



# Interannual climate variability seen in the Pliocene Model Intercomparison Project

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Abstract

Following proxy observations of weakened temperature gradients along the Equator in the early Pliocene, there has been much speculation about Pliocene climate variability. A major advance for our knowledge about the later Pliocene has been the coordination of modelling efforts through the Pliocene Model Intercomparison Project (PlioMIP). Here the changes in interannual modes of sea surface temperature variability will be presented across PlioMIP. Previously model ensembles have shown little consensus in the response of the El Niño–Southern Oscillation (ENSO) to imposed forcings – either for the past or future. The PlioMIP ensemble, however, shows surprising agreement with eight models simulating reduced variability and only one model indicating no change. The Pliocene’s robustly weaker ENSO also saw a shift to lower frequencies. Model ensembles focussed at a wide variety of forcing scenarios have not yet shown this level of coherency. Nonetheless the PlioMIP ensemble does not show a robust response of either ENSO flavour or sea surface temperature variability in the Tropical Indian and North Pacific Oceans. Existing suggestions of ENSO properties linked to changes in zonal temperature gradient, seasonal cycle and the elevation of the Andes Mountains are investigated, yet prove insufficient to explain the coherent response. The reason for this surprisingly coherent signal warrants further investigation.

1 Introduction

The Pliocene is an interesting time period for palaeoclimate research. It was the last time that atmospheric carbon dioxide concentrations were similar to their present, elevated values (Masson-Delmotte et al., 2013; Fedorov et al., 2013). Sea surface temperature estimates from geochemical proxies indicate that the Tropical Pacific saw reduced zonal (Wara et al., 2005) and meridional temperature gradients (Brierley et al., 2009) in the Early Pliocene. The gradient developed as time progressed, but was still significantly reduced by the mid-Pliocene (Dowsett et al., 2012). These reduced tem-

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perature gradients are thought to have consequences for the climate variability of the period. Most discussions have centred on the behaviour of the El Niño Southern Oscillation (ENSO) – an interannual mode of climate variability whose warm, El Niño phase has been used to help visualise a world with reduced zonal temperature gradients (e.g. Fedorov et al., 2006).

Efforts have been made from both the observational and modelling sides to understand the relationship between the reduced zonal temperature gradient in the Pliocene and ENSO. Watanabe et al. (2012) and Scroxton et al. (2011) present geological evidence for the existence of ENSO variability, whilst Watanabe et al. (2012) go further and suggest that any changes in period cannot have been too great. The idealised model study of Fedorov and Philander (2001) anticipates weaker ENSO amplitude with a weaker sea surface temperature (SST) gradient, whilst von der Heydt et al. (2011) disagree. Both Haywood et al. (2007) and Fedorov et al. (2010) simulate an extant ENSO in coupled climate models with reduced zonal SST gradients. Manucharyan and Fedorov (2014) manipulate the zonal SST gradient in a coupled climate model and describe a non-linear dependence of the climate variability on the resultant state. It has not yet been possible to drive substantial alterations in the Pliocene zonal SST gradient through a physically plausible mechanism that does not involve modifications to the fundamental physics in a climate model.

Zhang et al. (2012b) posit an elegant solution to these questions by describing the tropical climate of a simulation performed as part of the Pliocene Model Intercomparison Project (PlioMIP). They found that the NorESM-L model (Table 1) simulates a reduction of zonal SST gradient as well as a weakening of ENSO amplitude when forced by mid-Piacenzian boundary conditions. Here I describe the climate variability of all the model simulations that were performed as part of the same PlioMIP ensemble. In part, this work is motivated by a desire to explore whether this result holds across the ensemble; in part by academic curiosity about past climate variability.

The remainder of this article will first give a brief overview of the PlioMIP ensemble (Sect. 2.1) and detail the data analysis methods used (Sect. 2.2). The results section

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will describe Pliocene climate variability in the Tropical Pacific (Sect. 3.1), Indian (Sect. 3.2) and North Pacific (Sect. 3.3) oceans. I will then discuss the implications of these results and attempt to provide an explanation for them.

## 2 Methods

### 2.1 PlioMIP

The PlioMIP ensemble has been created to investigate the climate of the warm mid-Piacenzian. Here, I utilize only the contributions to experiment 2 which requires the use of coupled climate models (Haywood et al., 2013). Nine climate models from different institutions around the world have contributed simulations to the project (Table 1). Each model has performed a preindustrial simulation following the CMIP5 protocol. They also perform a simulation representing the average interglacial conditions between 3.264 to 3.025 million years ago (Haywood et al., 2011). This involves increasing the carbon dioxide concentrations (278 to 405 ppm) and altering the prescribed boundary conditions in line with the Pliocene Research Interpretation and Synoptic Mapping reconstructions (PRISM; Dowsett et al., 2012). A full description of the PlioMIP experiment protocols are given by Haywood et al. (2011) and the implementation of them for each model is given by the accompanying reference in Table 1.

The simulations in the PlioMIP ensemble have been previously investigated. As well as each of the single model descriptions (listed in Table 1), there have been coordinated studies across the whole ensemble. The large-scale features of the ensemble are described in Haywood et al. (2013). The simulations warm between 1.8 and 3.6 °C on the global mean. Two of the individual model papers describe aspects of the climate variability (Rosenbloom et al., 2013; Zheng et al., 2013). Additionally Zhang et al. (2012b) focus specifically on the ENSO simulation of NorESM-L. They find substantially weaker variability in the Niño 3.4 region, which they associate with the reduced equatorial SST gradient in the Pliocene simulation. An analysis of the climate variability

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across the PlioMIP ensemble using an unified methodology has not been previously performed.

## 2.2 Data analysis

Many differing approaches have been devised to study climate variability. Two distinct techniques to detect modes of variability are followed here: principal component analysis and index-driven metrics. Both techniques involve computing the average annual cycle and looking at anomalies from that. The lengths of records archived in the PlioMIP database is at best 200 years of sea surface temperatures; at worst 100 years are available (see Table 1). As these time series represent “equilibrium” states from the end of long simulations, they should have quasi-stable climates. The annual cycle was computed as the average of the whole available record for each model grid point and month of the year. Sea surface temperature anomalies (SSTA) from this annual cycle were computed.

The area-weighted average SSTA were computed over several standard regions. The Niño 3.4 region is used for the majority of the analysis, which is defined as 170–120° W, 5° S–5° N. Niño 3 (150–90° W, 5° S–5° N) and Niño 4 (160° E–150° W, 5° S–5° N) were also computed. They show similar response to the Niño 3.4 indices, so the following will concentrate on it. However, Ren and Jin (2011) devised a simple transformation of the Niño 3,  $N_3$ , and Niño 4,  $N_4$ , regions to track two different types of ENSO. These are discussed further in Sect. 3.1.4. This transformation creates two new indices,  $N_{CT}$  and  $N_{WP}$  characterising the cold tongue and warm pool types of El Niño respectively (Eq. (1); Ren and Jin, 2011).

$$\begin{cases} N_{CT} = N_3 - \alpha N_4 \\ N_{WP} = N_4 - \alpha N_3 \end{cases} \quad \alpha = \begin{cases} \frac{2}{5}, & N_3 N_4 > 0 \\ 0, & \text{otherwise} \end{cases} \quad (1)$$

Trenberth (1997) advocate using a 5 month running window to smooth ENSO SSTA indices. Some ensemble members exhibit high sub-seasonal variability (e.g. NorESM-L; Zhang et al., 2012b). Also the normalisation (described below) of the power spectra

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would be distorted by variations in the annual cycle. Therefore, smoothing is instead performed with a Lanczos low pass filter of eighteen months (UCAR, 2014).

The Dipole Mode Index is used to look at Indian Ocean variability (Saji et al., 1999; Cai et al., 2013). It is defined as the difference between area-weighted average temperature anomalies between the West (50–70° E, 10° S–10° N) and the South East (90–110° W, 10–0° S) Indian Ocean.

An alternate approach to define the Indian Ocean Dipole is as the second Empirical Orthogonal Function (EOF) of the Indian Ocean between 20° S and 20° N (Deser et al., 2010). Principal component analysis is used to compute the EOFs from the unsmoothed sea surface temperature anomalies using an area-weighted covariance approach (UCAR, 2014). The first two EOFs of the Tropical Pacific (140° E–80° W, 15° S–15° N) and North Pacific (110° E–90° W, 20–65° N) are also computed along with the global patterns. Power et al. (2013) found a robust future change in the ENSO structure by creating ensemble averaged EOF patterns. To facilitate this, the EOFs are computed on their native grid and then bilinearly interpolated onto the 2° × 2° grid of the ERSST observations (Smith et al., 2008). The EOF patterns are then normalised by the spatial standard deviation of the region (Power et al., 2013).

Spectral analysis is used to analyse the frequency properties of the interannual variability. The power spectra are computed from a smoothed and tapered periodogram (UCAR, 2014). To isolate the temporal properties from any amplitude changes, all time series are normalised by their own standard deviations. (Visualisation of SST time series would be normalised solely by the preindustrial standard deviation, as in Zhang et al. (2012b), to retain changes in amplitude.) This applies to both the SST indices and the principal components, both of which are low pass filtered at 18 months (as described above) to avoid signatures associated annual cycle changes. The dominant period (as shown in Tables 2 and 3) is defined as the period with the maximal power density in the spectrum. Ensemble average spectra are computed after this normalisation.

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## 3 Results

### 3.1 Tropical Pacific

The most well studied mode of interannual climate variability is the El Niño Southern Oscillation (ENSO). However, this does not mean it is yet well understood. Future projections of ENSO are model-dependent – with some models showing increasing variability and others showing reductions (Collins et al., 2010). There are a plethora of metrics and approaches to investigate how well models simulate the present-day properties of ENSO. Novel approaches have started to focus on the underlying physical processes, rather than the emergent climate variability that will be the focus of this work (e.g. Jin et al., 2006). The model representation of ENSO has improved in the CMIP5 compared to CMIP3 (Bellenger et al., 2013) with more models showing a realistic power spectrum (Flato et al., 2013). Power et al. (2013) was able to show the benefit of separating the various aspects of ENSO during model intercomparison by finding a robust pattern of change in 21st century simulations. Motivated by this result, combined with the fact that the PlioMIP models show a large diversity of ENSO simulation quality, each property is investigated independently of the others.

#### 3.1.1 Moments

The quality of the ENSO representation varies across the ensemble. COSMOS has substantially more variability than seen in observations (over six times as much as seen in ERSST). In this particular model, it is probably related to the excessive seasonality of an extensive cold tongue. Conversely, MIROC4m has only half as much variance as seen in observations. It is also worth pointing out that where the model versions have been included in the CMIP3 or CMIP5 ensembles, the statistics computed do not exactly match those of Bellenger et al. (2013). I suspect this arises predominantly from the shorter lengths of the preindustrial controls simulations analysed here (which have been chosen to correspond to the Pliocene data length where possible), with

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some contribution from the 1.5 year low-pass filter applied here. These differences do lend a cautionary aspect as they highlight the centennial scale variability in ENSO characteristics (Wittenberg, 2009) that could confound the results shown here.

The variance is the most discussed moment as it is a valuable metric for the amplitude of the Niño 3.4 distribution (although sometimes its square root is used instead such as by Taschetto et al., 2014). There is a near unanimous weakening ENSO amplitude in seen in the mid-Pliocene simulations (Fig. 1). The lone dissenting model is MRI, which itself shows no change. The ensemble average variance in the mid-Pliocene is roughly two-thirds the variance in the preindustrial control simulation (Table 2).

A less coherent pattern comes from looking at the higher moments of the distribution. It is observed that El Niños are generally stronger than their La Niña counterparts – equivalent to the Niño 3.4 distribution having a positive skew. This property is captured by two thirds of the ensemble members in the preindustrial conditions. However the ensemble is equivocal about changes in skew for the Pliocene simulations (Fig. 1). The kurtosis is a measure of spread of the distribution compared to a Gaussian. A positive kurtosis means more of the activity occurs near the mean (zero by definition in this case) rather than out in the tails. Again there is little consistency across the ensemble as to the change in kurtosis (Fig. 1).

These moments can combine to define other possible descriptions of ENSO. For example, Cai et al. (2014) detect an increase in intensity of the strongest El Niños in future simulations across most CMIP5 ensemble members, despite little agreement on trends in the moments investigated here. Their analysis looked at precipitation where this feature has oft been noted before such as in the hydrological cycle (Collins et al., 2013) and tropical cyclones (Knutson et al., 2010). Nonetheless it is not implausible for increasing skew or a reduction in kurtosis to alter the tail of the Niño 3.4 SST distribution, such that more intense El Niños would be seen despite a reduction in variance overall.

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### 3.1.2 Frequency

The only observational evidence of ENSO properties in the Pliocene relate to its frequency (Watanabe et al., 2012). Therefore the simulated changes in the frequency are studied in isolation. This is achieved by normalising the time series by their standard deviations (Sect. 2.2). This then allows the power spectra to be averaged across the ensemble (Fig. 2). The power at the annual and sub-annual period have been removed through the data analysis (Sect. 2.2; unlike An and Choi, 2013).

Eight of the nine models show the dominant period of Niño 3.4 SSTA becoming longer; only COSMOS exhibits a shift to a shorter period (Table 2). This behaviour is also shown in the ensemble average power spectra (Fig. 2). Both these spectra exhibit two peaks at around 3 years and a stronger one at lower frequencies. The lower frequency peak shifts substantially to longer periods, the significance of the smaller shift in the 3 year peak is debatable. A coherent lengthening of ENSO period is not robustly detected in future simulations. An and Choi (2013) notice a lengthening for the mid-Holocene, but only with the PMIP2 generation of models not for PMIP3.

### 3.1.3 Structure

The conventional approach to investigate changes in the SST pattern associated with an El Niño involve compositing of the events above a certain threshold (e.g. Haywood et al., 2007). I aim to study changes in the structure climate variability patterns in isolation from any changes in amplitude or period, both of which show a robust Pliocene response individually. This is achieved by following the methodology of Power et al. (2013) through EOF analysis and normalisation. For each model, the first EOF of the Tropical Pacific is significantly correlated to the Niño 3.4 ( $P < 0.05\%$ ). The first EOFs of the ensemble explain on average 60 % of the Tropical Pacific variance in the preindustrial simulations compared to 56 % of the variance in the Pliocene simulations on the ensemble average.

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The structure of ENSO is predominantly similar in the two time periods (Fig. 3). There is a shift towards the warmth of an El Niño expanding further across the Tropical Pacific. Statistical significance with such a small ensemble is hard to compute. Instead a simpler measure of coherency is used – a location is considered coherent if more than seven of the nine models show a response in the same direction (Fig. 3).

### 3.1.4 Flavour

A recent development has been the identification of two types of El Niño (Ashok et al., 2007), which probably represent end members from a range of behaviour. A variety of approaches exist to detect and disaggregate the two different flavours. For ease of computation, I opt for the technique of Ren and Jin (2011, Eq. 1). The variances of the cold tongue and warm pool types of ENSO show a large spread in quality of their simulation compared to observations (Fig. 4). This is also true of the wider CMIP5 ensemble (Taschetto et al., 2014).

The ensemble shows no robust change between the variances of the two flavours (Fig. 4). Instead the reduction in amplitude effects both flavours (only MRI-CGCM2 shows an increase in either form and even that is minimal). Both HadCM3 and NorESM-L switch the dominant flavour in the mid-Pliocene, but in opposite directions. An alternate method to address the same question is to analyse the second EOF of the Tropical Pacific. This also shows little coherent shift (not shown).

### 3.1.5 Summary

To summarise, the PlioMIP ensemble is remarkable amongst climate ensembles in that it shows a consistent message about changes in ENSO properties. ENSO in the Pliocene is simulated to have been weaker and less frequent. Changes in structure and flavour are less coherent however. The reason why the PlioMIP ensemble shows a signal, whilst other ensembles do not, is unclear. Compared to CMIP5 future scenario ensembles, this one has been run to equilibrium and has longer time series to analyse.

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However, that cannot be sufficient as other equilibrium ensembles do not see robust responses – for example the mid-Holocene (An and Choi, 2013) and the Last Glacial Maximum (Zheng et al., 2007).

### 3.2 Tropical Indian variability

Climate variability in the Tropical Indian Ocean has strong links to that in the Pacific, in part through their sharing of a single warm pool. The leading mode of Indian Ocean temperature variability is therefore a corollary to ENSO, although it is sometimes referred as to the Basin Mode (Deser et al., 2010). Around the turn of the century an additional mode of variability was identified – the Indian Ocean Dipole (IOD, Saji et al., 1999). The Indian Ocean Dipole can sometimes track ENSO, yet sometimes operate in isolation. It manifests through changes in the zonal gradient of the Indian Ocean and can be tracked by the Dipole Mode Index (DMI, Saji et al., 1999, Table 2). The quality of the preindustrial simulation of the Dipole Mode Index varies across the ensemble; as it does in the wider CMIP5 ensemble (Weller and Cai, 2013). Whilst several of the ensemble members show a change in amplitude of the Indian Ocean Dipole, there is no consensus in their sign (Table 2). The detection of a coherent signal may be hampered by biases in the climate models meaning that the DMI is not capturing the centres of actions of the Indian Ocean Dipole.

An alternate approach is to use principal component analysis to detect the Indian Ocean Dipole. A conventional methodology is to label the first EOF as the Basin Mode and the second EOF as the Dipole Mode (Deser et al., 2010). Because of potential model biases, the most appropriate mode is determined from visual comparison of the model-derived EOFs to those found from the detrended ERSST observations. The results of this analysis is shown in Table 3, along with both the variance and dominant period of the accompanying principal component timeseries. It is hard to pull out a coherent message about Pliocene Indian Ocean variability from this ensemble. Although it is interesting to note the potential for a change in the relative strengths of the two modes – as seen by CCSM4 for example (Table 3).

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### 3.3 North Pacific variability

The dominant form of sea surface temperature variability in the North Pacific is a low frequency mode termed the Pacific Decadal Oscillation (PDO, Mantua et al., 1997). The centre of action of this oscillation is the Kuroshio Extension around 40° N, although it has significant impacts remote from this region. The connection between the PDO and low frequency variations in ENSO is complex. Here, I follow Deser et al. (2010) and define the PDO as the first EOF of the North Pacific (20–70° N). The EOF patterns are again normalised by their spatial standard deviations, however in this instance both simulations are normalised by the preindustrial value. This retains information about the relative amplitude of the mode in the principal component time series. The variances of the PDO timeseries are shown in Fig. 5 along with the percentage change. There is little consistent response of the PDO amplitude to the imposing of the Pliocene conditions. The same is also true of the frequency and structure of the PDO when studied in isolation (not shown).

## 4 Discussion

Climate variability in three regions of the global ocean have been investigated in the PlioMIP ensemble. The nine constituent models often are equivocal about the changes in climate variability that might have occurred in the Pliocene. However, for the two most studied facets of the dominant mode of variability, they provide a coherent message: Pliocene ENSO was weaker and less frequent.

This raises three questions: (a) how does this response fit with the geological evidence (b) did ENSO become so weak that it lost its global reach, and (c) why does the PlioMIP ensemble show this response?

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## 4.1 Observations

Observational evidence for Pliocene climate variability is sparse. Hill et al. (2011) discuss changes in the Northern Annular Mode; something not analysed here. Scropton et al. (2011) use the distribution of single foraminifera in the Eastern Equatorial Pacific to infer the presence of ENSO in the Pliocene. Changes in that distribution are more likely to reflect a response in the amplitude of the underlying seasonal cycle that ENSO properties (Thirumalai et al., 2013). Nonetheless, the technique should be sufficient to detect presence or absence of ENSO. None of the climate models in PlioMIP ensemble (or elsewhere for that matter) show a lack of interannual variability in the Eastern Equatorial Pacific – in agreement with the results of Scropton et al. (2011). However, it is doubtful that one would expect a lack of variability in this region, even with a substantially reduced zonal SST gradient in the Equatorial Pacific (Manucharyan and Fedorov, 2014).

The only other observational study of Pliocene climate variability looks at fossil corals from the Philippines (Watanabe et al., 2012). The  $\delta^{18}\text{O}$  is sampled at regular intervals during the corals' roughly thirty year life span. A periodicity of approximately three years is found in the Pliocene corals, which lived between 3.5–3.8 million years ago. This again provides evidence of ENSO variability in the Tropical Pacific. Unfortunately due to natural, internal shifts in ENSO properties on multidecadal timescales (Wittenberg, 2009), it would be impossible to infer an increase in this period with any confidence from the 30 year records available.

## 4.2 Global dominance

ENSO is the focus of substantial interest amongst the palaeoclimate community, partly because it is the globally dominant mode of climate variability (Deser et al., 2010). It is projected to stay this way for the coming century (Christensen et al., 2013). In the PlioMIP ensemble, ENSO becomes weaker whilst other modes of climate variability do not appear to do so. Is it possible that ENSO becomes so weak that its teleconnected

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response is not large enough to overwhelm local climate variability? Such a situation was suggested by Brierley (2013) as existing around four million year ago at the time of the weakest zonal SST gradients in the Tropical Pacific (Fedorov et al., 2013).

Quantitatively detecting a change in global dominance is problematic. However, visual inspection of the leading global empirical orthogonal function can potentially show this. The two leading preindustrial EOFs of NorESM-L show strong signatures in both the North Pacific and the Niño region (Fig. 6). In fact, their principal component time-series show a marginally higher correlation to the PDO than ENSO – a feature not seen in observations. NorESM-L shows the strongest reduction in ENSO amplitude in the Pliocene across the ensemble (Table 2). For this model, the two leading global EOFs in the Pliocene show no signature in the Tropical Pacific (Fig. 6). A similar feature is not seen with the other PlioMIP models (not shown).

Whilst this analysis does not state that ENSO in the Pliocene was confined to the Tropical Pacific, it at least shows the potential for that to be the case. A lack of ENSO-style periodicities in regions remote from the tropical Pacific in the Pliocene may not denote a change in teleconnection patterns, but rather be indicative of a weaker amplitude for ENSO.

### 4.3 Cause

Why does the PlioMIP ensemble see a robust weakening and lengthening of ENSO? There have been many suggestions as to what controls the properties of ENSO. It is beyond the scope of this article to test them all or to investigate the ENSO changes from a process-based approach. However, several possibilities can be easily evaluated from the data analysed here.

Other model ensembles show a range of behaviours and often contain more members than analysed here. Perhaps it is purely coincidence that the PlioMIP ensemble presents such a coherent message? Using the subset of CMIP5 models by Taschetto et al. (2014) this probability can be estimated. The PlioMIP and CMIP5 models are categorised into three groups depending on their Niño 3.4 standard deviation change:

decrease, increase or no change. Here “no-change” is arbitrarily defined as having an absolute standard deviation of less than 0.05 °C, although the conclusion is robust to other threshold settings. The PlioMIP ensemble therefore forms a set of 9 models with 7 decreasing and 2 with no-change. The CMIP5 ensemble contains 27 models with 8 decreasing and no change in 11. There 5071 combinations showing a reduction at least as robust as the PlioMIP ensemble out of a total  $4.6 \times 10^6$ . This probability of only 0.01 % is likely somewhat of an underestimate, because the lengths of simulation analysed by Taschetto et al. (2014) are substantially shorter than here (Stevenson et al., 2012). Nonetheless, I consider the PlioMIP signal did not occur by chance and is a real feature that needs exploring.

The commonality of PlioMIP experiments is in their application of (nearly) identical forcing and boundary conditions changes (Haywood et al., 2011). The ENSO weakening and lengthening should therefore be a response to one of these boundary changes. Feng and Poulsen (2014) find a response of ENSO to alterations in the elevation of the Andes mountain range, using CCSM4 simulations. The PRISM reconstruction has lower Andes elevation in the Pliocene (Haywood et al., 2011). Unfortunately, Feng and Poulsen (2014) show that a lower Andes would result in a shorter ENSO period – something that even the CCSM4 PlioMIP simulation does not show (Table 2). Therefore, any Andean response must be masked by another factor (assuming it is not model dependent, which is as yet untested).

Much of the interest in Pliocene ENSO has been driven by the reduced zonal SST gradients seen in the mean climate state (Wara et al., 2005; Fedorov et al., 2013). It has been anticipated that ENSO properties are related to this gradient (e.g. Fedorov and Philander, 2001; Zhang et al., 2012b), although the precise nature of this relationship appears complicated (Manucharyan and Fedorov, 2014). Figure 7 shows the Equatorial SST in the Pacific (averaged between 5° S–5° N). It is clear that all ensemble members simulate an West–East gradient, although with a range of fidelity to the observed gradient. The warmth of the Pliocene can be seen by the fact that all models show at least 0.5 °C of warming at all longitudes. However there is no consistent

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preferential warming seen in the Eastern Equatorial Pacific, as seen in the palaeobservations (Dowsett et al., 2012; Fedorov et al., 2013). Therefore a consistent reduction in the zonal SST gradient cannot explain the robustly weaker and longer ENSO. A similar lack of explanation is borne out, when assessing whether the mean state becomes more El Niño like (Fig. 8; Collins and The CMIP Modelling Groups, 2005).

An alternate possibility that can be easily tested is the suggestion that the amplitude of ENSO is inversely correlated to the seasonal amplitude (Liu, 2002). This possibility has already been questioned by Braconnot et al. (2011) and the PlioMIP ensemble does not show support for it either. Whilst there are changes in the seasonal cycle of Niño 3, they are not coherent across the PlioMIP ensemble (Fig. 9).

## 5 Conclusions

The Pliocene Model Intercomparison Project (PlioMIP) has coordinated nine coupled climate models to simulate the conditions of the mid-Piacenzian. Through analysis of sea surface temperature, I have described the interannual climate variability seen across the ensemble. Although some models show changes in the Indian Ocean Dipole or Pacific Decadal Oscillation, there was little coherency throughout the ensemble in these changes. Four different properties of the El Niño Southern Oscillation have been isolated and the Pliocene changes in them investigated. There is no coherent change in the prevalence between the cold tongue and warm pool types of El Niño – in keeping with future projections (Taschetto et al., 2014). Analysis of the structure of ENSO finds an expansion of the phenomenon further polewards – unlike the pattern found by Power et al. (2013) for the future projections.

The finding that eight of the nine PlioMIP models show a reduction in Niño 3.4 variance and a shift towards lower frequencies was unexpected. The chance of such a coherent message arising by chance is slim, when considering the range of results in the CMIP5 future scenarion runs (Taschetto et al., 2014). Three possible suggestions were investigated in the hope of explaining this result: namely reductions in the annual

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mean zonal temperature gradient (Fedorov and Philander, 2001; Zhang et al., 2012b), Andean uplift (Feng and Poulsen, 2014), and seasonal cycle alterations. None was found to have much explanatory power and so the real reason for this result will require investigation. Recent developments in process-based analysis (e.g. Jin et al., 2006) that are currently being applied to modern simulations (Bellenger et al., 2013) may help find the answer to this conundrum.

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**Table 1.** Simulations contributing to the PlioMIP ensemble (see Sect. 2.1 and Haywood et al., 2013, for details). The resolution given relates to the deep Tropics for models with a varying ocean resolution.

Model	Institution Initials	Resolution		Record Length*	CMIP5	Reference
		Atmos.	Ocean			
CCSM4	NCAR	$0.9 \times 1.25^\circ$	$0.4 \times 1^\circ$	200, 100	Yes	Rosenbloom et al. (2013)
COSMOS	AWI	T31	$3 \times 1.8^\circ$	100, 150	No	Stepanek and Lohmann (2012)
FGOALS-g2	LASG	$2.8 \times 2.8^\circ$	$0.5 \times 1^\circ$	100, 100	Yes	Zheng et al. (2013)
GISS-E2-R	NASA/GISS	$2 \times 2.5^\circ$	$1 \times 1.25^\circ$	200, 200	Yes	Chandler et al. (2013)
HadCM3	UKMO	$2.5 \times 3.75^\circ$	$1.25 \times 1.25^\circ$	198, 200	Yes	Bragg et al. (2012)
IPSLCM5A	LSCE	$3.75 \times 1.9^\circ$	$0.5 \times 2^\circ$	200, 200	Yes	Contoux et al. (2012)
MIROC4m	JAMSTEC	T42	$0.5 \times 1.4^\circ$	200, 100	No	Chan et al. (2011)
MRI-CGCM2.3	MRI	T42	$0.5\text{--}2.5^\circ$	200, 200	No	Kamae and Ueda (2012)
NorESM-L	BCCR	T31	$1.5 \times 3^\circ$	200, 200	No	Zhang et al. (2012a)

\* Preindustrial followed by Pliocene simulation.

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**Table 2.** Properties of area averaged sea surface temperature indices (see Sect. 2.2 for details). The period listed below is that with the greatest spectral power. The preindustrial and Pliocene simulations, along with the difference between them, are denoted by piCtl, Plio and Diff, respectively.

Model	Niño 3.4 variance ( $^{\circ}\text{C}^2$ )			Niño 3.4 period (years)			DMI variance ( $^{\circ}\text{C}^2$ )		
	piCtl	Plio	Diff	piCtl	Plio	Diff	piCtl	Plio	Diff
CCSM4	1.04	0.71	−0.33	4.1	6.9	2.8	0.20	0.23	0.02
COSMOS	3.20	2.08	−1.16	6.5	4.2	−2.3	0.11	0.13	0.01
FGOALS-g2	0.37	0.20	−0.18	6.5	7.4	1.0	0.06	0.02	−0.03
GISS-E2-R	0.22	0.19	−0.03	3.4	4.3	0.8	0.05	0.05	−0.00
HadCM3	0.50	0.32	−0.18	3.8	4.2	0.3	0.11	0.14	0.03
IPSLCM5A	0.38	0.29	−0.08	3.3	5.1	1.8	0.13	0.08	−0.05
MIROC4m	0.22	0.09	−0.13	5.9	7.7	1.8	0.21	0.16	−0.04
MRI-CGCM2.3	0.44	0.45	0.01	2.2	2.5	0.2	0.03	0.03	0.00
NorESM-L	0.48	0.10	−0.38	4	4.1	0.1	0.17	0.06	−0.11
Ensemble Mean	0.76	0.49	−0.27	4.4	5.1	0.7	0.12	0.11	−0.01
ERSST Obs.	0.48	–	–	5.5	–	–	0.06	–	–

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**Table 3.** Properties of Empirical Orthogonal Functions (EOFs) of the Indian Ocean. The percentage of variance explained by each model's first two EOFs and the period with the most spectral power is shown. The correspondence of each EOF to physically meaningful modes of variability has been determined by visual inspection. The preindustrial and Pliocene simulations are denoted by piCtl and Plio, respectively.

Model	EOF	Mode		Var. Expl. (%)		Period (years)	
		piCtl	Plio	piCtl	Plio	piCtl	Plio
CCSM4	1st	Basin	Dipole	24.4	29.3	4.1	6.9
	2nd	Dipole	Basin	22.3	20.4	4.0	6.9
COSMOS	1st	Basin	Basin	47.3	37.7	7.1	4.1
	2nd	Dipole	Dipole	11.1	14.5	2.8	4.9
FGOALS-g2	1st	Basin	Basin	33.4	29.2	6.9	3.1
	2nd	Dipole	Dipole	12.8	10.4	6.5	4.0
GISS-E2-R	1st	Basin	Basin	27.2	24.6	3.9	2.9
	2nd	Dipole	Dipole	11.0	11.9	3.3	6.7
HadCM3	1st	Dipole	Dipole	18.6	18.1	5.0	4.2
	2nd	Dipole	Dipole	11.4	14.1	5.3	7.7
IPSLCM5A	1st	Basin	Basin	22.9	20.0	3.3	5.1
	2nd	Dipole	Dipole	11.9	13.2	6.5	7.7
MIROC4m	1st	Basin	Dipole	28.9	25.7	5.1	6.3
	2nd	Dipole	Dipole	16.0	13.7	3.9	2.9
MRI-CGCM2.3	1st	Basin	Basin	37.7	35.1	4.1	4.8
	2nd	Dipole	Dipole	8.8	8.9	7.4	5.0
NorESM-L	1st	Basin	Basin	21.0	20.2	4.0	4.5
	2nd	Dipole	Basin	14.3	12.7	5.1	2.0
ERSST Obs.	1st	Basin	–	39.1	–	3.6	–
	2nd	Dipole	–	11.8	–	5.6	–

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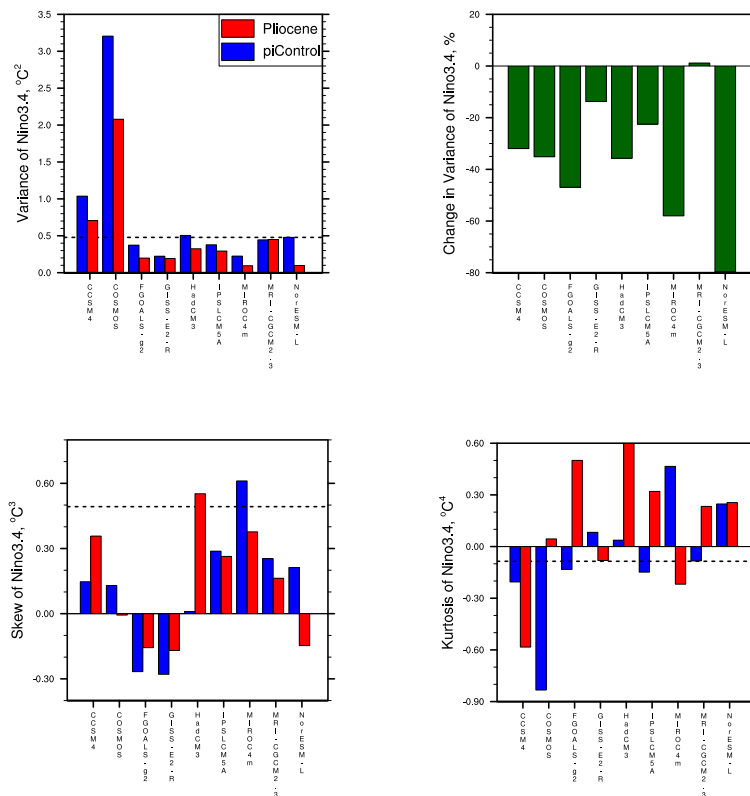
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**Figure 1.** The moments of the distributions of the Niño 3.4 sea surface temperature anomalies in the ensemble members. The first moment (i.e. the mean) is zero by definition. The variance shows the amplitude of the SST variations. If the skew is positive, then the El Niño tail is stronger than the corresponding La Niña state. If the kurtosis is greater than 0, then the distribution is more peaked than a Gaussian. The dashed line shows the moment of the ERSST observations.

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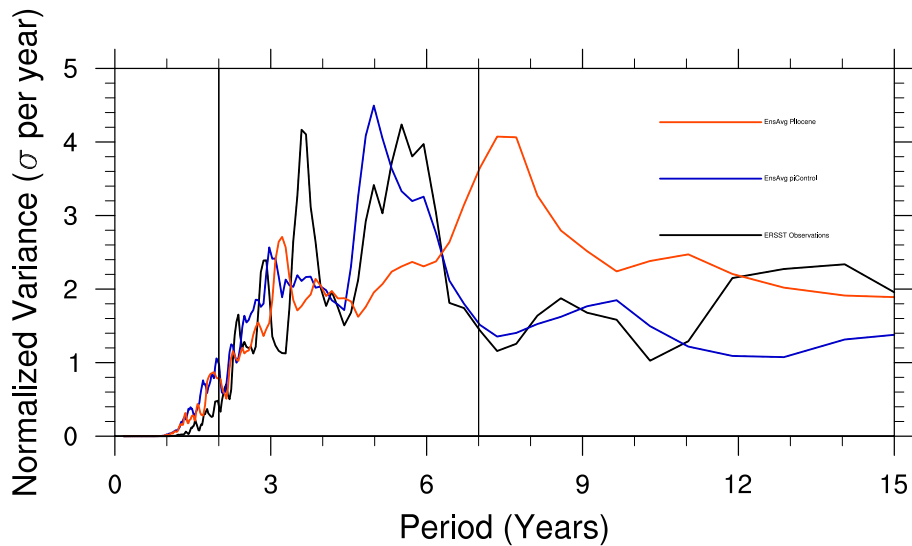
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**Figure 2.** The ensemble average power spectra of the Niño 3.4 sea surface temperature anomalies.

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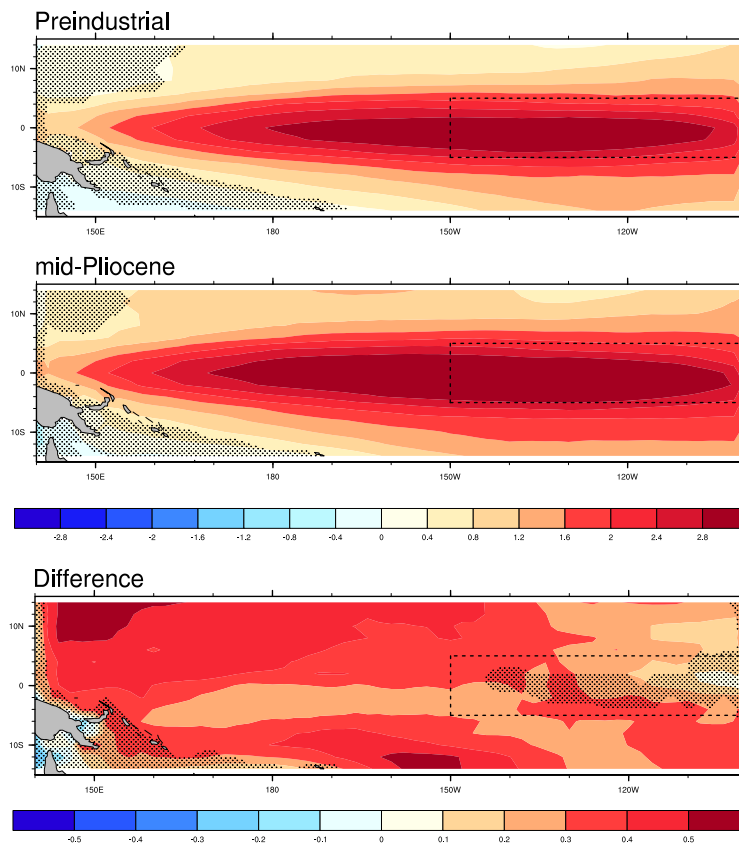
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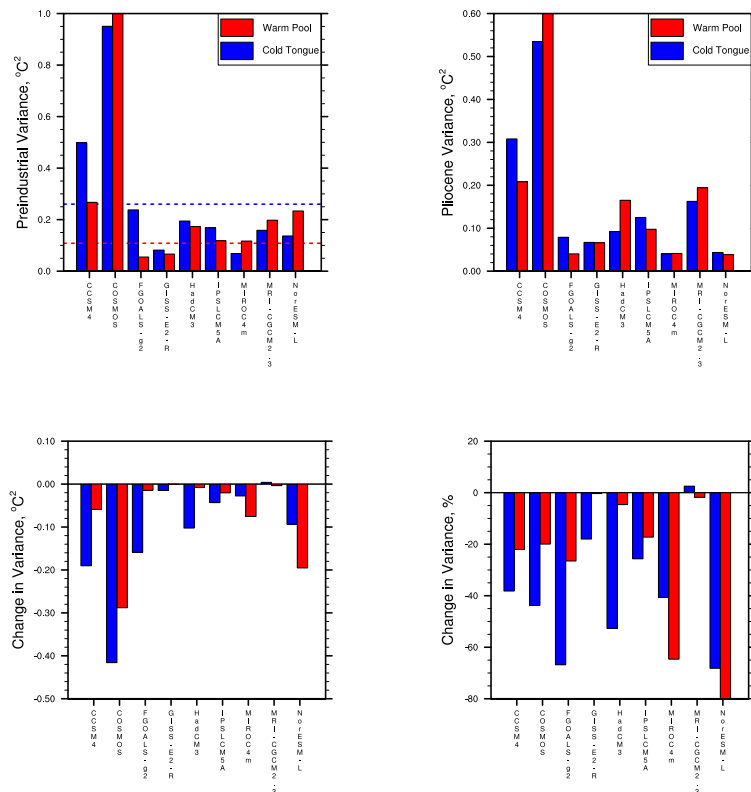
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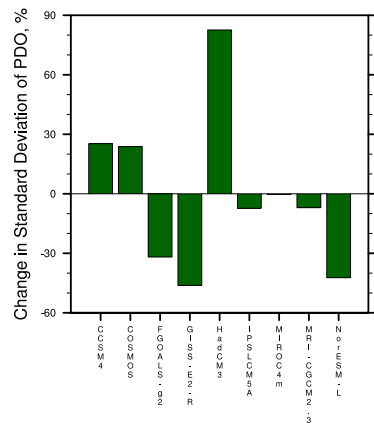
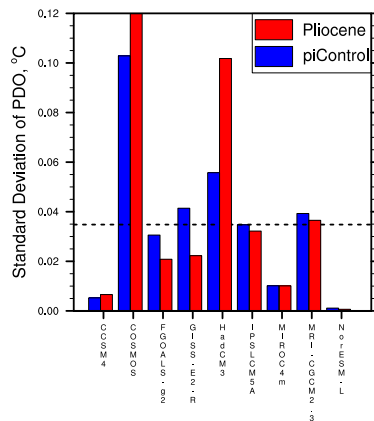
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**Figure 3.** The ensemble average empirical orthogonal functions computed for the tropical Pacific. All patterns are defined to be positive in the Niño 3.4 region (dotted line). Stippling indicates that less than eight of the nine models agree on the sign of the change.



**Figure 4.** Response of the two different flavours of ENSO considered separately. The flavours are disaggregated using the method of Ren and Jin (2011) into cold-tongue and warm-pool types. The red and blue dashed lines indicated the variance found in the ERSST observations.



**Figure 5.** The standard deviations of Pacific Decadal Oscillation – defined as the first EOF of the North Pacific. The dashed line shows the standard deviation of the PDO in the ERSST observations.

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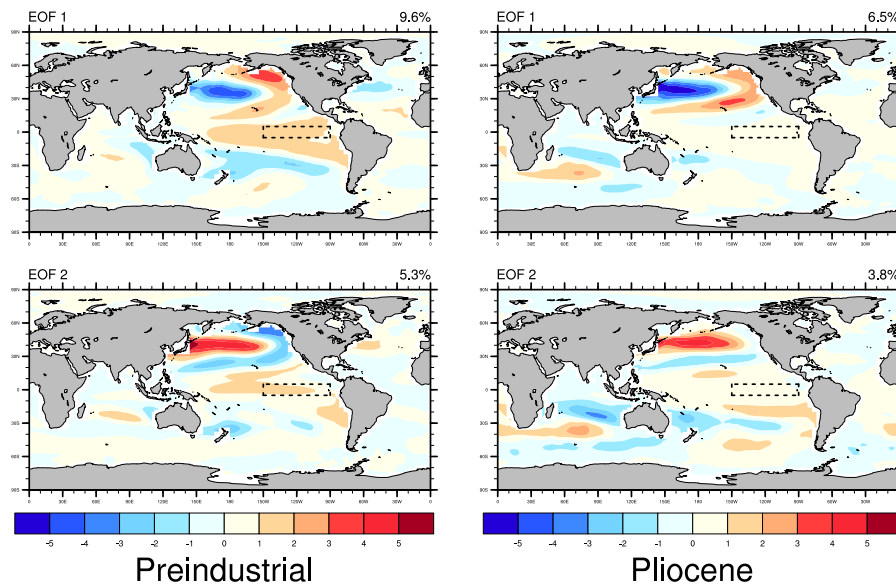
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**Figure 6.** The first two EOFs of the global SST anomalies in NorESM-L for both the preindustrial simulation (left) and the Pliocene simulation (right). The percentage of the variance that each EOF explains also shown.

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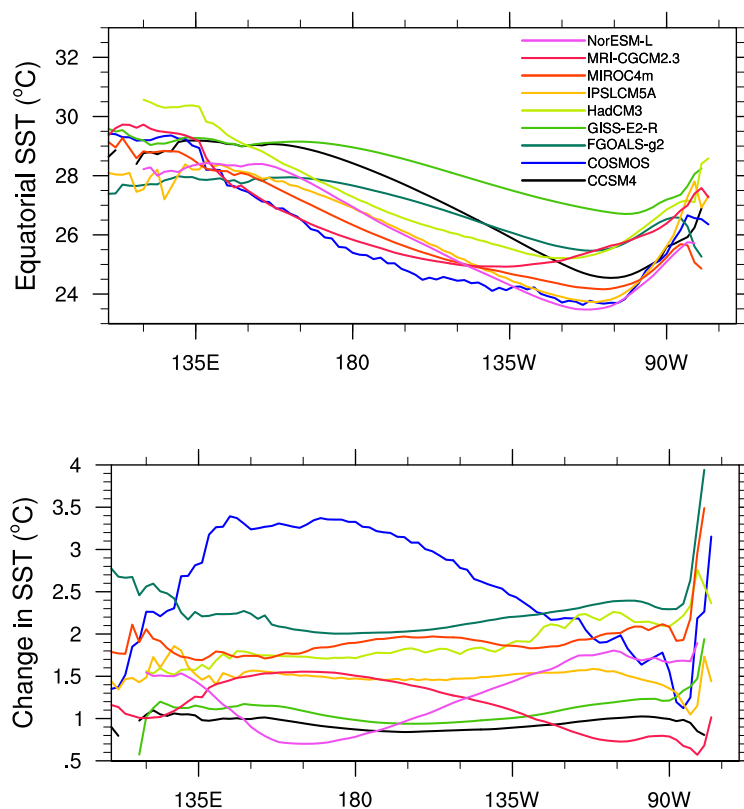
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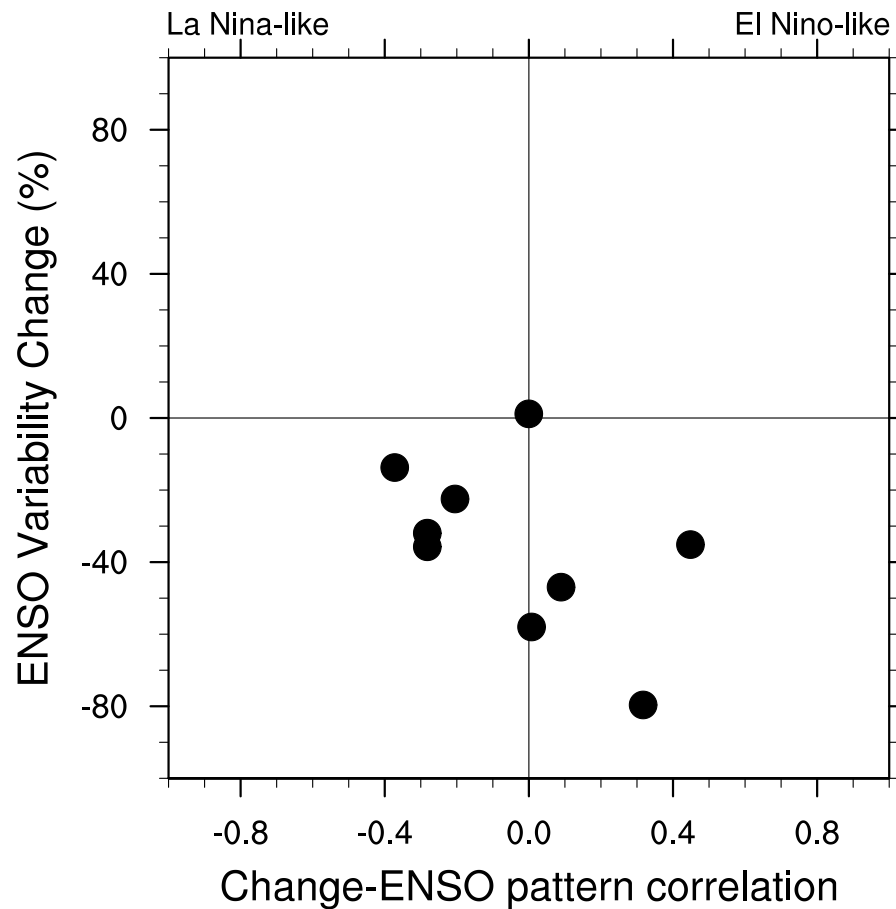
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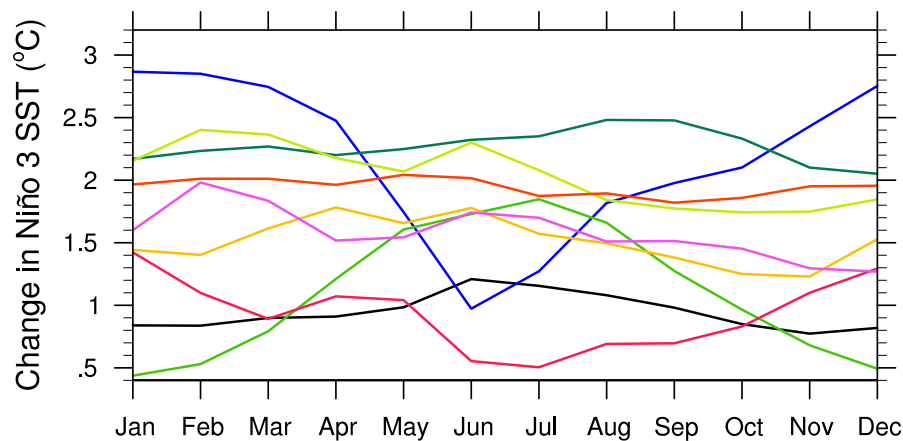
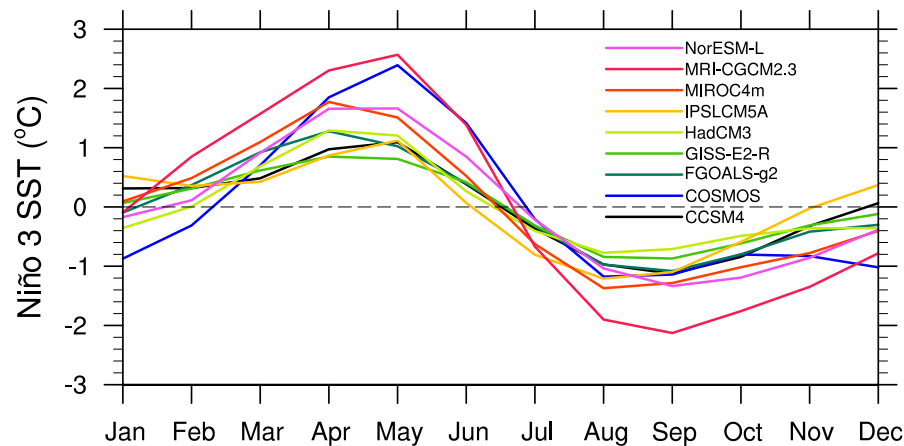


**Figure 7.** The annual mean Pacific temperature gradients along the Equator in the preindustrial simulations and the change in the Pliocene simulations.





**Figure 8.** A scatterplot of the change in ENSO amplitude (from Fig. 1) to the correlation of Pliocene mean state changes projected onto the model's leading tropical Pacific EOF.



**Figure 9.** The seasonal cycle of the Niño 3 SSTs and its changes.