3% in
24 the sapropel layers S1, S3 and S4, and to 1.5% in ~~the~~ pre-sapropelic layer of S5.

26 The terrigenous sand content is generally <4%, the silt/clay ratio <1.0. The data sets
27 show the same general pattern (Fig. 3a) with finer grained intervals ~~around 0–60,~~
28 ~~490–515, 550–570, 600–620 cm and~~ enclosing the sapropel layers. A
29 distinct coarse interval occurs at 132–135 ka ~~625–640 cm~~. Between 73 and 18 ka
30 ~~480 and 110 cm~~ we observe an upward coarsening trend. The high silt/clay ratio at
31 ~~ca. 17 ka (103–102 cm)~~ reflects a thin, light gray, not graded silt layer **consisting**
32 **almost entirely of quartz grains.**

34 The clay mineral assemblage of SL110 (Fig. 3b) is dominated by smectite that
35 fluctuates between ca. 30% ~~35%~~ and 65%. Maximum values occur around 3, 31–37,

1 46, 50–59, 65–69, 77–89, 99–116, and 127–132 ka (~~15, 210–260, 315, 340–400,~~
2 ~~460, 490–525, 545–585 and 600–625 cm~~). Minimum values occur at 132–135 ka
3 ~~625–640 cm~~. Illite shows an opposite distribution pattern with concentrations ranging
4 between 5 and 25%. Kaolinite concentration fluctuates between 20% and 30% and
5 shows only a few distinct maxima centred at 95, 116 and 133 ka, ~~540 cm, 590 cm~~
6 ~~and 630 cm~~, and minima at 0–5, 35, 87, 111 and 128–132 ka ~~0–30, 230–250, 520,~~
7 ~~570 and 600–620 cm~~. Chlorite is present in low concentrations of 7–11%, with a
8 pattern roughly resembling that of kaolinite. Palygorskite occurs in trace amounts
9 throughout the core, but jumps to 4% at 132–135 ka ~~625–640 cm~~.

12 6 Discussion

14 Sedimentation at SL110

16 According to our age model (Fig. 2, Table 2), sediment core SL110 has an
17 extrapolated basal age of ca. 137 ka ~~145 ka~~. The linear sedimentation rates are 2.2
18 to 9.8 ~~8.8~~ cm/ka (Fig. 3a). They are low during ~~late MIS 6 and~~ MIS 5, and distinctly
19 higher in ~~late MIS 6 and MIS 4 MIS 3~~ to MIS 1. Sedimentation rates for nearby core
20 MD84-641, ~~extracted calculated~~ from data by Fontugne and Calvert (1992), show
21 ~~less distinct changes a similar pattern~~ (Fig. 2) ~~and but~~ are distinctly lower than in
22 SL110 due to its further offshore position.

24 The non-carbonate sediment fraction of seafloor surface sediments in the region of
25 SL110 is dominantly derived by long-distance sediment transport from the Nile
26 (Venkatarathnam and Ryan, 1971; Stanley et al., 1997; 1998; Krom et al., 1999b;
27 Weldeab et al., 2002; Hamann et al., 2009). Storm waves and coastal currents erode
28 coarse sediments from the delta. Geostrophic and wave-induced near-shore and
29 shelf currents displace this material in a large counter-clockwise current along the
30 coast towards the Levantine Sea (Fig. 1) and deposit it in the shallow waters off
31 Israel. The fine fraction, in contrast, is derived from direct input from the Nile into the
32 sea and may be transported also further offshore and further to the north (e.g.
33 Stanley et al., 1997; 1998; Venkatarathnam and Ryan, 1971; Foucault and Mélières,
34 2000). Aeolian influx from the Sahara and riverine supply of small Near East rivers
35 and wadis are only of minor importance (Stanley et al., 1997; Sandler and Herut,

1 2000). The enhanced glacial sedimentation rates observed in core SL110 might be a
2 result of stronger currents and/or a shift of the currents and therewith of the high-
3 accumulation areas to a more seaward position due to a lower sea level. These
4 processes may also be responsible for the coarsening of the sediments as indicated
5 by the terrigenous sand content and the silt/clay ratios (Fig. 3a).

6
7 The increasing glacial sedimentation rates theoretically also could be explained by
8 increasing aridity resulting in an enhanced influx of aeolian dust. Endmember
9 modelling of the terrigenous silt fraction of the nearby core SL112 (892 m water
10 depth), which is located closer to the coast, indicates increased dust fluxes during the
11 LGM and the late glacial (Hamann et al., 2008). However, such an influx is not
12 documented in the clay mineral record of SL110, e.g. by the concentration of wind-
13 blown palygorskite (Fig. 3b). In addition, glacial sedimentation rates at site SL112
14 were lower when compared to the Holocene, thus supporting the interpretation of a
15 glacial-interglacial shift of the high-accumulation zone.

16 [Clay Mineral Assemblages](#)

17
18
19 Smectite in southeastern Levantine Sea sediments is generally regarded to be of
20 detrital origin and is the typical clay mineral in Nile-derived sediments
21 (Venkatarathnam and Ryan, 1971; Maldonado and Stanley, 1981; Stanley and
22 Wingerath, 1996a; Foucault and Mélières, 2000). According to the compilation by
23 Hamann et al. (2009), the Nile assemblage contains up to 85% smectite. Kaolinite is
24 the second important clay mineral with concentrations in the order of 20%. Illite
25 contents are <10%, and chlorite occurs in trace amounts. ~~The Nile incorporates
26 almost its entire suspension load close to its source (Adamson et al., 1980).~~

27
28 The riverine suspension load is reflected in the modern clay mineral assemblage in
29 seafloor surface sediments of the SE Levantine Sea. Accordingly, the clay mineral
30 assemblage off southern Israel consists of up to >70% smectite, 10–25% kaolinite,
31 10% illite and <10% chlorite (Hamann et al., 2009) and thus reflects the Nile source.
32 The uppermost samples of core SL110 with ca. 60% smectite, 22% kaolinite, 8% illite
33 and 9% chlorite fit well into this distribution pattern.

34
35 The composition of the aeolian clay mineral assemblage reaching the SE Levantine

1 Sea is variable, but generally characterized by high illite and kaolinite concentrations
2 (Chester et al., 1977). Palygorskite is a typical clay mineral in aeolian dust derived
3 from the Sahara (Foucault and Mélières, 2000; Scheuven et al., 2013).

4
5 The Nile incorporates almost its entire suspension load close to its source (Adamson
6 et al., 1980). The smectite comes with the Nile suspension load from the catchment
7 of the Blue Nile and Atbara rivers and originates from weathering of Cenozoic
8 volcanic rocks on the Ethiopian Plateau (e.g., Stanley and Wingerath, 1996a). The
9 White Nile south of the confluence of Sobat river, also discharging the Ethiopian
10 highlands, shows a considerably different clay mineral signature with significantly
11 lower smectite contents. Alluvial sediments from central Uganda have smectite
12 contents <20% (Nyakairu and Koeberl, 2001); no smectite was described from lake-
13 bottom sediments of Lake Victoria (Mothersill, 1976). Tropical weathering in the
14 equatorial region produces only small amounts of smectite (Garzanti et al., 2015).
15 Furthermore, the smectite content in SL110 correlates with the Fe content
16 documented by Revel et al. (2010) in delta sediments of the Nile, especially during
17 humid periods. Because Fe is unequivocally derived from the volcanic rocks of the
18 Ethiopian highlands, this also suggests that the overwhelming part of the smectite
19 comes from that source.

20
21 Sediments of the Egyptian wadis discharging into the Nile have clay mineral
22 assemblages with typically 40–55% kaolinite, 35–45% smectite and 5–15% illite
23 (Stanley and Wingerath, 1996b; Hamann et al., 2009; chlorite not determined, thus,
24 actual kaolinite concentrations may be somewhat lower). The kaolinite is mainly
25 derived from the erosion of kaolinite-rich Eocene, Palaeocene and Mesozoic
26 sediments and lateritic soils in the wadis (Stanley and Wingerath, 1996b; Bolle et al.,
27 2000). Theoretically, the White Nile could also deliver kaolinite during certain times of
28 enhanced sediment transport during humid periods (Garzanti et al. 2015). This would
29 result in a decrease in smectite at the same time due to dilution. This, however, is not
30 the case (Fig. 3b). Therefore we argue that the kaolinite accumulating in SL110
31 predominantly comes from the North African wadi regions.

32
33 ~~The composition of the aeolian clay mineral assemblage reaching the SE Levantine~~
34 ~~Sea is variable, but generally characterized by high illite and kaolinite concentrations~~
35 ~~(Chester et al., 1977). Palygorskite is a typical clay mineral in aeolian dust derived~~

1 ~~from the Sahara (Foucault and Mélières, 2000; Scheuven et al., 2013).~~

2
3 ~~The clay mineral assemblage in the seafloor surface sediments of the SE Levantine~~
4 ~~Sea off southern Israel consists of up to >70% smectite, 10–25% kaolinite, 10% illite~~
5 ~~and <10% chlorite (Hamann et al., 2009) and thus reflects the Nile source. The~~
6 ~~uppermost samples of core SL110 with ca. 60% smectite, 22% kaolinite, 8% illite and~~
7 ~~9% chlorite fit well into this distribution pattern.~~

8
9 According to the standard method, percentages of the clay minerals smectite, illite,
10 chlorite, kaolinite and palygorskite add up to 100%. A change in the abundance of
11 one mineral therefore causes also changes in the concentration of the others. Hence
12 we prefer to discuss clay mineral ratios. We use $Sm_r = \text{smectite}/(\text{illite}+\text{chlorite})$ as the
13 main proxy for Nile discharge provided by the Blue Nile and Atbara, and $Ka_r =$
14 $\text{kaolinite}/(\text{illite}+\text{chlorite})$ as an indicator for the contribution of the wadis.

15
16 The good correlation between the Ka_r and the Sm_r data (Fig. 4; $r^2 = 0.73$) confirms
17 that the kaolinite in SL110 is of fluvial origin (Venkatarathnam and Ryan, 1971;
18 Stanley and Wingerath, 1996a; Hamann et al., 2009), and not derived by aeolian
19 influx from the Sahara, as in other regions of the EMS (e.g. Chester et al., 1977;
20 Foucault and Mélières, 2000; Ehrmann et al., 2013). The correlation implies that
21 during times of enhanced Nile runoff and smectite transport from Ethiopia, the
22 erosion of kaolinite-rich sediments and soils in the wadis was generally also active,
23 and that during dry periods in Ethiopia the wadis were dry, too.

24
25 This confirms earlier findings based on the interpretation of satellite images and
26 sedimentological investigations, which showed that extensive drainage systems were
27 active in the eastern Sahara during humid phases, especially of Eemian and
28 Holocene time (e.g. Pachur and Hoelzmann, 2000; Rohling et al., 2002b; Osborne et
29 al., 2008). Furthermore, oxygen isotope data from speleothems of Soreq Cave (Bar-
30 Matthews et al., 2000) imply that rainfall extended beyond the Sahara.

31
32 The clay mineral data of SL110 document a highly variable smectite and kaolinite
33 input through time (Figs. 3b, 4) and show similar distribution patterns as the Fe
34 record from the Nile delta (Revel et al., 2010; Caley et al., 2011). Fe is supposed to
35 be derived from weathering of volcanic rocks in the **Atbara and Blue Nile** headwaters

1 in the Ethiopian highlands, and thus changing Fe contents may reflect changing
2 precipitation regimes. Furthermore, the smectite and kaolinite distribution patterns
3 are inversely correlated to the record of Saharan dust in the **central Aegean Sea core**
4 **GeoTü SL143** (Ehrmann et al., 2013).

5

6 Penultimate Glacial Period

7

8 The oldest part of the record displays minimum Sm_r and Ka_r values **at ca. 135-132 ka**
9 **of MIS 6 (Fig. 4). The lack of sediment derived from the Nile delta indicates that the**
10 **site was not within the range of the Nile deep-sea turbidite system postulated for MIS**
11 **6 at ca. 33°30' to 33°50'E (Ducassou et al. 2009). The clay mineral data point**
12 **pointing to extremely dry conditions for the penultimate glacial maximum ~~at ca. 142-~~
13 **~~132 ka of MIS 6 (core interval 646-624 cm, age uncertainties >135 ka because of~~
14 **~~extrapolation).~~** The extreme drought possibly correlates with Heinrich Event 11,
15 which is known as a very cold period in the Western Mediterranean Sea **and dated to**
16 **135-130 ka** (Martrat et al., 2014; Marino et al., 2015). Obviously, only minor amounts
17 of Nile suspension load reached the coring site off Israel. Instead, maximum
18 palygorskite concentrations (Fig. 3b) and high proportions of terrigenous silt and
19 sand (Fig. 3a) document an increased aeolian influx from the Sahara. This arid
20 phase seems to have been more severe than that of the Last Glacial Maximum. After
21 the maximum drought of MIS 6 the record of SL110 show a sharp transition to a
22 sapropelic event embedded in an AHP.****

23

24 Sapropel Formation

25

26 It is well known that sapropel formation was coupled to precession-related maxima in
27 northern hemisphere summer insolation causing a northward shift of the ITCZ, an
28 intensification of the African monsoon system and enhanced precipitation in the
29 Ethiopian highlands (e.g. Rossignol-Strick, 1983; Rohling, 1994; Emeis et al., 2003;
30 Rohling et al., 2015). In SL110, high Sm_r , Ka_r and low silt/clay ratios **occur mainly**
31 **during interglacial periods with negative $\delta^{18}O$ values and** indicate phases with
32 enhanced Nile and wadi discharge of fine-grained suspension to the Levantine Sea
33 accompanying formation of sapropels S5, S4, S3 and S1. The enhanced rainfall
34 under a warm climate caused stronger chemical weathering of the volcanic rocks and
35 thus smectite formation. The enhanced river runoff contributed to the stratification of

1 the water column and facilitated the transport of nutrients and weathering products to
2 the EMS promoting ~~enhanced surface water productivity at least~~ a shoaling of the
3 ~~nutricline and thus development of a deep chlorophyll maximum~~ (Rohling and
4 ~~Gieskes, 1989; Grelaud et al., 2012)~~ particularly in areas close to the Nile delta.
5 These processes fostered sustained anoxic conditions in the deeper basins and
6 preservation of TOC (Fig. 3a).

7

8 Both the clay mineral record of southeastern Levantine Sea core SL110, the Fe
9 record from the Nile delta (Revel et al., 2010; **Caley et al., 2011**) and the record of
10 Saharan dust influx to the Aegean Sea (Ehrmann et al., 2013) reveal that each phase
11 of sapropel formation occurred within a longer AHP (Fig. 4; AHPs numbered
12 according to the corresponding sapropels).

13

14 [The last Interglacial Period](#)

15

16 The Sm_r maximum linked to AHP 5 shows a sharp onset after the glacial maximum of
17 MIS 6, simultaneously with the start of the pre-sapropelic layer of S5 (Fig. 4). It
18 indicates a northward moving rain belt resulting in enhanced monsoonal rainfall in
19 Ethiopia and Blue Nile runoff as early as 132 ka at a time of increasing insolation but
20 under the influence of penultimate glacial boundary conditions. Also the Ka_r values
21 clearly increase with the beginning of the pre-sapropelic layer. However, they do not
22 exhibit such distinct maximum values as during the early phases of the younger
23 AHPs. Most of sapropel S5 and the termination of AHP 5, however, are not
24 preserved in SL110 due to the hiatus at ca. ~~126–119 ka~~ ~~122–115 ka~~. Furthermore,
25 because of the curved top of the sapropel and the technique of sampling the core in
26 parallel slices, we did not receive the typical sapropel signature but a mixed
27 sapropel/post-sapropel signal (Fig. 4).

28

29 ~~According to the Fe and the dust record (Revel et al., 2010; Ehrmann et al., 2013),~~
30 ~~AHP 4 started at >108 ka and lasted until 95 ka or 98 ka, respectively. The Sm_r and~~
31 ~~Ka_r data of SL110 indicate that AHP 4 started during insolation rise at 116 ka and~~
32 ~~ended at 99 ka, with sapropel formation occurring at 108–104.5 ka (Fig. 4).~~

33 ~~According to the Fe and the dust record, AHP 4 lasted from >105 ka to ca. 95 ka~~
34 ~~(Revel et al., 2010; Ehrmann et al., 2013). The Sm_r and Ka_r data of SL110 indicate~~

1 ~~that AHP 4 started during insolation rise at 113 ka and ended at 94 ka, with sapropel~~
2 ~~formation occurring at 105–99.5 ka (Fig. 4).~~

3

4 AHP 3 started just before the insolation maximum and lasted ca. ~~89–77 ka~~ ~~86–74 ka~~
5 based on the Sm_r and Ka_r maxima and the minima in the silt/clay ratio and the
6 terrigenous sand content. These ages are in good accordance with those obtained
7 from the dust record in the Aegean Sea (Ehrmann et al., 2013). Sapropel S3 has a
8 shorter duration and spans the time interval ~~86–81 ka~~ ~~82–77.5 ka~~.

9

10 Despite the hiatus in the SL110 record, by assuming a duration of ~~128–121 ka for S5~~
11 ~~(Grant et al., 2012)~~ ~~124–119 ka for S5~~ ~~(Bar-Matthews et al., 2000; Rohling et al.,~~
12 ~~2002b; Morigi, 2009; Osborne et al., 2010)~~, our data indicate that the dry phase
13 between AHP 5 and AHP 4 lasted only ca. 5 ka, compared with ca. 11 ka for the one
14 ~~between AHP 4 and AHP 3 ca. 6 ka, and the one between AHP 4 and AHP 3 only ca.~~
15 ~~8 ka~~. It is interesting to note that a similar duration of the dry phases can be deduced
16 from the pollen record of Tenaghi Philippon in Greece, although this remote site is
17 governed dominantly by northern climatic factors (Tzedakis, 2005; Pross et al.,
18 2015).

19

20 The last Glacial Period

21

22 After a further short dry period at ~~77–70 ka~~ ~~74–68 ka~~, Sm_r , Ka_r and silt/clay ratios
23 indicate humid phase AHP 2 and enhanced Nile runoff in early MIS 4 at ~~69–65 ka~~
24 ~~68–64 ka~~. It is of similar magnitude as the interglacial AHPs discussed before and
25 also starts during increasing insolation (Fig. 4). In contrast to the previous AHPs this
26 one is not accompanied by low $\delta^{18}O$ values. It is also not accompanied by a
27 sapropel, dark sediment colour or enhanced TOC values (Fig. 3a). However, it
28 coincides with the common occurrence of shallow infaunal *Uvigerina peregrina*. This
29 highly opportunistic species is adapted to the seasonal deposition of high amounts of
30 fresh phytodetritus (Koho et al., 2008). In analogy to nearby SL112, the occurrence
31 of this species at site SL110 during AHP 2 is likely fuelled by pronounced algal
32 blooms related to seasonally enhanced Nile runoff and nutrient supply (Schmiedl et
33 al., 2010). However, in core SL110 we have strong indications that AHP 2 did not
34 culminate in a sapropel formation. Sediment and nutrient delivery by the Nile was
35 cut-off by a drought linked to Heinrich Event 6 occurring just at the time of maximum

1 insolation (Fig. 4). This drought, possibly combined with stronger winds and better
2 ventilation of the deep waters, reduced sediment discharge into the Mediterranean
3 Sea and hampered stagnation of the deep water and sapropel formation.

4

5 Lourens et al. (1996 1966) described a thin sapropel layer S2, corresponding to AHP
6 2, south of Crete that occurs after the 60 ka insolation maximum. It was argued that
7 this sapropel is not visible any more in most sedimentary sequences of the eastern
8 Mediterranean Sea because of post-depositional oxidation of the organic matter
9 (Emeis et al., 2000; Löwemark et al., 2006).

10

11 It is not understood why our clay mineral signal has no counterpart in the Fe signal
12 from the Nile delta (Revel et al., 2010). Possibly Nile discharge to the Eastern
13 Mediterranean Sea at that time was not via the western Rosetta branch, but via the
14 eastern Damietta branch, and thus site MS27PT starved of sediment. It also cannot
15 be ruled out that a match is obscured because of different age models, insofar as the
16 Fe 73 ka peak corresponds to our well-dated 67 ka 68-ka clay mineral peak (Fig. 4).

17

18 The slightly enhanced Sm_r and Ka_r values at 59–50 ka 58–48-ka possibly document
19 the late phase of AHP 2 after the termination of Heinrich Event 6, reflecting the re-
20 occurrence of slightly enhanced Nile River runoff. Other indications for somewhat
21 increased temperature, humidity and Nile influx during this time interval come from
22 the Fe record in the Nile delta (Revel et al., 2010) and the $\delta^{18}O$ records of Soreq
23 Cave and of sediment cores 9509 and 9501 in the eastern Levantine Sea (Almogi-
24 Labin et al., 2009). The latter records show a marked drop in planktonic foraminiferal
25 $\delta^{18}O$ values between 58 and 49 ka BP. The signal of warm and freshened surface
26 waters was particularly expressed in core 9501 from the northern part of the basin
27 and coincided with D-O interstadial 14. The spread of deciduous and other trees
28 during the same time period likely supported the migration of anatomically modern
29 humans from Africa into the Levant (Müller et al., 2011). In the Sm_r data of core
30 SL110, this time interval documents only a moderate increase of Nile runoff (Fig. 4)
31 suggesting that the observed warming and freshening of the eastern Mediterranean
32 region was predominantly related to a high-latitude climate change.

33

34 Thus, the sediments of SL110 document that the long-term climate regime governed
35 by the African monsoon and characterized by pronounced AHPs during the last

1 interglacial period was replaced as of the occurrence of Heinrich Event 6 by a
2 distinctly different glacial climate system governed by short-term, millennial-scale
3 climate changes having their origin in the North Atlantic and Arctic regions. Former
4 studies showed that sudden freshwater influxes into the northern Atlantic Ocean
5 during the Younger Dryas and the Heinrich Events not only weakened the
6 thermohaline circulation and oceanic heat transport to the northern high latitudes
7 (e.g. Bond et al., 1993), but also had major consequences for the Mediterranean Sea
8 environment. Increased aridity and even North African mega-droughts during such
9 events can be seen in marine sediments of the North Atlantic Ocean (Mulitza et al.,
10 2008), speleothem data and lake level drops of Lake Lisan (palaeo-Dead Sea; Bar-
11 Matthews et al., 1999; Bartov et al., 2003) and of African lakes (Gasse et al., 2008;
12 Shanahan et al., 2015). Cold water influx from the North Atlantic Ocean to the
13 Mediterranean Sea intensified. Sea surface temperatures and deep-water
14 temperatures decreased and deep-water ventilation fortified in the northwestern
15 Mediterranean Sea (Cacho et al., 1999; 2000) and probably also in the EMS.
16 Enhanced northward flux of wind-transported sediment from the Sahara is
17 documented for both the Western and the Eastern Mediterranean Sea (Moreno et al.,
18 2002; Bout-Roumzeilles et al., 2007; Hamann et al., 2008). Strongly decreased Sm_r
19 and Ka_r values in core SL110 (Fig. 4) indicate reduced Nile sediment discharge
20 during Heinrich Events, and therewith major droughts in the headwaters of the **Atbara**
21 **and** Blue Nile and diminished erosion and river runoff. During these events the ITCZ
22 and the African Rain Belt moved southward. Both Lake Tana, the source of the Blue
23 Nile, and Lake Victoria, the source of the White Nile, desiccated during the Younger
24 Dryas and the Heinrich Event 1 (Lamb et al., 2007; Marshall et al., 2011). Minimum
25 Sm_r and Ka_r values as well as the occurrence of a silt layer of probably aeolian origin
26 (Fig. 3a) imply that Heinrich Event 1 was the driest interval since the MIS 6 glacial
27 maximum (cf. Stager et al., 2011).

28

29 The 65°N insolation curve shows a further low-amplitude maximum at ca. 37 ka. It is
30 also reflected by slightly enhanced smectite and kaolinite abundances in core SL110,
31 ranging **37–31 ka** ~~38–30 ka~~ and by the Fe record in the Nile delta (Revel et al.,
32 2010). No sapropel formation was recorded following the weak insolation maximum
33 at 37 ka. Thus, the two potential humid periods linked to the glacial insolation
34 maxima at 60 and 37 ka are only vaguely expressed in the EMS, because their

1 effects were much less distinct and because they were suppressed by the glacial
2 climate regime, particularly the impacts of cold and dry stadials.

4 [The last African Humid Period and the Holocene](#)

6 The clay mineral data of SL110 and of nearby core SL112 (Hamann et al., 2009)
7 document a gradual onset of AHP 1 at about 14 ka ~~15-ka~~ (Fig. 4), which coincides
8 with the overflow of Lake Tana, the source of the Blue Nile (Lamb et al., 2007;
9 Marshall et al., 2011). At the same time, Saharan dust influx to the EMS decreased
10 (Ehrmann et al., 2013). S1 at ca. 11–6.5 ka and the peak Blue Nile runoff at 10–9 ka
11 (Weldeab et al., 2014) are only poorly reflected in the clay mineral data. Precipitation
12 generally declined gradually after sapropel formation throughout the Holocene, the
13 Nile River runoff responded quasi-linearly to the changes in rainfall (Weldeab et al.,
14 2014; Blanchet et al., 2014), and the AHP 1 ended at ca. 5 ka (deMenocal et al.,
15 2000). However, the clay mineral data do not mirror this trend but show continued
16 increasing Sm_r and Ka_r values and Nile sediment discharge until ca. 2 ka. Sediment
17 discharge thus is decoupled from water runoff. Possibly, human activity altered the
18 natural signal. Saharan population increased shortly after 11 ka and reached a
19 maximum at 7.5 ka (Manning and Timpson, 2014). Major land-use changes occurred
20 with the advent of Neolithic farming some 8 ka ago (Williams, 2009).

22 [Feedback Effects](#)

24 The clay mineral and grain size data of SL110 indicate an enhanced riverine influx of
25 suspension from the Blue Nile and Atbara to the Mediterranean Sea during the AHPs
26 and therefore document enhanced weathering of the Ethiopian basalts, erosion of
27 soils and transport of the suspension load to the sea. This view is supported by
28 several investigations (Wehausen and Brumsack, 1998; Emeis et al., 2000; Revel et
29 al., 2010; Blanchet et al., 2014). However, it is in conflict with some other studies
30 (Adamson et al., 1980; Krom et al., 1999a; 2002; Box et al., 2011) that argued for a
31 lower Blue Nile sediment discharge during the warm and humid Holocene AHP and
32 higher fluxes during dry climate phases. They reasoned that at present the limited
33 vegetation cover on the Ethiopian highlands allows extensive erosion when the
34 summer monsoon rains fall. During increased monsoon activity, in contrast, longer

1 periods of rain, longer growing seasons and larger vegetation-covered areas reduced
2 soil erosion.

3

4 Our data indicate that an increase in precipitation and Nile suspension delivery
5 happened long before sapropel formation, during phases of increasing insolation.
6 However, in the case of AHP 4 Sm_r and Ka_r maxima occur just before and after
7 sapropel S4 formation, each accompanied by low silt/clay ratios also indicating
8 enhanced riverine suspension (Fig. 3a). Thus, maximum riverine sediment influx
9 does not coincide with maximum water discharge during sapropel formation. This is
10 possibly due to a feedback mechanism as postulated, for example, by Adamson et
11 al. (1980) and Krom et al. (2002). ~~Vegetation feedbacks were obviously not active in~~
12 ~~glacial time. Vegetation feedbacks were obviously less expressed during glacial~~
13 ~~conditions, particularly during the last glacial maximum. This is linked to the generally~~
14 ~~more arid boundary conditions during the last glacial in tropical and northern Africa~~
15 ~~(Gasse, 2000) and resulting sparse vegetation cover, particularly in the deflation~~
16 ~~areas of the Sahara (Tjallingii et al., 2008). The rapid climate changes linked to~~
17 ~~Dansgaard-Oeschger cycles and Heinrich Events find immediate response in the~~
18 ~~sediment delivery to the EMS.~~

19

20 The different timing between the Saharan dust record and the Nile discharge record
21 of the humid phases (Fig. 4) may reflect regional differences in the climatic history.
22 The northward migrating tropical rain belt may first have reached the Ethiopian
23 highlands, where the smectite signals has its origin, and with a time lag the central
24 Sahara, there inhibiting effective dust production. The differences may also be due to
25 different processes being at work. The Nile discharge is mainly driven by annual
26 floods, fluvial erosion and weathering. It therefore may have reacted immediately on
27 a northward shift of the ITCZ and the rain belt. The dust record, in contrast, is mainly
28 a function of aridity, wind strength and, probably most important, the availability of
29 deflatable material. The spreading and retreat of a vegetation cover plays an
30 important role and may react non-linearly and with a different time lag on a shift of
31 the ITCZ than river discharge (Claussen et al., 2013; Ehrmann et al., 2013). Thus,
32 recently Blanchet et al. (2014) could show for the end of AHP 1 that the vegetation
33 retreat from the Sahara was much faster than the decrease in Nile runoff during the
34 southward migration of the rain belt.

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7 Conclusions

– Clay minerals are suitable tools to reconstruct the late Quaternary Nile sediment discharge to the EMS. The abundance of smectite reflects the runoff of the Blue Nile and Atbara, whereas kaolinite abundance reflects the discharge from wadis. The correlation of smectite and kaolinite abundances implies that during times of enhanced Nile suspension discharge the wadis also were active sediment sources.

– High smectite and kaolinite abundances and low terrigenous silt/clay ratios indicate phases of enhanced precipitation and suspension influx from the Blue Nile, Atbara and the wadis to the Levantine Sea, the so-called African Humid Periods (AHPs), which are accompanied by the formation of sapropel layers. This largely contradicts the view that sediment delivery during humid periods was reduced because an extensive vegetation cover hampered soil erosion.

– Precipitation and suspension discharge started much earlier and ended later than sapropel formation. Maximum sediment discharge does not coincide with maximum water discharge during times of sapropel formation because of minor vegetation feedback.

– The time lag between the increase/decrease of Nile suspension discharge and the decrease/increase in the dust record (SL143) can be explained by a successive migration of the rain belt, possibly combined with a vegetation feedback.

– The dry periods between the interglacial AHPs were relatively short, each lasting only [ca. 5–10 ka](#) ~~ca. 6–8 ka~~.

– The sedimentary record of core SL110 demonstrates the interplay of two competing climate systems. The long-term climate regime was governed by the African monsoon and characterized by pronounced AHPs during the last interglacial period. This system was replaced in the last glacial by short-term, millennial-scale climate changes having their origin in the North Atlantic region. They are characterised by short but pronounced drought periods linked to Heinrich Events and more humid interstadials. Precipitation in the Nile catchment, river runoff and

1 sapropel formation can be hampered or even inhibited under the influence of the
2 North Atlantic climate system, particularly under glacial boundary conditions.

3

4

5 *Data Availability:* The raw data to this article ~~is available at~~
6 <http://dx.doi.org/10.1594/PANGAEA.848291>, are stored in the PANGAEA data base
7 at the Alfred Wegener Institute for Polar and Marine Research, Bremerhaven,
8 Germany (<http://doi.pangaea.de/10.1594/PANGAEA.848291>).

9

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17

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- 15

1 **TABLE CAPTIONS**

2

3 **Table 1.** Metadata for the investigated sediment core GeoTü SL110 and other cores
4 mentioned in the text (cf. Fig. 1).

5

6 **Table 2.** Data used for constructing the age model for the investigated sediment core
7 GeoTü SL110. Oxygen isotope ages are based on a graphical correlation with the
8 LC21 record (Grant et al., 2012) from the southern Aegean Sea. ~~Oxygen isotope~~
9 ~~ages follow Lisiecki and Raymo (2005). The age of the base of S5 is according to~~
10 ~~Bar Matthews et al. (2000), Rohling et al. (2002), Morigi (2009) and Osborne et al.~~
11 ~~(2010).~~

12

13

14 **FIGURE CAPTIONS**

15

16 **Figure 1.** Location map of the Nile River basin and the Eastern Mediterranean Sea.
17 **(a)** The Nile River and its main tributaries. Blue dashed lines indicate wadis. The
18 green dotted line shows the main outcrop of Cenozoic basalts in Ethiopia. Also
19 shown are the northern summer and winter limits of the African rain belt (red dashed
20 lines, ARB-S, ARB-W, after Blanchet et al., 2014). Black arrows in the Eastern
21 Mediterranean Sea indicate the surface currents (after Pickard and Emery, 1990). **(b)**
22 Map of the southeastern Levantine Sea with the location of the investigated sediment
23 core GeoTü SL110 and other cores mentioned in the text; **not shown are GeoTü**
24 **SL143 and LC21, which are positioned further to the north** (cf. Table 1). D.B. =
25 Damietta branch of the Nile in the delta, R.B. = Rosetta branch.

26

27 **Figure 2.** Depth-Age plot for the investigated sediment core GeoTü SL110 (red). The
28 age points are listed in Table 2. For comparison, corresponding data were **extracted**
29 **calculated** for core MD84-641 (blue, Fontugne and Calvert, 1992: **Table 1**). Marine
30 Isotope Stages (MIS) follow Lisiecki and Raymo (2005).

31

32

1 **Figure 3.** Basic data of core GeoTü SL110. **(a)** $\delta^{18}\text{O}$ of *G. ruber*, lightness (L^*), %
2 total organic carbon (TOC), ~~and~~ % sand and silt/clay ratio of the terrigenous
3 sediment fraction, ~~and linear sedimentation rates (LSR, cm/ka)~~. **(b)** clay mineral
4 percentages of smectite, illite, kaolinite, chlorite and palygorskite. Gray bars mark the
5 position of sapropel layers S1, S3, S4, S5 and the pre-sapropelic layer beneath S5
6 (light grey). Marine Isotope Stages (MIS) are indicated. ~~The top of S5 corresponds to~~
7 ~~a hiatus.~~

8

9 **Figure 4.** Combination of the June insolation at 65°N **(a)**, the Saharan dust record in
10 the central Aegean Sea core GeoTü SL143 **(b)**; Ehrmann et al., 2013), the Nile
11 discharge based on Fe data in the Nile delta **(c)**; Revel et al., 2010, age model by
12 Caley et al., 2011), and the Nile sediment discharge based on K_r **(d)** and Sm_r **(e)** at
13 site GeoTü SL110. Grey bars indicate the positions of sapropels S1, S3, S4, S5 and
14 of the pre-sapropelic layer (light grey) beneath S5 in core in SL110. Heinrich Events
15 (H) are flagged with arrows. Marine Isotope Stages (MIS) and African Humid Periods
16 (AHP) are indicated at the bottom.

Table 1. Metadata for the investigated sediment core GeoTü SL110 and other cores mentioned in the text (cf. Fig. 1).

Core	Latitude (°N)	Longitude (°E)	Water depth depth (m)	Reference
GeoTü SL110	32°38.95'	34°06.22'	1437	this study
GeoTü KL83	32°36.87'	34°08.89'	1431	Weldeab et al. (2002)
GeoTü SL112	32°44.52'	34°39.02'	892	Hamann et al. (2008, 2009)
GeoTü SL143	38°15.71'	25°06.19'	665	Ehrmann et al. (2013)
MD84-641	33°02'	32°38'	1375	Fontugne and Calvert (1992)
9501	34°32'	33°59'	980	Almogi-Labin et al. (2009)
9509	32°01'	34°16'	884	Almogi-Labin et al. (2009)
MS27PT	31°47.90'	29°27.70'	1389	Revel et al. (2010)
LC21	35°40'	26°35'	1522	Grant et al. (2012)

Table 2. Data used for constructing the age model for the investigated sediment core GeoTü SL110. Oxygen isotope ages are based on a graphical correlation with the LC21 record (Grant et al., 2012) from the southern Aegean Sea. ~~Oxygen isotope ages follow Lisiecki and Raymo (2005). The age of the base of S5 is according to Bar-Matthews et al. (2000), Rohling et al. (2002), Morigi (2009) and Osborne et al. (2010).~~

Depth cm	Age cal. ka BP	Data
0.00	0.00	Sediment surface
34.50	6.43	¹⁴ C AMS dating
64.50	11.53	¹⁴ C AMS dating
150.50	24.38	¹⁴ C AMS dating
189.11	29.29	LC21
234.50	33.92	¹⁴ C AMS dating
254.50	36.29	LC21
359.54	53.50	LC21
465.30	67.11	LC21
509.55	84.99	LC21
529.23	89.03	LC21
554.51	103.32	LC21
569.46	107.79	LC21
579.50	110.38	LC21
598.00	119.00	Hiatus end
598.00	126.00	Hiatus start
603.00	128.00	LC21
624.52	131.94	LC21

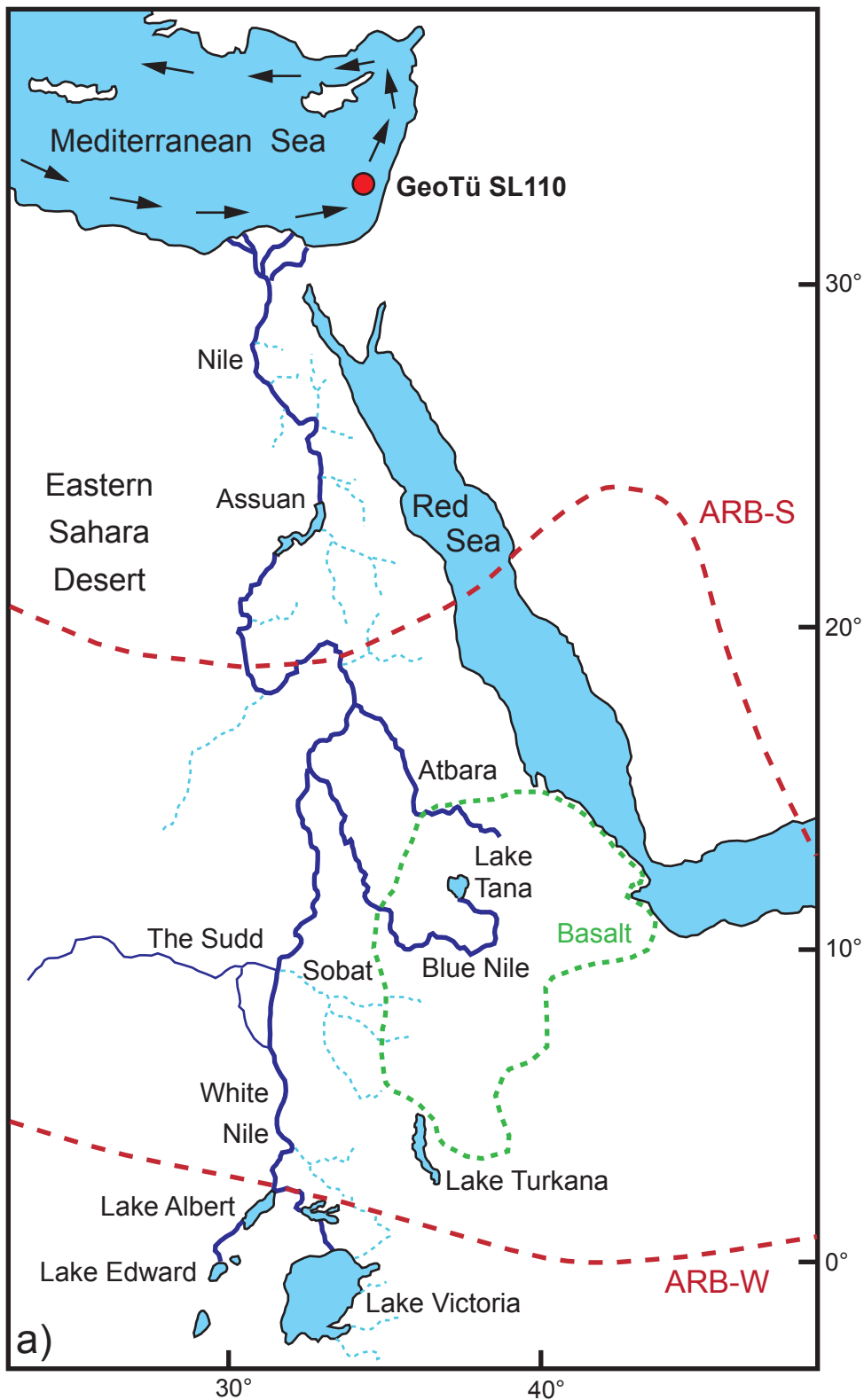
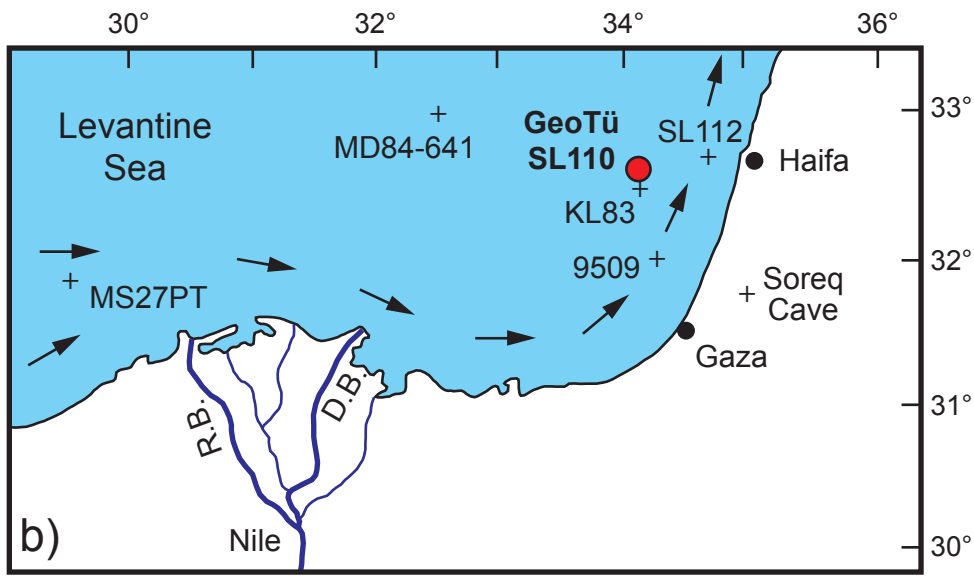


Figure 1
Ehrmann et al.

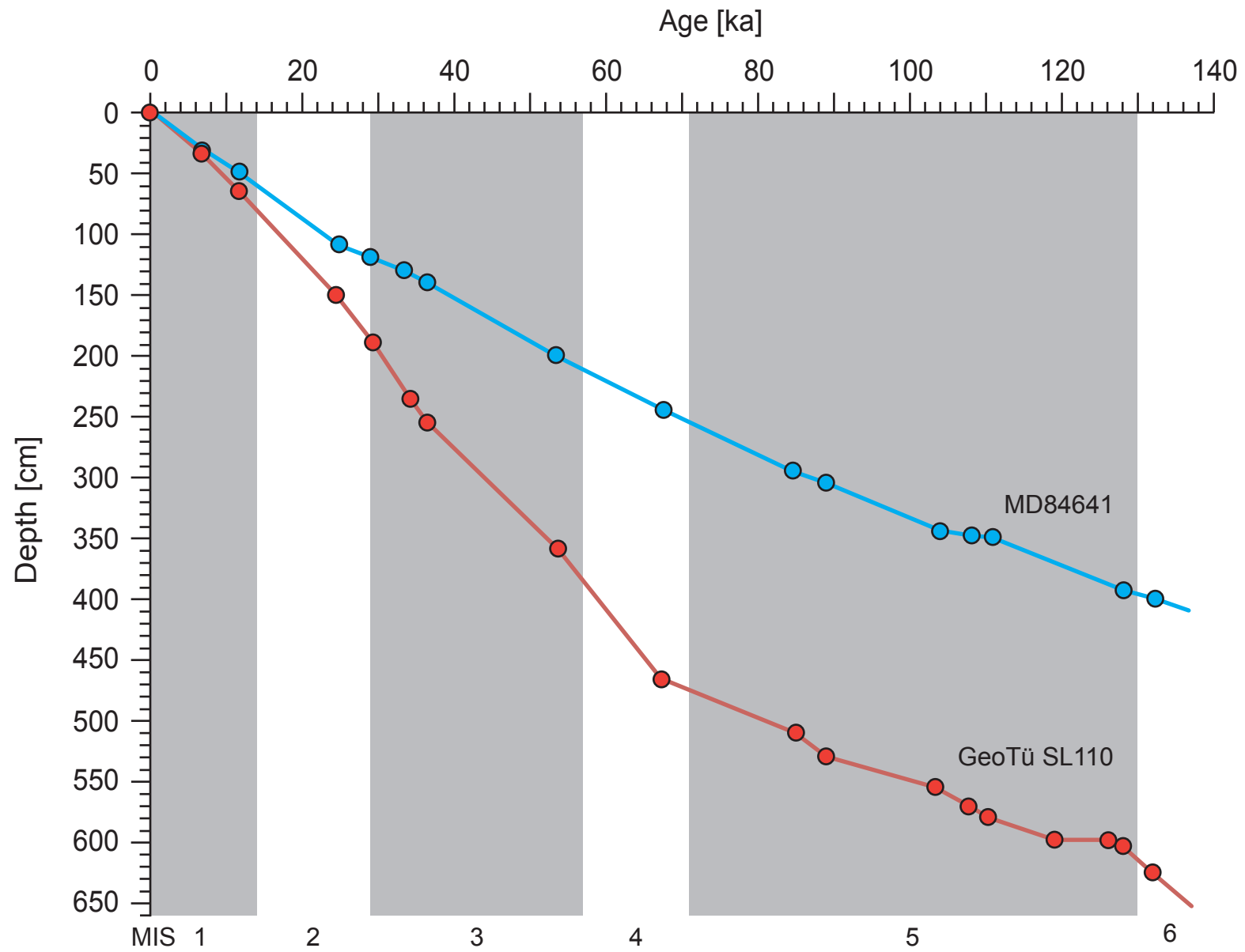


Figure 2 - Ehrmann et al.

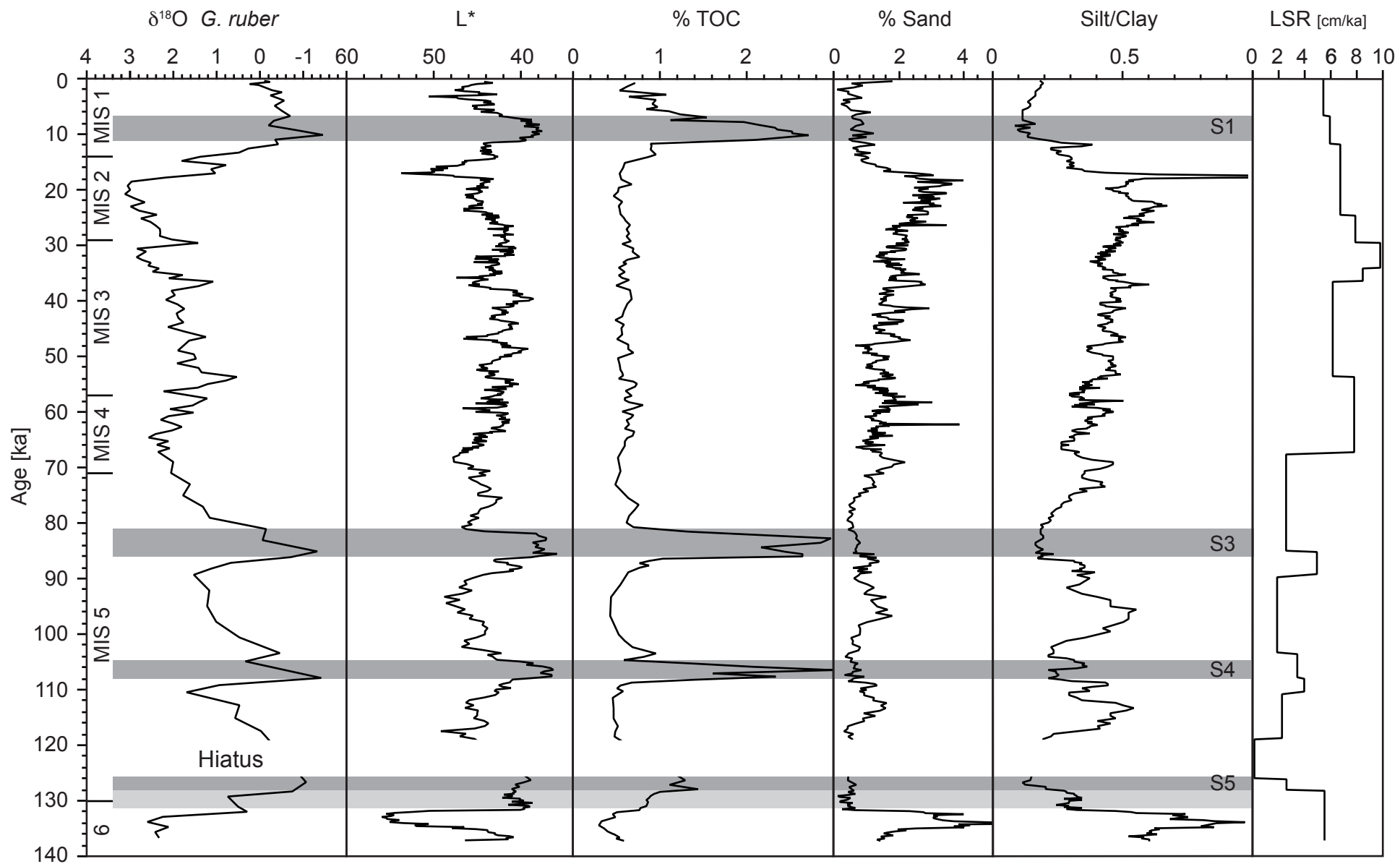


Figure 3a - Ehrmann et al.

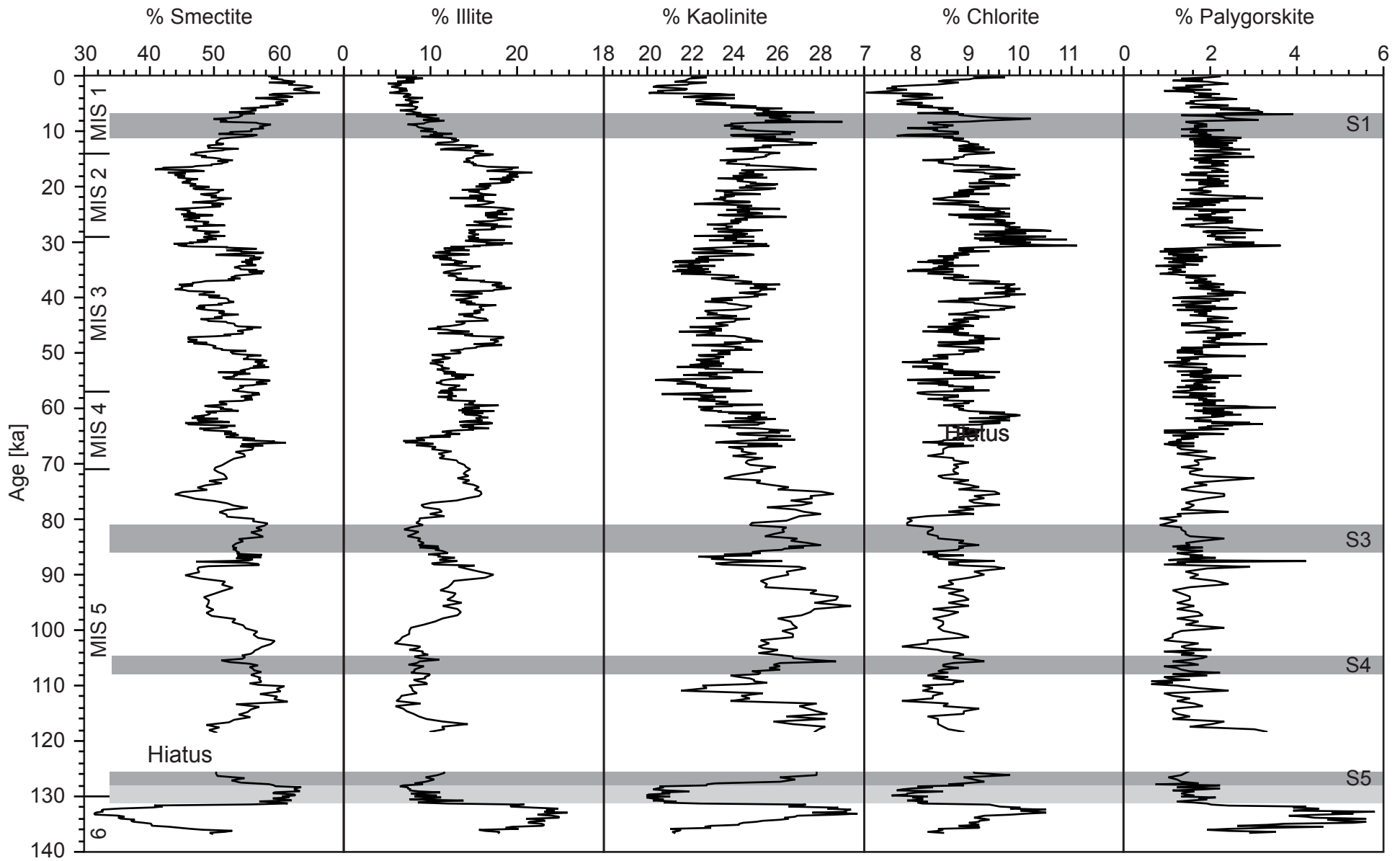


Figure 3b - Ehrmann et al.

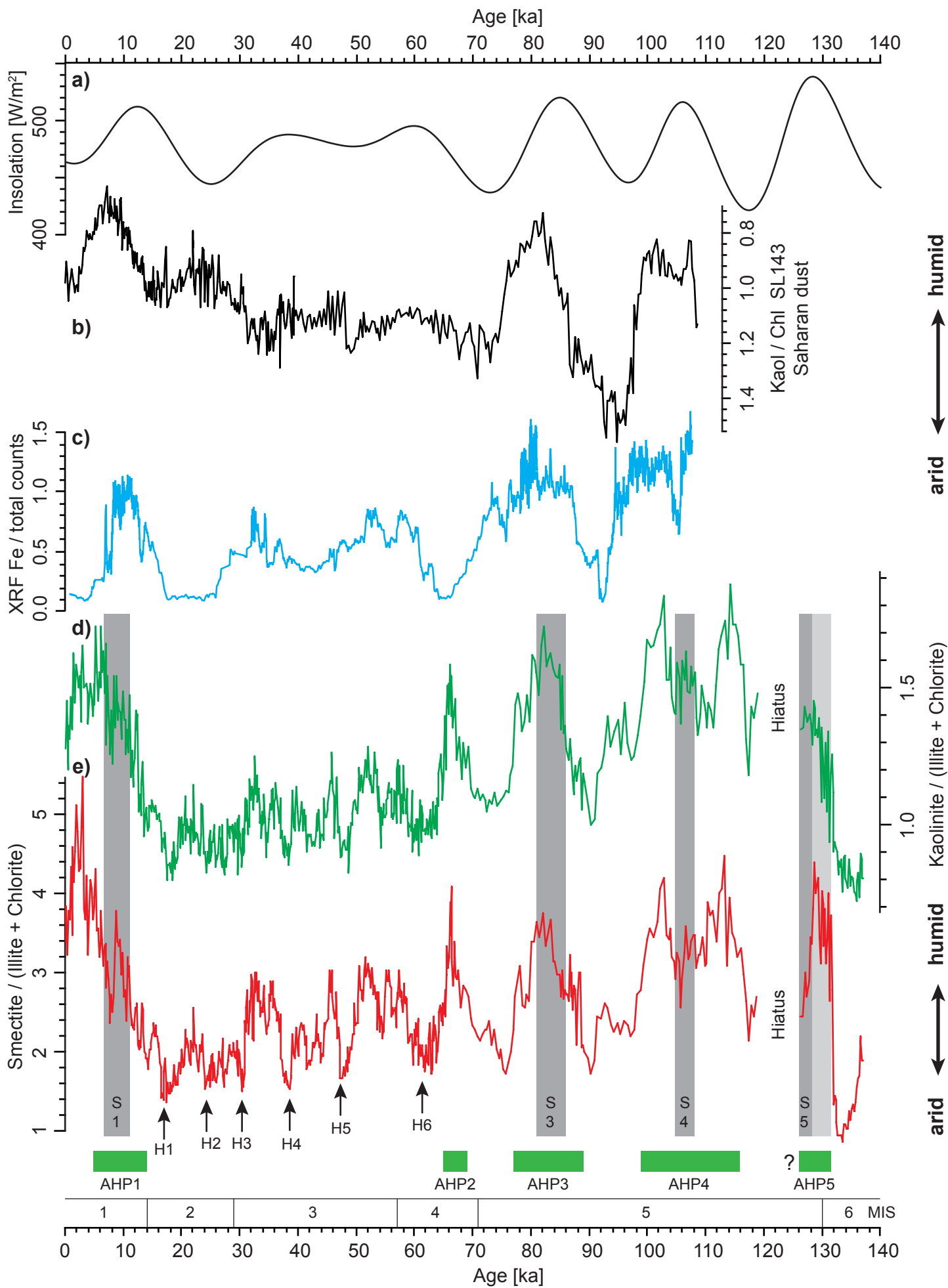


Figure 4 - Ehrmann et al.