

1 **A distal 140 ka sediment record of Nile discharge and**
2 **East African monsoon variability**

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21

1 **Abstract**

2

3 Clay mineral assemblages in a sediment core from the distal Nile discharge plume off
4 Israel have been used to reconstruct the late Quaternary Nile sediment discharge
5 into the Eastern Mediterranean Sea (EMS). The record spans the last ca. 140 ka.
6 Smectite abundances indicate the influence of the Blue Nile and Atbara that have
7 their headwaters in the volcanic rocks of the Ethiopian highlands. Kaolinite
8 abundances indicate the influence of wadis, which contribute periodically to the
9 suspension load of the Nile.

10

11 Due to the geographical position, the climate and the sedimentary framework of the
12 EMS is controlled by two climate systems. The long-term climate regime was
13 governed by the African monsoon that caused major humid periods with enhanced
14 sediment discharge at 132 to <126 ka (AHP5), 116 to 99 ka (AHP4), and 89 to 77 ka
15 (AHP3). They lasted much longer than the formation of the related sapropel layers
16 S5 (>2 ka), S4 (3.5 ka) and S3 (5 ka). During the last glacial period (MIS 4–2) the
17 long-term changes of the monsoonal system were superimposed by millennial-scale
18 changes of an intensified mid-latitude glacial system. This climate regime caused
19 short but pronounced drought periods in the Nile catchment, which are linked to
20 Heinrich Events and alternate with more humid interstadials.

21

22 The clay mineral record further implies that feedback mechanisms between
23 vegetation cover and sediment discharge of the Nile are detectable but of minor
24 importance for the sedimentary record in the southeastern Mediterranean Sea during
25 the investigated African Humid Periods.

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27

28 **1 Introduction**

29

30 The Nile in northern Africa is the longest river of the world and the dominant
31 sediment source for the Eastern Mediterranean Sea (Milliman and Syvitski, 1992). Its
32 drainage basin is about $3 \times 10^6 \text{ km}^2$ (Garzanti et al., 2015) and extends from the
33 equator to ca. 30°N (Fig. 1). The main tributaries of the Nile are the perennial White
34 Nile originating from Lake Victoria in tropical East Africa and the highly seasonal Blue

1 Nile and Atbara originating from the subtropical Ethiopian highlands. Downstream of
2 the confluence, the Nile flows for more than 2000 km through the hyperarid Sahara.

3
4 Because of this special geographical situation, the runoff of the Nile is sensitive to
5 climatic changes, and its discharge sediments are major recorders of the geological
6 and climatic conditions in the catchment areas. The Nile sediments mainly reflect the
7 intensity of rainfall in the headwaters, which has direct control on weathering, erosion
8 and transport of sediments (Krom et al., 2002; Revel et al., 2010; Box et al., 2011;
9 Garzanti et al., 2015). The present-day summer floods of the Nile are linked to a
10 northward movement of the Intertropical Convergence Zone (ITCZ) and the African
11 Rain Belt (ARB; Fig. 1) causing especially intense precipitation in July and August in
12 the Ethiopian highlands.

13
14 Prior to the emplacement of the Assuan High Dam in 1964, the Nile carried a
15 suspension load of about $120\text{--}160 \times 10^6 \text{ t yr}^{-1}$ to the Eastern Mediterranean Sea
16 (EMS) (Holeman, 1968; Milliman and Syvitski, 1992; Stanley and Wingerath, 1996a).
17 More than 95% of the material is derived from the Ethiopian highlands via Blue Nile
18 and Atbara. In contrast, the sediment discharge of the White Nile can be neglected
19 because it's load stored in Lake Victoria, Lake Albert and in the Sudd basin of South
20 Sudan (Adamson et al., 1980; Foucault and Stanley, 1989; Williams et al., 2006;
21 Williams et al., 2015; Garzanti et al., 2015). About 32.5% of the Nile suspension
22 consists of clay (Quelennec and Kruc, 1976). Most of the silt- and sand-sized
23 sediment accumulates in the Nile delta, on the Nile cone and along the
24 Mediterranean shelf, whereas the clay-sized sediment fraction is transported in
25 suspension by the surface currents of the Mediterranean Sea to the east and north
26 (Fig. 1; Venkatarathnam and Ryan, 1971; Foucault and Mélières, 2000; Hamann et
27 al., 2009).

28
29 It is well known that past changes in the amount of rainfall in the Ethiopian highlands
30 were caused by precession-driven shifts in the position of the ITCZ that influenced
31 the intensity and the spatial extent of the monsoon (e.g. Rossignol-Strick, 1983;
32 Emeis et al., 2003). During African Humid Periods (AHPs) the spreading of the
33 vegetation cover in North Africa resulted in a “green” Sahara (Renssen et al., 2006;
34 Kröpelin et al., 2008) and a reduced influx of Saharan dust into the Mediterranean
35 Sea (deMenocal et al., 2000; Ehrmann et al., 2013). The increased runoff of the Nile

1 River and of other North African river/wadi systems brought huge amounts of
2 freshwater and nutrients to the Mediterranean Sea, leading to enhanced productivity
3 in the surface water, stagnation in the deep marine basin and formation of sapropel
4 layers (Rossignol-Strick et al., 1982; Rohling, 1994; Emeis et al., 2003; Rohling et al.,
5 2015). Although numerous studies concentrated on the Holocene AHP, very little is
6 known about timing and intensity of former AHPs. Marine proxy data across sapropel
7 S5 suggest that during the Eemian period (MIS 5e) the African monsoon was
8 particularly strong and the African rain belt was shifted even further north than during
9 the early Holocene. This resulted in flooding into the EMS through wadi systems
10 along the wider North African margin (Rohling et al., 2002a; Osborne et al., 2008).

11
12 Different opinions exist on whether Nile sediment discharge is primarily controlled by
13 precipitation and runoff, and thus closely linked to changes in the monsoon intensity
14 (Wehausen and Brumsack, 2000; Emeis et al., 2000; Revel et al., 2010), or whether
15 feedback mechanisms play an important role, with vegetation protecting soils and
16 weathering products in the source area from being eroded (Adamson et al., 1980;
17 Krom et al., 2002; Box et al., 2011).

18
19 A comprehensive effort to reconstruct Nile discharge during the last 100 ka was
20 undertaken using the Fe content in sediments recovered from the proximal Nile delta
21 (Revel et al., 2010; Caley et al., 2011). Most other geochemical investigations
22 concentrated on shorter time spans and used different parameters such as K/Al,
23 Mg/Al, Ti/Al, Ba/Ca to characterise the Nile discharge (Wehausen and Brumsack,
24 2000; Box et al., 2011; Weldeab et al., 2014). Several studies include the isotopic
25 fingerprint of Nile sediments to reconstruct the Nile history (Krom et al., 1999a;
26 Weldeab et al., 2002; Box et al., 2011; Blanchet et al., 2014).

27
28 We choose a different approach of reconstructing the Nile River sediment supply
29 throughout the last some 140 ka by mainly investigating the content of typical Nile-
30 derived clay minerals in sediment core GeoTü SL110, which was recovered from the
31 distal Nile discharge plume off Israel (Fig. 1). Some 80% of the terrigenous sediment
32 fraction of the surface sediments near SL110 is derived from the Nile (Krom et al.,
33 1999b), and is characterised by a high smectite content (e.g. Venkatarathnam and
34 Ryan, 1971; Foucault and Mélières, 2000). By analysing a distal core from the plume,

1 we document the Nile discharge activity as a whole rather than the activity of only a
2 single channel in the delta.

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5 **2 Material**

6

7 The investigated sediment core GeoTü SL110 was recovered from the SE Levantine
8 Sea (Fig. 1, Table 1) during cruise M51/3 of the German research vessel “*Meteor*” in
9 2001 (Hemleben et al., 2003). It was retrieved at the Israeli continental slope off
10 Haifa from a water depth of 1437 m. The sediments of the 652 cm long core mainly
11 consist of muds with traces of pteropods and foraminifers. The core top contains
12 basically undisturbed surface sediments as indicated by the presence of an oxidised
13 layer. The muds show frequent diffuse colour changes, mainly between grey and
14 olive grey. Darker sapropel layers (greenish black, dark greenish grey, dark grey)
15 occur at 35–61 cm (S1, with a possible interruption at 44–45 cm), 500–514 cm (S3),
16 558–570 cm (S4) and 598–622 cm (S5, including pre-sapropelic layer). The core
17 does not show any indications for slumping, debris flows or turbidites but indicates
18 rather constant sedimentary conditions in the distal discharge plume of the Nile. No
19 tephra layers were detected.

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22 **3 Methods**

23

24 The lightness (L^*) of core SL110 was determined in 1 cm steps with a Minolta colour
25 spectrophotometer immediately after opening of the core.

26

27 The content of total organic carbon (TOC) was measured on 174 ground bulk
28 sediment samples with an Eltra METALYST-CS-1000-S after removal of carbonate
29 with HCl.

30

31 The core was sampled in 1 cm intervals for investigations of the grain size
32 composition of the terrigenous sediment fraction and of the clay mineral
33 assemblages. The samples were oxidized and disaggregated by means of 5%
34 hydrogen peroxide. Carbonate was dissolved by 10% acetic acid. We isolated the

1 terrigenous sand fraction by sieving the samples through a 63 μm mesh. The silt
2 fraction (2–63 μm) was separated from the clay fraction (<2 μm) in settling tubes.

3

4 The analyses of the clay mineral composition followed standard methods (e.g.
5 Ehrmann et al., 2007). We added MoS_2 as an internal standard to the clay
6 suspension. Texturally oriented clay mounts were solvated with ethylene–glycol
7 vapour at a temperature of 60°C and then X-rayed with a Rigaku MiniFlex system
8 ($\text{CoK}\alpha$ radiation; 30 kV; 15 mA). We analysed the samples in the range 3–40° 2θ
9 with a step size of 0.02° 2θ and a measuring time of 2 s step^{-1} . In addition, we
10 analysed the range 27.5–30.6° 2θ with a step size of 0.01° 2θ and a measuring time
11 of 4 s step^{-1} in order to better resolve the (002) peak of kaolinite and the (004) peak
12 of chlorite. We evaluated the diffractograms by using the MacDiff software
13 (Petschick, 2001). After adjusting the diffractograms to the MoS_2 peak at 6.15 Å, we
14 deconvoluted the peak doublets smectite/chlorite (17/14 Å), palygorskite/illite
15 (10.5/10 Å) and kaolinite/chlorite (3.58/3.54 Å). We based the semi-quantitative
16 estimations of the clay mineral abundances on the integrated peak areas of the
17 individual peaks and weighting factors (Biscaye, 1964; 1965; Weaver and Beck,
18 1977).

19

20

21 **4 Age Model**

22

23 For establishing a stable oxygen isotope stratigraphy we analysed tests from the
24 250–355 μm size fraction of the surface-dwelling foraminifer species *Globigerinoides*
25 *ruber* (white) of about 130 samples using a Kiel IV online carbonate preparation line
26 connected to a MAT 253 mass spectrometer. The values are reported in ‰ relative
27 to VPDB using the delta notation. The reproducibility was better than $\pm 0.06\text{‰}$ (1σ).
28 Age points were derived by a graphic correlation of our $\delta^{18}\text{O}$ record with the *G. ruber*
29 (white) $\delta^{18}\text{O}$ data of sediment core LC21 from the southern Aegean Sea (Grant et al.,
30 2012), aided by the software AnalySeries 2.0 (Paillard et al., 1996) (Table 2, Fig. 2).
31 The LC21 chronology was constructed by using ^{14}C dating and by tuning its $\delta^{18}\text{O}$
32 record to the U/Th-dated Soreq Cave speleothem $\delta^{18}\text{O}$ record (Grant et al., 2012).

33

34 Further age control comes from four ^{14}C -accelerator mass spectrometry (AMS)
35 datings performed by Beta Analytic Radiocarbon Dating Laboratory on well-

1 preserved shells of planktonic foraminifera (*G. ruber*, *G. bulloides*, *G. sacculifer*, *O.*
2 *universa*) that represent the age of the surface waters. We applied an eastern
3 Levantine Sea Delta-R of 3 ± 66 years (Marine Reservoir Database). The
4 radiocarbon ages were converted to calendar years using the Marine13 data base
5 (Reimer et al., 2013; Table 2).

6
7 We identified a hiatus at 598 cm, at the top of the S5 interval, based on several
8 arguments. The lower part of the dark interval between 622 and 604 cm is
9 characterised by only about 0.8% TOC (Fig. 3a), by the occurrence of a relatively
10 diverse benthic foraminiferal assemblage containing *G. orbicularis*, *G. translucens*
11 and *C. laevigata*, and by relatively high $\delta^{18}\text{O}$ values (Fig. 3a). Despite the dark colour,
12 this indicates a pre-sapropelic interval at the transition between Marine Isotope
13 Stages (MIS) 6 and 5 rather than a full sapropel (Schmiedl et al., 2003). The TOC
14 concentrations start to rise to ca. 1.4% at 603 cm. We therefore assume the base of
15 sapropel S5 at 603 cm, with an age of 128.0 ka (Grant et al., 2012). The TOC values
16 drop abruptly to 0.5% at 596 cm. Thus, the TOC concentrations remain below the
17 typical values of 2–3% in S1, S3, S4 (Fig. 3a) and in S5 of nearby core KL83
18 (Weldeab et al., 2003). Also the analysis of the benthic foraminifera showed an
19 extremely thin anoxic phase in SL110, compared to a 27 cm thick S5 in KL83
20 (Schmiedl et al., 2003). Thus, most of S5 is missing in SL110 due to a hiatus that
21 was probably caused by a slump as indicated by the sharp and curved lithological
22 boundary at 598 cm. By comparing our $\delta^{18}\text{O}$ record with standard curves and the
23 record of MD84-641 (Table 1, Kallel et al., 2000) and considering an age of 128–121
24 ka for S5 (Grant et al., 2012) we argue for an extrapolated age of the hiatus from
25 126–119 ka. In nearby core GeoTü KL83 (Table 1), Weldeab et al. (2002, 2003)
26 identified a hiatus spanning the period 110–70 ka, with the base of the hiatus at 13
27 cm above S5.

28
29 For consistency reasons, we also adapted the age model of core GeoTü SL143
30 (Ehrmann et al., 2013; Fig. 4) by using basal ages of 86.0 ka and 108.0 ka for
31 sapropels S3 and S4, respectively, according to the LC21 record (Grant et al., 2012)

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1 **5 Results**

2

3 The raw data of our investigations on sediment core GeoTü SL110 are presented in
4 Fig. 3, which also shows the positions of the sapropel layers S1, S3, S4 and S5.

5

6 The $\delta^{18}\text{O}$ data of *G. ruber* go back to MIS 6 (Fig. 3a). The Last Glacial Maximum has
7 values of ca. 3‰, whereas the minimum values during interglacials MIS 1 and MIS 5
8 are around -1.5‰ .

9

10 The lightness data (L^* , Fig. 3a) reflect subtle colour changes within the core. The
11 darkest intervals correlate with the sapropels (L^* ca. 37). However, also large parts of
12 the other muds are relatively dark (L^* ca. 40–45). The lightest interval occurs during
13 the MIS 6 interval 132–135 ka.

14

15 TOC contents generally fluctuate between 0.4 and 0.7%, but increase to 2.5–3% in
16 the sapropel layers S1, S3 and S4, and to 1.5% in the pre-sapropelic layer of S5.

17

18 The terrigenous sand content is generally <4%, the silt/clay ratio <1.0. The data sets
19 show the same general pattern (Fig. 3a) with finer grained intervals enclosing the
20 sapropel layers. A distinct coarse interval occurs at 132–135 ka. Between 73 and 18
21 ka we observe an upward coarsening trend. The high silt/clay ratio at ca. 17 ka
22 reflects a thin, light gray, not graded silt layer consisting almost entirely of quartz
23 grains.

24

25 The clay mineral assemblage of SL110 (Fig. 3b) is dominated by smectite that
26 fluctuates between ca. 30% and 65%. Maximum values occur around 3, 31–37, 46,
27 50–59, 65–69, 77–89, 99–116, and 127–132 ka. Minimum values occur at 132–135
28 ka. Illite shows an opposite distribution pattern with concentrations ranging between
29 5 and 25%. Kaolinite concentration fluctuates between 20% and 30% and shows only
30 a few distinct maxima centred at 95, 116 and 133 ka, and minima at 0–5, 35, 87, 111
31 and 128–132 ka. Chlorite is present in low concentrations of 7–11%, with a pattern
32 roughly resembling that of kaolinite. Palygorskite occurs in trace amounts throughout
33 the core, but jumps to 4% at 132–135 ka.

34

35

1 **6 Discussion**

2

3 Sedimentation at SL110

4

5 According to our age model (Fig. 2, Table 2), sediment core SL110 has an
6 extrapolated basal age of ca. 137 ka. The linear sedimentation rates are 2.2 to 9.8
7 cm/ka (Fig. 3a). They are low during MIS 5, and distinctly higher in late MIS 6 and
8 MIS 4 to MIS 1. Sedimentation rates for nearby core MD84-641, extracted from data
9 by Fontugne and Calvert (1992), show less distinct changes (Fig. 2) and are distinctly
10 lower than in SL110 due to its further offshore position.

11

12 The non-carbonate sediment fraction of seafloor surface sediments in the region of
13 SL110 is dominantly derived by long-distance sediment transport from the Nile
14 (Venkatarathnam and Ryan, 1971; Stanley et al., 1997; 1998; Krom et al., 1999b;
15 Weldeab et al., 2002; Hamann et al., 2009). Storm waves and coastal currents erode
16 coarse sediments from the delta. Geostrophic and wave-induced near-shore and
17 shelf currents displace this material in a large counter-clockwise current along the
18 coast towards the Levantine Sea (Fig. 1) and deposit it in the shallow waters off
19 Israel. The fine fraction, in contrast, is derived from direct input from the Nile into the
20 sea and may be transported also further offshore and further to the north (e.g.
21 Stanley et al., 1997; 1998; Venkatarathnam and Ryan, 1971; Foucault and Mélières,
22 2000). Aeolian influx from the Sahara and riverine supply of small Near East rivers
23 and wadis are only of minor importance (Stanley et al., 1997; Sandler and Herut,
24 2000). The enhanced glacial sedimentation rates observed in core SL110 might be a
25 result of stronger currents and/or a shift of the currents and therewith of the high-
26 accumulation areas to a more seaward position due to a lower sea level. These
27 processes may also be responsible for the coarsening of the sediments as indicated
28 by the terrigenous sand content and the silt/clay ratios (Fig. 3a).

29

30 The increasing glacial sedimentation rates theoretically also could be explained by
31 increasing aridity resulting in an enhanced influx of aeolian dust. Endmember
32 modelling of the terrigenous silt fraction of the nearby core SL112 (892 m water
33 depth), which is located closer to the coast, indicates increased dust fluxes during the
34 LGM and the late glacial (Hamann et al., 2008). However, such an influx is not
35 documented in the clay mineral record of SL110, e.g. by the concentration of wind-

1 blown palygorskite (Fig. 3b). In addition, glacial sedimentation rates at site SL112
2 were lower when compared to the Holocene, thus supporting the interpretation of a
3 glacial-interglacial shift of the high-accumulation zone.

4

5 Clay Mineral Assemblages

6

7 Smectite in southeastern Levantine Sea sediments is generally regarded to be of
8 detrital origin and is the typical clay mineral in Nile-derived sediments
9 (Venkatarathnam and Ryan, 1971; Maldonado and Stanley, 1981; Stanley and
10 Wingerath, 1996a; Foucault and Mélières, 2000). According to the compilation by
11 Hamann et al. (2009), the Nile assemblage contains up to 85% smectite. Kaolinite is
12 the second important clay mineral with concentrations in the order of 20%. Illite
13 contents are <10%, and chlorite occurs in trace amounts.

14

15 The riverine suspension load is reflected in the modern clay mineral assemblage in
16 seafloor surface sediments of the SE Levantine Sea. Accordingly, the clay mineral
17 assemblage off southern Israel consists of up to >70% smectite, 10–25% kaolinite,
18 10% illite and <10% chlorite (Hamann et al., 2009) and thus reflects the Nile source.
19 The uppermost samples of core SL110 with ca. 60% smectite, 22% kaolinite, 8% illite
20 and 9% chlorite fit well into this distribution pattern.

21

22 The composition of the aeolian clay mineral assemblage reaching the SE Levantine
23 Sea is variable, but generally characterized by high illite and kaolinite concentrations
24 (Chester et al., 1977). Palygorskite is a typical clay mineral in aeolian dust derived
25 from the Sahara (Foucault and Mélières, 2000; Scheuven et al., 2013).

26

27 The Nile incorporates almost its entire suspension load close to its source (Adamson
28 et al., 1980). The smectite comes with the Nile suspension load from the catchment
29 of the Blue Nile and Atbara rivers and originates from weathering of Cenozoic
30 volcanic rocks on the Ethiopian Plateau (e.g., Stanley and Wingerath, 1996a). The
31 White Nile south of the confluence of Sobat river, also discharging the Ethiopian
32 highlands, shows a considerably different clay mineral signature with significantly
33 lower smectite contents. Alluvial sediments from central Uganda have smectite
34 contents <20% (Nyakairu and Koeberl, 2001); no smectite was described from lake-
35 bottom sediments of Lake Victoria (Mothersill, 1976). Tropical weathering in the

1 equatorial region produces only small amounts of smectite (Garzanti et al., 2015).
2 Furthermore, the smectite content in SL110 correlates with the Fe content
3 documented by Revel et al. (2010) in delta sediments of the Nile, especially during
4 humid periods. Because Fe is unequivocally derived from the volcanic rocks of the
5 Ethiopian highlands, this also suggests that the overwhelming part of the smectite
6 comes from that source.

7
8 Sediments of the Egyptian wadis discharging into the Nile have clay mineral
9 assemblages with typically 40–55% kaolinite, 35–45% smectite and 5–15% illite
10 (Stanley and Wingerath, 1996b; Hamann et al., 2009; chlorite not determined, thus,
11 actual kaolinite concentrations may be somewhat lower). The kaolinite is mainly
12 derived from the erosion of kaolinite-rich Eocene, Palaeocene and Mesozoic
13 sediments and lateritic soils in the wadis (Stanley and Wingerath, 1996b; Bolle et al.,
14 2000). Theoretically, the White Nile could also deliver kaolinite during certain times of
15 enhanced sediment transport during humid periods (Garzanti et al. 2015). This would
16 result in a decrease in smectite at the same time due to dilution. This, however, is not
17 the case (Fig. 3b). Therefore we argue that the kaolinite accumulating in SL110
18 predominantly comes from the North African wadi regions.

19
20 According to the standard method, percentages of the clay minerals smectite, illite,
21 chlorite, kaolinite and palygorskite add up to 100%. A change in the abundance of
22 one mineral therefore causes also changes in the concentration of the others. Hence
23 we prefer to discuss clay mineral ratios. We use $Sm_r = \text{smectite}/(\text{illite}+\text{chlorite})$ as the
24 main proxy for Nile discharge provided by the Blue Nile and Atbara, and $Ka_r =$
25 $\text{kaolinite}/(\text{illite}+\text{chlorite})$ as an indicator for the contribution of the wadis.

26
27 The good correlation between the Ka_r and the Sm_r data (Fig. 4; $r^2 = 0.73$) confirms
28 that the kaolinite in SL110 is of fluvial origin (Venkatarathnam and Ryan, 1971;
29 Stanley and Wingerath, 1996a; Hamann et al., 2009), and not derived by aeolian
30 influx from the Sahara, as in other regions of the EMS (e.g. Chester et al., 1977;
31 Foucault and Mélières, 2000; Ehrmann et al., 2013). The correlation implies that
32 during times of enhanced Nile runoff and smectite transport from Ethiopia, the
33 erosion of kaolinite-rich sediments and soils in the wadis was generally also active,
34 and that during dry periods in Ethiopia the wadis were dry, too.

35

1 This confirms earlier findings based on the interpretation of satellite images and
2 sedimentological investigations, which showed that extensive drainage systems were
3 active in the eastern Sahara during humid phases, especially of Eemian and
4 Holocene time (e.g. Pachur and Hoelzmann, 2000; Rohling et al., 2002b; Osborne et
5 al., 2008). Furthermore, oxygen isotope data from speleothems of Soreq Cave (Bar-
6 Matthews et al., 2000) imply that rainfall extended beyond the Sahara.

7

8 The clay mineral data of SL110 document a highly variable smectite and kaolinite
9 input through time (Figs. 3b, 4) and show similar distribution patterns as the Fe
10 record from the Nile delta (Revel et al., 2010; Caley et al., 2011). Fe is supposed to
11 be derived from weathering of volcanic rocks in the Atbara and Blue Nile headwaters
12 in the Ethiopian highlands, and thus changing Fe contents may reflect changing
13 precipitation regimes. Furthermore, the smectite and kaolinite distribution patterns
14 are inversely correlated to the record of Saharan dust in the central Aegean Sea core
15 GeoTü SL143 (Ehrmann et al., 2013).

16

17 Penultimate Glacial Period

18

19 The oldest part of the record displays minimum Sm_r and Ka_r values at ca. 135-132 ka
20 of MIS 6 (Fig. 4). The lack of sediment derived from the Nile delta indicates that the
21 site was not within the range of the Nile deep-sea turbidite system postulated for MIS
22 6 at ca. 33°30' to 33°50'E (Ducassou et al. 2009). The clay mineral data point to
23 extremely dry conditions for the penultimate glacial maximum. The extreme drought
24 possibly correlates with Heinrich Event 11, which is known as a very cold period in
25 the Western Mediterranean Sea and dated to 135-130 ka (Martrat et al., 2014;
26 Marino et al., 2015). Obviously, only minor amounts of Nile suspension load reached
27 the coring site off Israel. Instead, maximum palygorskite concentrations (Fig. 3b) and
28 high proportions of terrigenous silt and sand (Fig. 3a) document an increased aeolian
29 influx from the Sahara. This arid phase seems to have been more severe than that of
30 the Last Glacial Maximum. After the maximum drought of MIS 6 the record of SL110
31 show a sharp transition to a sapropelic event embedded in an AHP.

32

33

1 Sapropel Formation

2

3 It is well known that sapropel formation was coupled to precession-related maxima in
4 northern hemisphere summer insolation causing a northward shift of the ITCZ, an
5 intensification of the African monsoon system and enhanced precipitation in the
6 Ethiopian highlands (e.g. Rossignol-Strick, 1983; Rohling, 1994; Emeis et al., 2003;
7 Rohling et al., 2015). In SL110, high Sm_r , Ka_r and low silt/clay ratios occur mainly
8 during interglacial periods with negative $\delta^{18}O$ values and indicate phases with
9 enhanced Nile and wadi discharge of fine-grained suspension to the Levantine Sea
10 accompanying formation of sapropels S5, S4, S3 and S1. The enhanced rainfall
11 under a warm climate caused stronger chemical weathering of the volcanic rocks and
12 thus smectite formation. The enhanced river runoff contributed to the stratification of
13 the water column and facilitated the transport of nutrients and weathering products to
14 the EMS promoting a shoaling of the nutricline and thus development of a deep
15 chlorophyll maximum (Rohling and Gieskes, 1989; Grelaud et al., 2012) particularly
16 in areas close to the Nile delta. These processes fostered sustained anoxic
17 conditions in the deeper basins and preservation of TOC (Fig. 3a).

18

19 Both the clay mineral record of southeastern Levantine Sea core SL110, the Fe
20 record from the Nile delta (Revel et al., 2010; Caley et al., 2011) and the record of
21 Saharan dust influx to the Aegean Sea (Ehrmann et al., 2013) reveal that each phase
22 of sapropel formation occurred within a longer AHP (Fig. 4; AHPs numbered
23 according to the corresponding sapropels).

24

25 The last Interglacial Period

26

27 The Sm_r maximum linked to AHP 5 shows a sharp onset after the glacial maximum of
28 MIS 6, simultaneously with the start of the pre-sapropelic layer of S5 (Fig. 4). It
29 indicates a northward moving rain belt resulting in enhanced monsoonal rainfall in
30 Ethiopia and Blue Nile runoff as early as 132 ka at a time of increasing insolation but
31 under the influence of penultimate glacial boundary conditions. Also the Ka_r values
32 clearly increase with the beginning of the pre-sapropelic layer. However, they do not
33 exhibit such distinct maximum values as during the early phases of the younger
34 AHPs. Most of sapropel S5 and the termination of AHP 5, however, are not
35 preserved in SL110 due to the hiatus at ca. 126–119 ka. Furthermore, because of the

1 curved top of the sapropel and the technique of sampling the core in parallel slices,
2 we did not receive the typical sapropel signature but a mixed sapropel/post-sapropel
3 signal (Fig. 4).

4

5 According to the Fe and the dust record (Revel et al., 2010; Ehrmann et al., 2013),
6 AHP 4 started at >108 ka and lasted until 95 ka or 98 ka, respectively. The Sm_r and
7 Ka_r data of SL110 indicate that AHP 4 started during insolation rise at 116 ka and
8 ended at 99 ka, with sapropel formation occurring at 108–104.5 ka (Fig. 4).

9

10 AHP 3 started just before the insolation maximum and lasted ca. 89–77 ka based on
11 the Sm_r and Ka_r maxima and the minima in the silt/clay ratio and the terrigenous
12 sand content. These ages are in good accordance with those obtained from the dust
13 record in the Aegean Sea (Ehrmann et al., 2013). Sapropel S3 has a shorter duration
14 and spans the time interval 86–81 ka.

15

16 Despite the hiatus in the SL110 record, by assuming a duration of 128–121 ka for S5
17 (Grant et al., 2012), our data indicate that the dry phase between AHP 5 and AHP 4
18 lasted only ca. 5 ka, compared with ca. 11 ka for the one between AHP 4 and AHP 3.
19 It is interesting to note that a similar duration of the dry phases can be deduced from
20 the pollen record of Tenaghi Philippon in Greece, although this remote site is
21 governed dominantly by northern climatic factors (Tzedakis, 2005; Pross et al.,
22 2015).

23

24 The last Glacial Period

25

26 After a further short dry period at 77–70 ka, Sm_r , Ka_r and silt/clay ratios indicate
27 humid phase AHP 2 and enhanced Nile runoff in early MIS 4 at 69–65 ka. It is of
28 similar magnitude as the interglacial AHPs discussed before and also starts during
29 increasing insolation (Fig. 4). In contrast to the previous AHPs this one is not
30 accompanied by low $\delta^{18}O$ values. It is also not accompanied by a sapropel, dark
31 sediment colour or enhanced TOC values (Fig. 3a). However, it coincides with the
32 common occurrence of shallow infaunal *Uvigerina peregrina*. This highly
33 opportunistic species is adapted to the seasonal deposition of high amounts of fresh
34 phytodetritus (Koho et al., 2008). In analogy to nearby SL112, the occurrence of this
35 species at site SL110 during AHP 2 is likely fuelled by pronounced algal blooms

1 related to seasonally enhanced Nile runoff and nutrient supply (Schmiedl et al.,
2 2010). However, in core SL110 we have strong indications that AHP 2 did not
3 culminate in a sapropel formation. Sediment and nutrient delivery by the Nile was
4 cut-off by a drought linked to Heinrich Event 6 occurring just at the time of maximum
5 insolation (Fig. 4). This drought, possibly combined with stronger winds and better
6 ventilation of the deep waters, reduced sediment discharge into the Mediterranean
7 Sea and hampered stagnation of the deep water and sapropel formation.

8

9 Lourens et al. (1996) described a thin sapropel layer S2, corresponding to AHP 2,
10 south of Crete that occurs after the 60 ka insolation maximum. It was argued that this
11 sapropel is not visible any more in most sedimentary sequences of the eastern
12 Mediterranean Sea because of post-depositional oxidation of the organic matter
13 (Emeis et al., 2000; Löwemark et al., 2006).

14

15 It is not understood why our clay mineral signal has no counterpart in the Fe signal
16 from the Nile delta (Revel et al., 2010). Possibly Nile discharge to the Eastern
17 Mediterranean Sea at that time was not via the western Rosetta branch, but via the
18 eastern Damietta branch, and thus site MS27PT starved of sediment. It also cannot
19 be ruled out that a match is obscured because of different age models, insofar as the
20 Fe 73 ka peak corresponds to our well-dated 67 ka clay mineral peak (Fig. 4).

21

22 The slightly enhanced Sm_r and Ka_r values at 59–50 ka possibly document the late
23 phase of AHP 2 after the termination of Heinrich Event 6, reflecting the re-occurrence
24 of slightly enhanced Nile River runoff. Other indications for somewhat increased
25 temperature, humidity and Nile influx during this time interval come from the Fe
26 record in the Nile delta (Revel et al., 2010) and the $\delta^{18}O$ records of Soreq Cave and
27 of sediment cores 9509 and 9501 in the eastern Levantine Sea (Almogi-Labin et al.,
28 2009). The latter records show a marked drop in planktonic foraminiferal $\delta^{18}O$ values
29 between 58 and 49 ka BP. The signal of warm and freshened surface waters was
30 particularly expressed in core 9501 from the northern part of the basin and coincided
31 with D-O interstadial 14. The spread of deciduous and other trees during the same
32 time period likely supported the migration of anatomically modern humans from
33 Africa into the Levant (Müller et al., 2011). In the Sm_r data of core SL110, this time
34 interval documents only a moderate increase of Nile runoff (Fig. 4) suggesting that

1 the observed warming and freshening of the eastern Mediterranean region was
2 predominantly related to a high-latitude climate change.

3

4 Thus, the sediments of SL110 document that the long-term climate regime governed
5 by the African monsoon and characterized by pronounced AHPs during the last
6 interglacial period was replaced as of the occurrence of Heinrich Event 6 by a
7 distinctly different glacial climate system governed by short-term, millennial-scale
8 climate changes having their origin in the North Atlantic and Arctic regions. Former
9 studies showed that sudden freshwater influxes into the northern Atlantic Ocean
10 during the Younger Dryas and the Heinrich Events not only weakened the
11 thermohaline circulation and oceanic heat transport to the northern high latitudes
12 (e.g. Bond et al., 1993), but also had major consequences for the Mediterranean Sea
13 environment. Increased aridity and even North African mega-droughts during such
14 events can be seen in marine sediments of the North Atlantic Ocean (Mulitza et al.,
15 2008), speleothem data and lake level drops of Lake Lisan (palaeo-Dead Sea; Bar-
16 Matthews et al., 1999; Bartov et al., 2003) and of African lakes (Gasse et al., 2008;
17 Shanahan et al., 2015). Cold water influx from the North Atlantic Ocean to the
18 Mediterranean Sea intensified. Sea surface temperatures and deep-water
19 temperatures decreased and deep-water ventilation fortified in the northwestern
20 Mediterranean Sea (Cacho et al., 1999; 2000) and probably also in the EMS.
21 Enhanced northward flux of wind-transported sediment from the Sahara is
22 documented for both the Western and the Eastern Mediterranean Sea (Moreno et al.,
23 2002; Bout-Roumazelles et al., 2007; Hamann et al., 2008). Strongly decreased Sm_r
24 and Ka_r values in core SL110 (Fig. 4) indicate reduced Nile sediment discharge
25 during Heinrich Events, and therewith major droughts in the headwaters of the Atbara
26 and Blue Nile and diminished erosion and river runoff. During these events the ITCZ
27 and the African Rain Belt moved southward. Both Lake Tana, the source of the Blue
28 Nile, and Lake Victoria, the source of the White Nile, desiccated during the Younger
29 Dryas and the Heinrich Event 1 (Lamb et al., 2007; Marshall et al., 2011). Minimum
30 Sm_r and Ka_r values as well as the occurrence of a silt layer of probably aeolian origin
31 (Fig. 3a) imply that Heinrich Event 1 was the driest interval since the MIS 6 glacial
32 maximum (cf. Stager et al., 2011).

33

34 The 65°N insolation curve shows a further low-amplitude maximum at ca. 37 ka. It is
35 also reflected by slightly enhanced smectite and kaolinite abundances in core SL110,

1 ranging 37–31 ka and by the Fe record in the Nile delta (Revel et al., 2010). No
2 sapropel formation was recorded following the weak insolation maximum at 37 ka.
3 Thus, the two potential humid periods linked to the glacial insolation maxima at 60
4 and 37 ka are only vaguely expressed in the EMS, because their effects were much
5 less distinct and because they were suppressed by the glacial climate regime,
6 particularly the impacts of cold and dry stadials.

7

8 The last African Humid Period and the Holocene

9

10 The clay mineral data of SL110 and of nearby core SL112 (Hamann et al., 2009)
11 document a gradual onset of AHP 1 at about 14 ka (Fig. 4), which coincides with the
12 overflow of Lake Tana, the source of the Blue Nile (Lamb et al., 2007; Marshall et al.,
13 2011). At the same time, Saharan dust influx to the EMS decreased (Ehrmann et al.,
14 2013). S1 at ca. 11–6.5 ka and the peak Blue Nile runoff at 10–9 ka (Weldeab et al.,
15 2014) are only poorly reflected in the clay mineral data. Precipitation generally
16 declined gradually after sapropel formation throughout the Holocene, the Nile River
17 runoff responded quasi-linearly to the changes in rainfall (Weldeab et al., 2014;
18 Blanchet et al., 2014), and the AHP 1 ended at ca. 5 ka (deMenocal et al., 2000).
19 However, the clay mineral data do not mirror this trend but show continued
20 increasing Sm_r and Ka_r values and Nile sediment discharge until ca. 2 ka. Sediment
21 discharge thus is decoupled from water runoff. Possibly, human activity altered the
22 natural signal. Saharan population increased shortly after 11 ka and reached a
23 maximum at 7.5 ka (Manning and Timpson, 2014). Major land-use changes occurred
24 with the advent of Neolithic farming some 8 ka ago (Williams, 2009).

25

26 Feedback Effects

27

28 The clay mineral and grain size data of SL110 indicate an enhanced riverine influx of
29 suspension from the Blue Nile and Atbara to the Mediterranean Sea during the AHPs
30 and therefore document enhanced weathering of the Ethiopian basalts, erosion of
31 soils and transport of the suspension load to the sea. This view is supported by
32 several investigations (Wehausen and Brumsack, 1998; Emeis et al., 2000; Revel et
33 al., 2010; Blanchet et al., 2014). However, it is in conflict with some other studies
34 (Adamson et al., 1980; Krom et al., 1999a; 2002; Box et al., 2011) that argued for a
35 lower Blue Nile sediment discharge during the warm and humid Holocene AHP and

1 higher fluxes during dry climate phases. They reasoned that at present the limited
2 vegetation cover on the Ethiopian highlands allows extensive erosion when the
3 summer monsoon rains fall. During increased monsoon activity, in contrast, longer
4 periods of rain, longer growing seasons and larger vegetation-covered areas reduced
5 soil erosion.

6
7 Our data indicate that an increase in precipitation and Nile suspension delivery
8 happened long before sapropel formation, during phases of increasing insolation.
9 However, in the case of AHP 4 Sm_r and Ka_r maxima occur just before and after
10 sapropel S4 formation, each accompanied by low silt/clay ratios also indicating
11 enhanced riverine suspension (Fig. 3a). Thus, maximum riverine sediment influx
12 does not coincide with maximum water discharge during sapropel formation. This is
13 possibly due to a feedback mechanism as postulated, for example, by Adamson et
14 al. (1980) and Krom et al. (2002). Vegetation feedbacks were obviously less
15 expressed during glacial conditions, particularly during the last glacial maximum. This
16 is linked to the generally more arid boundary conditions during the last glacial in
17 tropical and northern Africa (Gasse, 2000) and resulting sparse vegetation cover,
18 particularly in the deflation areas of the Sahara (Tjallingii et al., 2008). The rapid
19 climate changes linked to Dansgaard-Oeschger cycles and Heinrich Events find
20 immediate response in the sediment delivery to the EMS.

21
22 The different timing between the Saharan dust record and the Nile discharge record
23 of the humid phases (Fig. 4) may reflect regional differences in the climatic history.
24 The northward migrating tropical rain belt may first have reached the Ethiopian
25 highlands, where the smectite signals has its origin, and with a time lag the central
26 Sahara, there inhibiting effective dust production. The differences may also be due to
27 different processes being at work. The Nile discharge is mainly driven by annual
28 floods, fluvial erosion and weathering. It therefore may have reacted immediately on
29 a northward shift of the ITCZ and the rain belt. The dust record, in contrast, is mainly
30 a function of aridity, wind strength and, probably most important, the availability of
31 deflatable material. The spreading and retreat of a vegetation cover plays an
32 important role and may react non-linearly and with a different time lag on a shift of
33 the ITCZ than river discharge (Claussen et al., 2013; Ehrmann et al., 2013). Thus,
34 recently Blanchet et al. (2014) could show for the end of AHP 1 that the vegetation

1 retreat from the Sahara was much faster than the decrease in Nile runoff during the
2 southward migration of the rain belt.

3

4

5 **7 Conclusions**

6

7 – Clay minerals are suitable tools to reconstruct the late Quaternary Nile sediment
8 discharge to the EMS. The abundance of smectite reflects the runoff of the Blue Nile
9 and Atbara, whereas kaolinite abundance reflects the discharge from wadis. The
10 correlation of smectite and kaolinite abundances implies that during times of
11 enhanced Nile suspension discharge the wadis also were active sediment sources.

12

13 – High smectite and kaolinite abundances and low terrigenous silt/clay ratios indicate
14 phases of enhanced precipitation and suspension influx from the Blue Nile, Atbara
15 and the wadis to the Levantine Sea, the so-called African Humid Periods (AHPs),
16 which are accompanied by the formation of sapropel layers. This largely contradicts
17 the view that sediment delivery during humid periods was reduced because an
18 extensive vegetation cover hampered soil erosion.

19

20 – Precipitation and suspension discharge started much earlier and ended later than
21 sapropel formation. Maximum sediment discharge does not coincide with maximum
22 water discharge during times of sapropel formation because of minor vegetation
23 feedback.

24

25 – The time lag between the increase/decrease of Nile suspension discharge and the
26 decrease/increase in the dust record (SL143) can be explained by a successive
27 migration of the rain belt, possibly combined with a vegetation feedback.

28

29 – The dry periods between the interglacial AHPs were relatively short, each lasting
30 only ca. 5–10 ka.

31

32 – The sedimentary record of core SL110 demonstrates the interplay of two
33 competing climate systems. The long-term climate regime was governed by the
34 African monsoon and characterized by pronounced AHPs during the last interglacial
35 period. This system was replaced in the last glacial by short-term, millennial-scale

1 climate changes having their origin in the North Atlantic region. They are
2 characterised by short but pronounced drought periods linked to Heinrich Events and
3 more humid interstadials. Precipitation in the Nile catchment, river runoff and
4 sapropel formation can be hampered or even inhibited under the influence of the
5 North Atlantic climate system, particularly under glacial boundary conditions.

6

7

8 *Data Availability:* The raw data to this article are stored in the PANGAEA data base
9 at the Alfred Wegener Institute for Polar and Marine Research, Bremerhaven,
10 Germany (<http://doi.pangaea.de/10.1594/PANGAEA.848291>).

11

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19

20

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- 8

1 **TABLE CAPTIONS**

2

3 **Table 1.** Metadata for the investigated sediment core GeoTü SL110 and other cores
4 mentioned in the text (cf. Fig. 1).

5

6 **Table 2.** Data used for constructing the age model for the investigated sediment core
7 GeoTü SL110. Oxygen isotope ages are based on a graphical correlation with the
8 LC21 record (Grant et al., 2012) from the southern Aegean Sea.

9

10

11 **FIGURE CAPTIONS**

12

13 **Figure 1.** Location map of the Nile River basin and the Eastern Mediterranean Sea.

14 **(a)** The Nile River and its main tributaries. Blue dashed lines indicate wadis. The
15 green dotted line shows the main outcrop of Cenozoic basalts in Ethiopia. Also
16 shown are the northern summer and winter limits of the African rain belt (red dashed
17 lines, ARB-S, ARB-W, after Blanchet et al., 2014). Black arrows in the Eastern
18 Mediterranean Sea indicate the surface currents (after Pickard and Emery, 1990). **(b)**
19 Map of the southeastern Levantine Sea with the location of the investigated sediment
20 core GeoTü SL110 and other cores mentioned in the text; not shown are GeoTü
21 SL143 and LC21, which are positioned further to the north (cf. Table 1). D.B. =
22 Damietta branch of the Nile in the delta, R.B. = Rosetta branch.

23

24 **Figure 2.** Depth-Age plot for the investigated sediment core GeoTü SL110 (red). The
25 age points are listed in Table 2. For comparison, corresponding data were extracted
26 for core MD84-641 (blue, Fontugne and Calvert, 1992: Table 1). Marine Isotope
27 Stages (MIS) follow Lisiecki and Raymo (2005).

28

29

1 **Figure 3.** Basic data of core GeoTü SL110. **(a)** $\delta^{18}\text{O}$ of *G. ruber*, lightness (L^*), %
2 total organic carbon (TOC), and % sand and silt/clay ratio of the terrigenous
3 sediment fraction, and linear sedimentation rates (LSR, cm/ka). **(b)** clay mineral
4 percentages of smectite, illite, kaolinite, chlorite and palygorskite. Gray bars mark the
5 position of sapropel layers S1, S3, S4, S5 and the pre-sapropelic layer beneath S5
6 (light grey). Marine Isotope Stages (MIS) are indicated.

7

8 **Figure 4.** Combination of the June insolation at 65°N **(a)**, the Saharan dust record in
9 the central Aegean Sea core GeoTü SL143 **(b;** Ehrmann et al., 2013), the Nile
10 discharge based on Fe data in the Nile delta **(c;** Revel et al., 2010, age model by
11 Caley et al., 2011), and the Nile sediment discharge based on K_a_r **(d)** and Sm_r **(e)** at
12 site GeoTü SL110. Grey bars indicate the positions of sapropels S1, S3, S4, S5 and
13 of the pre-sapropelic layer (light grey) beneath S5 in core in SL110. Heinrich Events
14 (H) are flagged with arrows. Marine Isotope Stages (MIS) and African Humid Periods
15 (AHP) are indicated at the bottom.

Table 1. Metadata for the investigated sediment core GeoTü SL110 and other cores mentioned in the text (cf. Fig. 1).

Core	Latitude (°N)	Longitude (°E)	Water depth depth (m)	Reference
GeoTü SL110	32°38.95'	34°06.22'	1437	this study
GeoTü KL83	32°36.87'	34°08.89'	1431	Weldeab et al. (2002)
GeoTü SL112	32°44.52'	34°39.02'	892	Hamann et al. (2008, 2009)
GeoTü SL143	38°15.71'	25°06.19'	665	Ehrmann et al. (2013)
MD84-641	33°02'	32°38'	1375	Fontugne and Calvert (1992)
9501	34°32'	33°59'	980	Almogi-Labin et al. (2009)
9509	32°01'	34°16'	884	Almogi-Labin et al. (2009)
MS27PT	31°47.90'	29°27.70'	1389	Revel et al. (2010)
LC21	35°40'	26°35'	1522	Grant et al. (2012)

Table 2. Data used for constructing the age model for the investigated sediment core GeoTü SL110. Oxygen isotope ages are based on a graphical correlation with the LC21 record (Grant et al., 2012) from the southern Aegean Sea.

Depth cm	Age cal. ka BP	Data
0.00	0.00	Sediment surface
34.50	6.43	¹⁴ C AMS dating
64.50	11.53	¹⁴ C AMS dating
150.50	24.38	¹⁴ C AMS dating
189.11	29.29	LC21
234.50	33.92	¹⁴ C AMS dating
254.50	36.29	LC21
359.54	53.50	LC21
465.30	67.11	LC21
509.55	84.99	LC21
529.23	89.03	LC21
554.51	103.32	LC21
569.46	107.79	LC21
579.50	110.38	LC21
598.00	119.00	Hiatus end
598.00	126.00	Hiatus start
603.00	128.00	LC21
624.52	131.94	LC21

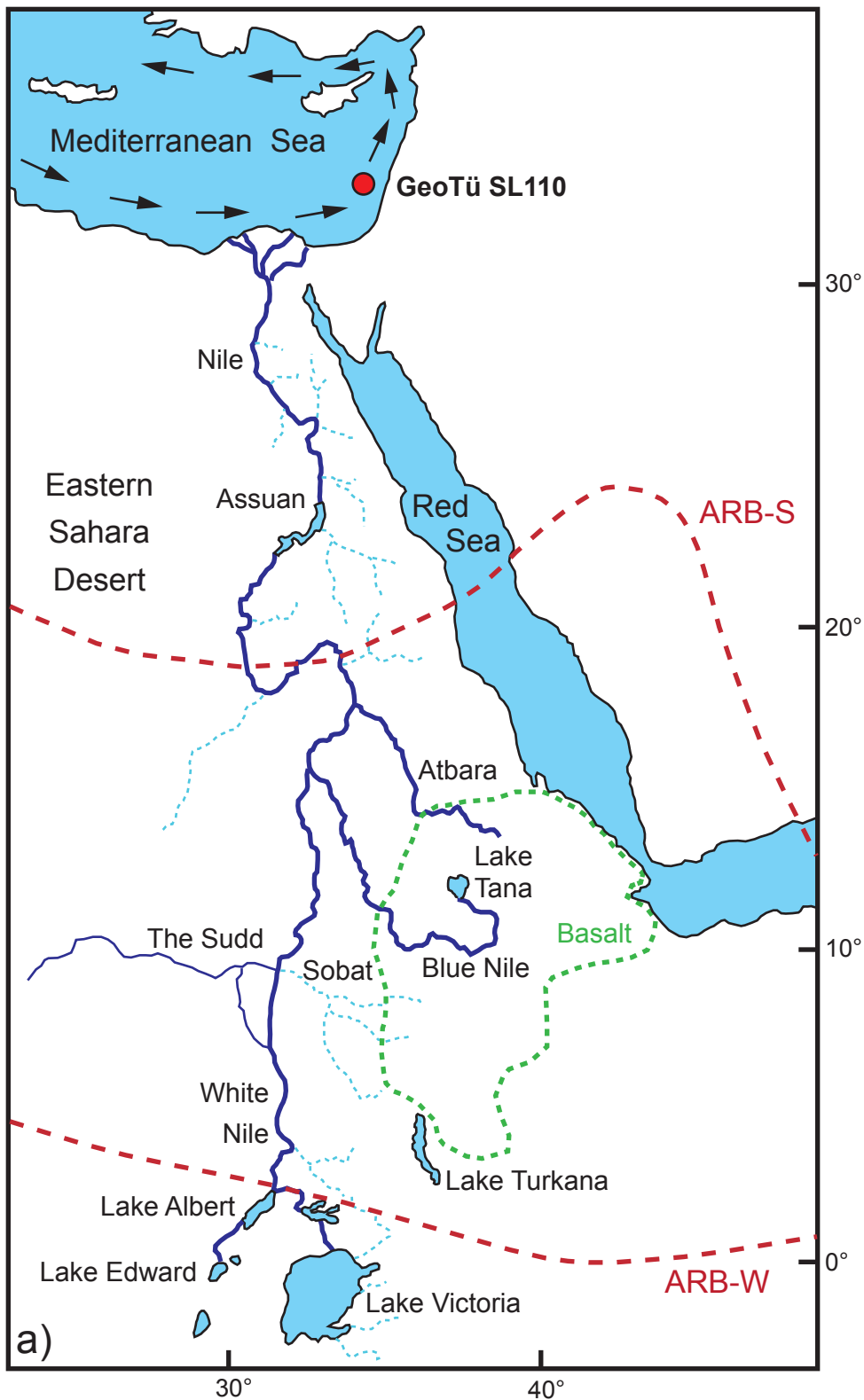
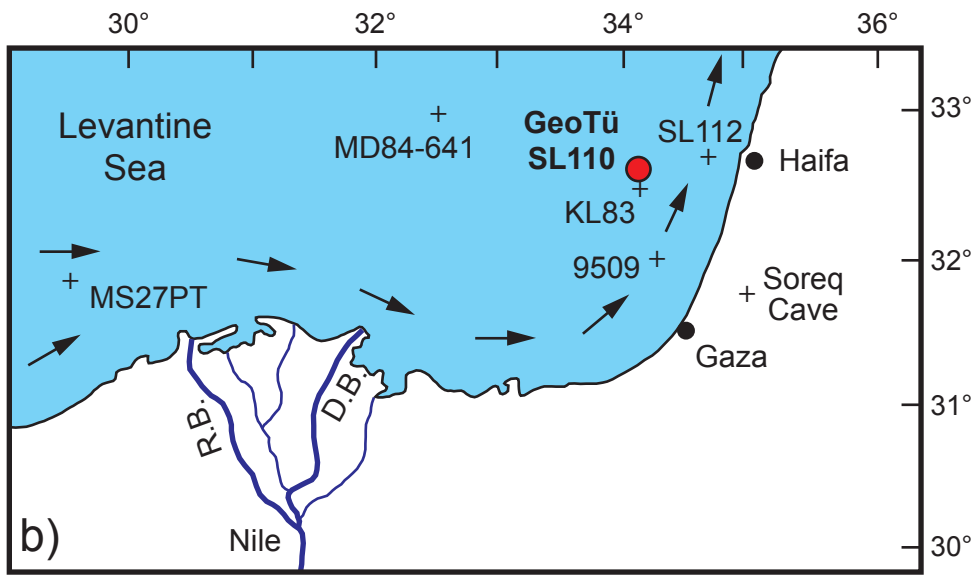


Figure 1
Ehrmann et al.

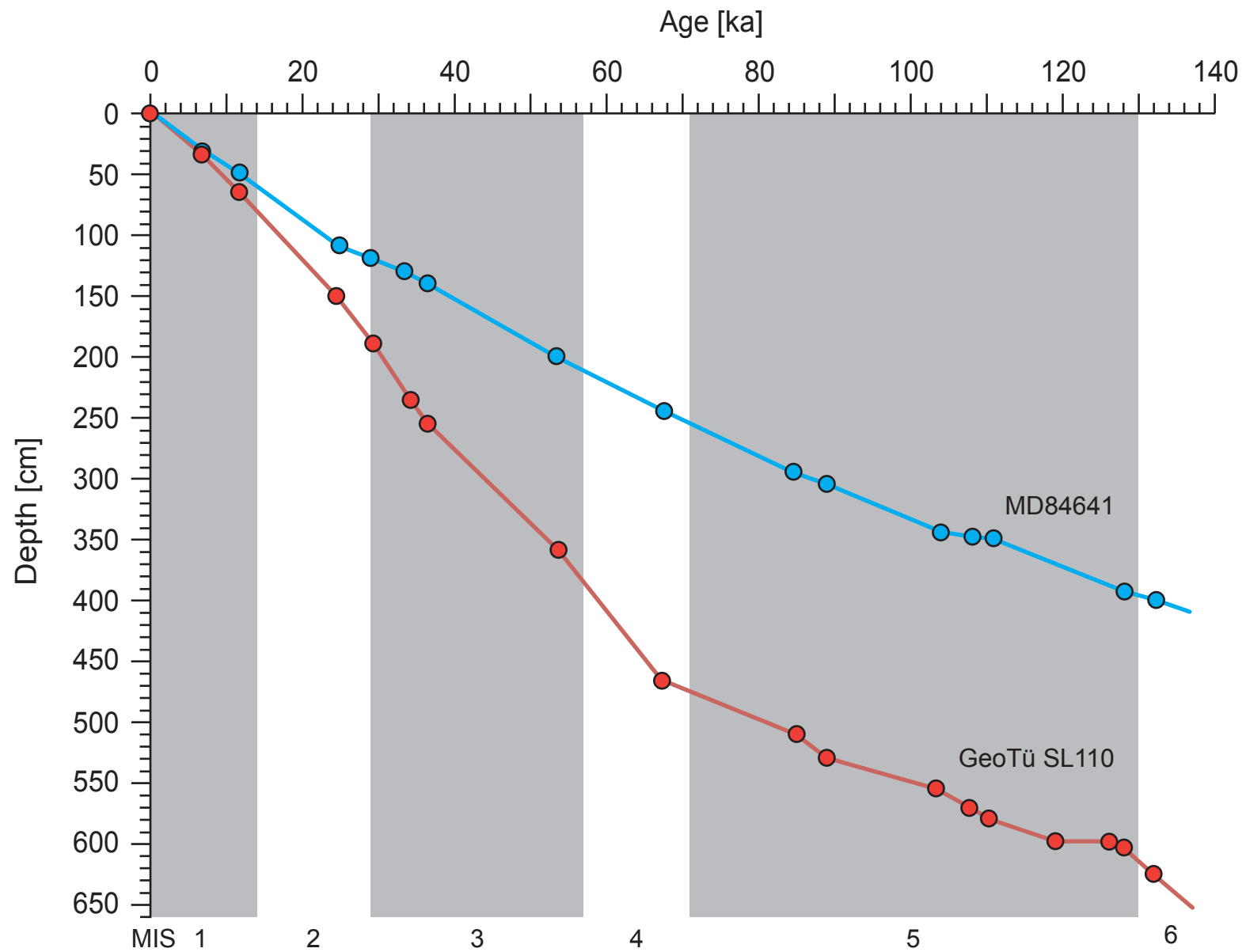


Figure 2 - Ehrmann et al.

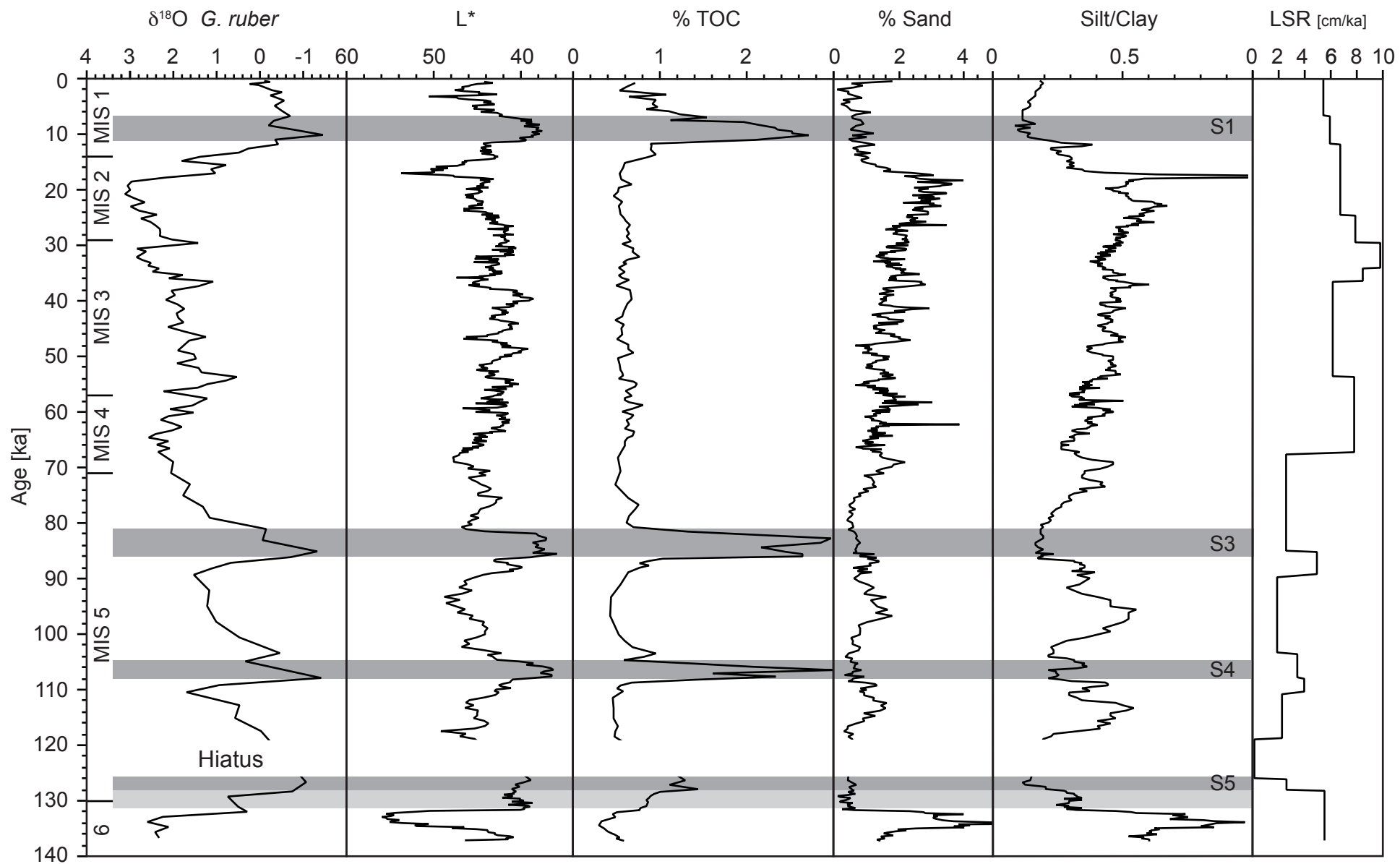


Figure 3a - Ehrmann et al.

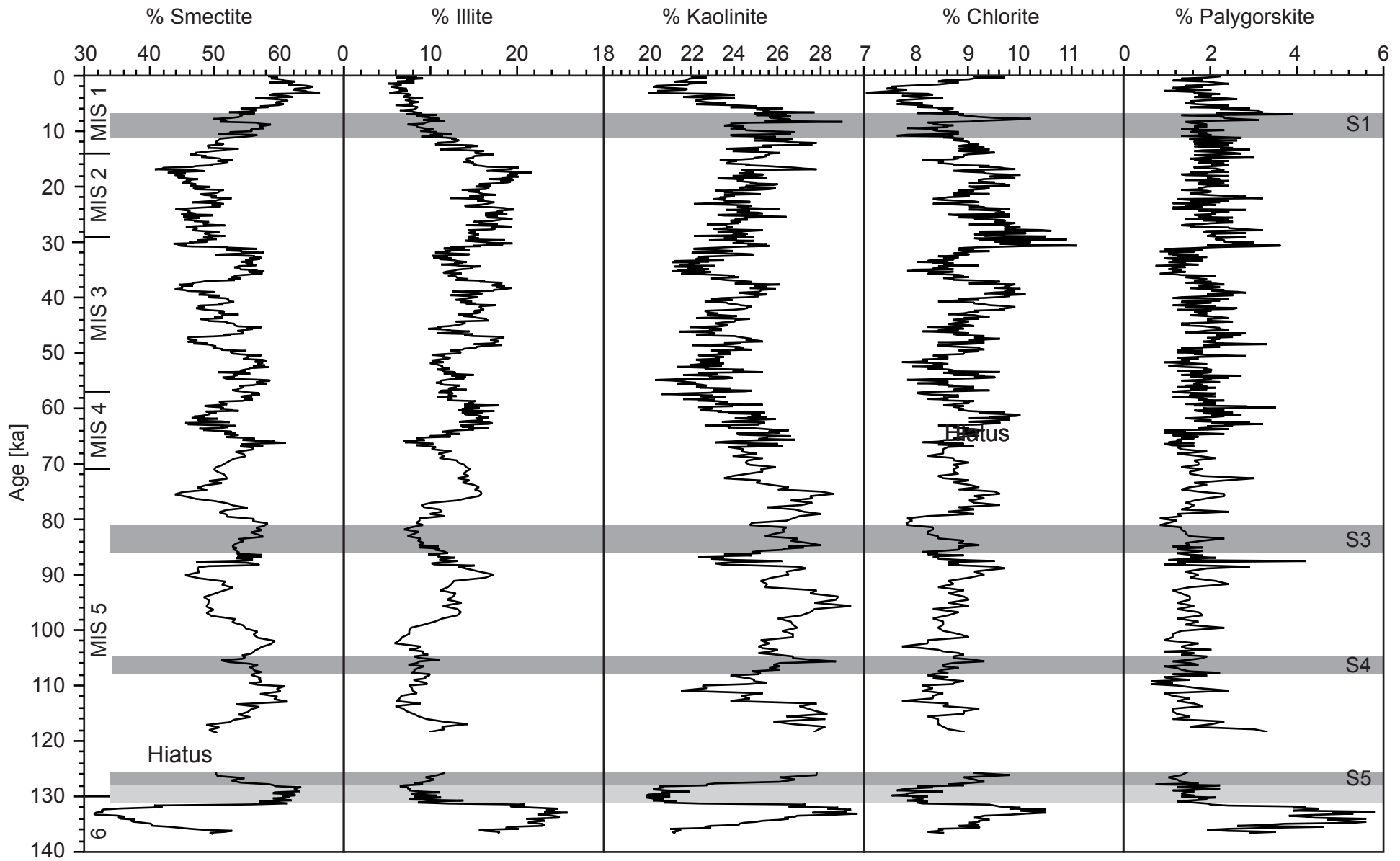


Figure 3b - Ehrmann et al.

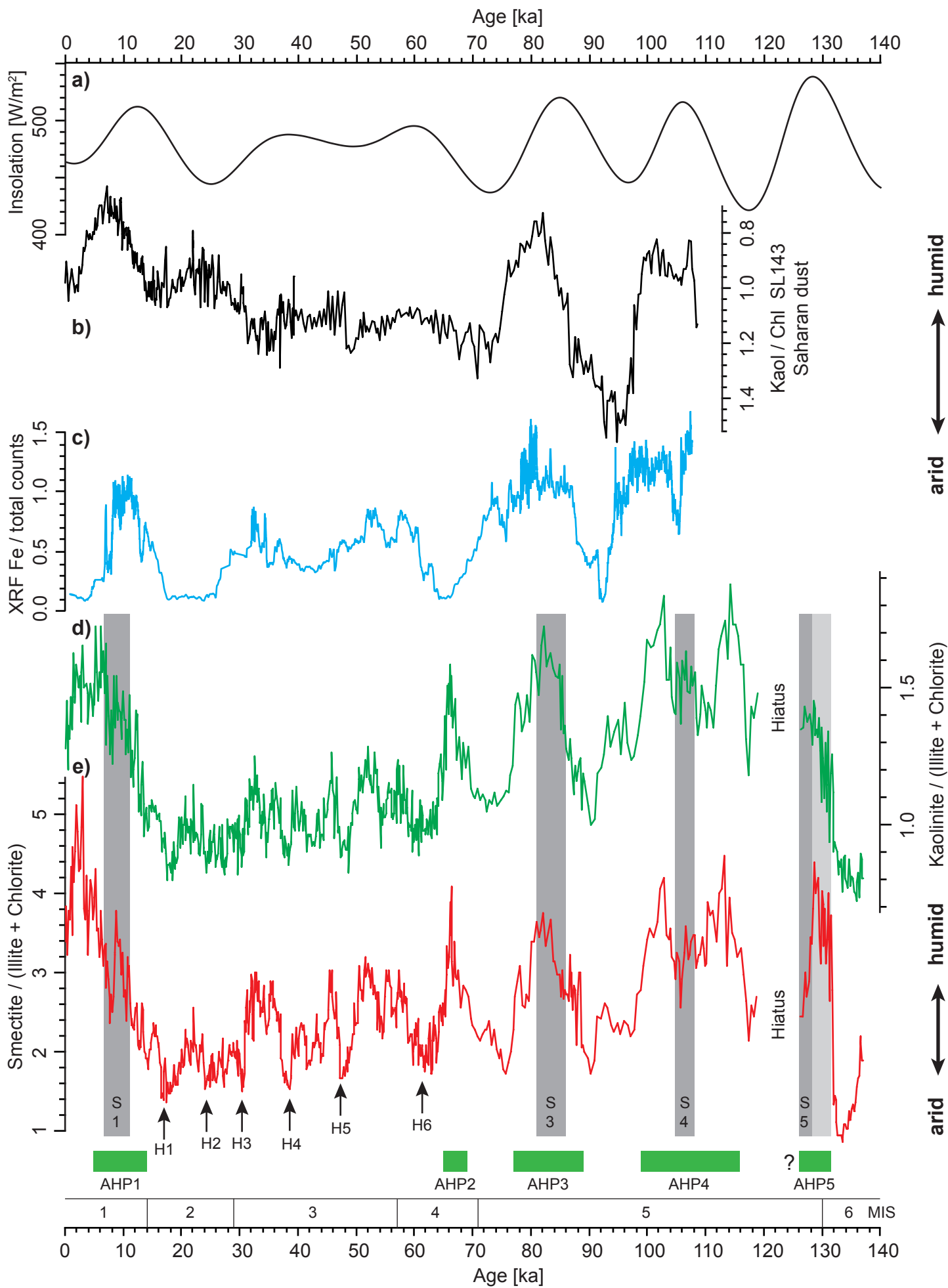


Figure 4 - Ehrmann et al.